



CICENNS: a 300-kg CsI(Na) Detector for Coherent Elastic Neutrino-Nucleus Scattering

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on behalf of the CICENNS Collaboration

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Coherent Elastic Neutrino-Nucleus Scattering (CEvNS)

PHYSICAL REVIEW D

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1 MARCH 1974

Coherent effects of a weak neutral current

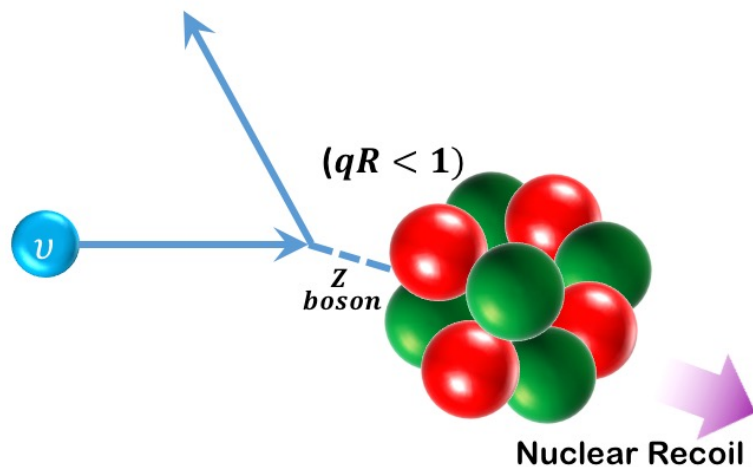
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National Accelerator Laboratory, Batavia, Illinois 60510

and Institute for Theoretical Physics, State University of New York, Stony Brook, New York 11790

(Received 15 October 1973; revised manuscript received 19 November 1973)

- $qR < 1$
- Z-exchange between neutrino and nucleus
- Nucleus recoils as a whole
- Coherent up to $E_\nu \sim 50$ MeV

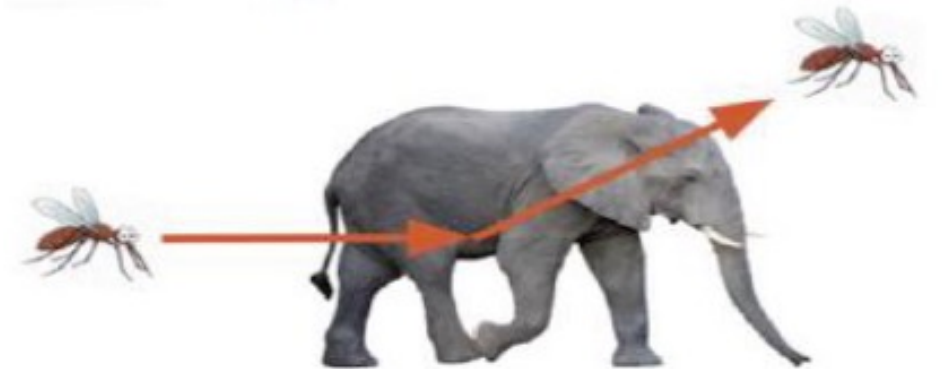


$$\frac{d\sigma}{dT} = \frac{G_F^2}{4\pi} Q_W^2 M \left(1 - \frac{MT}{2E_\nu^2}\right) F(Q^2)^2$$

$$\sigma \propto Q_W^2 \propto (N - (1 - 4\sin^2 \theta_W)Z)^2$$

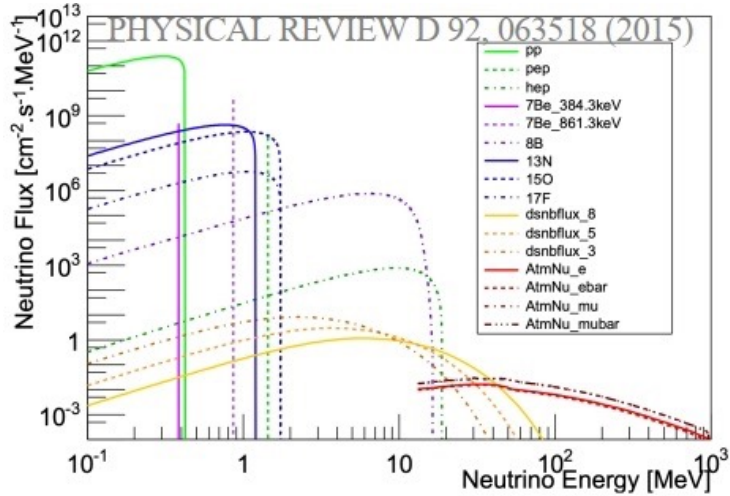
$$\implies \boxed{\sigma \propto N^2}$$

$$T_R \approx \frac{E_\nu^2 (\text{MeV})}{A} \text{ keV}$$



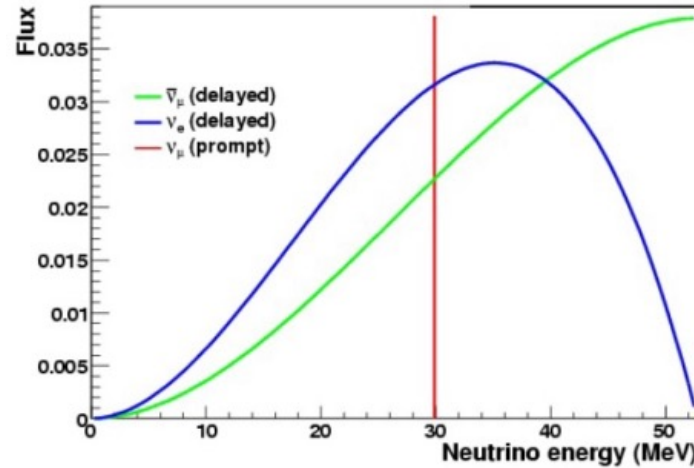
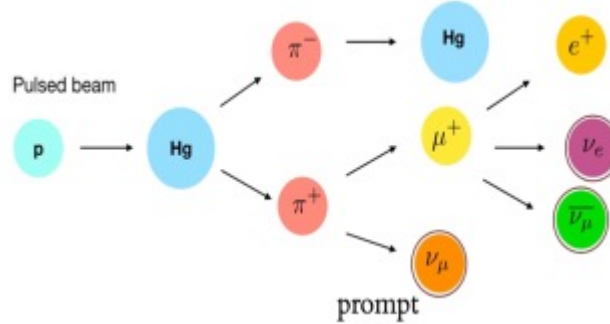
Neutrino sources for CEνNS detection

The Sun



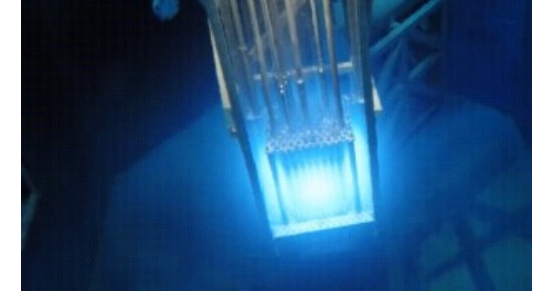
- $\sim 5 \times 10^6 / \text{s/cm}^2$ for ^8B neutrino
- Mostly $E_\nu < 10 \text{ MeV}$
- Not observed yet

Spallation Neutron Source

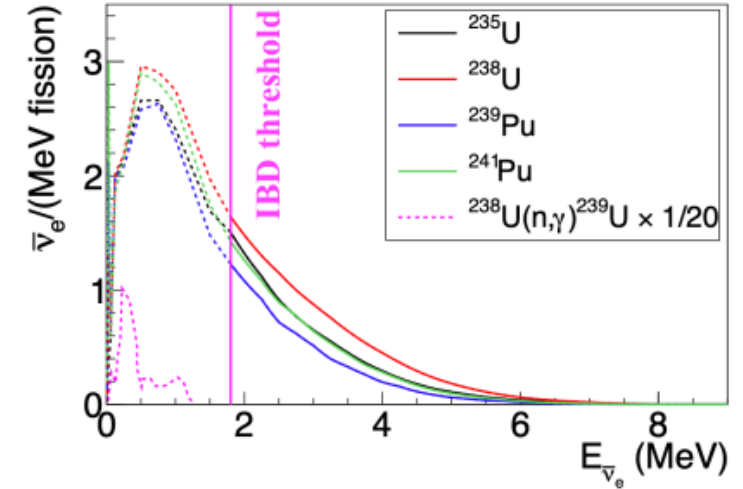


- $\sim 1 \times 10^7 / \text{s/cm}^2$ @ 20 m
- $E_\nu \sim 50 \text{ MeV}$
- Observed in 2017

Nuclear Reactor



Reactor Antineutrino Spectrum

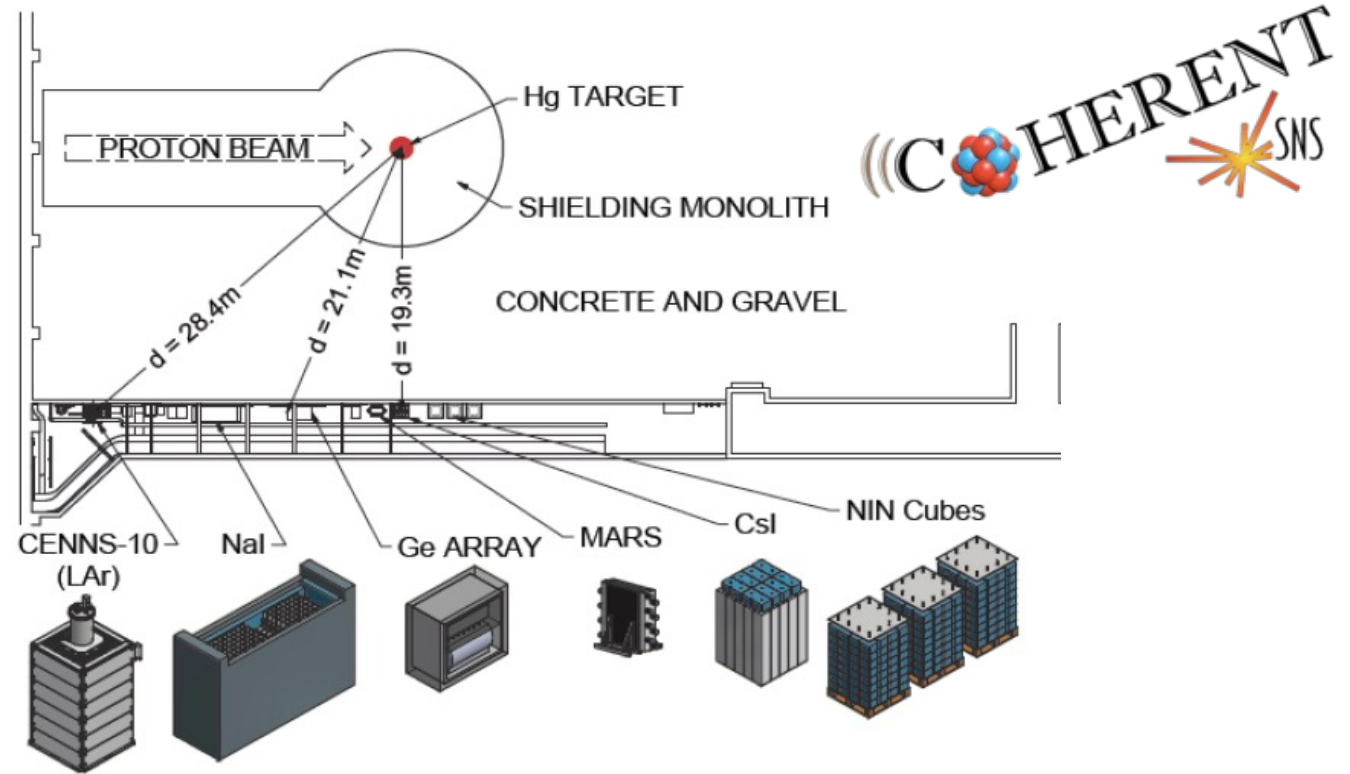


- $\sim 1 \times 10^{13} / \text{s/cm}^2$ @ 25m with 3 GW power
- Mostly $E_\nu < 6 \text{ MeV}$
- Not observed yet

First experimental observation by COHERENT



SCIENCE 357, no. 6356,
1123-1126 (2017)

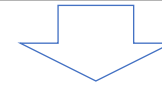
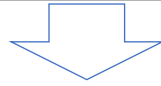


- 14.6 kg CsI(Na) detector
- 308.1 day live-time (beam-on)
- 134 ± 22 events observed
- 6.7σ observation

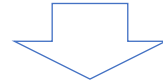
CICENNS experiment

Development of 300-kg CsI(Na) detector for CE ν NS

- Obtain sufficient CE ν NS events with substantial target mass
- Establish a new field of studying unexplored physics using CE ν NS



Detection of ~ 30 MeV neutrinos from pion/muon decays at rest
(at spallation neutron source)



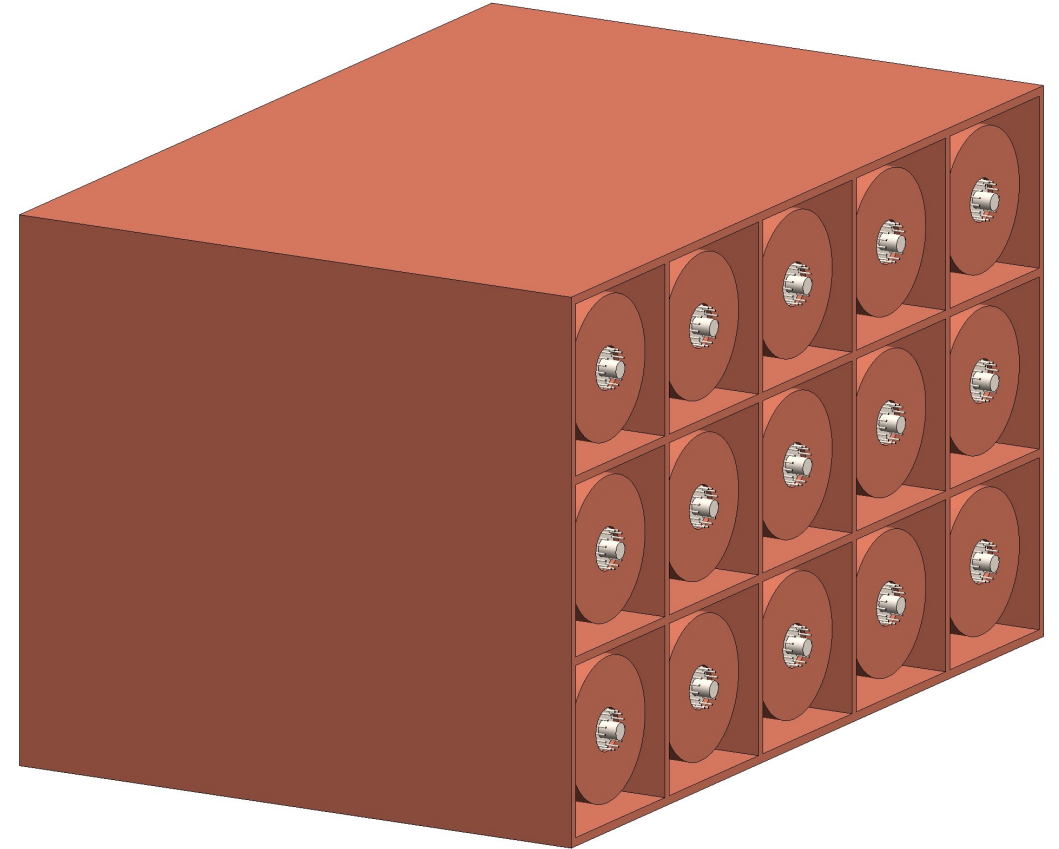
- **Precise measurements**
 - CE ν NS cross section \rightarrow weak couplings at low momentum transfer
 - Mean radius of neutron distribution inside nucleus
- **New physics searches**
 - Non-standard neutrino interactions
 - A new region of neutrino magnetic moment
 - Efficient search for sterile neutrino oscillation by neutral current
 - New particle searches: dark photons or axion-like particles

Drawings of detector and shielding



CsI(Na) crystal detector

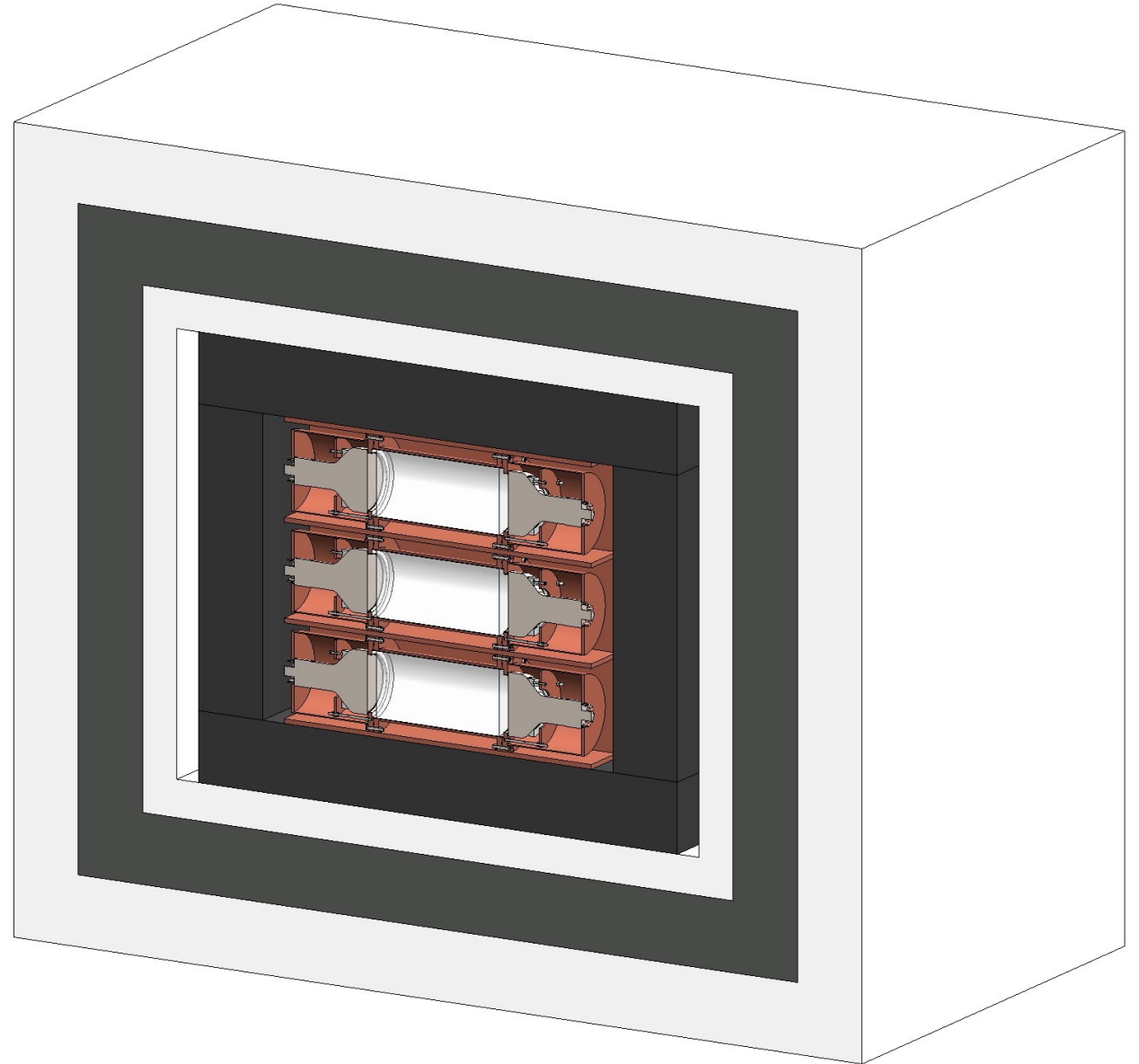
- 20-kg CsI(Na) × 15
- 140 mm (Dia.) × 287 mm (L) each
- Two 5-inch R877-100 PMTs (super-bialkali)
- OFHC copper encapsulation



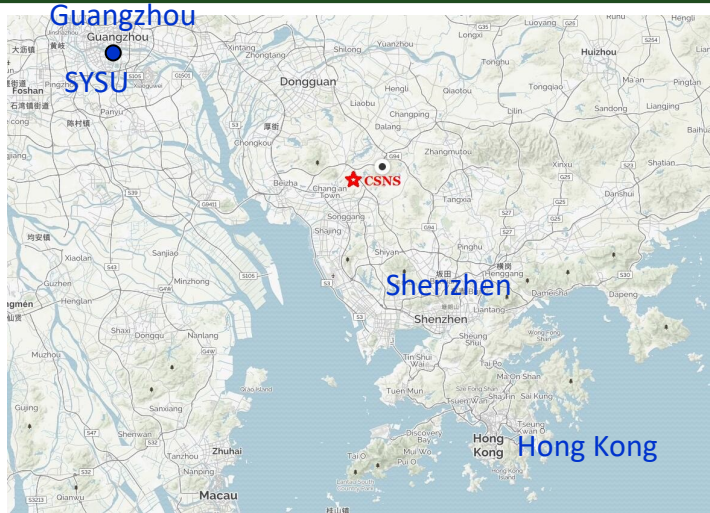
Drawings of shielding

Veto and shielding (from inside to outside)

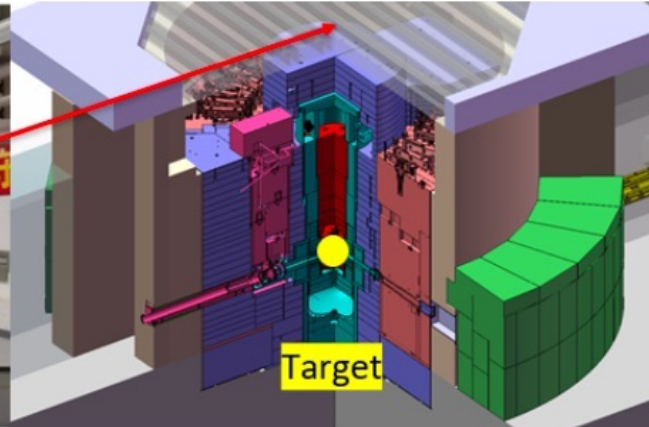
- 15-cm thick plastic scintillator as veto
- 8-cm thick inner HDPE
- 15-cm thick lead
- 15-cm thick outer HDPE



China Spallation Neutron Source (CSNS)



8.2 m above the beam target (7.2 m thick: steel, 1 m thick: concrete)



- 1.6 GeV proton beam
- 140 kW (→ 500 kW)
- 25 Hz repetition rate

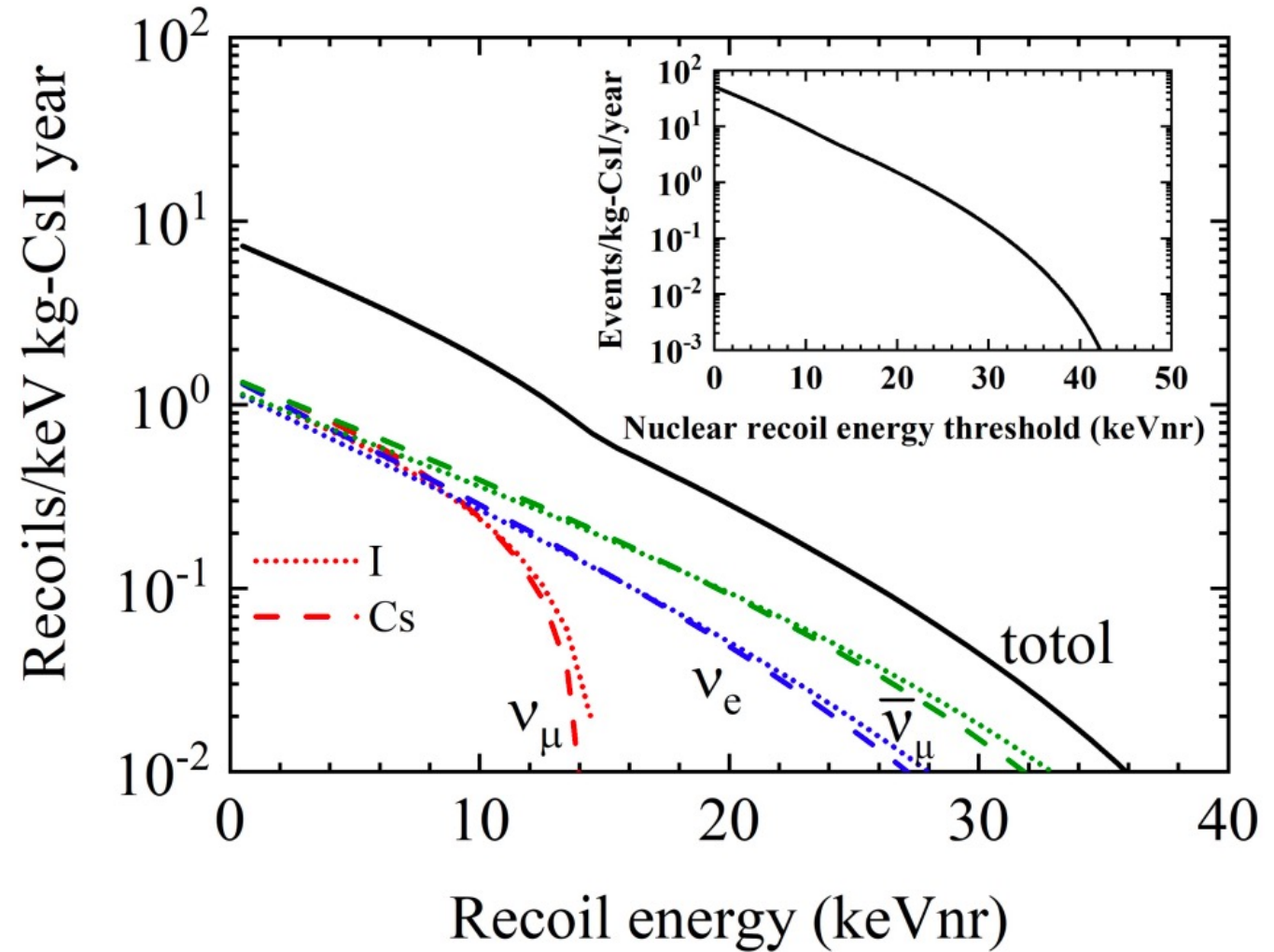
(Taken from C. Su, Q. Liu and T. Liang, arXiv:2303.13423 (2023))

Neutrino flux at 10.5 m: 2.0×10^7 /cm²/s (~40% of COHERENT)

→ ~2500 observed events with 300 kg (~300 events with 14.6 kg at COHERENT)

Expected CEvNS events of CICENNS at CSNS

- 14.6 kg CsI(Na) crystal for ~ 300 CEvNS events (PRL 129, 081801, 2022)
- 300 kg CsI crystals for $\sim 2,500$ CEvNS events (neutrino flux at CSNS: $\sim 40\%$ of ONL) if the energy threshold is $5 \text{ keV}_{\text{nr}}$
- for $\sim 4,000$ CEvNS events if the energy threshold can be lowered to $2 \text{ keV}_{\text{nr}}$



Backgrounds

- Beam related fast neutron
 - The most significant background
 - On-site measurement on-going, joint effort from SYSU and UCAS
- PMT dark count pile-up
- Radioactivity in detector materials
 - Crystal internal background from ^{87}Rb , ^{137}Cs , ...
 - Background from PMTs and surrounding materials
- Environmental radioactivity
- Cosmic muon induced neutrons and gammas

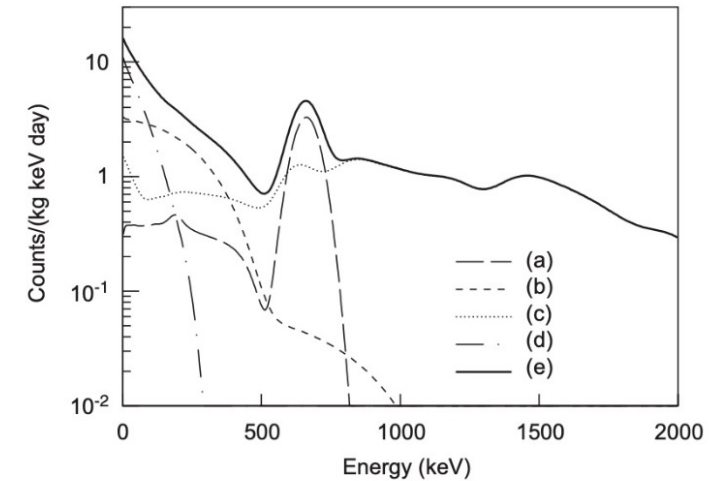
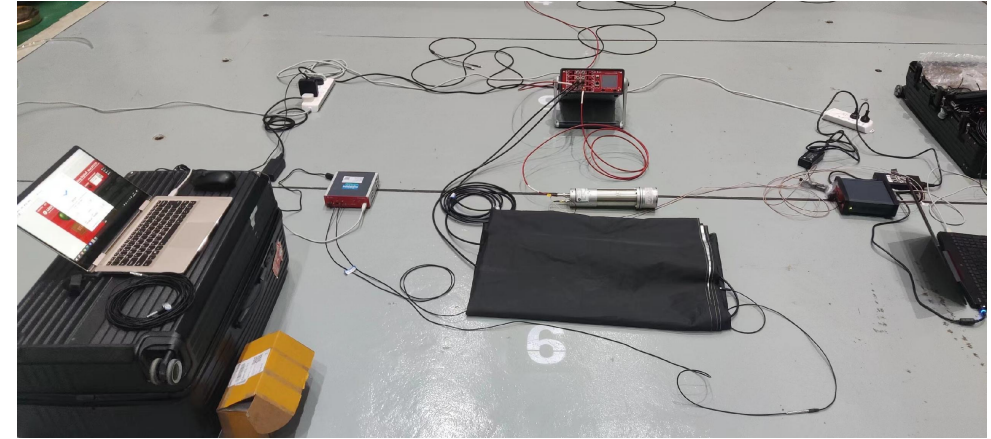


Figure 18: Background spectra obtained using GEANT4 simulation for a 20 kg CsI(Na) crystal with contaminations of 10 mBq/kg ^{137}Cs , 30 mBq/kg ^{134}Cs , and 10 ppb ^{87}Rb : (a) $^{137}\text{Ba}^*$; (b) ^{137}Cs ; (c) ^{134}Cs ; (d) ^{87}Rb ; and (e) total summed spectrum.

(Ref: KIMS Collaboration, Nucl. Instr. Meth. A 571, 644 (2007))

Full Geant4-based MC on-going

Key features of CICENNS design

- Lower internal radioactivity background
 - 20 ppb ^{87}Rb in COHERENT \rightarrow ~ 1 ppb ^{87}Rb achieved in CICENNS (collaboration with SICCAS)
- Lower dark count pile-up background
 - two PMTs on both ends of CsI(Na) crystal, ~ 3 orders of magnitude reduction (ref COSINE-100)
- Active veto
 - Plastic scintillator veto surrounding crystal detectors, mostly to veto beam related fast neutron, and also to identify cosmic muon track
- Better light collection of crystal encapsulation
- Multiple detector modules for further background rejection

Isotopes	^{238}U (ppt)	^{232}Th (ppt)	^{87}Rb (ppb)	^{137}Cs (mBq/kg).	^{134}Cs (mBq/kg).
CICENNS	10 ± 10	30 ± 10	~ 0.5	??	??
KIMS	0.75 ± 0.23	0.38 ± 0.07	1.3 ± 0.4	6.3 ± 0.7	14.1 ± 1.1
COHERENT	< 1000	< 1000	~ 20	$28 \pm 3.$	26 ± 2

Sensitivity: flavored CEvNS cross section and weak couplings

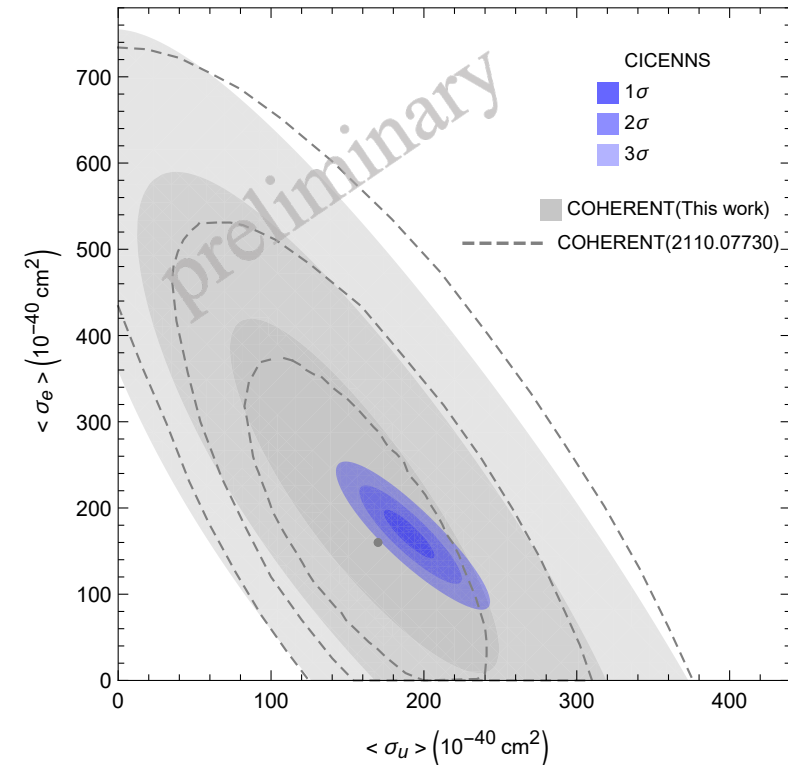
The CEvNS rate is a clean SM prediction

$$\frac{d\sigma}{dT} = \frac{G_F^2 M Q_W^2}{2\pi \cdot 4} F^2(Q) \left(2 - \frac{MT}{E_\nu^2} \right)$$

small nuclear uncertainties (±4.8%)

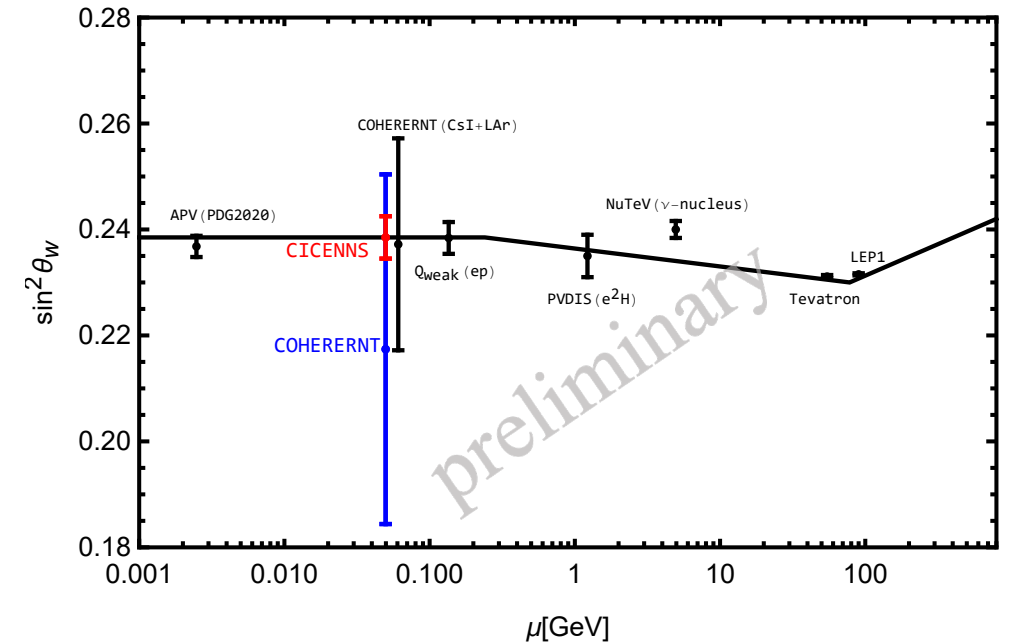
At momentum transfer $Q=50 \text{ MeV}/c$

- COHERENT: $\sin^2 \theta_W = 0.220_{-0.026}^{+0.028}$ ($\pm 10\%$)
- CICENNS: $\sim 1\%$
- * SM prediction: 0.23857(5)



Uncertainty of CEvNS cross section:

- COHERENT: +18/-15%
- CICENNS: 6.5%



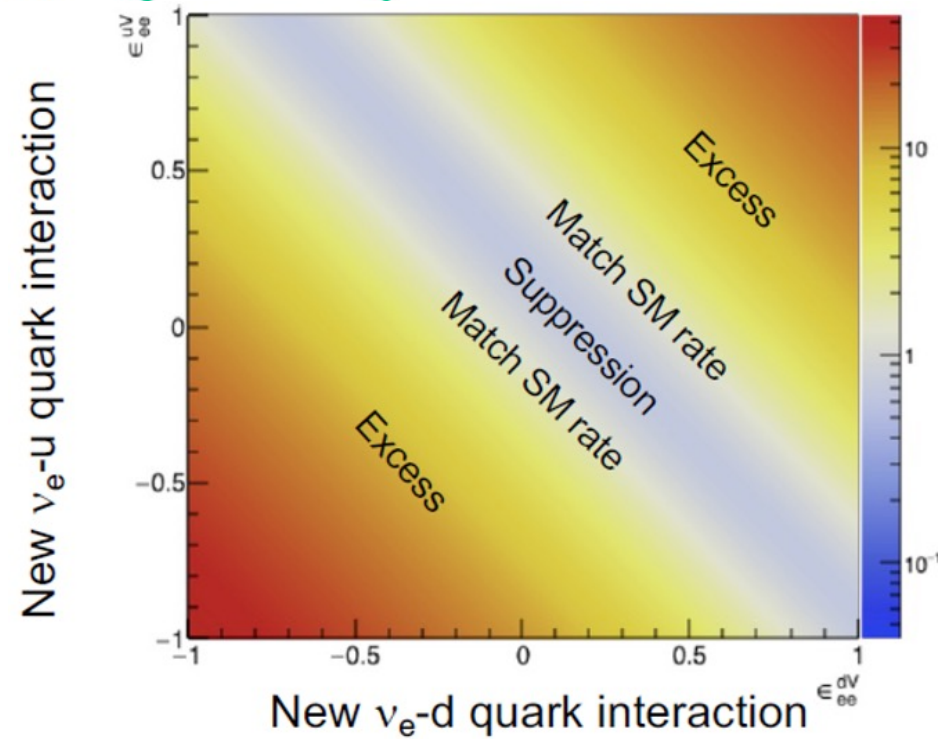
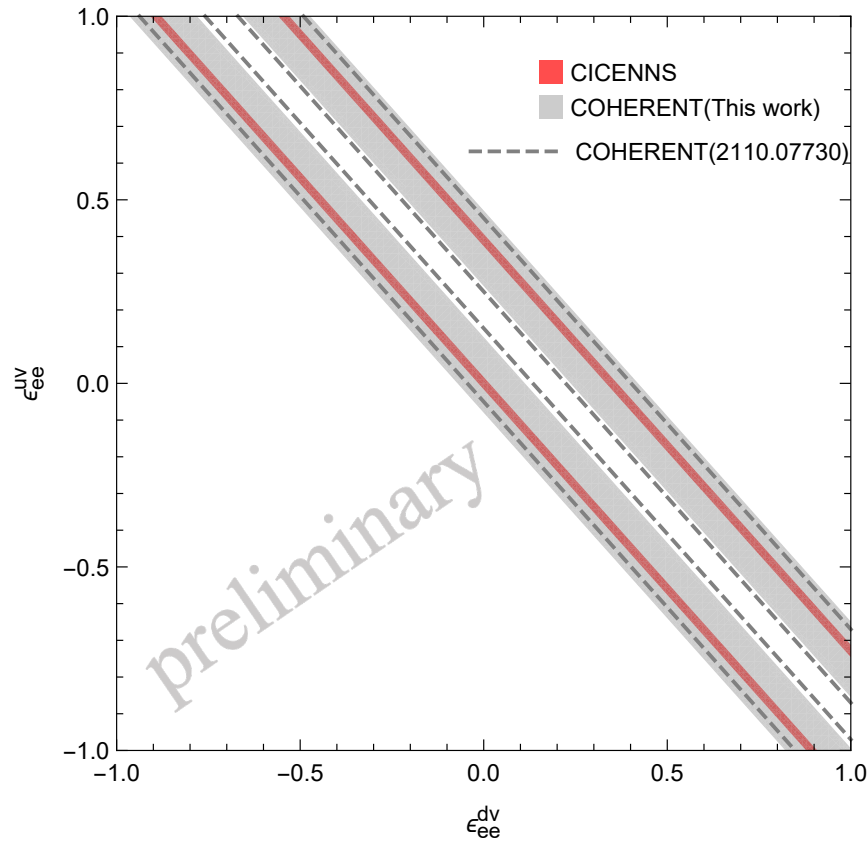
Sensitivity: Non-standard neutrino interaction

Look for a CEvNS **excess** or **deficit** wrt SM expectation

$$\mathcal{L}_{\nu H}^{NSI} = -\frac{G_F}{\sqrt{2}} \sum_{\substack{q=u,d \\ \alpha,\beta=e,\mu,\tau}} [\bar{\nu}_\alpha \gamma^\mu (1 - \gamma^5) \nu_\beta] \times (\varepsilon_{\alpha\beta}^{qL} [\bar{q} \gamma_\mu (1 - \gamma^5) q] + \varepsilon_{\alpha\beta}^{qR} [\bar{q} \gamma_\mu (1 + \gamma^5) q])$$

Csl

Ratio
wrt SM



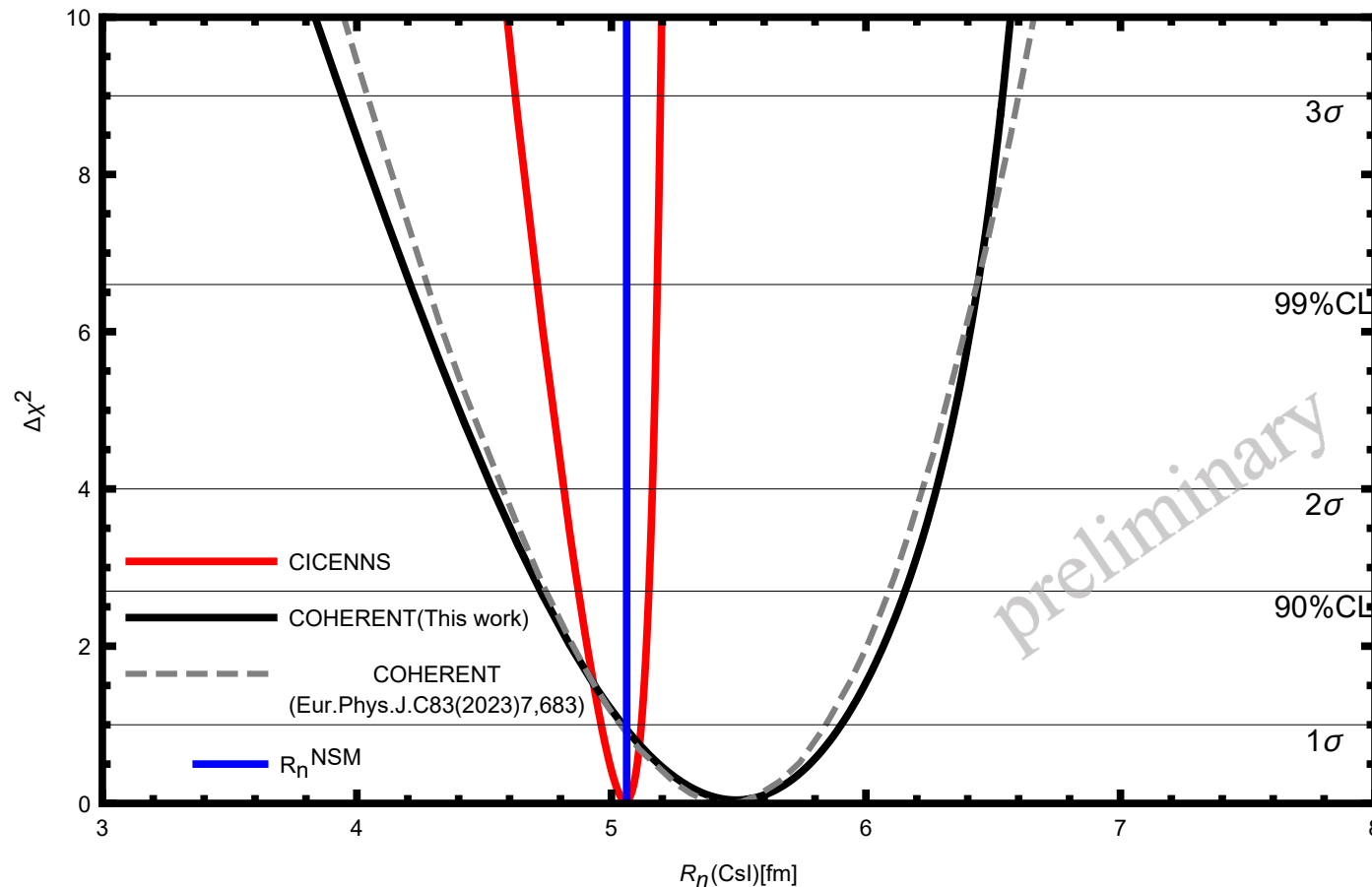
Sensitivity: mean radius of neutron distribution inside nucleus

Observed CEvNS recoil energy spectrum

→ mean radius of neutron distribution inside nucleus
(Cs or I: ~5 fm)

Uncertainty of neutron form factor

• COHERENT: ~8%
→ CICENNS: ~1%



Timeline

- Oct. 2022: CICENNS effort initiated at SYSU
- Dec. 2022: Procurement started
- Jan. 2023 – present
 - Physics design and engineer design
 - MOU signed between SYSU and SICCAS
 - R&D on CsI(Na) crystal purification
 - CsI(Na) mass production
 - PMT & electronics test
 - Full Geant4-based simulation of optical property and background
 - Technical Design Report in preparation
- End of 2024: Encapsulated CsI(Na) crystal detectors delivered to SYSU
- Mar. 2025: Completion of detector assembly
- Jun. 2025: Testing and commissioning at SYSU
- Jul. 2025: Start to deploy CICENNS at CSNS and taking data

Summary

- A 300-kg CsI(Na) crystal detector for CEνNS measurement at CSNS has been proposed, and is being built now.
- Precise test of SM at low momentum transfer and new physics search BSM are being pursued based on large statistics of CEνNS events.
- We wish to finalize the detector design and a TDR, and complete the detector construction in 2025.

We cordially invite your joining this effort!

Thank you!