#### 第三届地下和空间粒子物理与宇宙物理前沿问题研讨会

# 大统一理论在地下和空间实验探测中联合检验

#### 周也铃 杭高院 2024-5-8



#### 2024年5月7日-11日

#### 基础物理与数学科学学院

**School of Fundamental Physics and Mathematical Sciences** 





#### 地下实验探测: 大型中微子实验



#### JUNO 江门地下中微子观测站 20kt FV ~ $7 \times 10^{33}$ proton



DUNE, run in 2030?















### 空间实验探测:引力波观测



#### Pulsar-Timing Arrays (PTAs)











$$\nabla \cdot \mathbf{E} = 4\pi\rho$$
$$\nabla \cdot \mathbf{B} = 0$$
$$\nabla \times \mathbf{E} = -\frac{\partial B}{\partial t}$$
$$\nabla \times \mathbf{B} = 4\pi\mathbf{J} + \frac{\partial E}{\partial t}$$

James C. Maxwell

#### 强相互作用理论(QCD)

#### 电弱标准模型

#### $SU(2)_L \times U(1)_Y$



Glashow, Salam, Weinberg



### 大统一理论简介



# Grand 大 **U**nified 统一 Theory 理论 可能存在的 其他相互作用力 (GUT)



通往大统一之路



The scale where three gauge couplings are unified, denoted as  $M_{\rm GUT}$  in this talk





通往大统一之路

- Georgi-Glashow model (1974) SU(5) $\bigcirc$
- 24 规范波色子 +15 手征费米子 + 若干Higgs粒子 0



规范统—: non-SUSY ⇒ SUSY  $\bigcirc$ 



#### $-(u_R^2)^c$ $(d_R^1)^c$ $(u_R^3)^c$ $-d_{L}^{1}$ 0 $-(u_R^3)^c$ $-d_{L}^{2}$ $(u_R^1)^c$ $(d_R^2)^c$ $\begin{array}{cccc} -(d_R^1)^c & 0 & -u_L^3 \\ u_L^2 & u_L^3 & 0 \\ d_L^2 & d_L^3 & -(e_R)^c \end{array}$ $(d_R^3)^c$ $e_L$ $(d_R^2)^c$ $-d_L^3$ $\overline{\sqrt{2}}$ $u_L^1$ $(e_R)^c$ 0 $\psi_a \sim \bar{\mathbf{5}}$ $w^{ab} \sim 10$

Langacker, Luo, 91



通往大统一之路

More realistic SU(5)0

$$SU(5) \times U(1)_{B-L} \equiv G_{51}$$

• Flipped  $SU(5) \times U(1)_X \equiv G_{51}^{\text{flip}}$ 

Rujula, Georgi, Glashow (1980); Barr, (1982); Derendinger, Kim, Nanopoulos (1984); Antoniadis, Ellis, Hagelin, Nanopoulos (1989)

Pati-Salam 0

$$(1973), SU(4)_c \times SU(2)_L \times SU(2)_R \equiv G_{422} (4, 2, 1) : \psi_L = \begin{pmatrix} u^1 & u^2 & u^3 & \nu \\ d^1 & d^2 & d^3 & e \end{pmatrix}_L^{,} (\bar{4}, 1, 2) : \psi_R = \begin{pmatrix} u^1 & u^2 & u^3 & \nu \\ d^1 & d^2 & d^3 & e \end{pmatrix}_R^{c} \\ \text{JTs} & \text{Fritzsch, Minkowski (1975)} \\ \text{Ve can be} & \mathbf{16} = \bar{\mathbf{5}} + \mathbf{10} + \mathbf{1} = (\mathbf{4}, \mathbf{2}, \mathbf{1}) + (\bar{\mathbf{4}}, \mathbf{1}, \mathbf{2}) \\ \text{SO}(10) & SU(5) & SU(4)_c \times SU(2)_L \times SU(2)_R \end{aligned}$$

*SO*(10) Gl

All the above embedded 比如增加新粒子来解释中微子的质量

中微子质量通过B-L的自发破缺来获得

$$u \leftrightarrow d, \nu \leftrightarrow e$$



SO(10) 大统一的唯象学



Fermion masses and mixing

### Unwanted topological defects: monopoles and domain walls

In any breaking chains, inflation has to been introduced to inflate unwanted defects

 $G_{422} = SU(4)_C \times SU(2)_L \times SU(2)_R$  $G_{51} = SU(5) \times U(1)_X$  $G_{51}^{\text{flip}} = SU(5)_{\text{flip}} \times U(1)_{\text{flip}}$  $Z_2^C$ :  $\psi_L \leftrightarrow \psi_R^c$  $G_x = G_{421}$  or  $G_{3221}$  $G_{3221} = SU(3)_C \times SU(2)_L \times SU(2)_R \times U(1)_{R-L}$  $G_{421} = SU(4)_C \times SU(2)_L \times U(1)_V$  $G_{3211} = SU(3)_C \times SU(2)_L \times U(1)_R \times U(1)_{R-L}$  $G'_{3211} = SU(3)_C \times SU(2)_L \times U(1)_V \times U(1)_X$ 

#### $G_{\rm SM} = SU(3)_C \times SU(2)_L \times U(1)_Y$

King, Pascoli, Turner, YLZ, 2005.13549



质子衰变

#### ◎ 大统一引入传递重子数破坏的新粒子,这些粒子退耦后诱导出重子数破坏的有效算符

# Dim-6算符 主要通过规范波色子传导 $\frac{1}{\Lambda_1^2} \left[ (\overline{u_R^c} \gamma^\mu Q) (\overline{d_R^c} \gamma_\mu L) + (\overline{u_R^c} \gamma^\mu Q) (\overline{e_R^c} \gamma_\mu Q) \right] \\ + \frac{1}{\Lambda_2^2} \left[ (\overline{d_R^c} \gamma^\mu Q) (\overline{u_R^c} \gamma_\mu L) + (\overline{d_R^c} \gamma^\mu Q) (\overline{\nu_R^c} \gamma_\mu Q) \right]$ $\pi^0 e^+$ $\tau_{\pi^0 e^+} \sim 10^{35} \,\mathrm{years} \times \left(\frac{\Lambda_{\mathrm{pd}}}{10^{16} \,\mathrm{GeV}}\right)$ Dim-5算符 主要针对超对称大统一模型 $\frac{c_1}{M_T} (\overline{Q^c}Q)(\tilde{L}\tilde{Q}) + \frac{c_2}{M_T} (\overline{u_R^c}d_R)(\tilde{e}_R\tilde{u}_R) + \cdots$



 $\tau_{K^+\bar{\nu}} \propto (\Lambda_{\rm pd} \cdot \Lambda_{\rm SUSY} \cdot Y_{\rm Yukawa})^2$ 









#### $\simeq 2 \times 10^{10} \text{ g} \times (2 + 8)/18 \text{ mol/g} \times N_A$ /mol

≃ 7 × 10<sup>33</sup> 个质子

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#### JUNO的潜力





#### 下一代中微子实验的潜力



![](_page_12_Picture_2.jpeg)

### SO(10)中的质子衰变

![](_page_13_Figure_1.jpeg)

SO(10)	$\stackrel{\rm defect}{\longrightarrow}_{\rm Higgs}$	$G_2$	$\stackrel{\text{defect}}{\longrightarrow}_{\text{Higgs}}$	$G_1$	$\stackrel{\rm defect}{\longrightarrow}_{\rm Higgs}$	$G_{ m SM}$
II1:	$\xrightarrow{\mathrm{m}}$ <b>210</b>	$G_{422}$	$\xrightarrow{\mathrm{m}}$ <b>45</b>	$G_{3221}$	$\xrightarrow{s}{\overline{126}}$	
II2:	$\xrightarrow{\mathrm{m,s}}$ 54	$G^C_{422}$	$\xrightarrow{\mathrm{m}}$ 210	$G^C_{3221}$	$\xrightarrow{s,w}$	
II3:	$\xrightarrow{\mathrm{m,s}}54$	$G^C_{422}$	$\stackrel{\mathrm{m,w}}{\longrightarrow}$ 45	$G_{3221}$	$\xrightarrow{s}{\overline{126}}$	
II4:	$\xrightarrow{\mathrm{m,s}}210$	$G^{C}_{3221}$	$\xrightarrow[]{W}{45}$	$G_{3221}$	$\xrightarrow{s}{\overline{126}}$	
II5:	$\xrightarrow{\mathrm{m}}210$	$G_{422}$	$\xrightarrow{\mathrm{m}}$ <b>45</b>	$G_{421}$	$\xrightarrow{\frac{1}{s}}{\frac{1}{126}}$	
II6:	$\xrightarrow{\mathrm{m,s}}54$	$G^C_{422}$	$\xrightarrow{\mathrm{m}}$ <b>45</b>	$G_{421}$	$\xrightarrow{s}{\overline{126}}$	
II7:	$\stackrel{\mathrm{m,s}}{\longrightarrow}$ 54	$G^C_{422}$	$\stackrel{\mathrm{w}}{\longrightarrow}$ 210	$G_{422}$	$\xrightarrow{\text{m}}{126.45}$	
II8:	$\xrightarrow{\mathrm{m}}$ <b>45</b>	$G_{3221}$	$\xrightarrow{\mathrm{m}}$ <b>45</b>	$G_{3211}$	$\xrightarrow{s}{\overline{126}}$	
II9:	$\xrightarrow{\mathrm{m,s}}210$	$G^{C}_{3221}$	$\xrightarrow{\mathrm{m,w}}$ <b>45</b>	$G_{3211}$	$\xrightarrow{s}{\overline{126}}$	
II10:	$\xrightarrow{\mathrm{m}}$ 210	$G_{422}$	$\xrightarrow{\mathrm{m}}$ 210	$G_{3211}$	$\xrightarrow{s}{\overline{126}}$	
II11:	$\xrightarrow{\mathrm{m,s}}$ <b>54</b>	$G^C_{422}$	$\stackrel{\mathrm{m,w}}{\longrightarrow}$ 210	$G_{3211}$	$\xrightarrow{s}{126}$	
II12:	$\xrightarrow{\mathrm{m}}$ <b>45</b>	$G_{421}$	$\xrightarrow{\mathrm{m}}$ <b>45</b>	$G_{3211}$	$\xrightarrow{s}{126}$	

## King, Pascoli, Turner, **YLZ**, 2106.15634

![](_page_13_Picture_4.jpeg)

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#### SO(10) 中的费米子质量与味混合 所有物质场 { $Q = (u_L, u_R), u_R, d_R, L = (\nu_L, e_L), e_R, \nu_R$ } 都安排在SO(10)的**16**重态中 $\bigcirc$ $16 \times 16 = 10 + 126 + 120$

- •为了得到规范不变的可重整化的Yukawa耦合,三个Higgs场可以引进来 10<sub>H</sub>,  $\overline{126}_{H}$ , 120<sub>H</sub> 这其中,标准模型中的Higgs是这些Higgs混合出来的最轻的那一个
- ◎ 基本的Yukawa耦合项可以写为

由此得到的

$$Y = Y_{10}^* \mathbf{16} \cdot \mathbf{16} \cdot \mathbf{10}_H + Y_{126}^* \mathbf{16} \cdot \mathbf{16} \cdot \overline{\mathbf{126}}_H + Y_{120}^* \mathbf{16} \cdot \mathbf{16} \cdot \mathbf{120}_H + h.c.$$
  
中费米子场的Yukawa耦合相互关联 Dutta, Mimura, Mohapatra, 0412105  

$$Y_d = r_1(h + f + ih') \qquad Y_u = h + r_2 f + i r_3 h'$$
  

$$Y_e = r_1(h - 3f + i c_e h') \qquad Y_v = h - 3r_2 f + i c_v h'$$
  

$$M_{v_R} = m_{v_R} f$$
  
Yukawa/mass matrices in SO(10)  
× 3 matrices  $h \propto Y_{v_R} = f \propto Y_{\overline{v_R}} \qquad h' \propto Y_{\overline{v_R}}$ 

$$Y = Y_{10}^* \mathbf{16} \cdot \mathbf{16} \cdot \mathbf{10}_H + Y_{126}^* \mathbf{16} \cdot \mathbf{16} \cdot \overline{\mathbf{126}}_H + Y_{120}^* \mathbf{16} \cdot \mathbf{16} \cdot \mathbf{120}_H + h.c.$$
  
+ 费米子场的Yukawa耦合相互关联 Dutta, Mimura, Mohapatra, 0412105  

$$Y_d = r_1(h + f + ih') \qquad Y_u = h + r_2 f + i r_3 h'$$
  

$$Y_e = r_1(h - 3f + i c_e h') \qquad Y_v = h - 3r_2 f + i c_v h'$$
  

$$M_{v_R} = m_{v_R} f$$
  
Yukawa/mass matrices in SO(10)  
× 3 matrices  $h \propto Y_{e_R} = f \propto Y_{e_R}$ 

$$3 \times 3$$
 matrices  $h \propto Y_{10}$ 

![](_page_14_Picture_9.jpeg)

![](_page_14_Picture_10.jpeg)

#### 中微子实验对大统一味混合的影响

8		Normal Ore	dering (best fit)	Inverted Ordering $(\Delta \chi^2)$		
		bfp $\pm 1\sigma$	$3\sigma$ range	bfp $\pm 1\sigma$	$3\sigma$ r	
without SK atmospheric data	$\sin^2 \theta_{12}$	$0.304^{+0.013}_{-0.012}$	$0.269 \rightarrow 0.343$	$0.304^{+0.012}_{-0.012}$	0.269 -	
	$\theta_{12}/^{\circ}$	$33.44_{-0.74}^{+0.77}$	$31.27 \rightarrow 35.86$	$33.45_{-0.74}^{+0.77}$	31.27 -	
	$\sin^2 \theta_{23}$	$0.573\substack{+0.018\\-0.023}$	$0.405 \rightarrow 0.620$	$0.578^{+0.017}_{-0.021}$	0.410 -	
	$\theta_{23}/^{\circ}$	$49.2^{+1.0}_{-1.3}$	$39.5 \rightarrow 52.0$	$49.5^{+1.0}_{-1.2}$	39.8 -	
	$\sin^2 \theta_{13}$	$0.02220^{+0.00068}_{-0.00062}$	$0.02034 \rightarrow 0.02430$	$0.02238^{+0.00064}_{-0.00062}$	0.02053 -	
	$\theta_{13}/^{\circ}$	$8.57^{+0.13}_{-0.12}$	$8.20 \rightarrow 8.97$	$8.60^{+0.12}_{-0.12}$	8.24 -	
	$\delta_{ m CP}/^{\circ}$	$194^{+52}_{-25}$	$105 \to 405$	$287^{+27}_{-32}$	192 -	
	$\frac{\Delta m^2_{21}}{10^{-5} \ {\rm eV^2}}$	$7.42^{+0.21}_{-0.20}$	$6.82 \rightarrow 8.04$	$7.42^{+0.21}_{-0.20}$	6.82 -	
	$\frac{\Delta m^2_{3\ell}}{10^{-3}~{\rm eV}^2}$	$+2.515_{-0.028}^{+0.028}$	$+2.431 \rightarrow +2.599$	$-2.498^{+0.028}_{-0.029}$	-2.584 -	

![](_page_15_Figure_3.jpeg)

![](_page_15_Figure_4.jpeg)

![](_page_15_Picture_5.jpeg)

### SO(10) 中的费米子质量与味混合

![](_page_16_Figure_1.jpeg)

2 6 8

# Take-away message:

 $M_1 \sim (2, 5) \times 10^{11} \text{ GeV}$  $M_3 \sim (1, 3) \times 10^{13} \text{ GeV}$  $m_{\beta\beta} \sim 10^{-2} \text{ eV}$ 

中微子质量次序倾向于正等级

#### Fu, King, Marsili, Pascoli, Turner, YLZ, arXiv:2209.00021

![](_page_16_Picture_9.jpeg)

![](_page_16_Picture_10.jpeg)

#### SO(10) 中的费米子质量与味混合

![](_page_17_Figure_1.jpeg)

![](_page_17_Picture_2.jpeg)

⇒物质-反物质不对称

$$\eta_B^{\rm CMB} = 6.15 \times 10^{-10}$$

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大统一辐射引力波

#### 宇宙弦引力波

- 绝大部分大统一模型都包含一个 U(1)<sub>B-L</sub> 的规范对称性
- U(1)规范对称性的自发破缺可以产生稳定的宇宙弦
- 长弦通过intersection和intercommution(俗称"打结")产生弦圈圈
- ◎ 弦圈圈振荡,辐射出引力波

![](_page_18_Picture_6.jpeg)

Another mechanism: GW via GUT phase transition?

![](_page_18_Picture_9.jpeg)

Vanchurin, Olum, Vilenkin, 0511159

![](_page_18_Figure_11.jpeg)

![](_page_18_Picture_12.jpeg)

#### 宇宙弦引力波的典型频谱

![](_page_19_Figure_1.jpeg)

Plateau $\Omega_{\rm GW} h^2 \sim 5 \times 10^{-5} \sqrt{G \mu}$ 

 $\propto \frac{M_{B-L}}{M_{\text{Planck}}} \quad \text{in GUTs}$ 

 $G = M_{\text{Planck}}^{-2}$  万有引力常数

Blanco-Pillado, Olum 1709.02693

![](_page_19_Picture_6.jpeg)

### Gauge unification correlates GUT scale with intermediate scales

![](_page_20_Figure_1.jpeg)

 M<sub>i</sub> 是第 i 个中

 间对称性 G<sub>i</sub> 破

 缺对应的规范

 波色子的质量

![](_page_20_Picture_3.jpeg)

### Gauge unification correlates GUT scale with intermediate scales

![](_page_21_Figure_1.jpeg)

 M<sub>i</sub> 是第 i 个中

 间对称性 G<sub>i</sub> 破

 缺对应的规范

 波色子的质量

![](_page_21_Picture_3.jpeg)

![](_page_22_Figure_0.jpeg)

![](_page_22_Figure_1.jpeg)

![](_page_22_Picture_4.jpeg)

### SO(10)预言的宇宙弦引力波频谱

![](_page_23_Figure_1.jpeg)

![](_page_23_Figure_3.jpeg)

![](_page_23_Figure_4.jpeg)

#### 2023年 PTA 对引力波的观测结果

#### On 28 Jun 2023

![](_page_24_Picture_2.jpeg)

![](_page_24_Picture_4.jpeg)

The **NANOGrav** 15 yr Data Set: Evidence for a Gravitational-wave Background 2306.16213

The second data release from the **European Pulsar Timing Array** III. Search for gravitational wave signals 2306.16214

Search for an isotropic gravitational-wave background with the Parkes Pulsar Timing Array 2306.16215

Searching for the nano-Hertz stochastic gravitational wave background with the **Chinese Pulsar Timing Array** Data Release I 2306.16216

If cosmic GW background is observed, then what is the origin?

![](_page_24_Picture_10.jpeg)

![](_page_25_Figure_0.jpeg)

![](_page_26_Picture_0.jpeg)

大统一还会产生磁单极,如果产生磁单极的能标(往 往就是大统一能标本身) 跟宇宙弦能标比较接近时...

弦圈圈会衰变,数密度在宇宙后期被压低  $n(l,t)e^{-\Gamma_d l t}$ 

![](_page_26_Figure_3.jpeg)

![](_page_26_Figure_6.jpeg)

$$\Gamma_d = \frac{\mu}{2\pi} e^{-\pi\kappa}$$

![](_page_26_Figure_8.jpeg)

A GUT inflation separating the GUT breaking and B-L breaking in the time scale is required.

Antusch, Hinze, Saad, Steiner, 2307.04595

![](_page_26_Picture_12.jpeg)

![](_page_26_Picture_13.jpeg)

### SUSY SO(10)可提升B-L破缺能标

![](_page_27_Figure_1.jpeg)

#### Also enhances $p \to K^+ \bar{\nu}$

![](_page_27_Picture_3.jpeg)

### SUSY SO(10) 用于解释 NANOGrav15 的数据?

#### 可以, 但是更多的唯象学 限制需要考虑进来

![](_page_28_Figure_2.jpeg)

![](_page_28_Picture_4.jpeg)

![](_page_29_Figure_1.jpeg)

$$\begin{split} G_{51}^{\text{flip},\text{S}} &\equiv SU(5) \times U(1)_{\chi} \times \text{SUSY} \\ & (\mathbf{24},0) \ \Big| \ \text{broken at } M_{\text{GUT}} \\ G_{3211}^{\text{S}} &\equiv SU(3)_c \times SU(2)_L \times U(1)_y \times U(1)_{\chi} \times \text{SUSY} \\ & (\overline{\mathbf{10}}, -\frac{1}{2}) \supset (\mathbf{1}, \mathbf{1}, 1, -\frac{1}{2}) \ \Big| \ \text{broken at } M \\ G_{\text{MSSM}} &\equiv SU(3)_c \times SU(2)_L \times U(1)_Y \times \text{SUSY} , \\ & \downarrow \\ & G_{\text{SM}} \equiv SU(3)_c \times SU(2)_L \times U(1)_Y . \end{split}$$

![](_page_29_Picture_4.jpeg)

![](_page_29_Picture_5.jpeg)

![](_page_30_Figure_1.jpeg)

$$\begin{split} G_{51}^{\text{flip},\text{S}} &\equiv SU(5) \times U(1)_{\chi} \times \text{SUSY} \\ & (\mathbf{24},0) \ \Big| \ \text{broken at } M_{\text{GUT}} \\ \\ \mathcal{I}_{3211}^{\text{S}} &\equiv SU(3)_c \times SU(2)_L \times U(1)_y \times U(1)_{\chi} \times \text{SUSY} \\ & (\overline{\mathbf{10}}, -\frac{1}{2}) \supset (\mathbf{1}, \mathbf{1}, 1, -\frac{1}{2}) \ \Big| \ \text{broken at } N \\ & G_{\text{MSSM}} \equiv SU(3)_c \times SU(2)_L \times U(1)_Y \times \text{SUSY} , \\ & \downarrow \\ & G_{\text{SM}} \equiv SU(3)_c \times SU(2)_L \times U(1)_Y . \end{split}$$

![](_page_30_Picture_3.jpeg)

31

![](_page_31_Figure_1.jpeg)

$$\begin{split} G_{51}^{\text{flip},\text{S}} &\equiv SU(5) \times U(1)_{\chi} \times \text{SUSY} \\ (\mathbf{24},0) \qquad \downarrow \text{ broken at } M_{\text{GUT}} \\ \mathcal{I}_{3211}^{\text{S}} &\equiv SU(3)_c \times SU(2)_L \times U(1)_y \times U(1)_{\chi} \times \text{SUSY} \\ (\overline{\mathbf{10}}, -\frac{1}{2}) \supset (\mathbf{1}, \mathbf{1}, 1, -\frac{1}{2}) \qquad \downarrow \text{ broken at } M \\ G_{\text{MSSM}} &\equiv SU(3)_c \times SU(2)_L \times U(1)_Y \times \text{SUSY} , \\ &\downarrow \\ G_{\text{SM}} &\equiv SU(3)_c \times SU(2)_L \times U(1)_Y . \end{split}$$

![](_page_31_Picture_3.jpeg)

![](_page_31_Picture_4.jpeg)

![](_page_32_Figure_1.jpeg)

$$\begin{split} G_{51}^{\text{flip},\text{S}} &\equiv SU(5) \times U(1)_{\chi} \times \text{SUSY} \\ & (\mathbf{24},0) \ \Big| \ \text{broken at } M_{\text{GUT}} \\ \\ G_{3211}^{\text{S}} &\equiv SU(3)_c \times SU(2)_L \times U(1)_y \times U(1)_{\chi} \times \text{SUSY} \\ & (\overline{\mathbf{10}}, -\frac{1}{2}) \supset (\mathbf{1}, \mathbf{1}, 1, -\frac{1}{2}) \ \Big| \ \text{broken at } N \\ \\ G_{\text{MSSM}} &\equiv SU(3)_c \times SU(2)_L \times U(1)_Y \times \text{SUSY} , \\ & \downarrow \\ \\ G_{\text{SM}} &\equiv SU(3)_c \times SU(2)_L \times U(1)_Y . \end{split}$$

![](_page_32_Picture_3.jpeg)

![](_page_32_Picture_4.jpeg)

![](_page_33_Figure_1.jpeg)

$$\begin{split} G_{51}^{\text{flip,S}} &\equiv SU(5) \times U(1)_{\chi} \times \text{SUSY} \\ & (\mathbf{24},0) \ \Big| \ \text{broken at } M_{\text{GUT}} \\ \\ S_{3211} &\equiv SU(3)_c \times SU(2)_L \times U(1)_y \times U(1)_{\chi} \times \text{SUSY} \\ & (\overline{\mathbf{10}}, -\frac{1}{2}) \supset (\mathbf{1}, \mathbf{1}, 1, -\frac{1}{2}) \ \Big| \ \text{broken at } N \\ \\ G_{\text{MSSM}} &\equiv SU(3)_c \times SU(2)_L \times U(1)_Y \times \text{SUSY} , \\ & \downarrow \\ \\ G_{\text{SM}} &\equiv SU(3)_c \times SU(2)_L \times U(1)_Y . \end{split}$$

#### B-L能标可以很自然 得趋近于大统一能标

![](_page_33_Picture_4.jpeg)

![](_page_33_Picture_5.jpeg)

![](_page_34_Figure_0.jpeg)

![](_page_34_Figure_1.jpeg)

GWs in flipped SU(5) King, Leontaris, YLZ, 2311.11857

![](_page_34_Picture_4.jpeg)

![](_page_35_Figure_0.jpeg)

![](_page_35_Picture_3.jpeg)

![](_page_36_Figure_0.jpeg)

![](_page_36_Picture_1.jpeg)