



Jiangmen Underground Neutrino Observatory

Status and Prospects

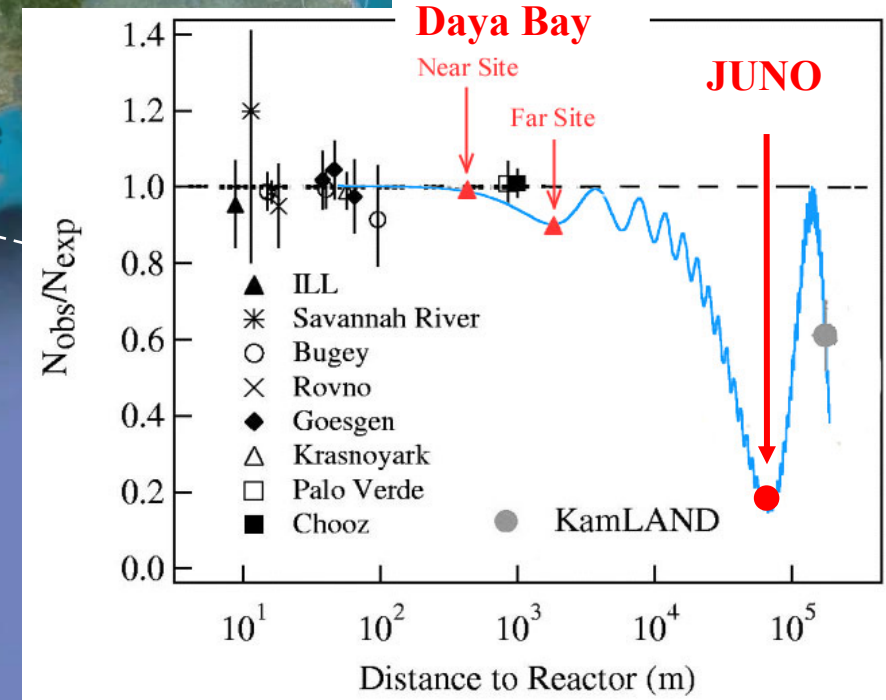
Xiaonan Li
IHEP, CAS, Beijing

- **Facility and Detector**
- **Installation status**
- **Performance and Physics Prospects**
- **Summary**



The Location of JUNO Facility

NPP	Daya Bay	Huizhou	Lufeng	Yangjiang	Taishan
Status	Operational	Planned	Planned	Operational	Operational
Power	17.4 GW	17.4	17.4 GW	17.4 GW	18.4 GW/2



J. Phys. G 43: 030401 (2016) (arXiv:1507.05613)

- Project firstly approved in China in 2013 and later in other countries. Construction started in January of 2015
- Collaboration established in 2014, now >700 collaborators from 74 institutions in 17 countries/regions

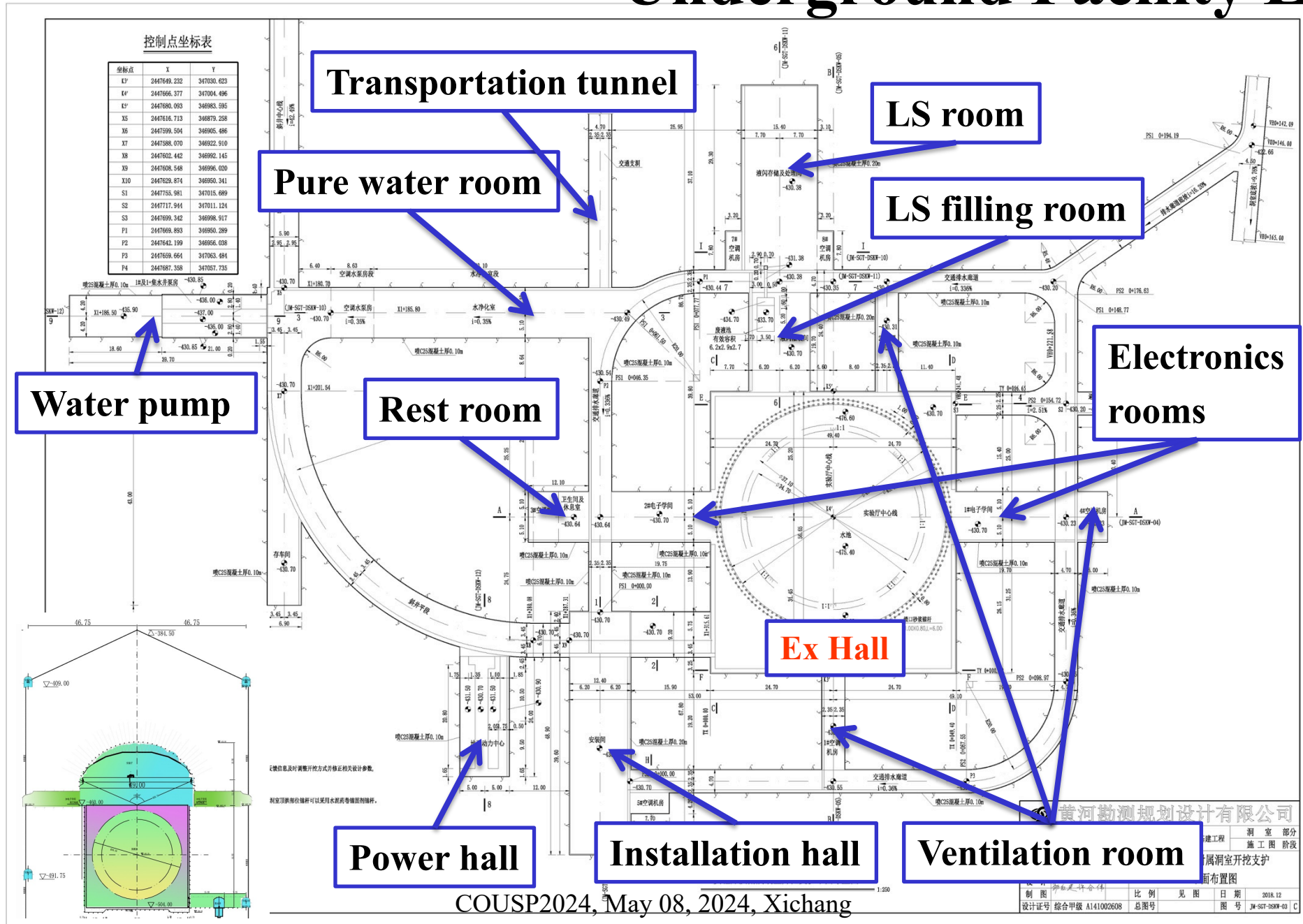


Facility and Detector



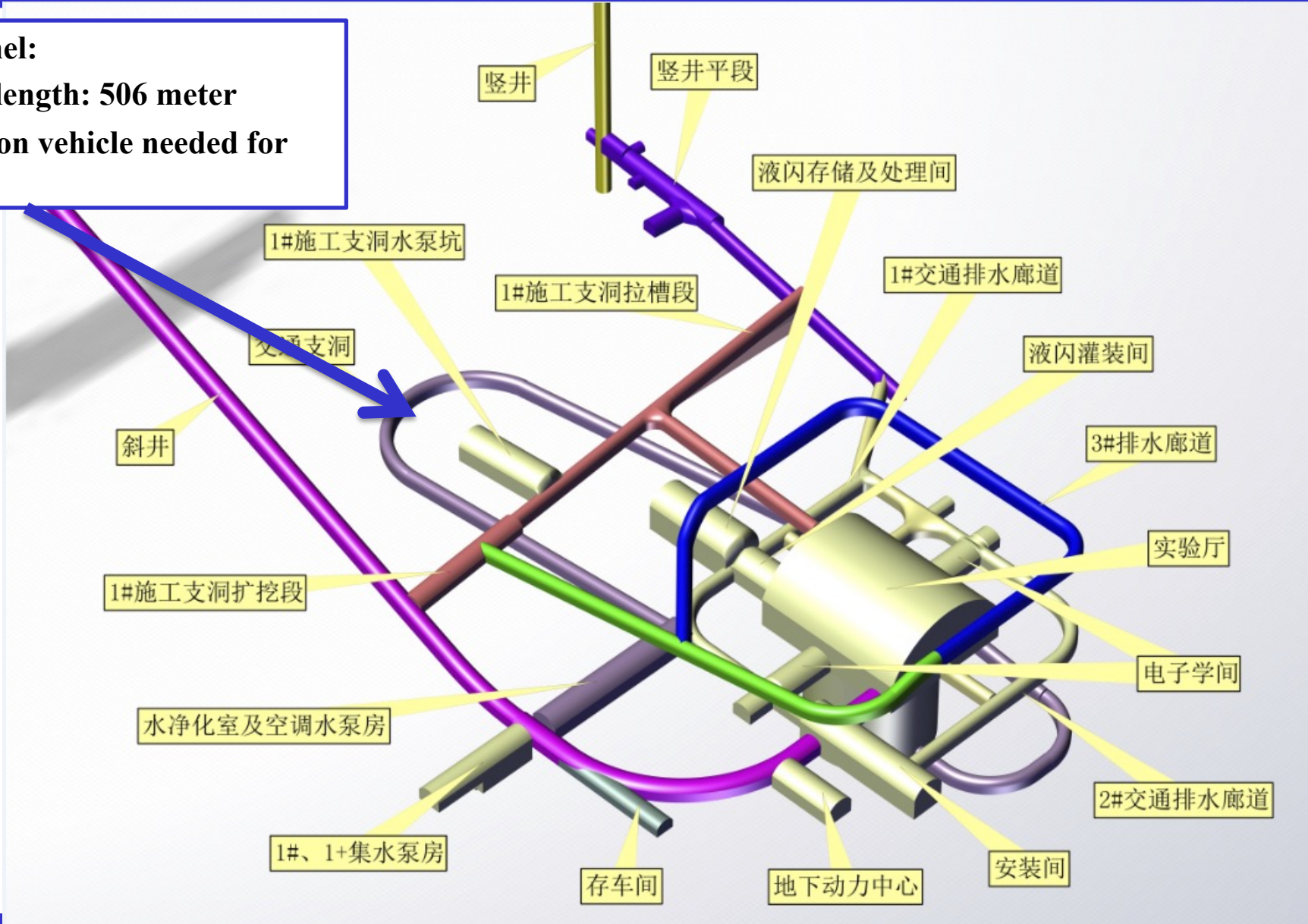


Underground Facility Layout



COUSP2024, May 08, 2024, Xichang

Transportation tunnel:
Slope: 3% @ 12%; length: 506 meter
Special transportation vehicle needed for installation



- Surface assembly building
- LAB storage (5kt)
- LS purification and mixing
- Water purification system
- Chilling water system
- Computing and power station
- Office, dorm and canteen



- **Dig down is generally difficult, especially in the presence of water and faults**
 - Largest water flow rate seen $\sim 600\text{m}^3/\text{h}$
- **More water than expected:**
 - Geological survey shows that the rock resistivity is low
 - Bore holes found no water, but iron ore with a low grade
 - Water are mostly from cracks and faults, not directly from the surface
- **Mitigations:**
 - Water drainage tunnel over and around exp. Hall
 - Drill holes to inject cement to seal water, and then drill holes for blasting
 - Elevate the hall by 35 m, shift the hall location to avoid cracks
 - Add a tunnel at the top of the hall to release the water pressure
- **Delayed by ~ 4 years**



Underground Transportation



For people, materials and equipment

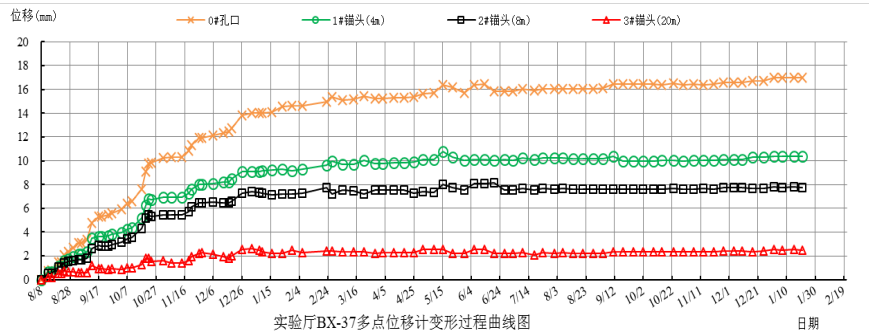
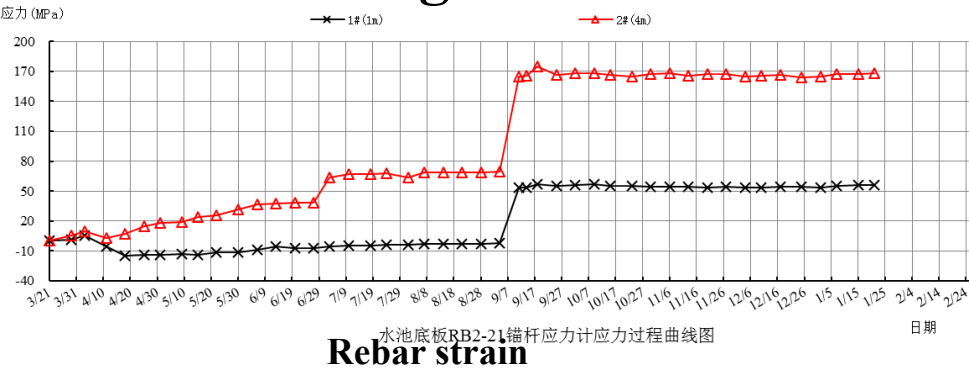
- ◆ Rock temperature ~ 31 °C, Exp. hall reduced and controlled to ~21°C
- ◆ VAC provided fresh air of 40 kM³/hour for underground cavern.



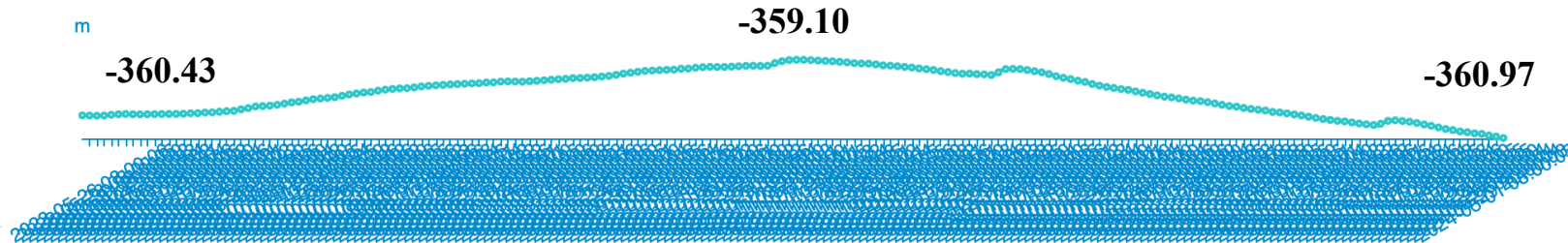
- Water pool HDPE(High Density PE) liner, thickness of 5mm



- **Stability monitoring of cavern structure**
 - Number of cavern stability monitoring sensors: 156
 - **Underground water level**



	Displacement	Rebar strain	Cable strain	Water pressure
Experimental hall	12	23	4	3
Water pool	21	58	8	19
Installation hall	3	3		



- Target mass of 20 kt liquid scintillator → × 20 KamLAND, × 40 Borexino
- Energy resolution < 3%@ 1MeV → 1200 PE/MeV
- Life time: >30 years

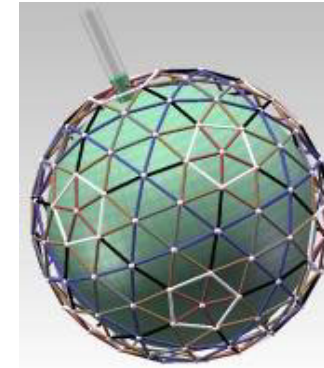
	KamLAND	JUNO	PE Ratio
Photon Statistics	250 p.e./MeV	1200 p.e./MeV	5
PMT coverage	34%	75%	2.2
LS transparency	~12 m	> 20 m	~0.9
Light yield(anthracene)	30%	45%	~1.5
Detection Eff.(QE×CE)	~15%	30%	~2

- **Technology R&D since 2009:**

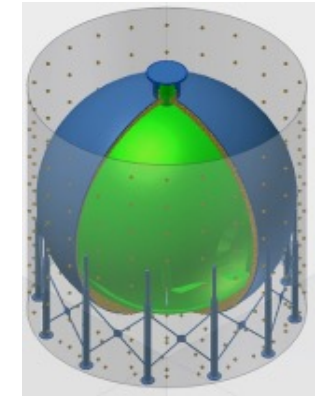
- Transparent & high light yield liquid scintillator
- High detection efficiency 20" PMTs
- Radiopurity U/Th/K < 10^{-17} g/g for 20 kt LS

- **Detector Design:**

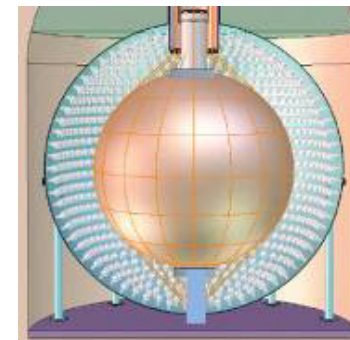
- Central target container: acrylic or balloon ?
- Mechanical structure: steel frame or steel tank ?
- Buffer layer: Water or Mineral oil ?



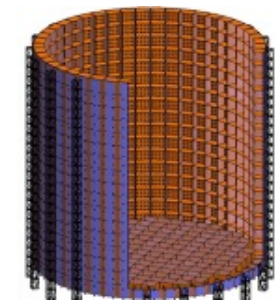
Steel frame+ Acrylic tank



Steel tank+ Acrylic tank



Steel Tank + Balloon

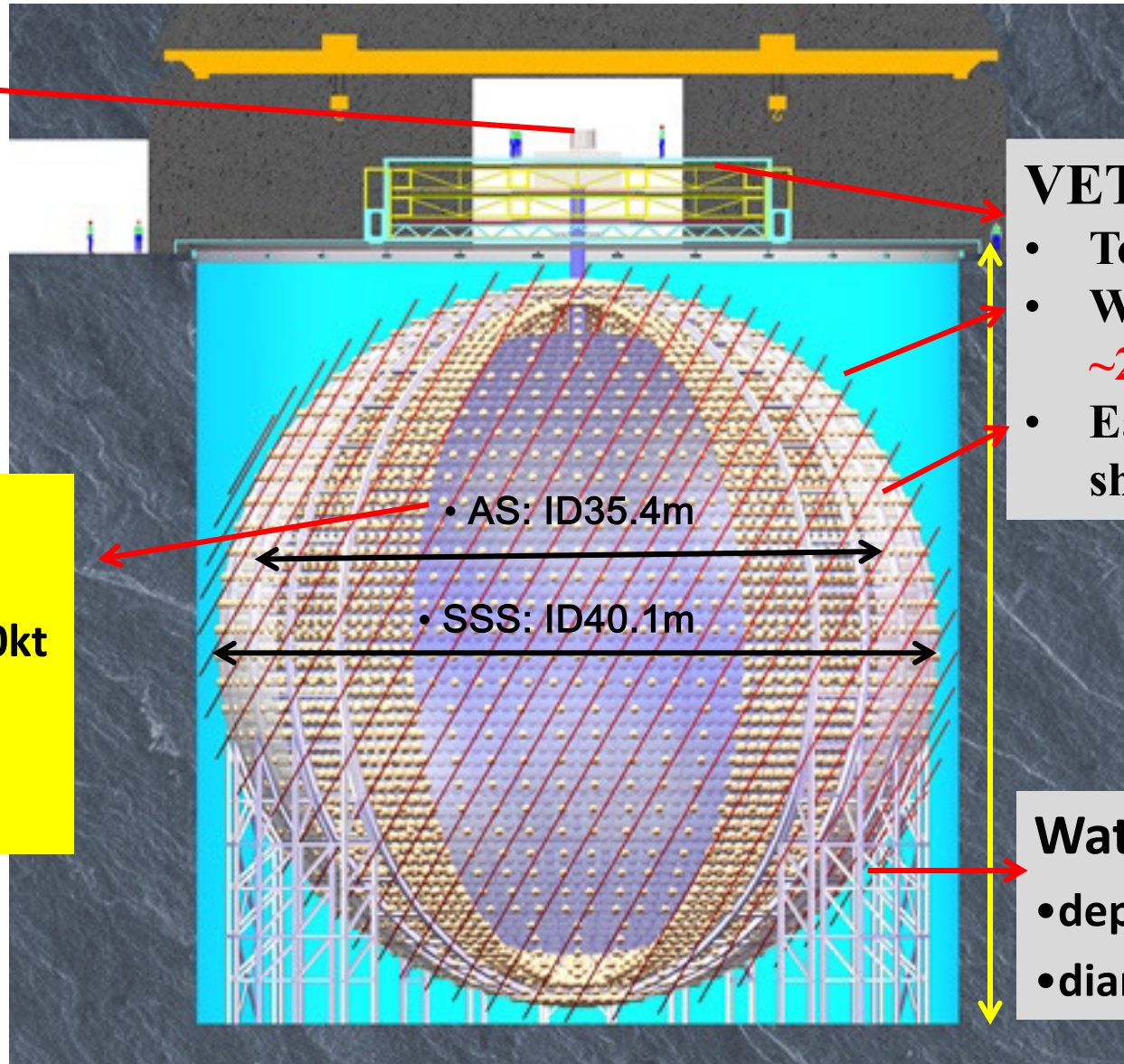


Steel Tank + Acrylic blocks

Calibration

Central detector

- Steel structure (SSS)
- Acrylic sphere (AS) +20kt Liquid scintillator
- ~17612 20" PMT
- ~25600 3" PMT



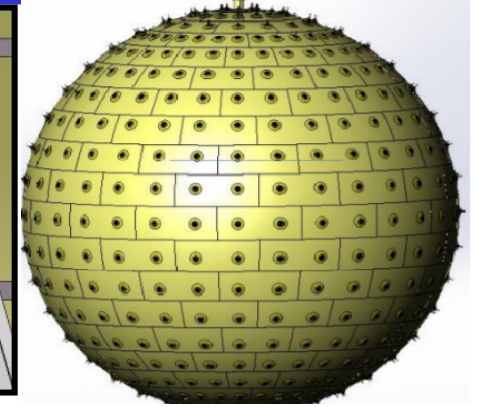
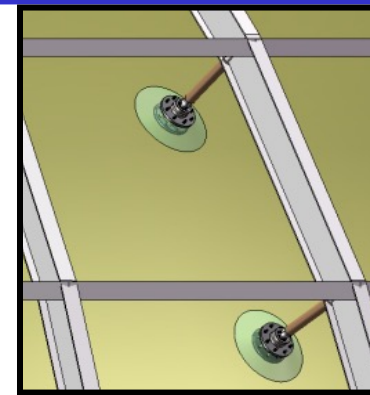
VETO system

- Top Tracker
- Water VETO with ~2400 20" PMT
- Earth Magnetic Field shielding coils

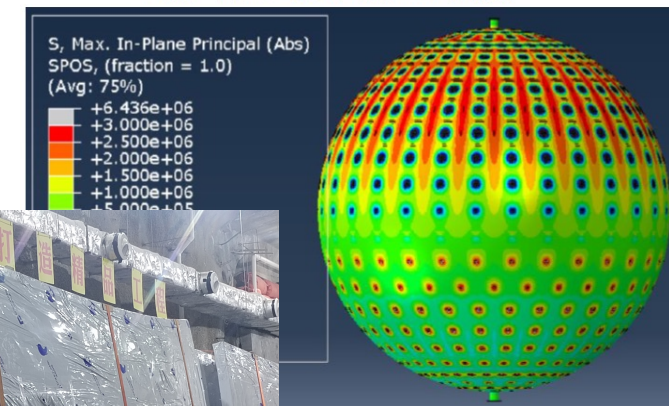
Water Pool

- depth: 44m
- diameter: 43.5m

- Acrylic sphere was chosen over balloon for safety, lower backgrounds and longevity
- Structure: SS frame with supporting bars to hold the acrylic sphere and to mount PMTs
 - 263 Acrylic panels: $\sim 8\text{m} \times 3\text{m} \times 12\text{ cm}$
 - Thermally pressed and machined to the spherical shape
 - Panels are bonded through PMMA polymerization
 - 1.7m long steel anchor bolts welded to the concrete steel structure to hold the SS frame against the buoyancy
- Main issues:
 - Mechanical precision for PMT clearance $\sim 3\text{ mm}$
 - No welding, clearance between screw and hole $< 1\text{mm}$
 - Thermal expansion controlled at $21^\circ\text{C} \pm 1^\circ\text{C}$
 - Earth quake safety with liquid-solid coupling, level 7
 - Acrylic quality and strength:
 - ◆ Effects of aging, creep, crazing, etc. $< 20\%$ loss of strength
 - ◆ Bonding: fast and good quality, effects $< 20\%$ loss of strength
 - ◆ Stress of the supporting rod $< 3.5\text{ MPa}$



工况：自重+水压（内4.5m高度）

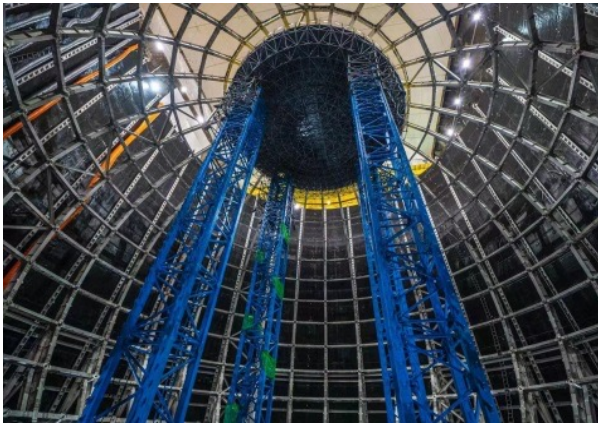
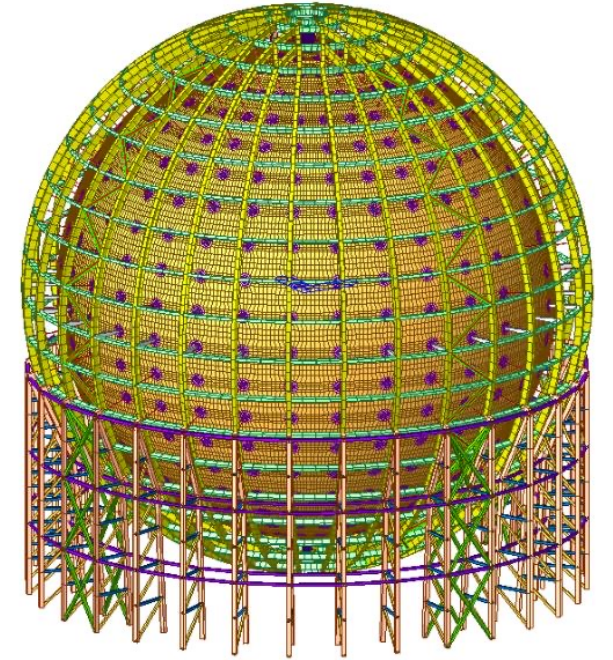


应力情况

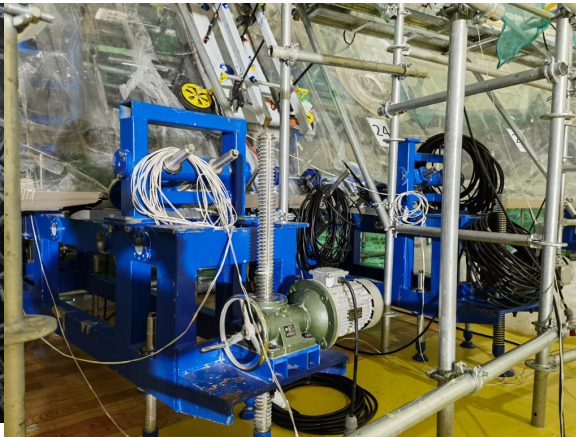
arXiv: 2311.17314



- SS structure completed except bottom 4 layers
- Acrylic panel production completed
 - A special production line for low backgrounds (< 1 ppt U/Th/K)
 - Shaping, sanding/polishing, cleaning, machining and protection of panels by $50\mu\text{m}$ PE, while maintaining high transparency ($>96\%$) and low surface background (< 5 ppt U/Th)
- Acrylic sphere construction on-going
 - Lifting platform: frequently change the diameter and height
 - Acrylic sphere built from the top, 16/23 layers finished, defects repaired
 - SS bars connecting the acrylic and SS, sensors for stress monitoring



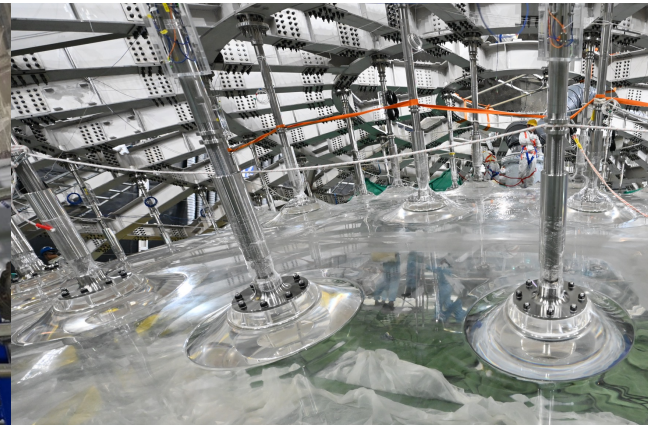
Hydraulic Lifting platform



Annealing



Sanding and polishing



Bars between SS and acrylic

VETO: Water Pool and Top Tracker



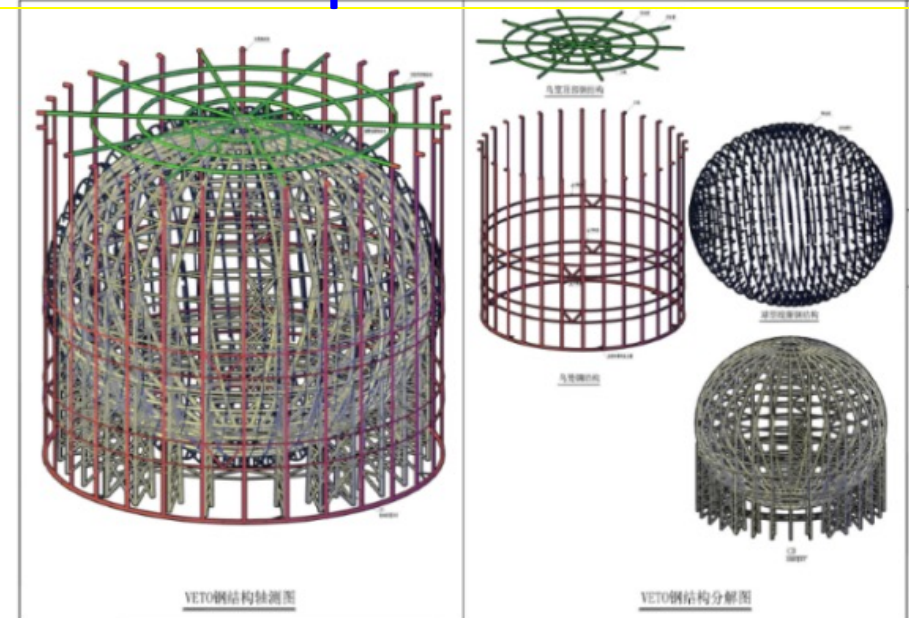
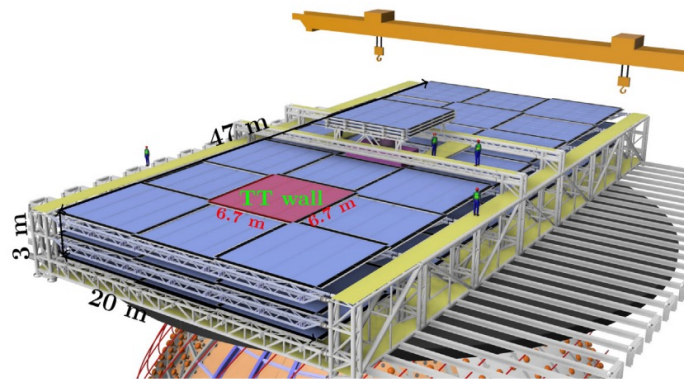
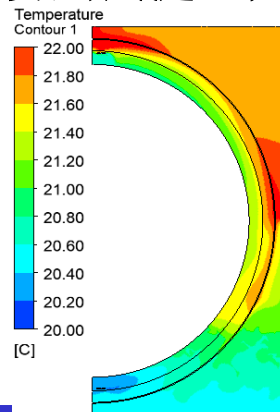
- ◆ **Design:**
 - ⇒ 700m overburden → $R_\mu = 0.004 \text{ Hz/m}^2$, $\langle E_\mu \rangle = 207 \text{ GeV}$
 - ⇒ 35 kt water to shield backgrounds from the rock
 - ⇒ 2400 20" PMTs in water to tag & reconstruct cosmic muons
 - ⇒ Top tracker: refurbished OPERA scintillators
- ◆ **Water pool lining: 5mm HDPE to keep the clean water and to stop Rn from the rock**
- ◆ **Earth magnetic field compensation coil**

100t/h pure water production for U/Th/K $< 10^{-14} \text{ g/g}$ and Rn $< 10 \text{ mBq/m}^3$, attenuation length $> 40\text{m}$, temperature controlled to $(21 \pm 1) \text{ }^\circ\text{C}$

报告：郭聪，江门中微子实验低氡超纯水研制



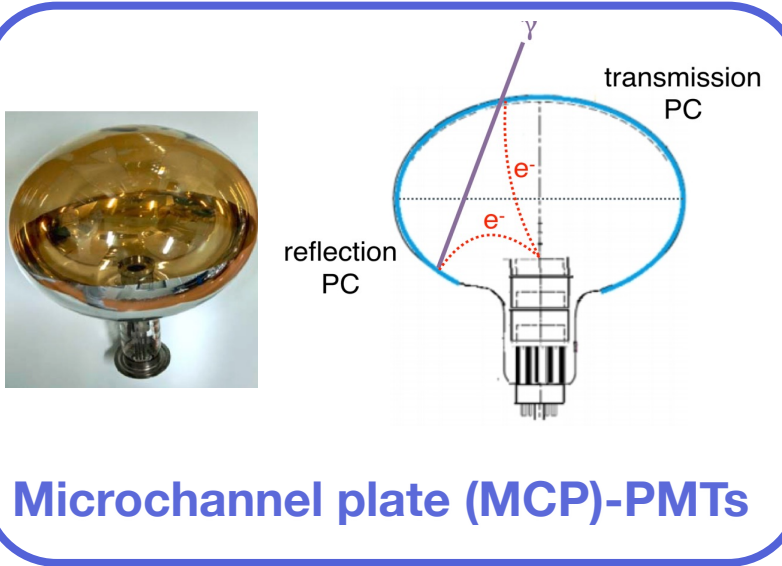
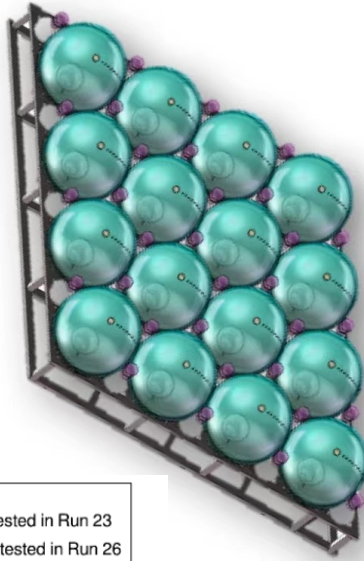
•Ultrapure water system



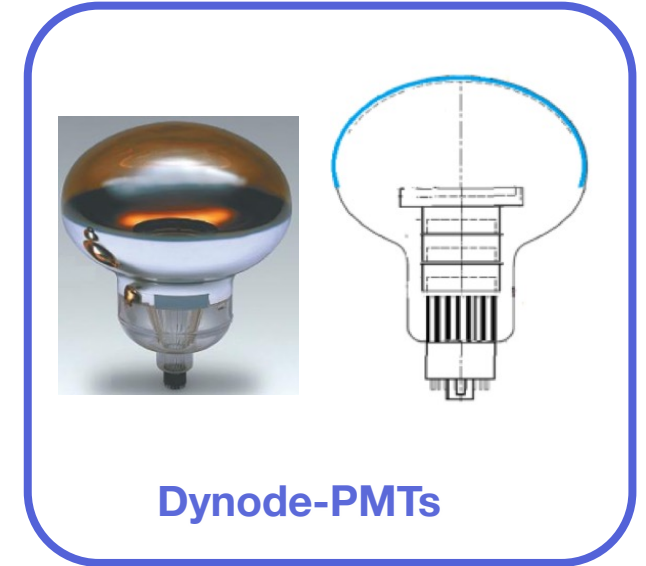
- uses large 20-inch PMTs as its main light-detection device:

- MCP-PMT(NNVT): 15000
- Dynode PMT(Hamamatsu): 5000

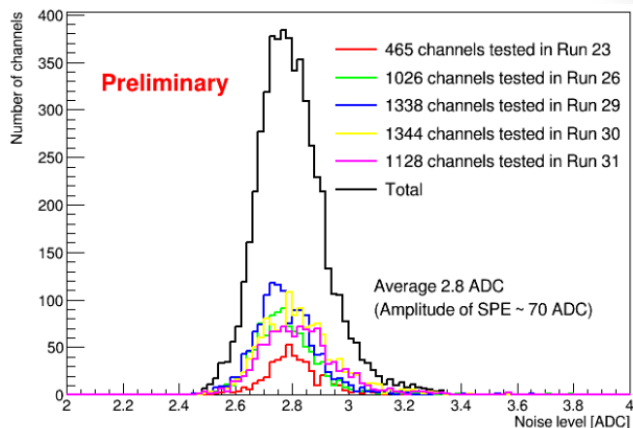
clearance of 3 mm
Coverage: 75%



Microchannel plate (MCP)-PMTs

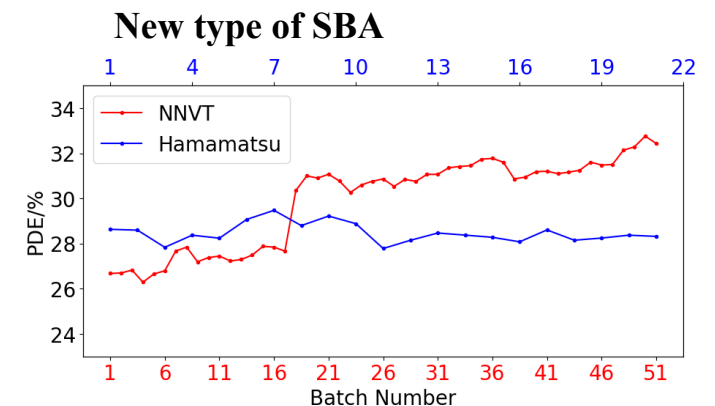


Dynode-PMTs



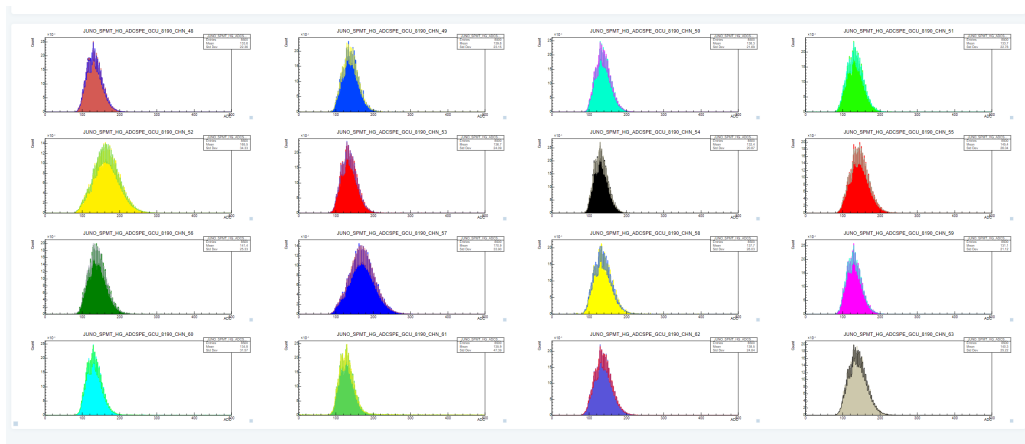
Noise level is ~ 0.05 PE: good
grounding and shielding

Higher collection efficiency
High efficiency photocathode, low backgrounds glass, ...
High production yield, automatic mass production

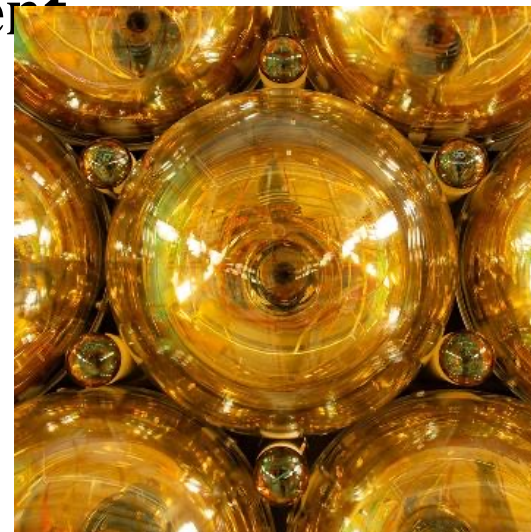


Both reach QE x CE ~ 30%!

- Mass production and waterproofing of 26,000 3-inch PMTs (XP72B22) from HZC Photonics
- Goal: 3% more light, higher dynamic range for muons, uniformity and linearity calibration for large PMTs, ...
- Two eye system look at same event with different systematics
 - Aid to position reconstruction and muon track reconstruction
 - Aid to supernova neutrino measurement



First commissioning data: SPE spectra of a group of 16 PMTs



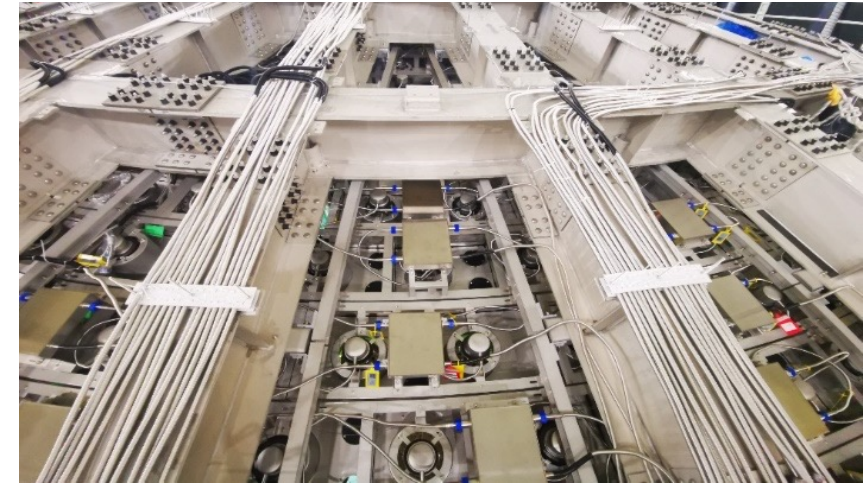
NIM A 1005 (2021) 165347

XP72B22



A custom design for JUNO!

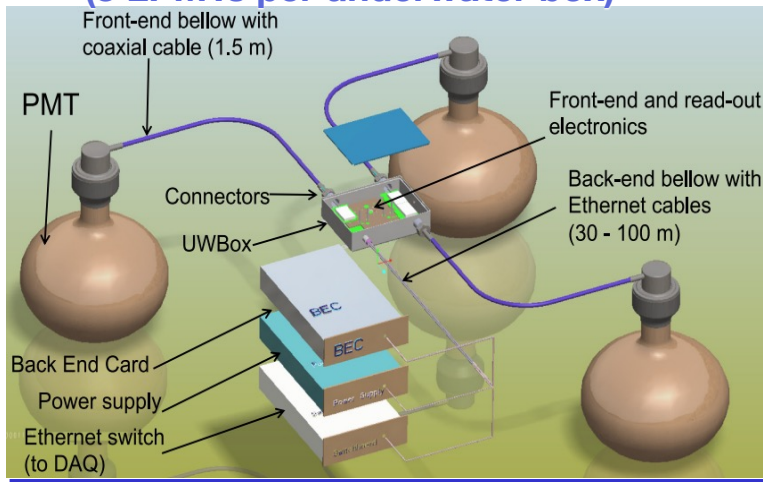
- **LPMT and sPMT frontend electronics are housed in SS underwater boxes**
 - **Power the PMTs, decouple signal from HV, process the analog signals and send the information to surface via digital CAT5 cables**
- **Cables are in the SS corrugated pipes**



胡俊： Design and production of readout electronics for the 20-inch PMT

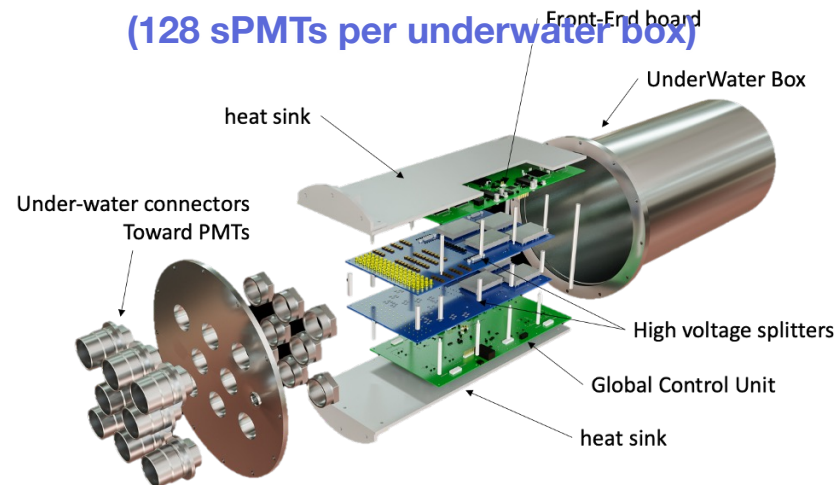
LPMT System

(3 LPMTs per underwater box)



sPMT System

(128 sPMTs per underwater box)

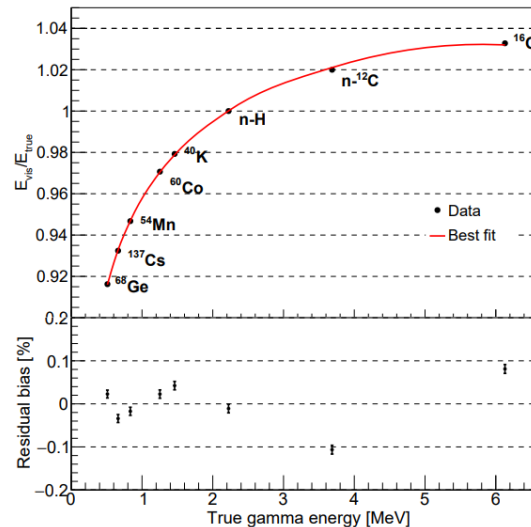


Power Modules, BEC and others

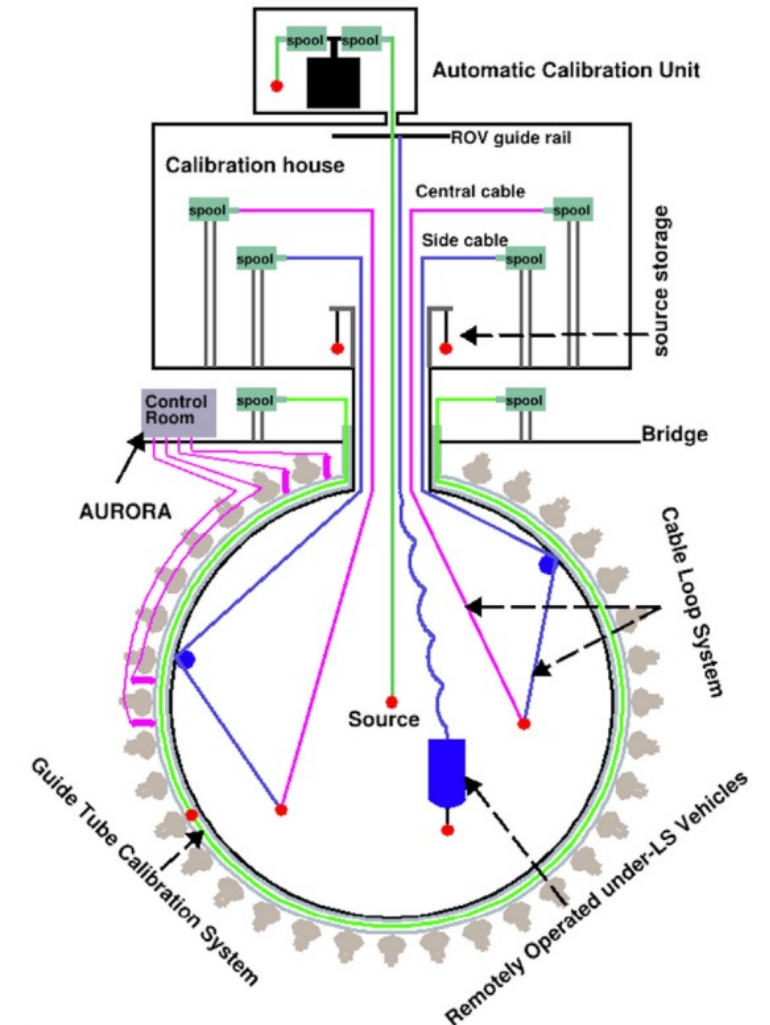


- Have an comprehensive calibration program:
 - 1D: Automated Calibration Unit (ACU) deploys radioactive and laser (1ns, keV-TeV range) sources along the central axis
 - 2D: Cable Loop System (CLS) to scan vertical planes
 - 2D: Guide Tube to scan the outer surface of the central detector (where the CLS cannot reach)
 - 3D: Remotely Operated Vehicle (ROV) operating inside the LS to scan the full volume

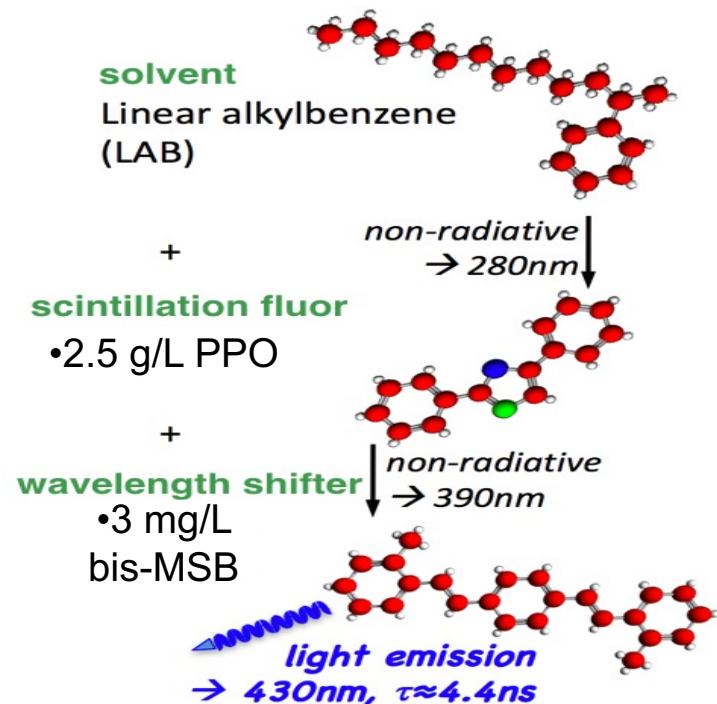
Calibrate energy scale and non-linearity to better than 1% using γ peaks and cosmogenic ^{12}B beta spectrum



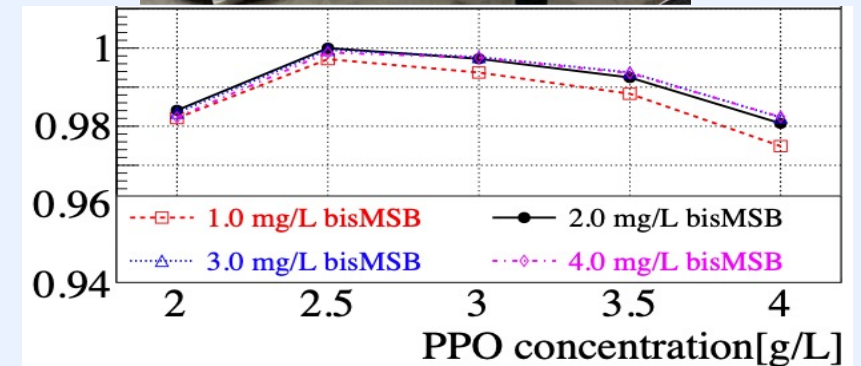
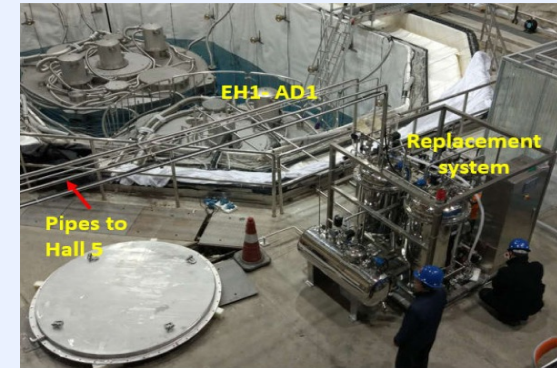
[JHEP 2021, 04 \(2021\)](#)



- **Main requirements: high light yield, high transparency and high radiopurity**
- **Recipe: Based on Daya Bay experience with dedicated optimization for higher light yield.**
 - LAB + 2.5g/L PPO + 3 mg/L bis-MSB
- **No dope**



Recipe optimized with a decommissioned Daya Bay detector whose results were extrapolated to JUNO size using a new optical model



• [NIMA 998, 164823](#)
(2021)

Recent measurements: LAB attenuation length > 24 m, LS attenuation length > 20 m

LS Production and Purification

- Four purification plants + LS Mixing + QA/QC + high purity N₂ and water production plant
- to guarantee radio-purity and transparency

俞伯祥: High pure nitrogen plant of JUNO
 孙希磊, 江门中微子实验的超纯液闪



•5000 m³ LAB storage tank



•1) Al₂O₃ for optical transparency

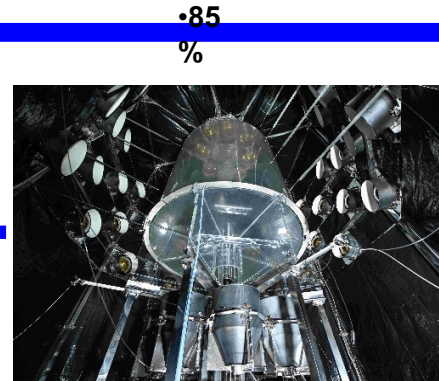
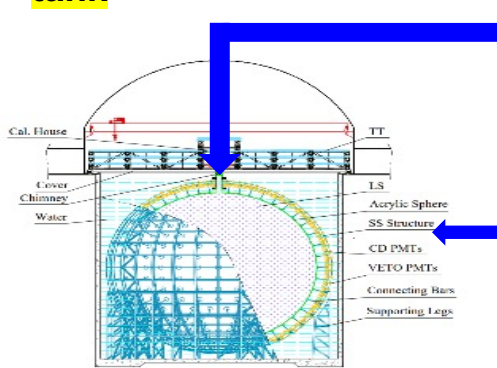


•2) Distillation for radiopurity



•Mixing LAB with PPO and bis-MSB

Aboveground



•OSIRIS to monitor the LS quality



•4) Gas stripping to remove Rn and O₂



•3) Water extraction to remove radioactive impurities

•1800 m SS pipes to underground

Underground

• Full chain of equipment were successfully commissioning and ready for final filling



Radiopurity and Cleanness Control

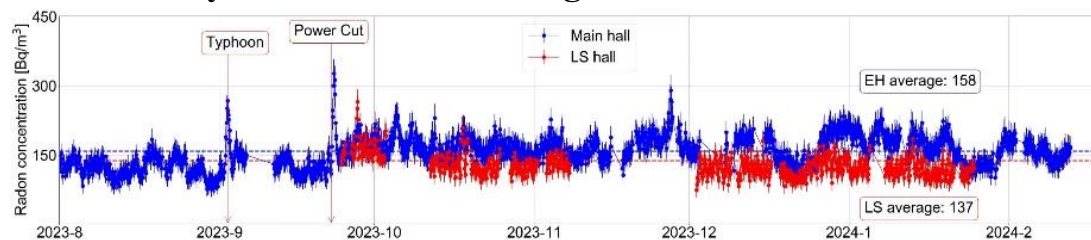
- Radiopurity control of raw material:

- ✓ Careful material screening
- ✓ Meticulous Monte Carlo Simulation
- ✓ Accurate detector production handling

- Radon and cleanness control of hall environment :

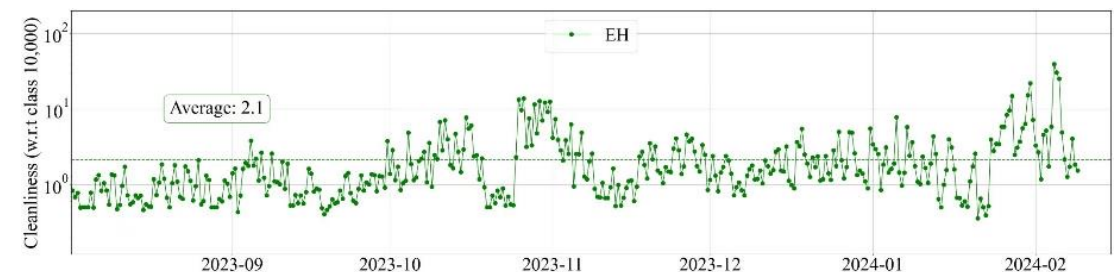
- High efficient VAC and component/personnel access control into exp. hall
- Radon concentration: ~160 Bq/m³ in the EH, ~140 Bq/m³ in the LS hall
- Cleanliness: class 20,000

崔晨阳: Study on dust fallout during installation

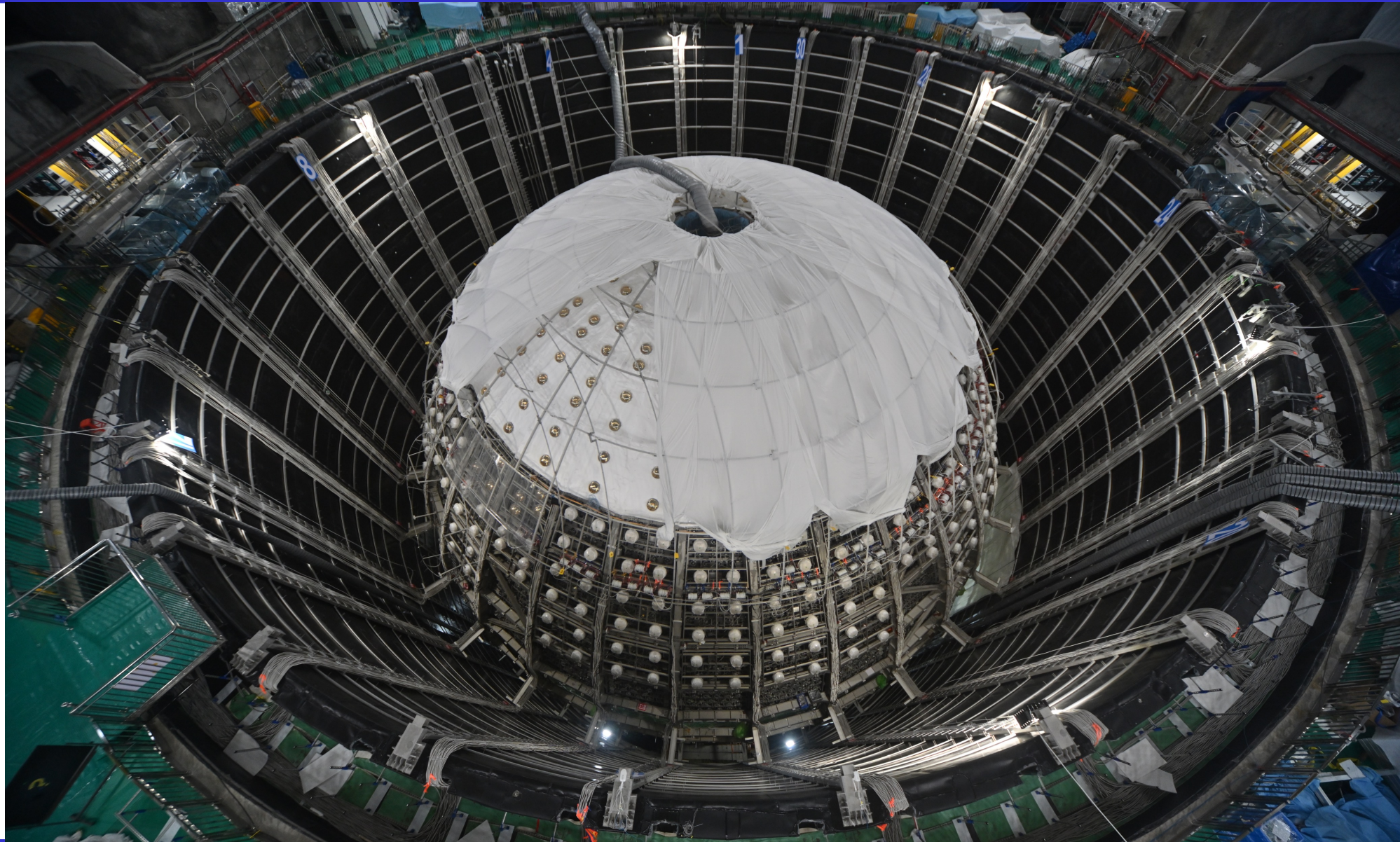


• [JHEP 2021, 102 \(2021\)](#)

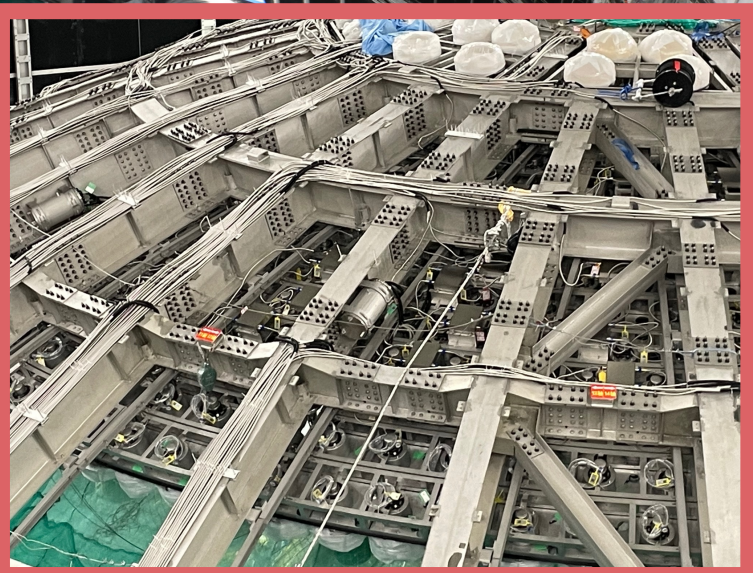
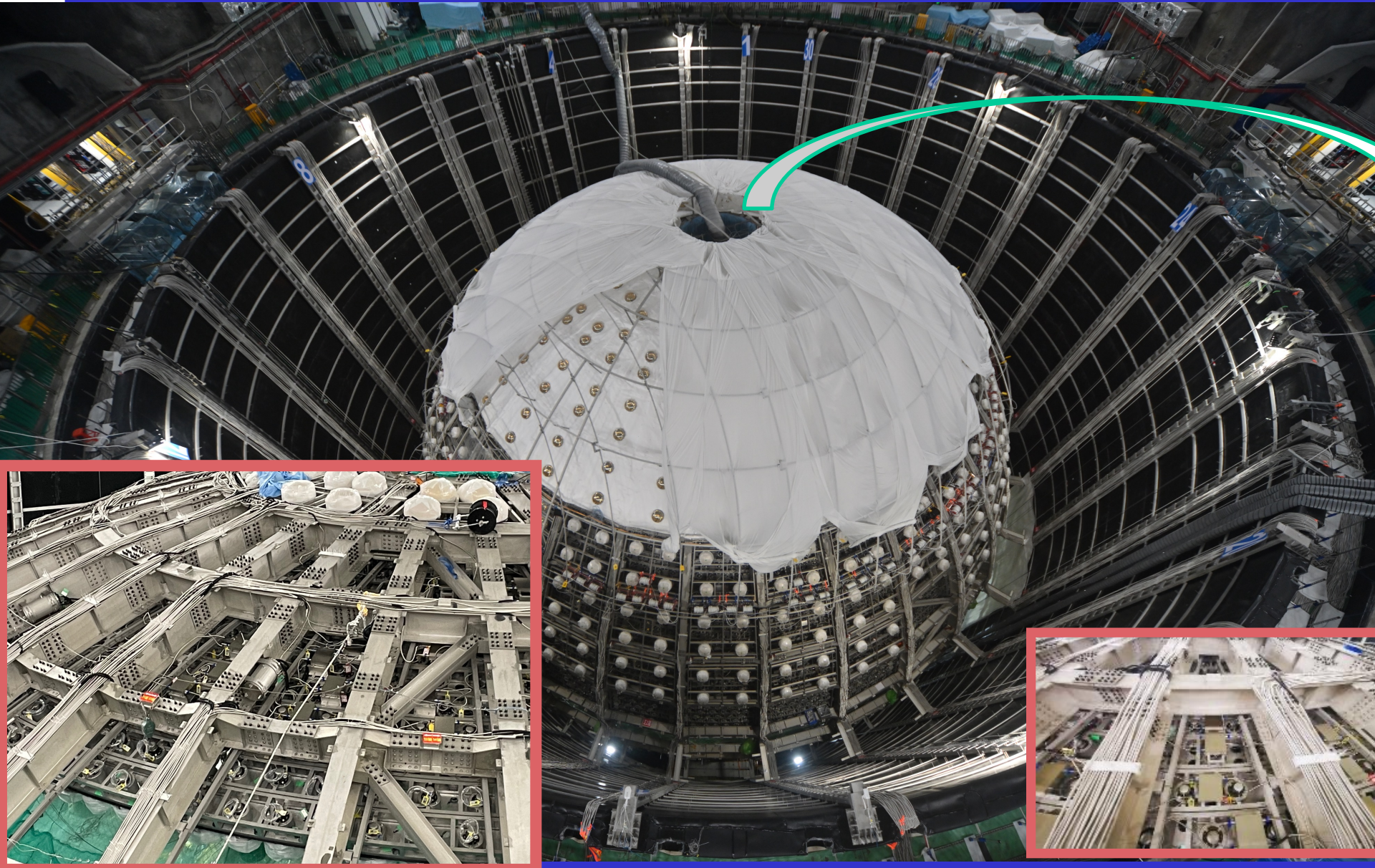
Singles (R < 17.2 m, E > 0.7 MeV)	Design [Hz]	Change [Hz]	Comment
LS	2.20	0	
Acrylic	3.61	-3.2	10 ppt -> 1 ppt
Metal in node	0.087	+1.0	Copper -> SS
PMT glass	0.33	+2.47	Schott -> NNVT/Ham
Rock	0.98	-0.85	3.2 m -> 4 m
Radon in water	1.31	-1.25	200 mBq/m ³ -> 10 mBq/m ³
Other	0	+0.52	Add PMT readout, calibration sys
Total	8.5	-1.3	Better than spec. by 15%



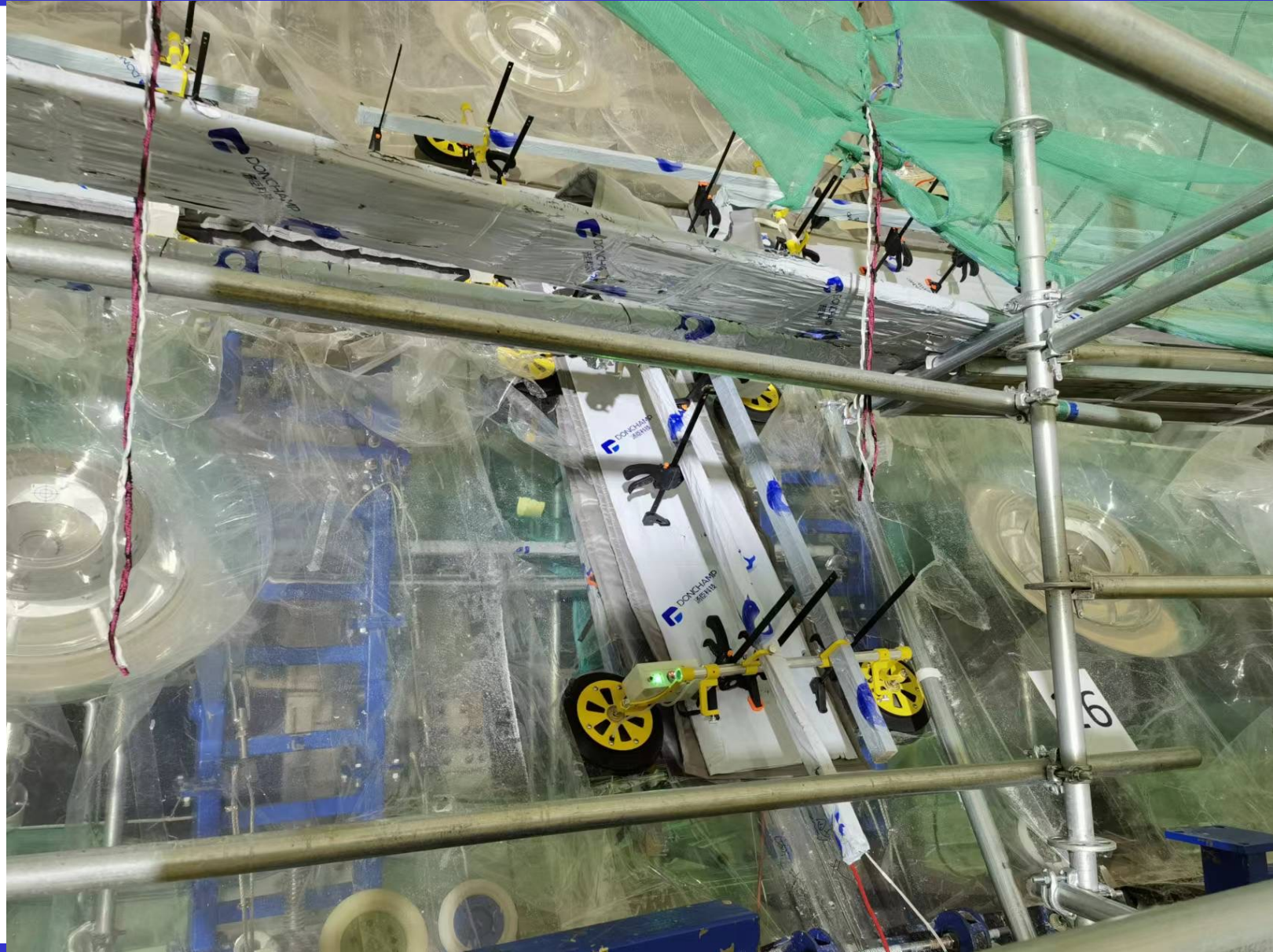
Construction Status





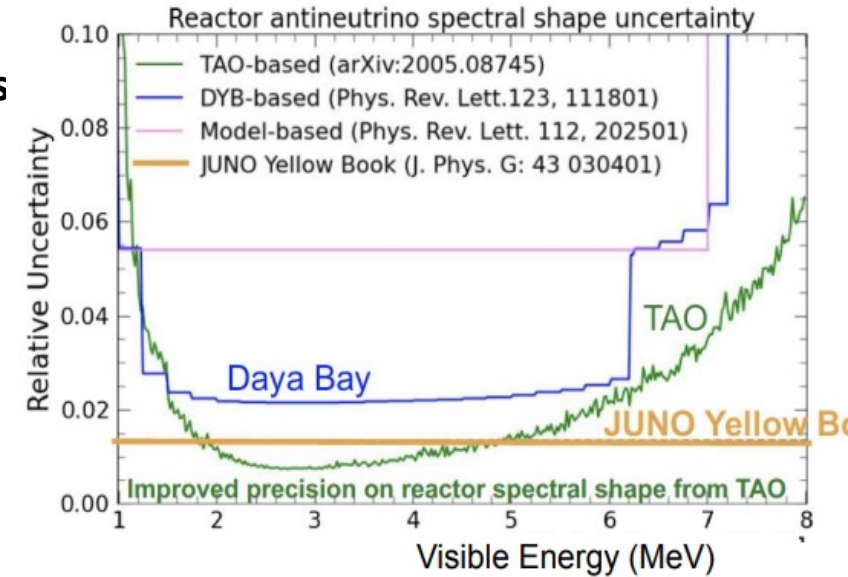




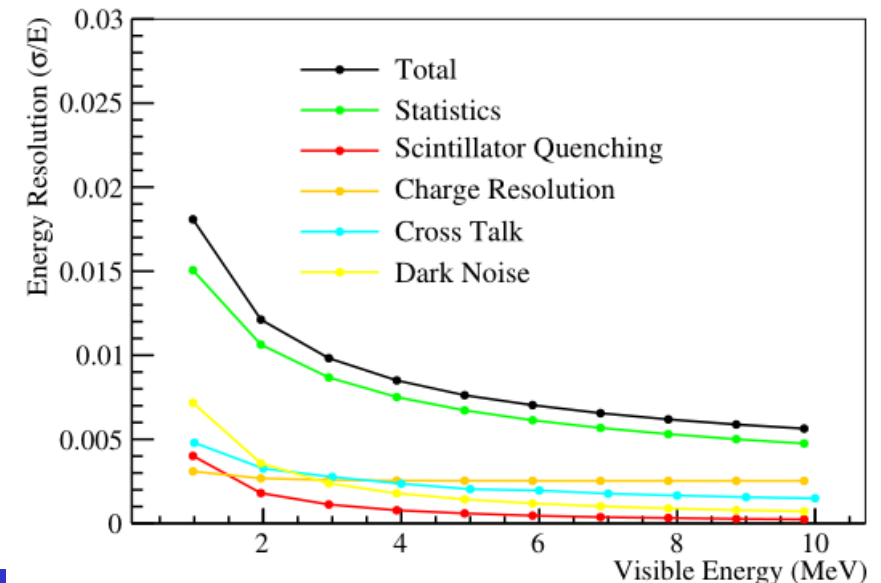
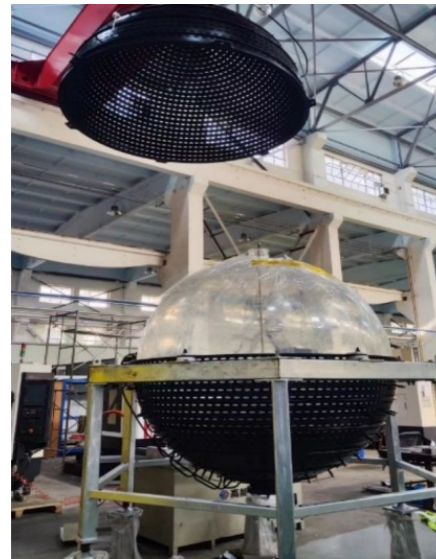
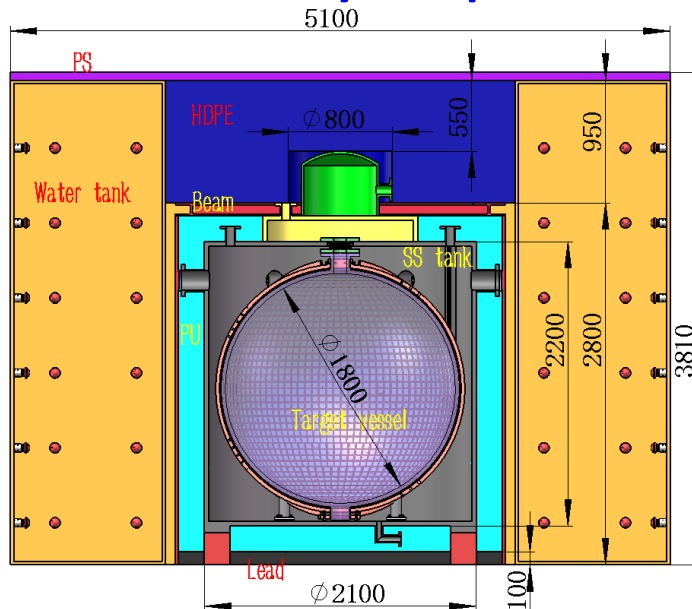




- **To measure the reactor neutrino spectrum as a reference to JUNO**
 - better resolution to reduce fine structure effects and spectrum uncertainties
 - Improve nuclear database, search for sterile neutrinos
- **Idea: Gd-loaded LS @-50°C + SiPM**
 - 700k neutrinos/year@44m from the core(4.6 GW), with ~10% bkg.
 - Energy resolution: ~1.5%/√E, 4500 p.e./MeV
 - 10 m² SiPM (>94% coverage) w/ PDE > 50%
 - Operating at -50 °C to reduce the SiPM dark rate by 10³ to 100 Hz/mm²
 - 2.8 ton(1t fiducial volume) new type of Gd-LS for -50 °C
- **Component production mostly completed**



谢宇广报告: TAO
张杰: MicroTCA' Application in TAO



Physics and Prospects

- Bkg. in LS can reach U/Th/K < 10^{-15} g/g for **MH**
- 10^{-17} g/g is feasible for **solar ν** and $0\nu\beta\beta$ **decays**
- From measured PMT, LS and acrylic properties, the energy resolution will be < **3%@1MeV**, based on full simulation, calibration and reconstruction

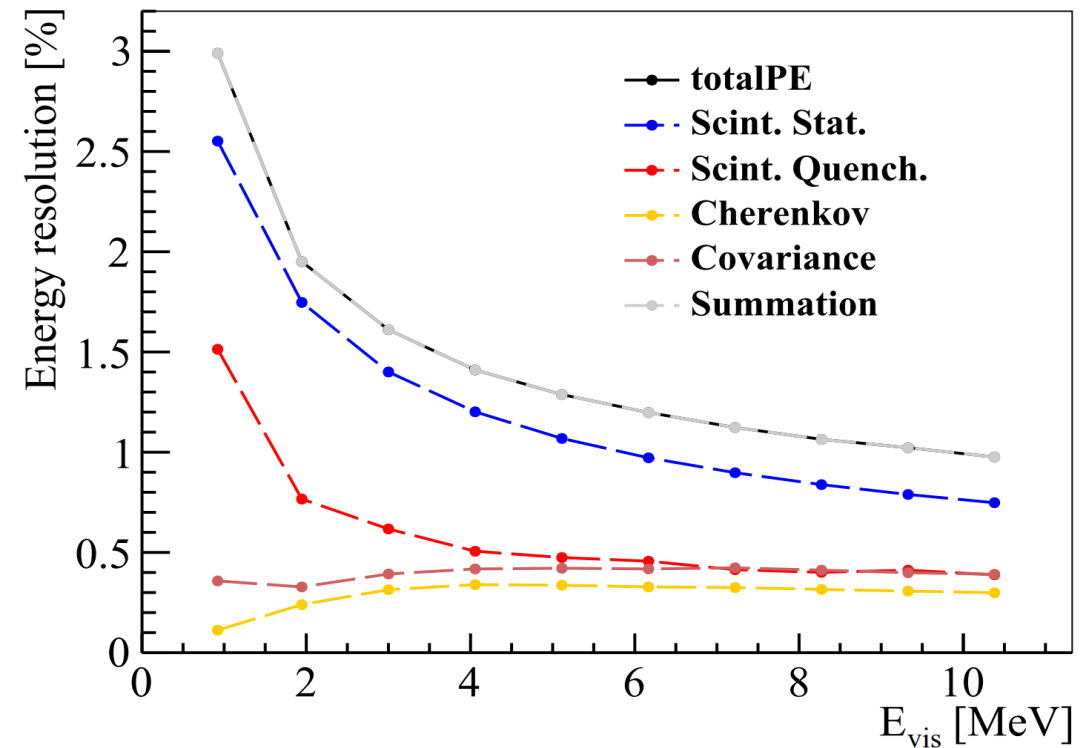
$$\frac{\sigma}{E_{vis}} = \sqrt{\left(\frac{2.614\%}{\sqrt{E_{vis}}}\right)^2 + 0.64\%^2 + \left(\frac{1.205\%}{E_{vis}}\right)^2}$$

↓ **Photon statistics**
↓ **Constant term**
↓ **Annihilation-induced γ s Dark noise**

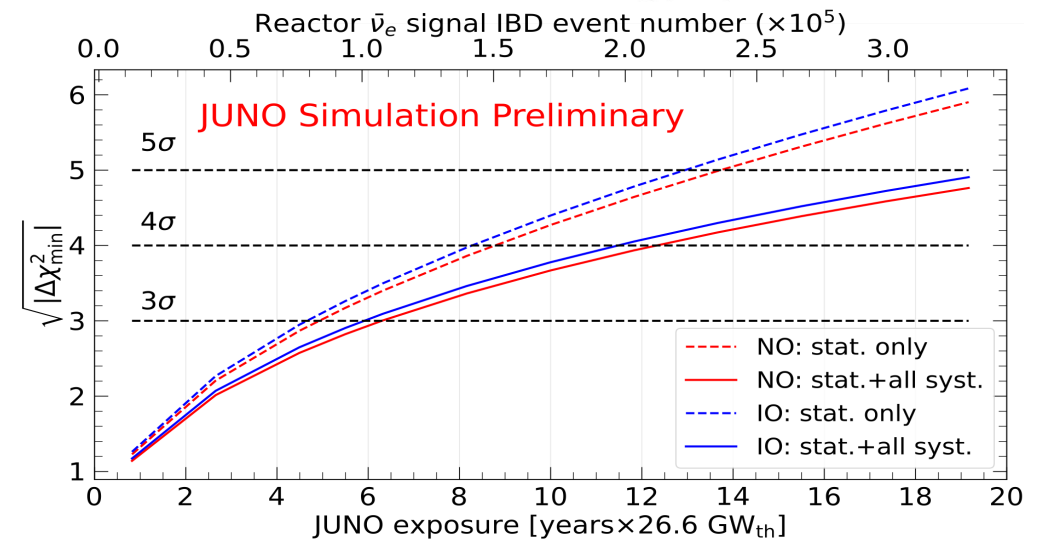
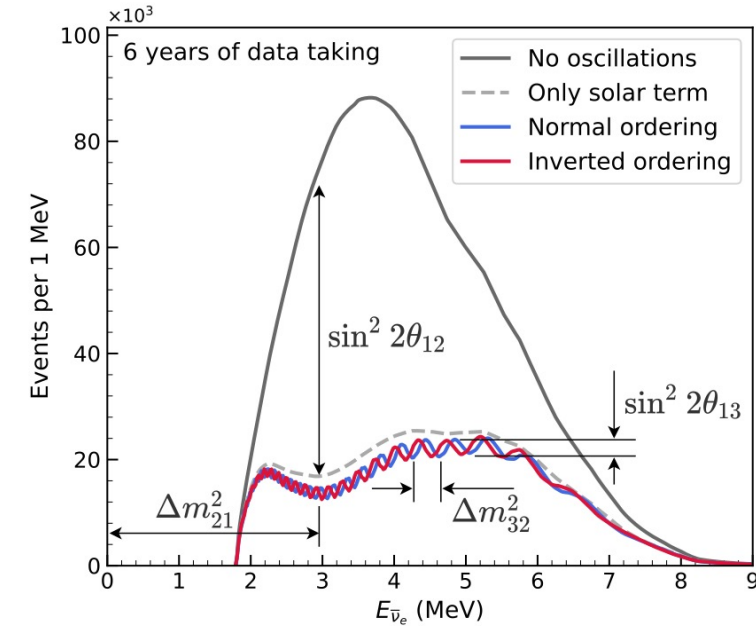
• Main changes vs design (*JHEP03(2021)004*):

- Photon detection eff.: 27% → 30%
- (EPJC 82 (2022) 12, 1168)
- New PMT optical model: +8%
- (EPJC 82 329 (2022))
- New central detector geometry and LS: 3%

• Total photon statistics: ~1660/MeV

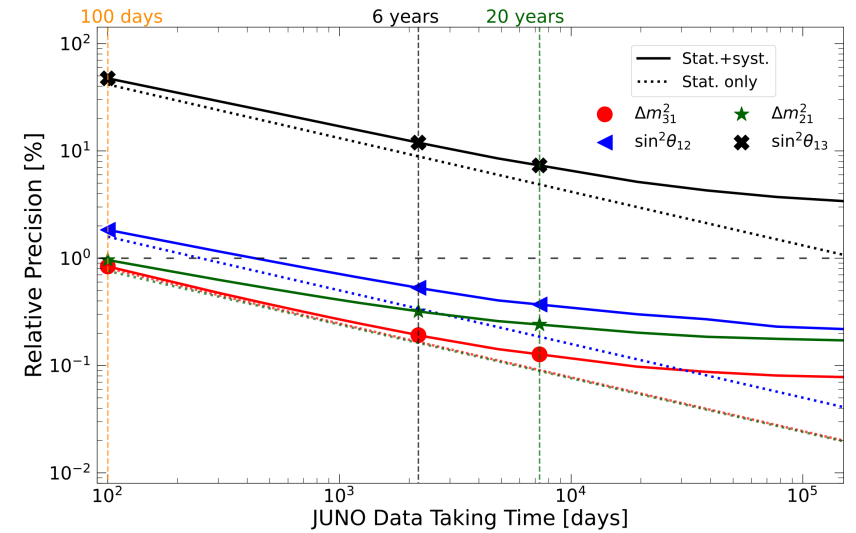


- Determine mass ordering
 - Exploit interference effects in the fine structure of the oscillated spectrum
 - 3σ sensitivity within ~ 6 years
 - Independent of θ_{23} and δ_{CP}
- Atmospheric neutrino is under consideration



程婕, Oscillation Physics Potential with Reactor Neutrino at JUNO

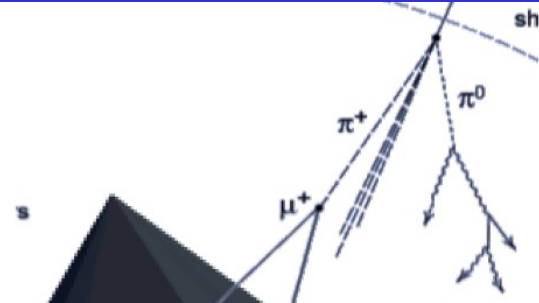
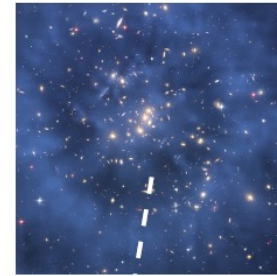
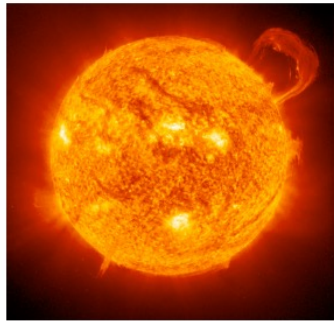
- Measure $\sin^2\theta_{12}$, Δm_{21}^2 and Δm_{31}^2 to better than 0.5% in 6 years
 - New era of precision for neutrino mass & mixing models
 - Model independent tests of the three-neutrino oscillation framework (U_{PMNS} unitarity tests)
 - Narrow down parameter space for $0\nu\beta\beta$ searches



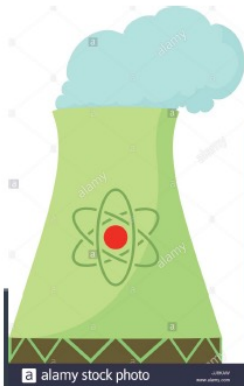
Roughly one order of magnitude improvement over existing precision for 3 parameters!

Parameter	$\sin^2\theta_{12}$	Δm_{21}^2	Δm_{31}^2	$\sin^2\theta_{13}$
Current Precision*	4.2%	2.4%	1.5%	3.2%
JUNO 6 years	0.5%	0.3%	0.2%	12.1%

A Multi-Purpose Neutrino Observatory



崔晨阳, Physics Potential of Detecting Solar Neutrino



Solar ν
 $O(10 - 1000)/\text{day}$

SuperNova ν
 5k in 10^5 s (10 kpc)

Wimp
 (dark matter)
 ?

Atmospheric ν
 10-20/day

Reactor ν
 $\sim 60 - 80/\text{day}$

Geo- ν
 $\sim 1 - 2/\text{day}$



辛钊, Prospect for the detection of the Geo-Neutrino

- **Although the project delayed, overcome a lot of challenges, the installation is under progress**
- **Commissioning shows that the performance of (part of) components/detector is better than the requirement**
- **Expect to beginning filling this year**
- **Mass ordering expected to determine in the decade**
- **Anticipating more exciting physics results**

十年磨一剑

THANKS