

# Symposium on Frontiers of Underground Physics

Hunting for (low-mass) particle dark matter - some naïve and personal thoughts

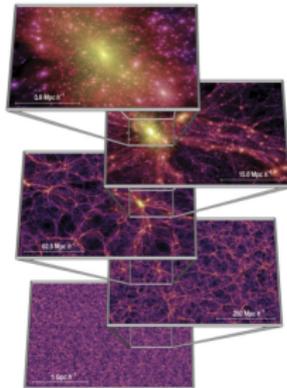
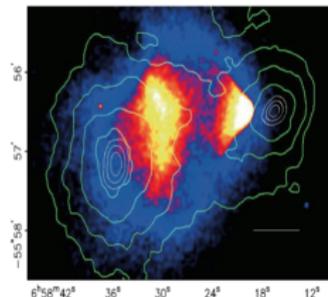
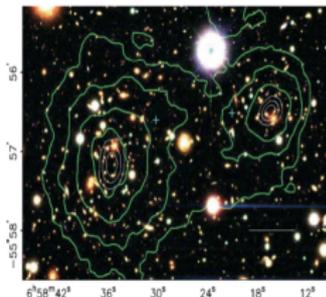
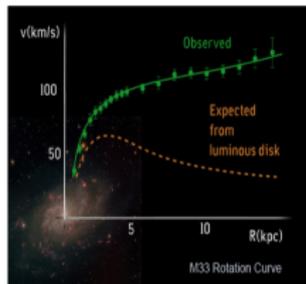
# Outline

- 1 DM “review” (kind of)
  - DM should be there: astrophysicist and cosmologist
  - Particle physicists: where is the damn DM?
- 2 The challenges to DM direct detection community
  - Theoretical challenges
- 3 The opportunities for DM particle physicists
  - New thoughts/actions worldwide
  - My thoughts
- 4 The progress of the ALETHEIA project
  - ALETHEIA Introduction
  - ALETHEIA prototype detector: the 30g-V1 LHe
  - ALETHEIA prototype detector: TPB coating on a PTFE chamber
  - ALETHEIA prototype detector: (preliminary) SiPMs tests at 4 K.
  - Ongoing tests at CIAE
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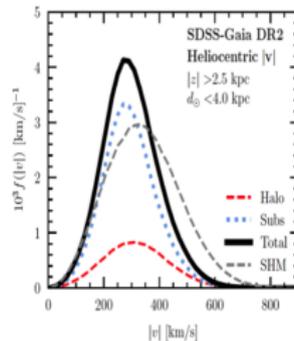
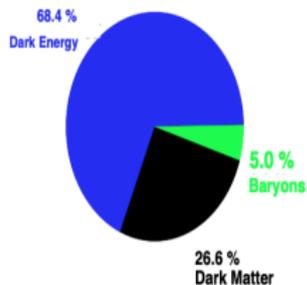
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# Astrophysical and cosmological evidence of DM existence.



Constituents of today's universe, Planck 2018



# Astrophysicists and cosmologists observed DM.

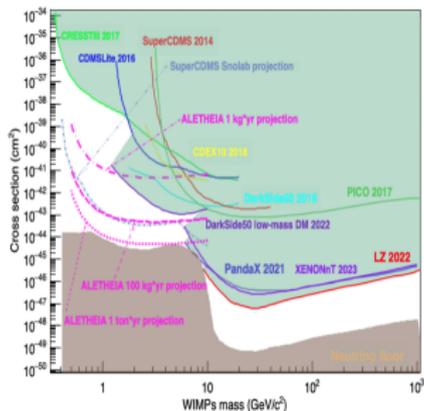
Astrophysicists/Cosmologists celebrating their decent publications



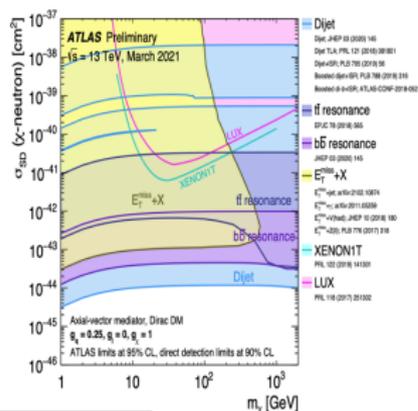
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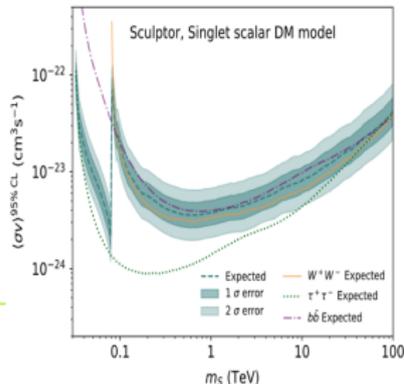
# No any convincing DM signals from each and every hunting strategy.



arXiv: 2210.01220



ATL-PHYS-PROC-2022-003



# Particle physicists working in DM.

## Particle physicists working in DM

Ten years ago



Now



Did not discover dark matter yet, discovered "dark life" instead

## Research on DM: Astrophysicists V.S. particle physicists

- Left picture: Astrophysicists working in DM.  
Right picture: Particle physicists working in DM.

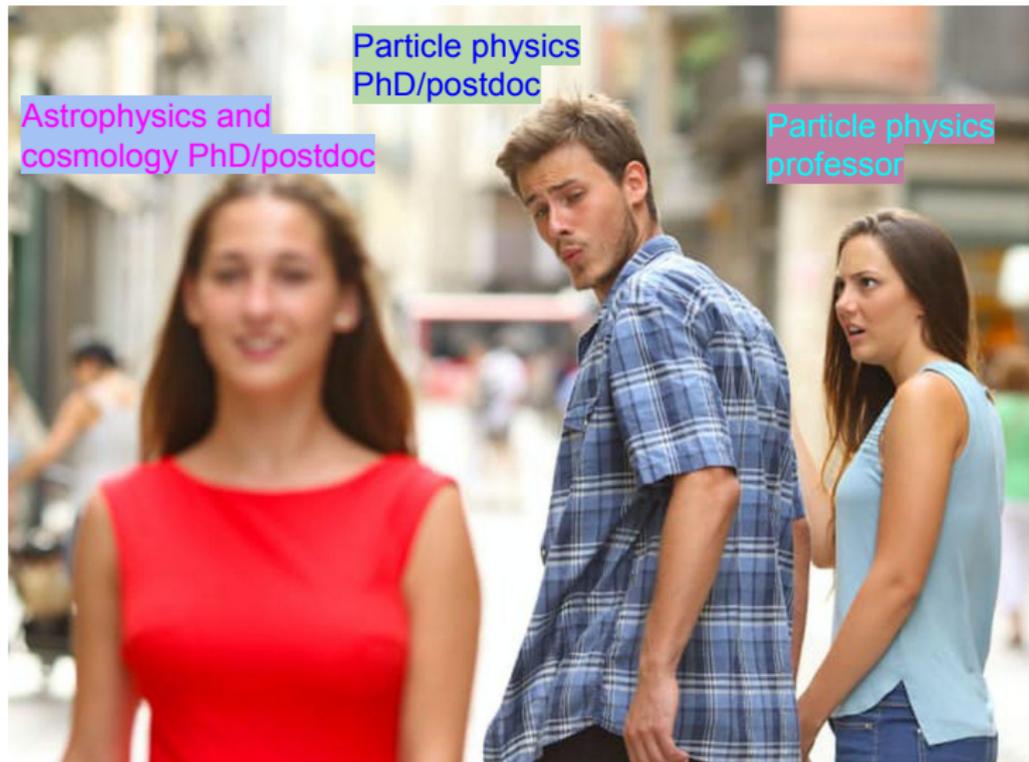
### ASTRONOMY



### HIGH ENERGY PHYSICS



# Particle physicist professors working in DM (Based on a true story.)



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than the electroweak scale. Thus, if supersymmetry is to be relevant to electroweak-symmetry breaking, the signs of supersymmetry must become visible near the electroweak scale,  $E \lesssim \mathcal{O}(\text{TeV})$ . If a supersymmetric particle spectrum does not become manifest by the time accelerator experiments reach such a scale, then the solution to the fine-tuning problem must be sought elsewhere. Thus supersymmetry is an interesting gamble. If it proves to be a winning bet, the rewards will be staggering. If it proves to be irrelevant to electroweak physics, then the absence of a viable treatment for the fine-tuning problem will be felt.

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- Some physicists do not agree that "WIMPs are dead", Jonathan Feng, Dan Hooper, Michael Peskin, Yufeng Zhou ...

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**ALETHEIA**, R&D stage

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**ALETHEIA**, R&D stage
- ALETHEIA progressed significantly since 2020: demonstrated the viability of single-phase LHe TPCs;  
R&D underway: dual-phase LHe TPCs.

# Report of the Topical Group on Particle Dark Matter for Snowmass 2021, arXiv: 2209.07426, 1/2

Name	Technology	Target	Active Mass	Experiment Location	Start Ops	End Ops
<b>Currently Running or Under Construction</b>						
LZ	TPC	LXe	7,000 kg	SURF	2021	2026
PandaX-4T	TPC	LXe	4,000 kg	CJPL	2021	2025
XENONnT	TPC	LXe	7,000 kg	LGNS	2021	2025
DEAP-3600	Scintillator	LAr	3,300 kg	SNOLAB	2016	2025
Darkside-20k	TPC	LAr	50 t	LNGS	2027	2035
DAMA/LIBRA	Scintillator	NaI	250 kg	LNGS	2003	
ANAIS-112	Scintillator	NaI	112 kg	Canfranc	2017	2022
SABRE PoP	Scintillator	NaI	5 kg	LNGS	2021	2022
COSINE-200	Scintillator	NaI	200 kg	YangYang	2022	2025
CDEX-10	Ionization (77K)	Ge	10 kg	CJPL	2016	
EDELWEISS III (High Field)	Cryo Ionization / HV	Ge	33 g	LSM	2019	
SuperCDMS CUTE	Cryo Ionization / HV	Ge/Si	5 kg/1 kg	SNOLAB	2020	2022
SuperCDMS SNOLAB	Cryo Ionization / HV	Ge/Si	11 kg/3 kg	SNOLAB	2023	2028
CRESST-III (HW Tests)	Bolometer Scintillation	CaWO4		LNGS	2020	
PICO-40	Bubble Chamber	C3F8	35 kg	SNOLAB	2020	
NEWS-G	Gas Drift	CH4		SNOLAB	2020	2025
DAMIC-M prototype	CCD Skipper	Si	18 g	LSM	2022	2023
DAMIC-M	CCD Skipper	Si	1 kg	LSM	2024	2025
SENSEI	CCD Skipper	Si	2 g	Fermilab	2019	2020
SENSEI	CCD Skipper	Si	100 g	SNOLAB	2021	2023

## arXiv: 2209.07426, 2/2

Name	Technology	Target	Active Mass	Experiment Location	Start Ops	End Ops
<b>Planned</b>						
SABRE (North)	Scintillator	NaI	50 kg	LNGS	2022	2027
SABRE (South)	Scintillator	NaI	50 kg	SUPL	2022	2027
COSINE-200 South Pole	Scintillator	NaI	200 kg	South Pole	2023	
COSINUS	Bolometer Scintillator	NaI		LNGS	2023	
Darwin / XLZD (US LXe G3)	TPC	LXe	50,000 kg	undetermined	2028	2033
ARGO	TPC or Scintillator	LAr	300 t	SNOLAB	2030	2035
CDEX-100 / 1T	Ionization (77K)	Ge	100-1000 kg	CJPL	202X	
PICO-500	Bubble Chamber	C3F8	430 kg	SNOLAB	2021	
<b>Concept or R&amp;D</b>						
Oscura	CCD Skipper	Si	10 kg Si	SNOLAB	2025	2028
SBC	Bubble Chamber	LAr	1 t	SNOLAB	2028	
SNOWBALL	Supercooled Liquid H2O					
DarkSide-LowMass	TPC	LAr	1.5 t			
ALETHEIA	TPC	He		China Inst. At. Energy		
TESSERACT	Cryo TES	LHe, SiO <sub>2</sub> , Al <sub>2</sub> O <sub>3</sub> , GaAs		undetermined	2026	
CYGN0	Gas Directional	He + CF <sub>4</sub>	0.5 - 1 kg	LNGS	2024	
CYGNUS	Gas Directional	He + SF <sub>6</sub> /CF <sub>4</sub>		Multiple sites		
Windchime	Accelerometer array			Multiple sites		
MAGNETO-χ	Cryogenic MMC	Diamond, Sapphire, etc.				

# Sub-GeV DM new techniques, Dan McKinsey at PKU in 2019

WIMPs	Light DM, existing techniques	This talk (sub-GeV new techniques)
DAMA/LIBRA	Argon S2-only	LXe/LAr bubble chamber
COSINE	DarkSide-LowMass	Snowball chamber
DarkSide-50	LUX: Xe Migdal effect	Xenon S2-only
XENON1T	DAMIC	Graphene
PICO	NEWS-G	Internally amplified Ge
DarkSide-20k	SuperCDMS	Color centers
PandaX-4T	CRESST	Scintillating crystals (GaAs, CsI, NaI)
LUX/ZEPLIN		Polar crystals
XENONnT		Diamond
DARWIN		Superconductors
		Superfluid helium

This is not a fully exhaustive list; apologies if your favorite new technique is not covered!

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## A DM search W/O bias: ER & NR co-existing should not be missed.

- There exist tens of elementary particles in the SM, reasonable to hypothesize there are more than one type of DM particle; some generate ER, some NR  $\Rightarrow$  ER and NR could co-exist in a DM detector's same dataset.

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- For more details, arXiv: 2302.12406.

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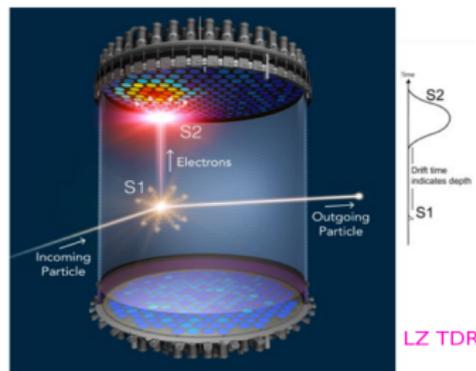
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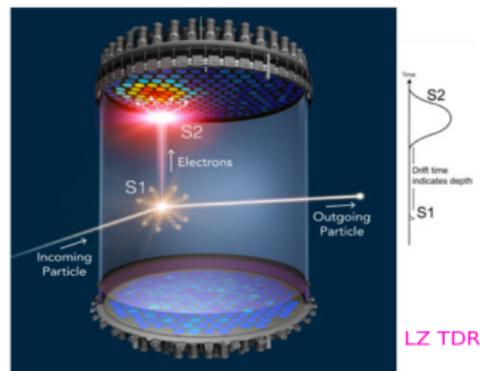
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LZ TDR

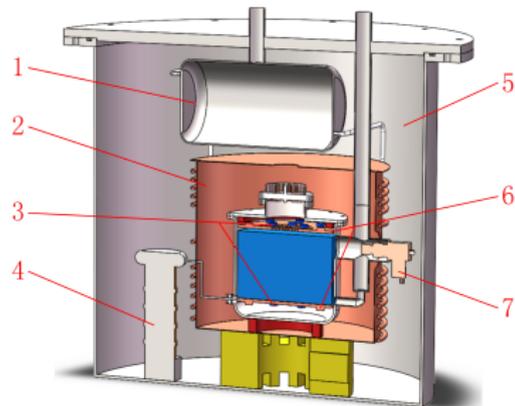
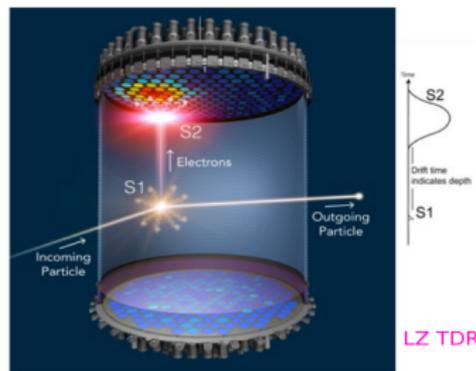
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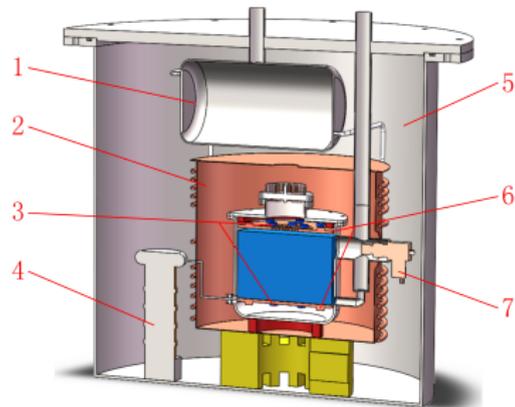
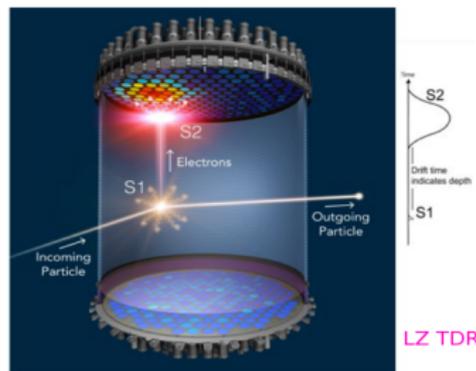
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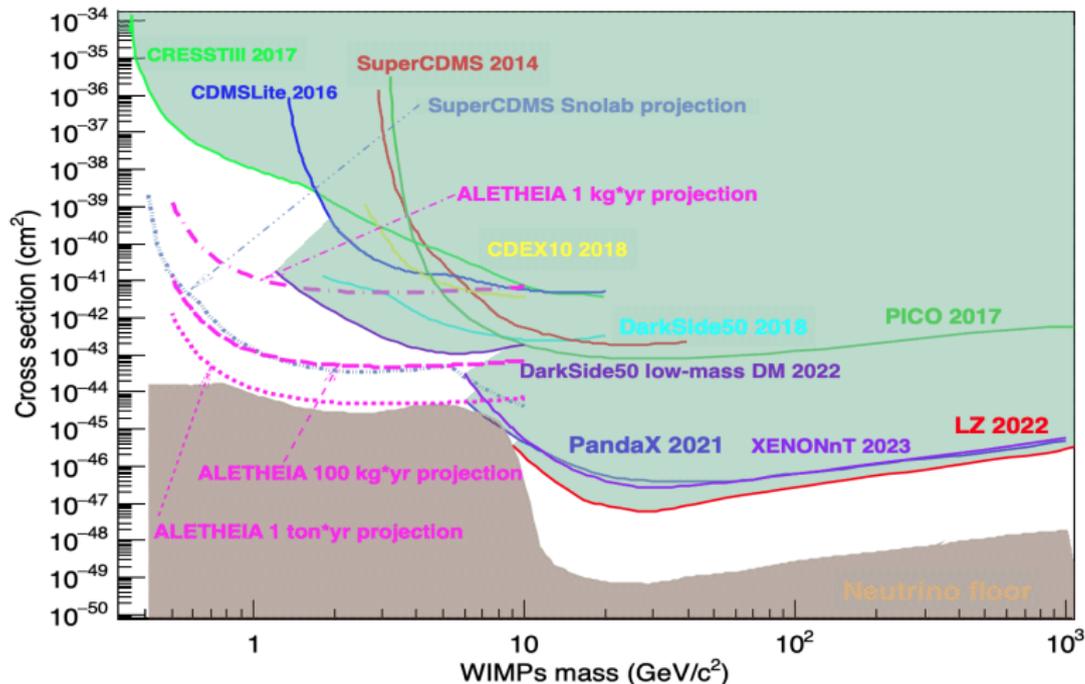
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# ALETHEIA NR channel: Projected sensitivities

- 1 ton\*yr ALETHEIA can “touch down” the  $^8\text{B}$  solar  $\nu$  fog (Assuming IBF, 50% Eff.).



# ALETHEIA review, Oct 2019.

Dark matter (WIMPs) direct detection Mini - Workshop, Oct 14-Oct 16, 2019, CHEP, PKU, CHINA.



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- *“It is possible that liquid helium could enable especially low backgrounds because of its powerful combination of intrinsically low radioactivity, ease of purification, and charge/light discrimination capability.”*

## ALETHEIA collaborators so far

5 institutions (increasing), ~ 20 members

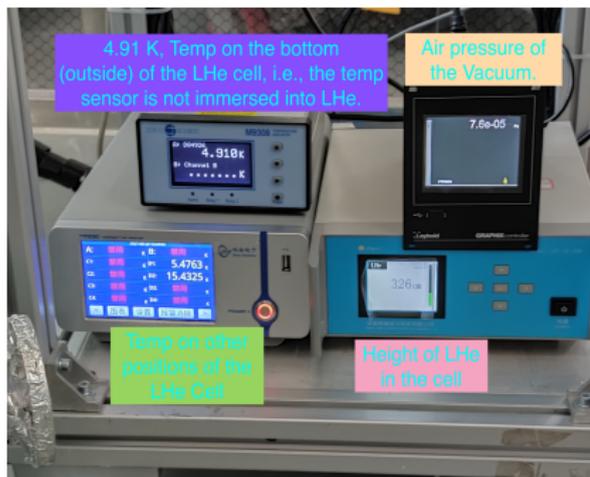
- CIAE (China Institute of Atomic Energy), ~ 10 researchers.
- Peking University, 1 + 2 (?) researchers.
- University of South China, 1 + 1(?) researchers.
- China Southern Power Grid Electric Power Research Institute, 5 researchers.
- SCRI (Shanghai Cable Research Institute), 3 researchers.

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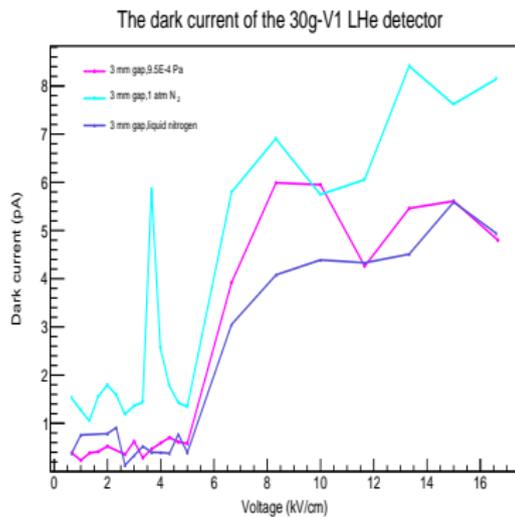
# The R&D of the 30g-V1 LHe prototype.

- Left picture: the detector successfully cooled to 4 K.



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- Left picture: the detector successfully cooled to 4 K.
- Right plot: dark current is less than 10 pA under several circumstances.

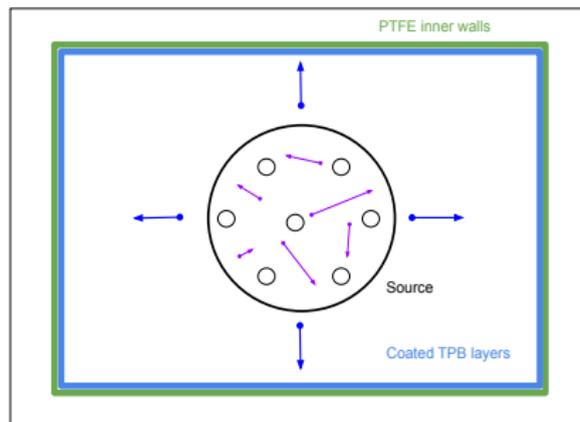


# Outline

- 1 DM “review” (kind of)
  - DM should be there: astrophysicist and cosmologist
  - Particle physicists: where is the damn DM?
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- 4 The progress of the ALETHEIA project
  - ALETHEIA Introduction
  - ALETHEIA prototype detector: the 30g-V1 LHe
  - **ALETHEIA prototype detector: TPB coating on a PTFE chamber**
  - ALETHEIA prototype detector: (preliminary) SiPMs tests at 4 K.
  - Ongoing tests at CIAE
- 5 Summary

## LHe light peaked 80 nm, TPB to convert into visible light.

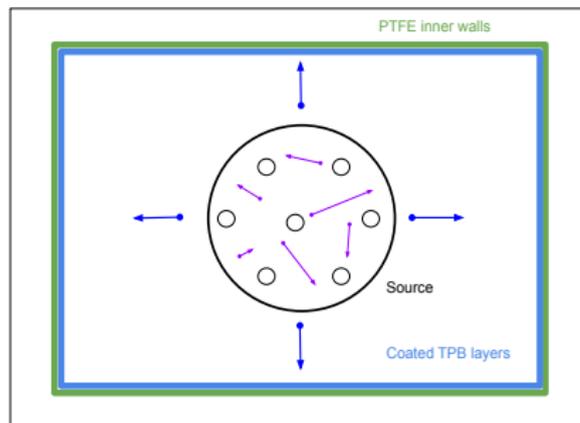
- Left picture: the principle of TPB coating.



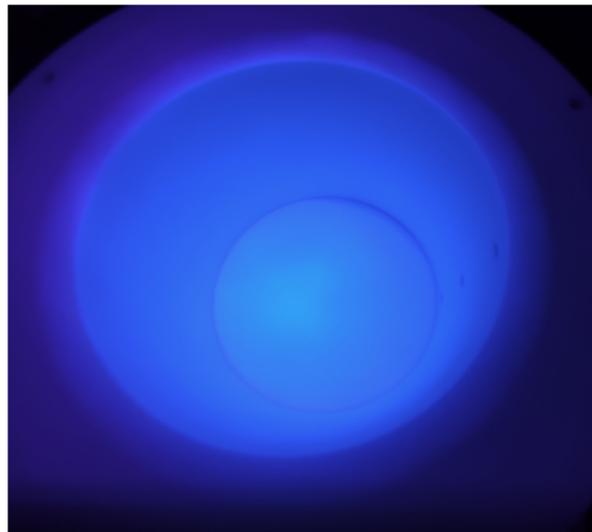
- TPB molecules move inside of the source.
- TPB molecules escape from the source then fly toward the inner walls of the cylindrical PTFE cells.

## LHe light peaked 80 nm, TPB to convert into visible light.

- Left picture: the principle of TPB coating.
- Right plot: top view of the coated 10-cm size PTFE chamber.
- Published: Acta Phys. Sin. Vol. 71, No. 22 (2022) 229501

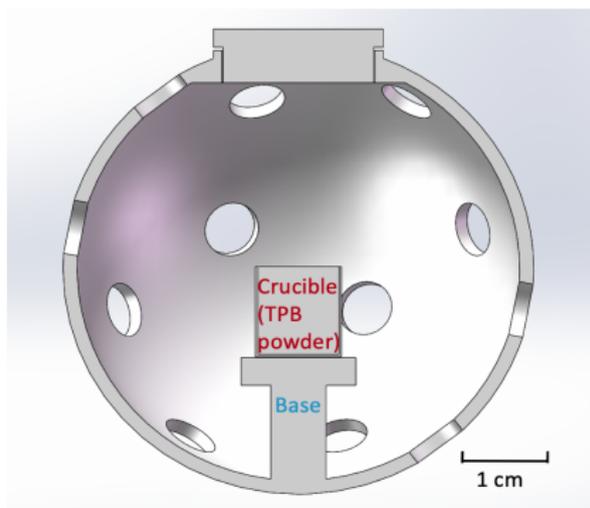


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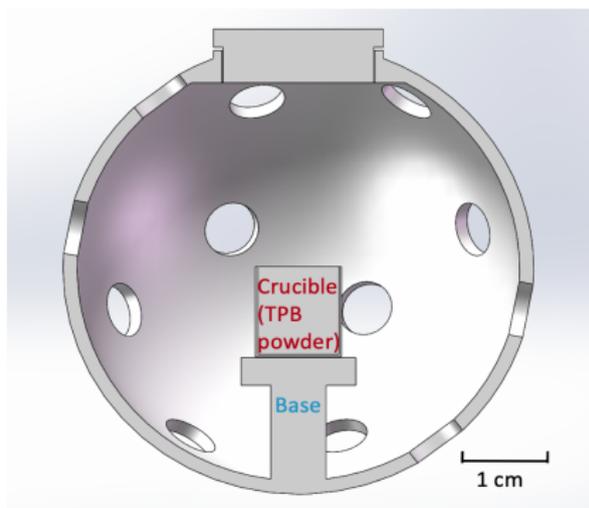
## The coating source.

- Left picture: The source's drawing.



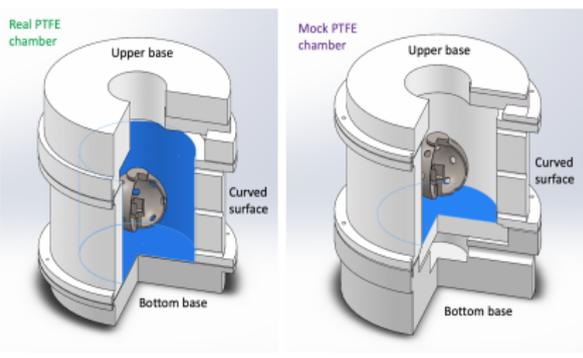
## The coating source.

- Left picture: The source's drawing.
- Right plot: The image of the source.



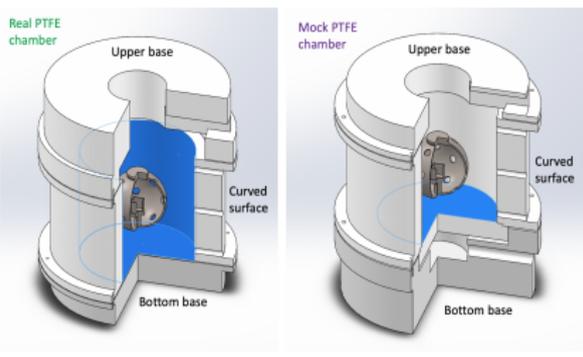
## Coating process.

- Left picture: Coating into steps.



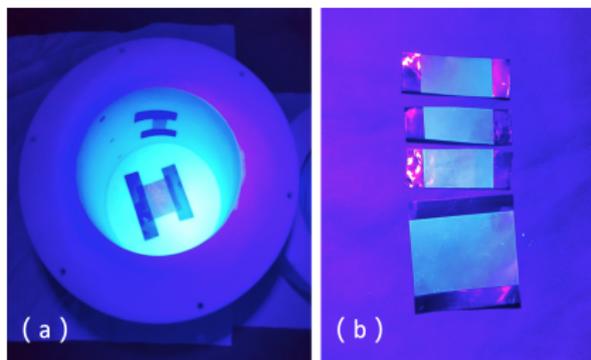
## Coating process.

- Left picture: Coating into steps.
- Right plot: real time monitoring on TPB thickness.



## Figure out the TPB coating thickness.

- Left picture: sample films inside of the chamber.



## Figure out the TPB coating thickness.

- Left picture: sample films inside of the chamber.
- Right plot: calculate TPB's thickness based on the mass difference.

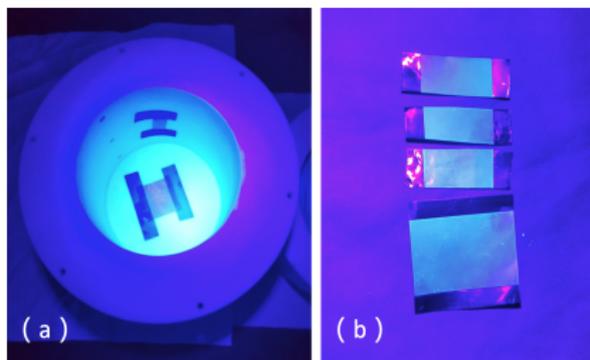


Table 1. TPB coating thickness calculation based on the mass difference before and after coating on aluminum plates

Sample #	Plate area (cm <sup>2</sup> )	Plate position	Mass increase (mg)	thickness (μm)
1	2	Chamber top	$0.75 \pm 0.02$	$3.48 \pm 0.11$
2	2	Chamber top	$0.46 \pm 0.04$	$2.13 \pm 0.17$
3	2	Curved surface	$0.87 \pm 0.04$	$4.03 \pm 0.16$
4	6	Chamber bottom	$2.54 \pm 0.02$	$3.92 \pm 0.03$

## Figure out the TPB coating thickness.

- Left picture: sample films inside of the chamber.
- Right plot: calculate TPB's thickness based on the mass difference.
- The third method to figure out TPB's thickness is based on the TPB mass consumed, 0.2 g.
- All of the three methods returned consistent thickness.

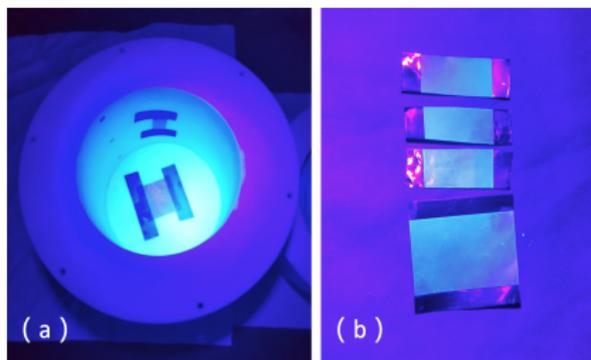
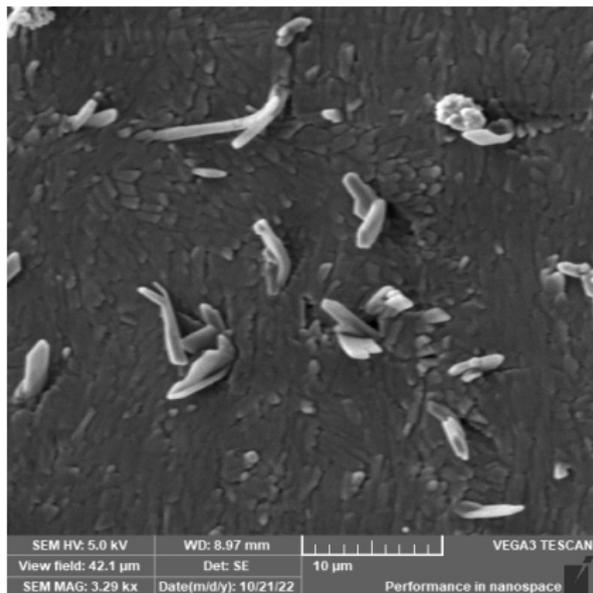


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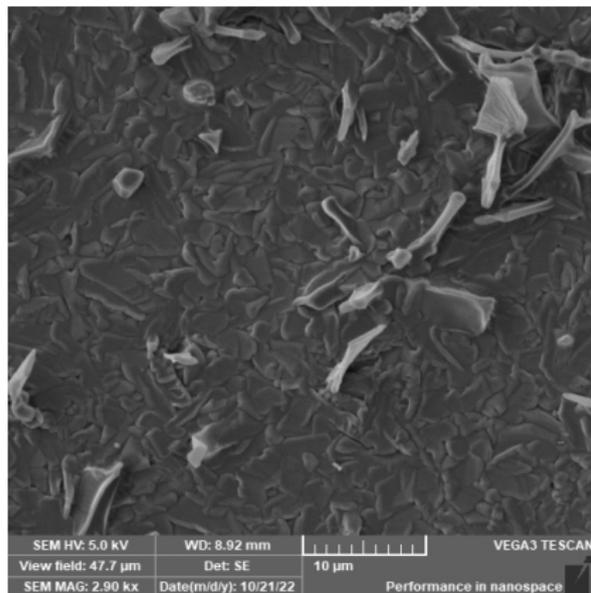
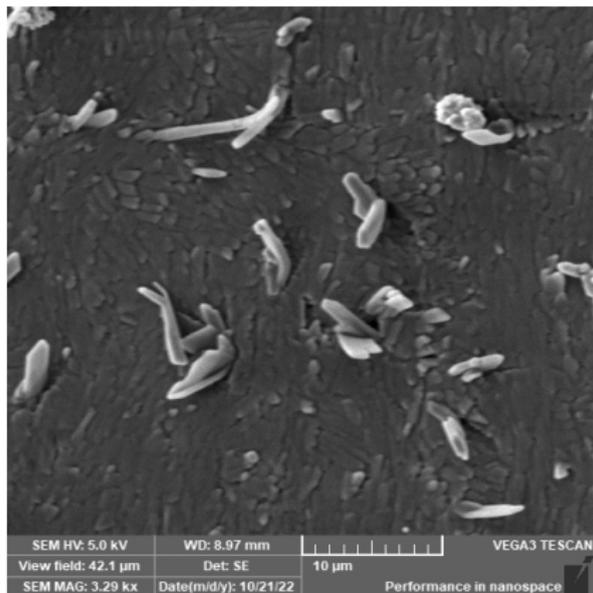
## TPB coating film, exposed at 4 K.

- Left picture: SEM scanning image on TPB coated film experienced at 4 K.



## TPB coating film, exposed at 4 K.

- Left picture: SEM scanning image on TPB coated film experienced at 4 K.
- Right plot: SEM scanning image on TPB coated film W/O cryogenic experience.
- Published in JINST, 2022 JINST 17 P12001.

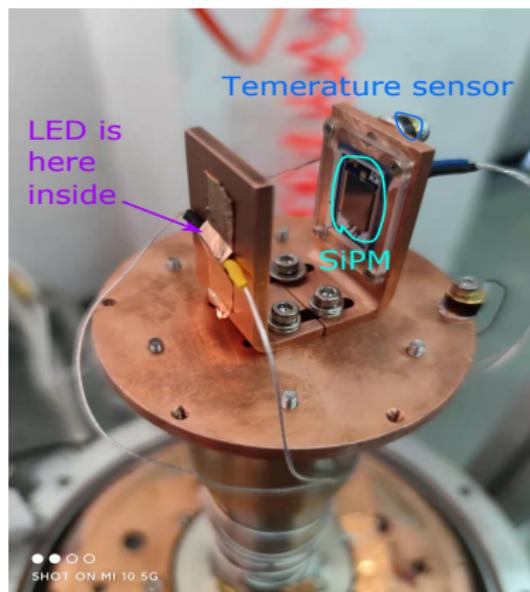


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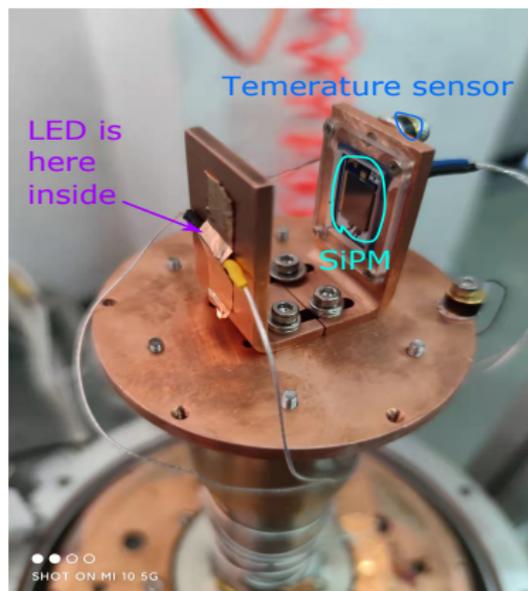
## SiPMs tests at 4 K, with a LED

- Left picture: experimental setup (Inside of the G-M cryocooler).

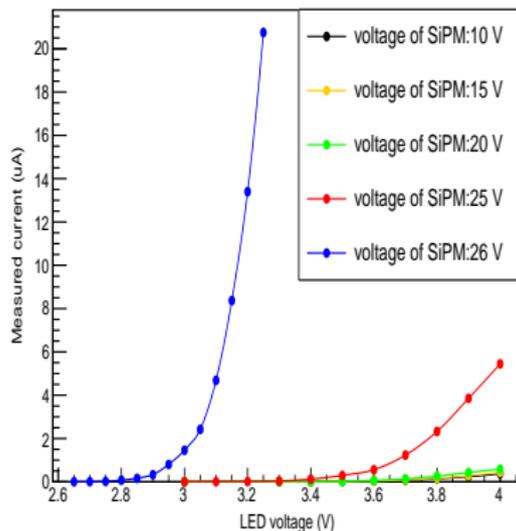


## SiPMs tests at 4 K, with a LED

- Left picture: experimental setup (Inside of the G-M cryocooler).
- Right plot: Preliminary results.

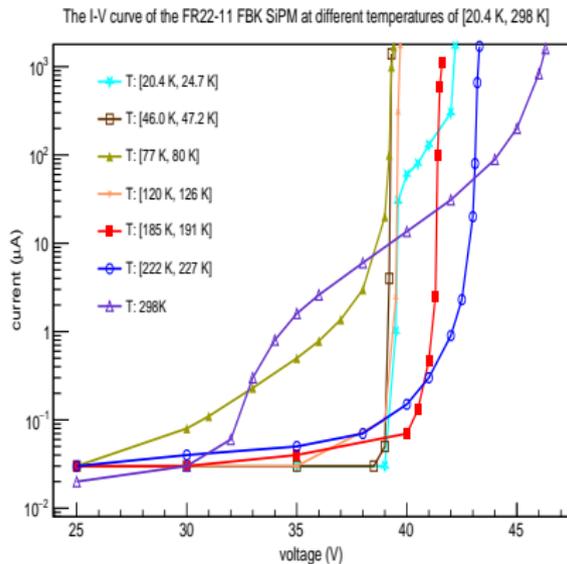


A 450 nm LED lights up an FBK SiPM at 4.8 K



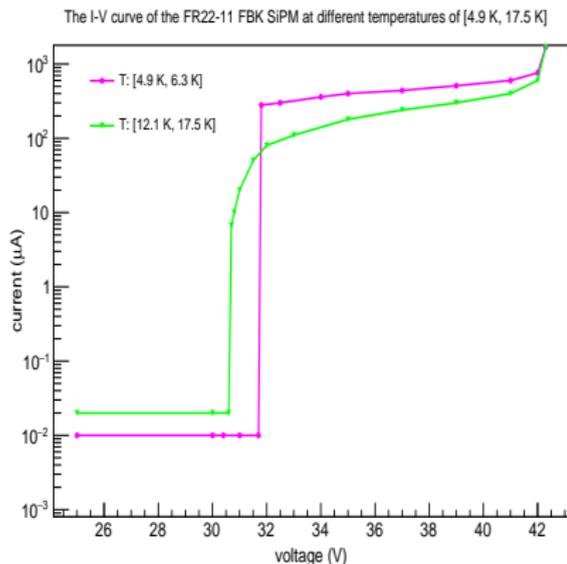
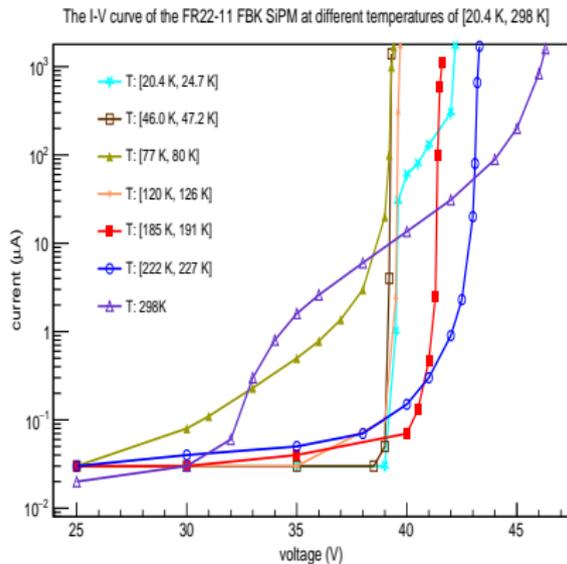
## SiPMs test at 4 K, IV curve measurement

- Left picture: SiPMs IV curve tests, 20 K - RT.



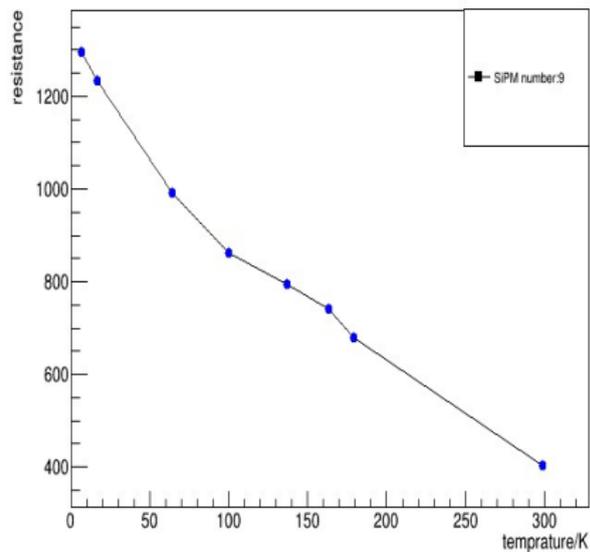
## SiPMs test at 4 K, IV curve measurement

- Left picture: SiPMs IV curve tests, 20 K - RT.
- Right plot: SiPMs IV curve tests, (4 - 20) K, 10 V plateau existed.



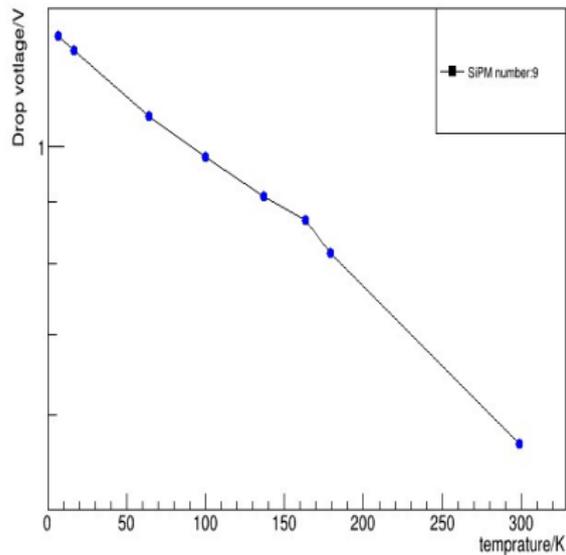
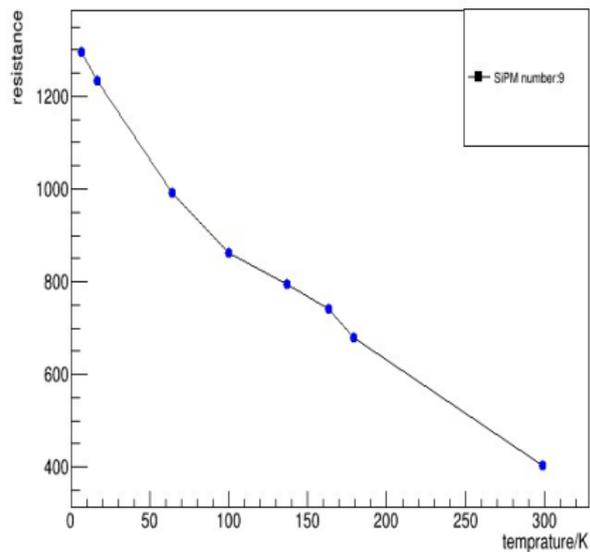
## FBK SiPM: resistance VS temp, and Drop voltage VS Temp.

- Left picture: FBK SiPM, resistance VS temp.



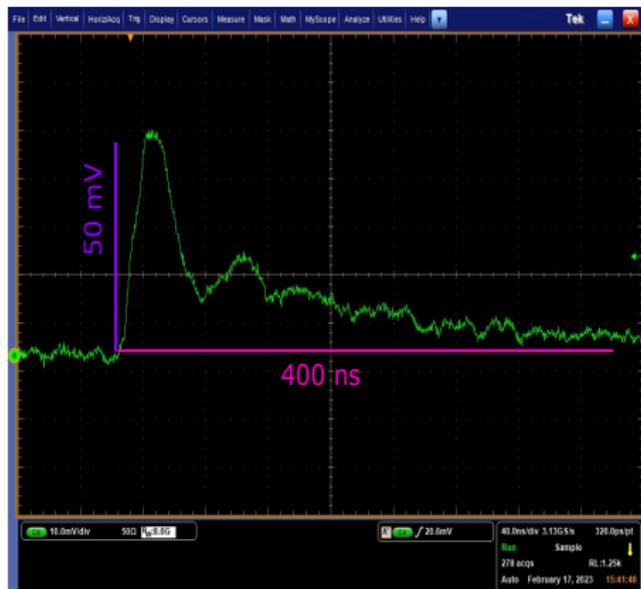
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- Left picture: FBK SiPM, resistance VS temp.
- Right plot: FBK SiPM, voltage VS temp.



## FBK SiPM @ 4 K, typical analog signal.

- 38 FBK SiPMs tested. Most (36/38) of FBK SiPMs are functional at 4 K. More detailed: Eur. Phys. J. Plus (2023) 138:128.



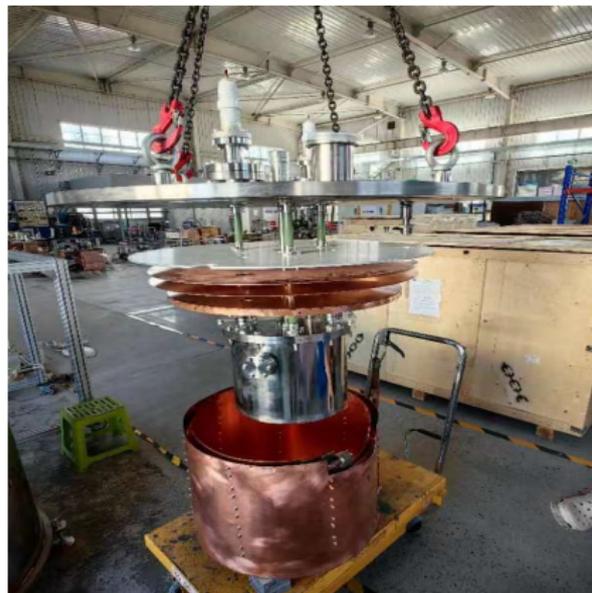
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A 10-cm chamber, ~ 100 g LHe prototype detector is assembling.

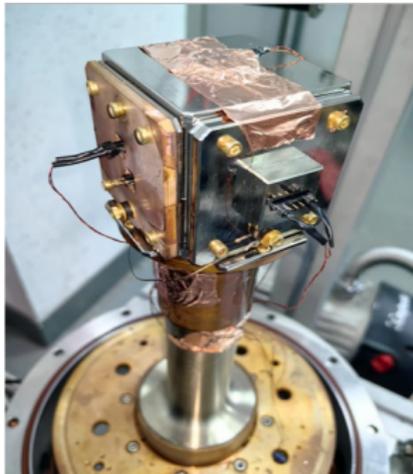


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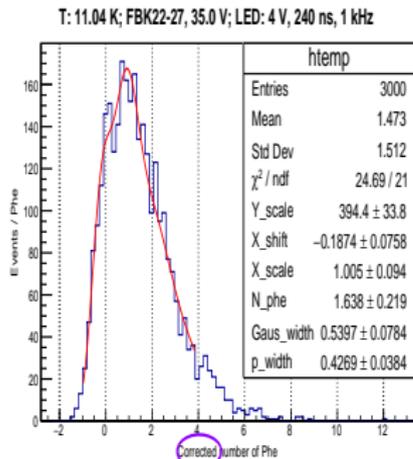
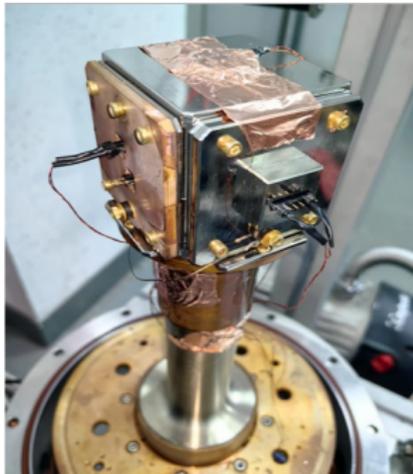
## FBK SiPM PDE tests at 10 K

- Left picture: An integrating sphere with a SiPM, an LED, and a photodiode on the sphere on it.



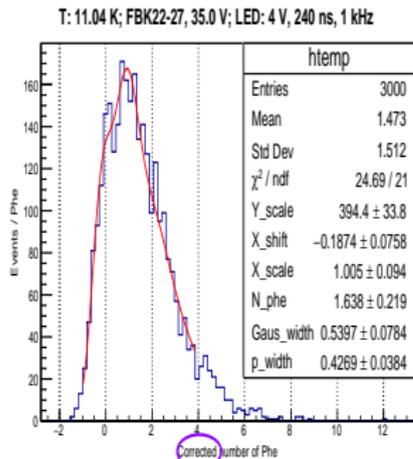
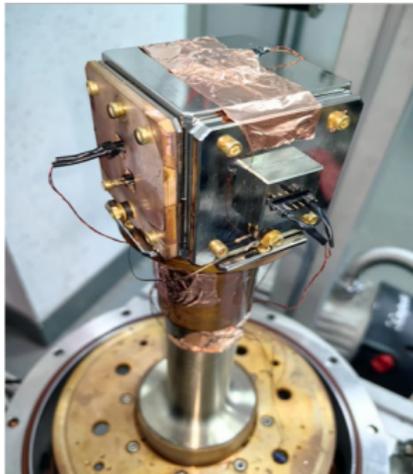
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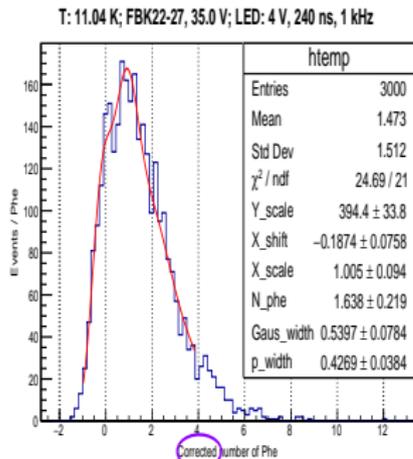
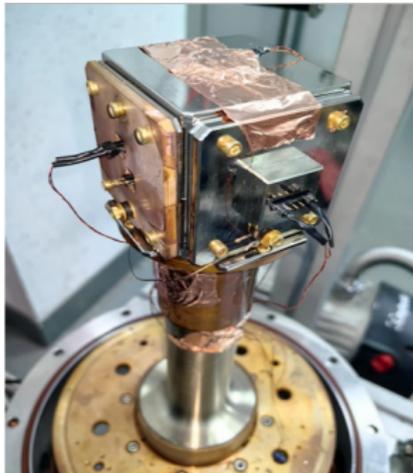
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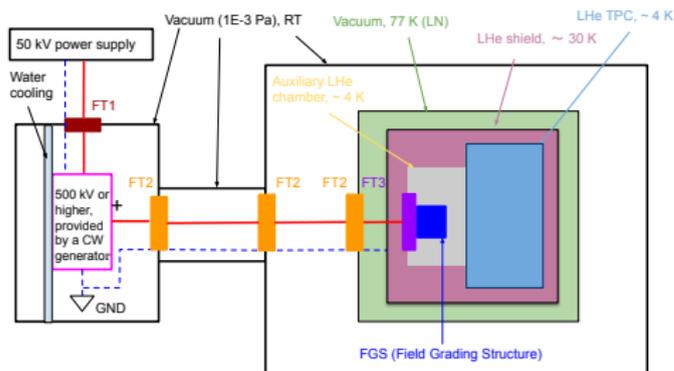
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- Preliminary results: PDE (Photon Detection Efficiency)  $\sim 50\%$  at 10 K and OV+5 V, consistent with DS-20k estimated at LAr temperature, 40%.



# Transmitting ~ MV (Million Volts) into an LHe TPC, 2310.12504

- 10 kV/cm drift field is trade-off to get reasonable drift speed (2 m/s) and fraction of ion-e separation (~50%); 1m size TPC (~ 100 kg LHe) requires 1 MV.



FT1: 50 kV, RT, one side is air, another is vacuum.

FT2: 500 kV or higher, both sides are RT or 77K and vacuum, no need to seal.

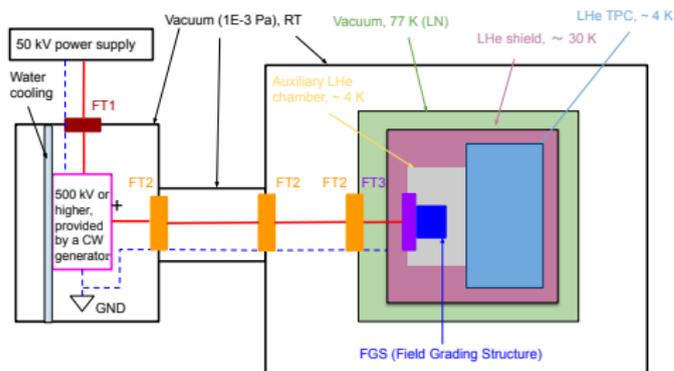
FT3: 500 kV or higher, one side is vacuum and ~ 30 K, another is LHe and 4 K, seal vacuum from LHe.

FGS: Immersed in 4 K LHe.

— : Thin wall tubing (304 SS); - - - : Grounding cable or cathode.

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- Left plot: the preliminary scheme. Right plot: an electrode capable of delivering 100 kV is house-made at CIAE. Testing underway.



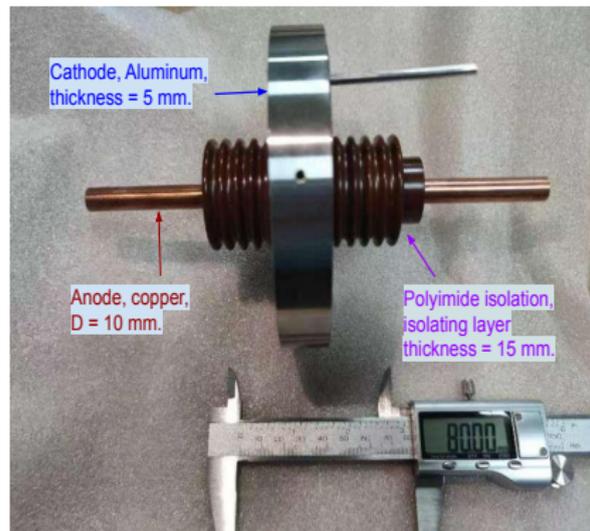
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## A novel NR calibrating method with the COMIMAC facility, 2310.12496

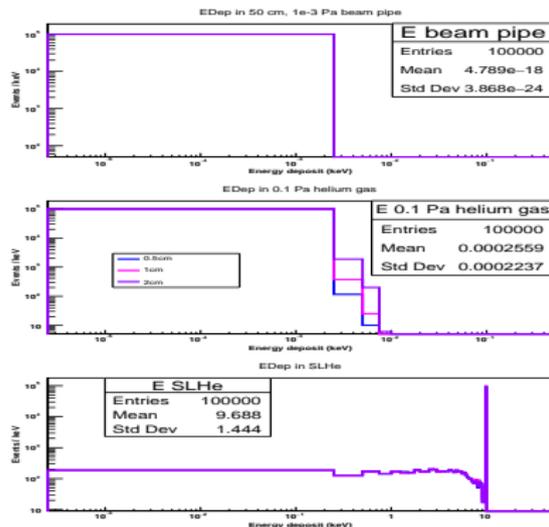
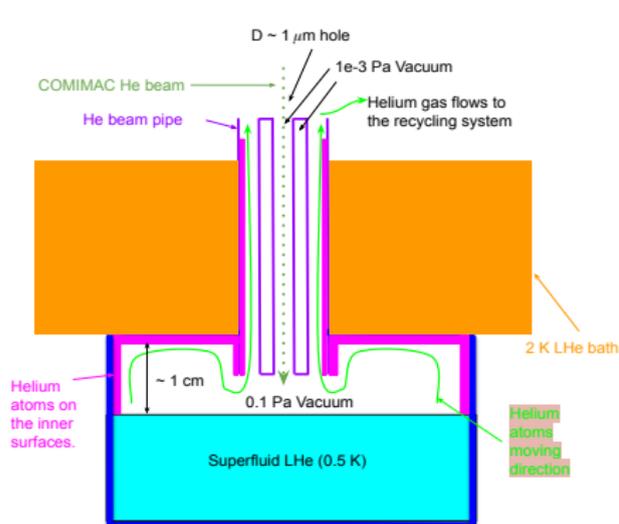
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- Conventional NR calibrations are difficult in providing (i)  $\sim 1 \text{ keV}$  neutrons, (ii) mono-energetic neutrons (accelerator neutrons are not truly mono-energetic).
- The COMIMAC facility, provides helium beam, being implemented in NR calibration for helium gas detector.



# Summary

- DM sector might have more than one elemental particle. DM signals not necessary to show up as NR recoil only: ER-only and ER&NR coexistence also possible.
- ALETHEIA project is supposed to only have single-digit number of ER and NR backgrounds with a 1 ton\*yr exposure, therefore, be sensitive to any kinds of DM signal combinations.
- We demonstrated the viability of a single-phase LHe TPC. The R&D on a dual-phase LHe TPC is underway.

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