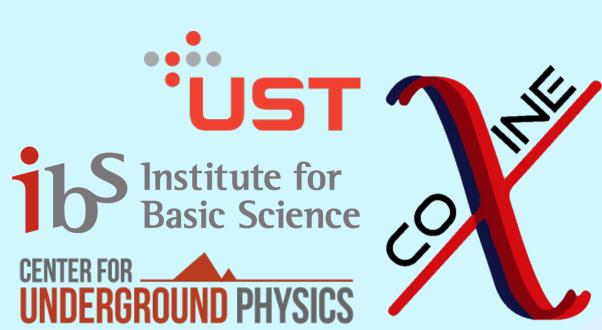
# Low-Threshold Analysis for Low-Mass WIMP Search with COSINE-100

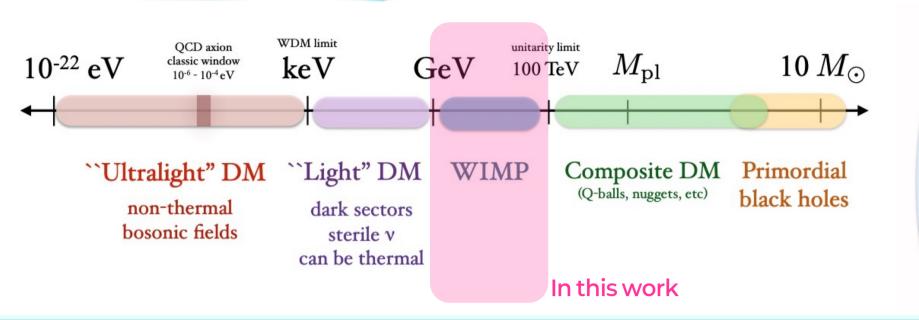


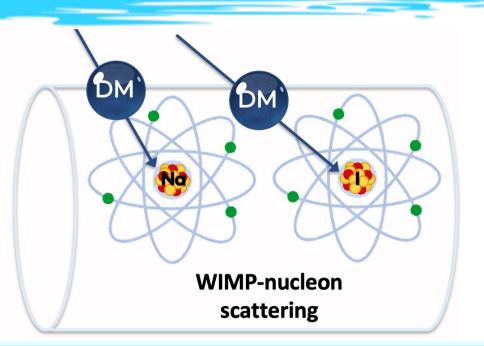
Won Kyung Kim<sup>1,2</sup> On behalf of the COSINE-100 Collaboration

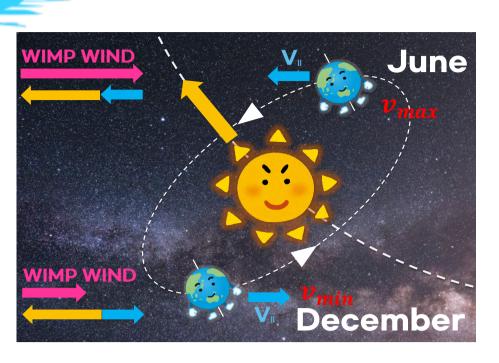
<sup>1</sup>Center for Underground Physics, Institute for Basic Science (IBS), Daejeon, Republic of Korea <sup>2</sup>IBS School, University of Science and Technology (UST), Daejeon, Republic of Korea

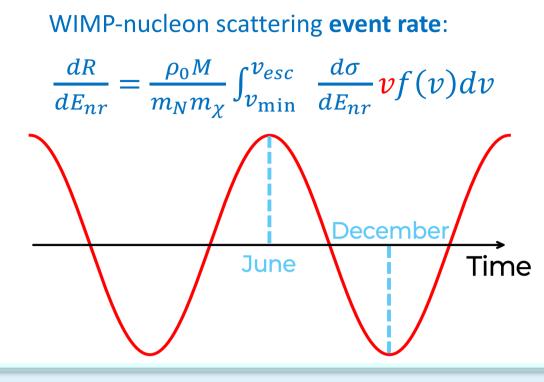


## **Direct Search for Dark Matter**



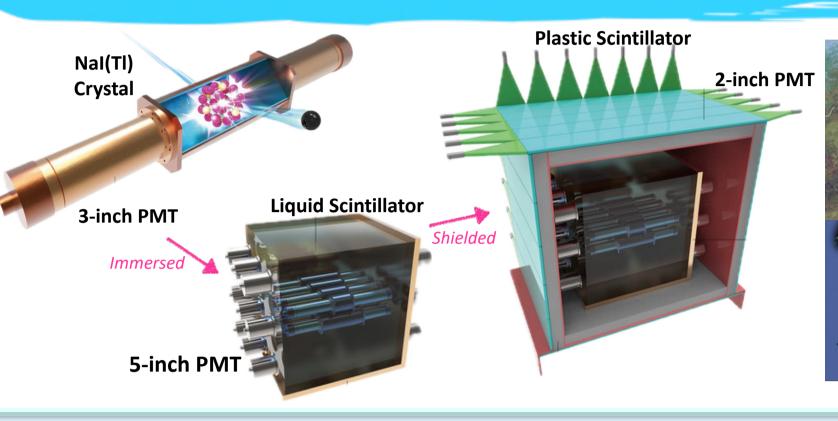


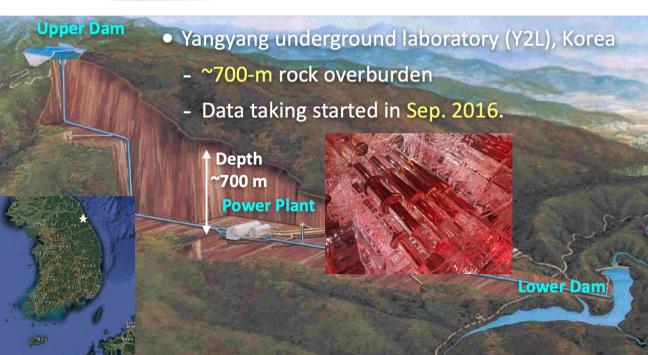


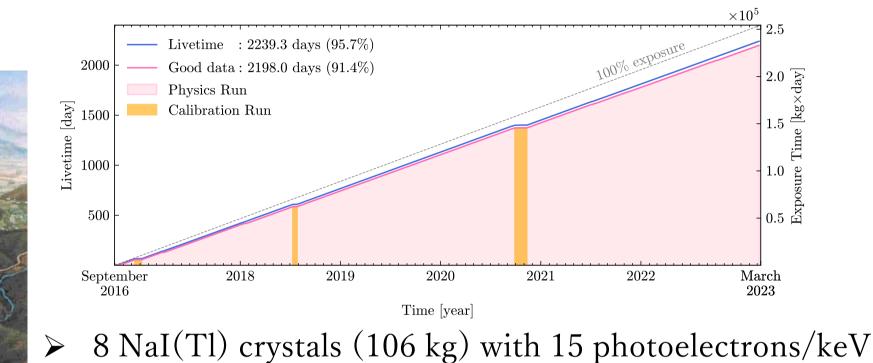


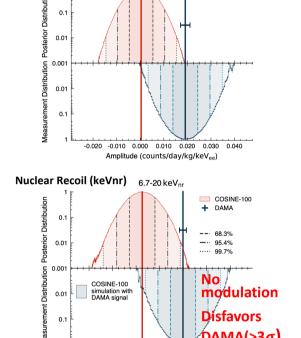
- WIMP-nucleon scattering
- Low energy threshold for low-mass dark matter search
- > Search from annual modulation

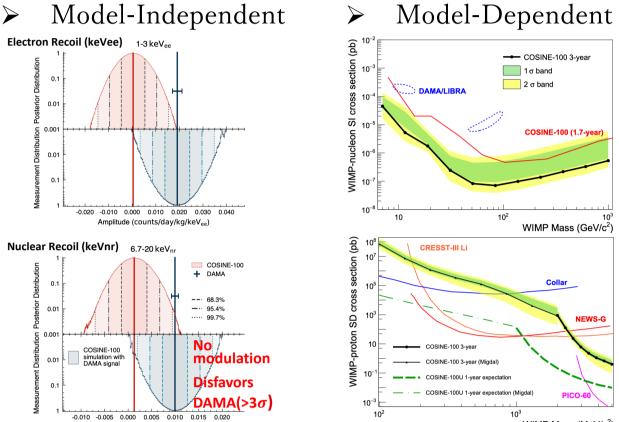
## **COSINE-100 Experiment**







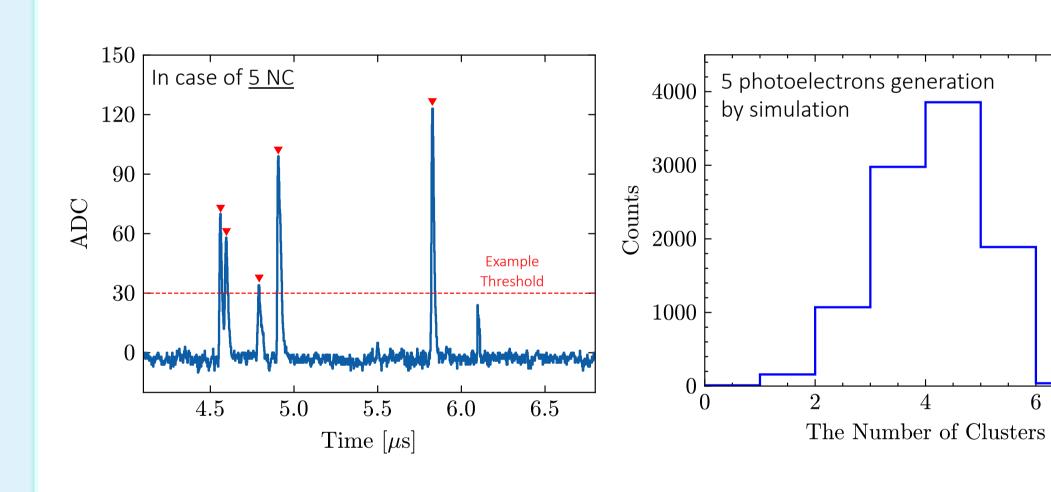




About 6.5 years of data

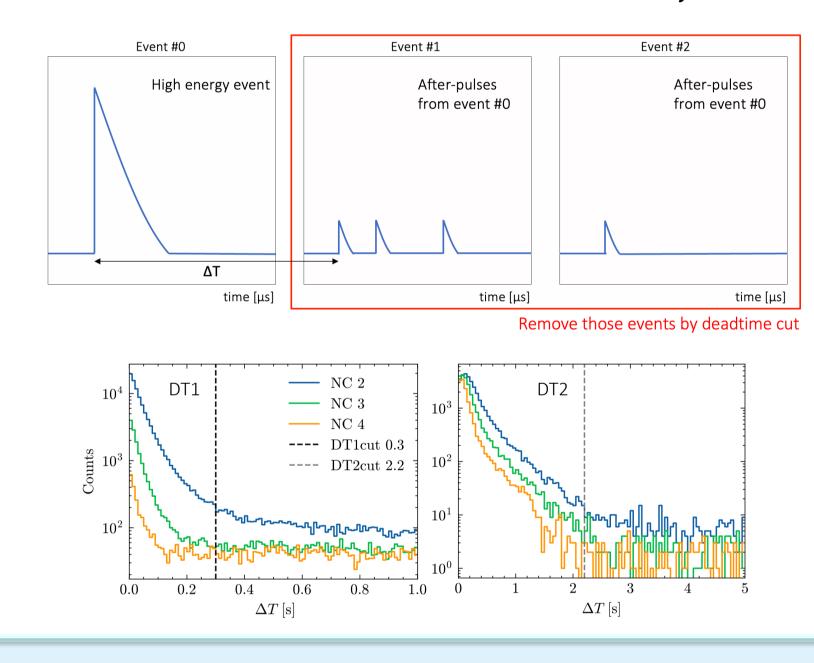
## **Event Selection**

- > Definition of **cluster**: Isolated pulse above threshold
- ➤ Analysis data: 2, 3, 4 number of cluster (**NC**)
- $\triangleright$  1 NC ≈ 1 Photoelectron ≈ 0.067 keV



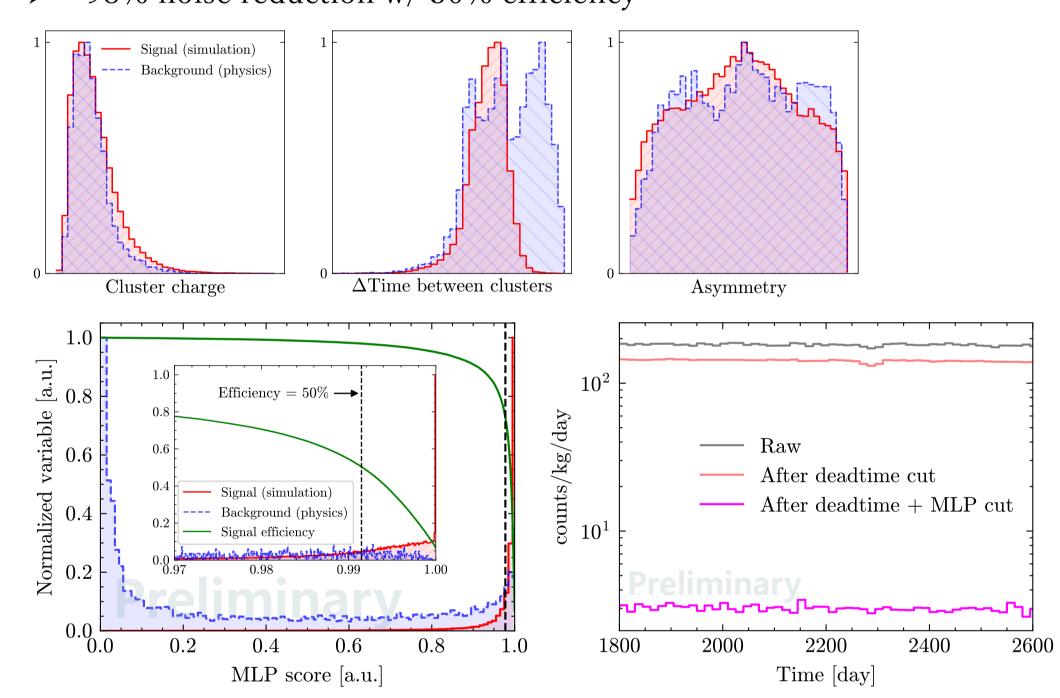
#### Deadtime cut

- Two time criteria for saturation
- 35.5% noise reduction w/ 97% efficiency

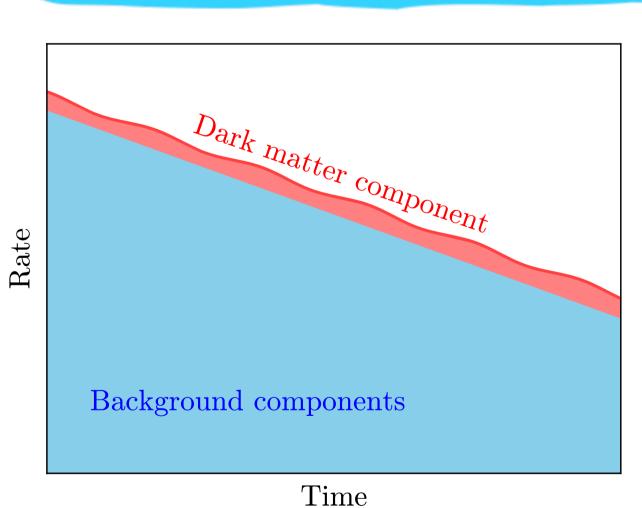


## MLP (Multi Layer Perceptron) training

- > Signal: Simulated data Background: COSINE-100 data
- 98% noise reduction w/ 50% efficiency



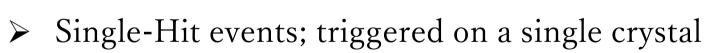
## **Extraction of Modulation Signal & Detector Sensitivity**

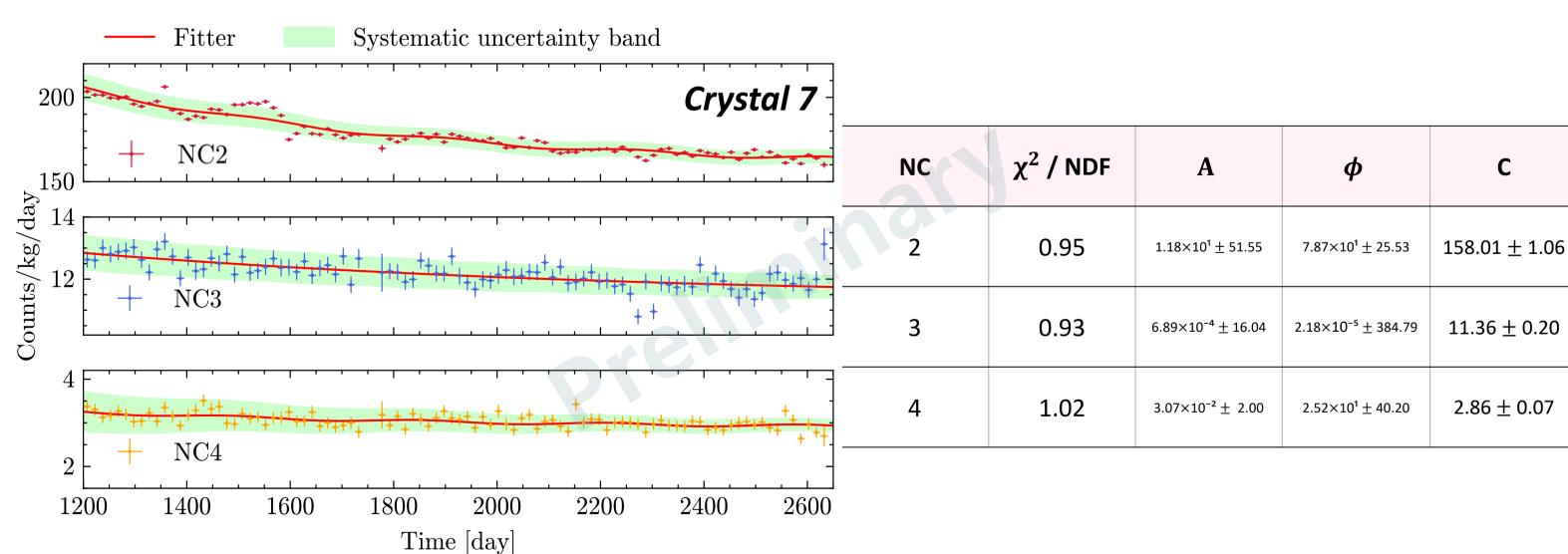


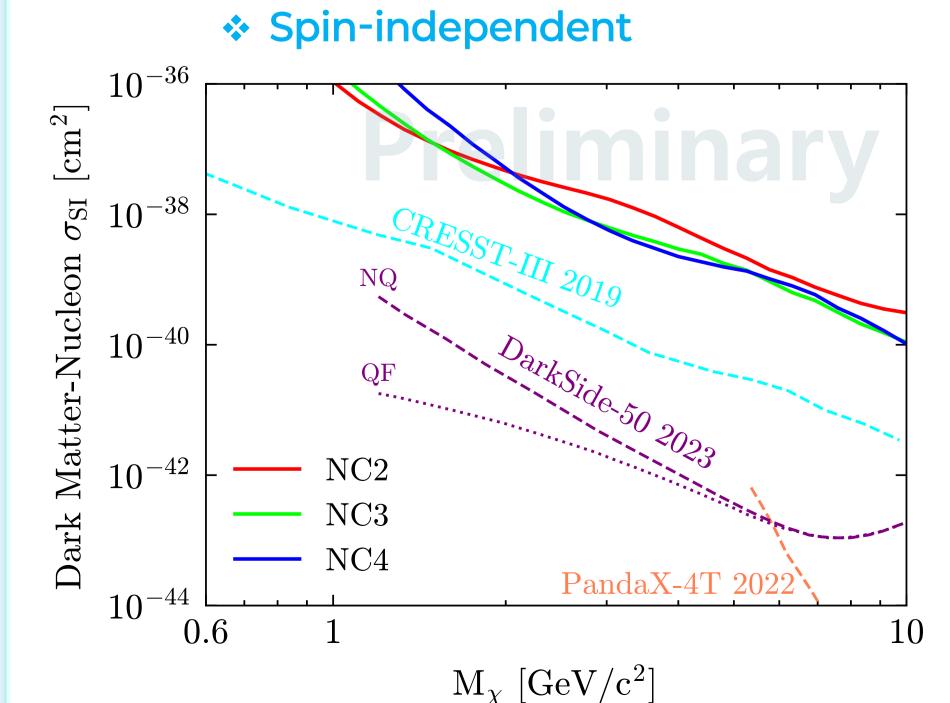
> Annual modulation fit with,

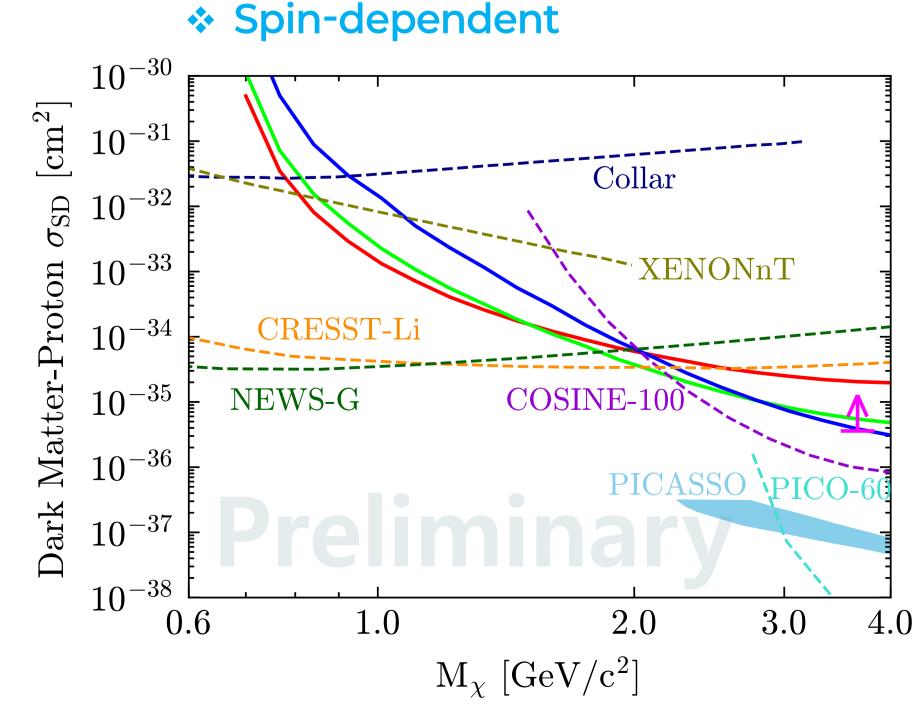
$$f(t) = A \cos(\frac{2\pi}{T}(t - \phi)) + e^{p_0 + p_1 t} + C$$

- A: modulation amplitude
- φ : phase
- T: 1 year = 365.25 days
- $p_0, p_1$ : background (exponential component)
- C: constant baseline









- Sensitivity estimation with pseudo data which based on real data
- 10,000 pseudo data generation & fitting to get an average 90% confidence limit
- > We should conduct simultaneous fit for every crystal and NC
- Systematic is not yet considered (going to be considered)
- $\sigma_{SD} \leq 3.57 \text{ pb } (90\% \text{ C.L.}) @ \text{M}\chi = 3.65 \text{ GeV}$
- > Our analysis shows the potential to explore regions of the low-mass WIMP parameter space

## Reference

- [1] G. Adhikari et al., "An experiment to search for dark-matter interactions using sodium iodide detectors", Nature 564, 83–86 (2018). [2] C. Ha et al., "Initial performance of the COSINE-100 experiment", Eur. Phys. J. C 78, 107 (2018). [3] G. H. Yu et al., "Improved background modeling for dark matter search with COSINE-100", Eur. Phys. J. C 85, 32 (2025).
- [4] C. Savage et al., "Compatibility of DAMA/LIBRA dark matter detection with other searches", JCAP 04, 010 (2009). [5] S. Derenzo et al., "Direct detection of sub-GeV dark matter with scintillating targets", Phys. Rev. D 96, 016026 (2017). [6] S. M. Lee et al., "COSINE-100 full dataset challenges the annual modulation signal of DAMA/LIBRA", arXiv:2409.13226 (2024).

[7] G. H. Yu et al., "Limits on WIMP dark matter with NaI(Tl) crystals in three years of COSINE-100 data", arXiv:2501.13665 (2025).