

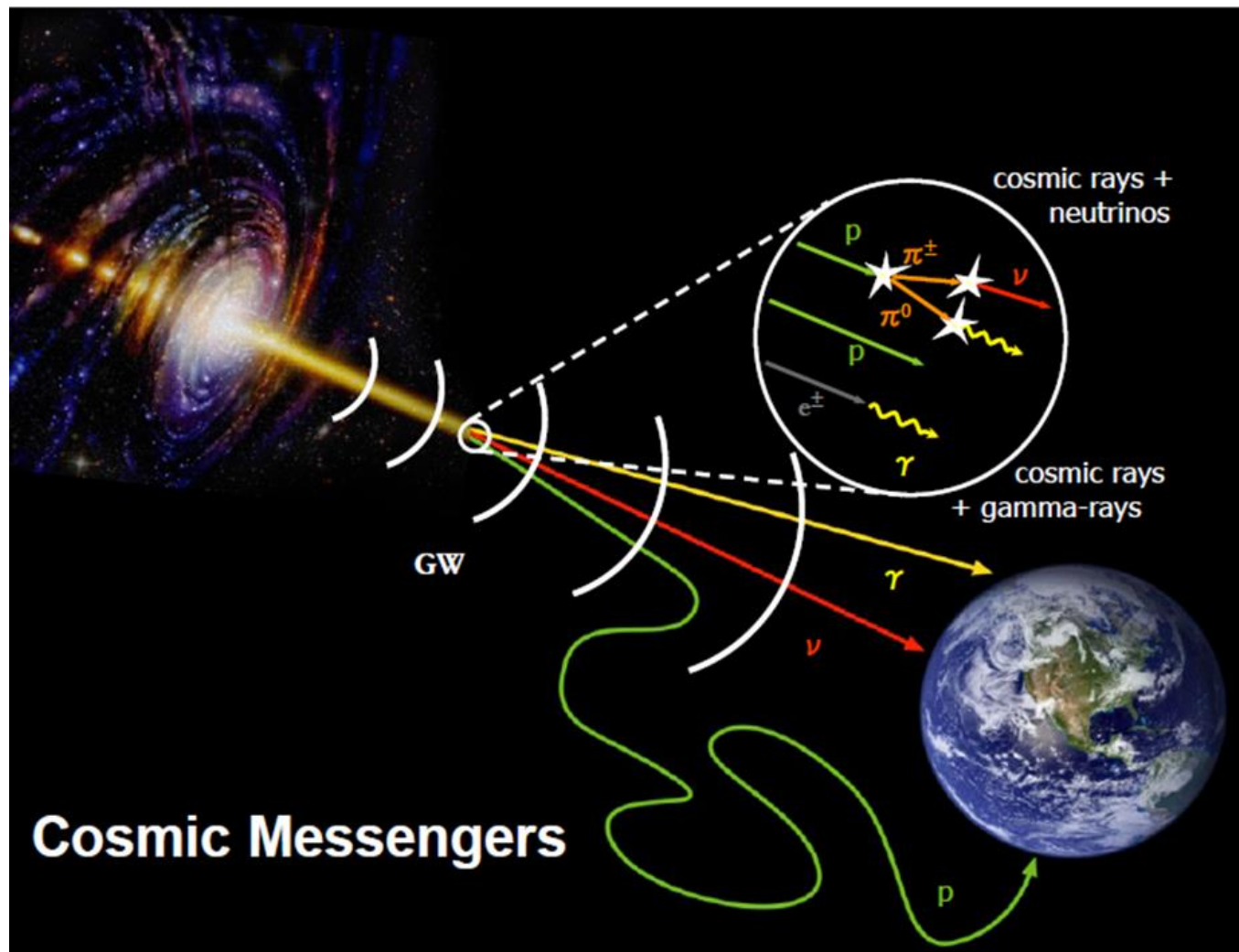


# Prototype Test of a Water Cherenkov Detector Designed Based on SWGO Lake Concept

Ziqi Huang

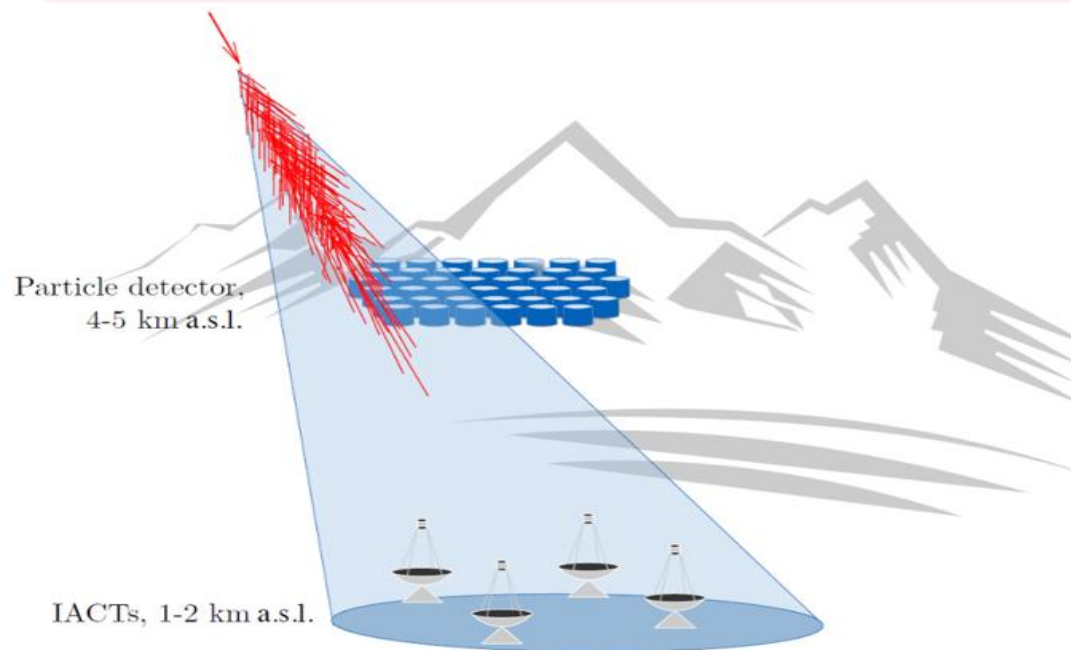
Shandong University & Institute of High Energy Physics, CAS

# Searching for the Origin of UHECRs



- ◎ The Origin of Ultra-High-Energy Gamma-Ray (UHECR) : A “Mystery of the Century”.
- ◎ A recognized grand challenge highlighted by the U.S. National Research Council and Science magazine.
- ◎ Multi-messenger Astrophysics: Gravitational-wave, Neutrino, Cosmic ray and Gamma ray.

# Ground-based gamma-ray astronomy

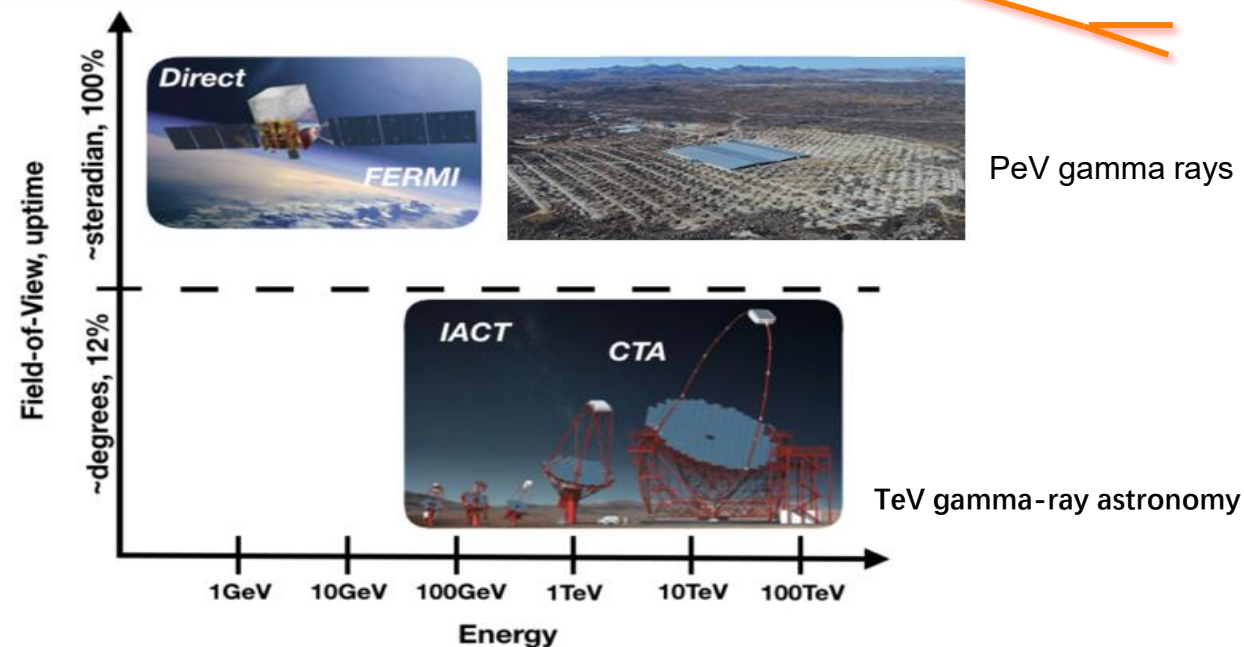


## ☉ Imaging Atmospheric Cherenkov Telescope (IACT) arrays:

( H.E.S.S., VERITAS, MAGIC, CTA ...)

- Angular resolution:  $0.05^\circ - 0.1^\circ$ ;
- Duty cycle:  $\sim 10\%$ ;
- field of view:  $< 5^\circ$ ;
- Energy threshold:  $< 100 \text{ GeV}$ ;

➔ Mainly focused on deep observation.



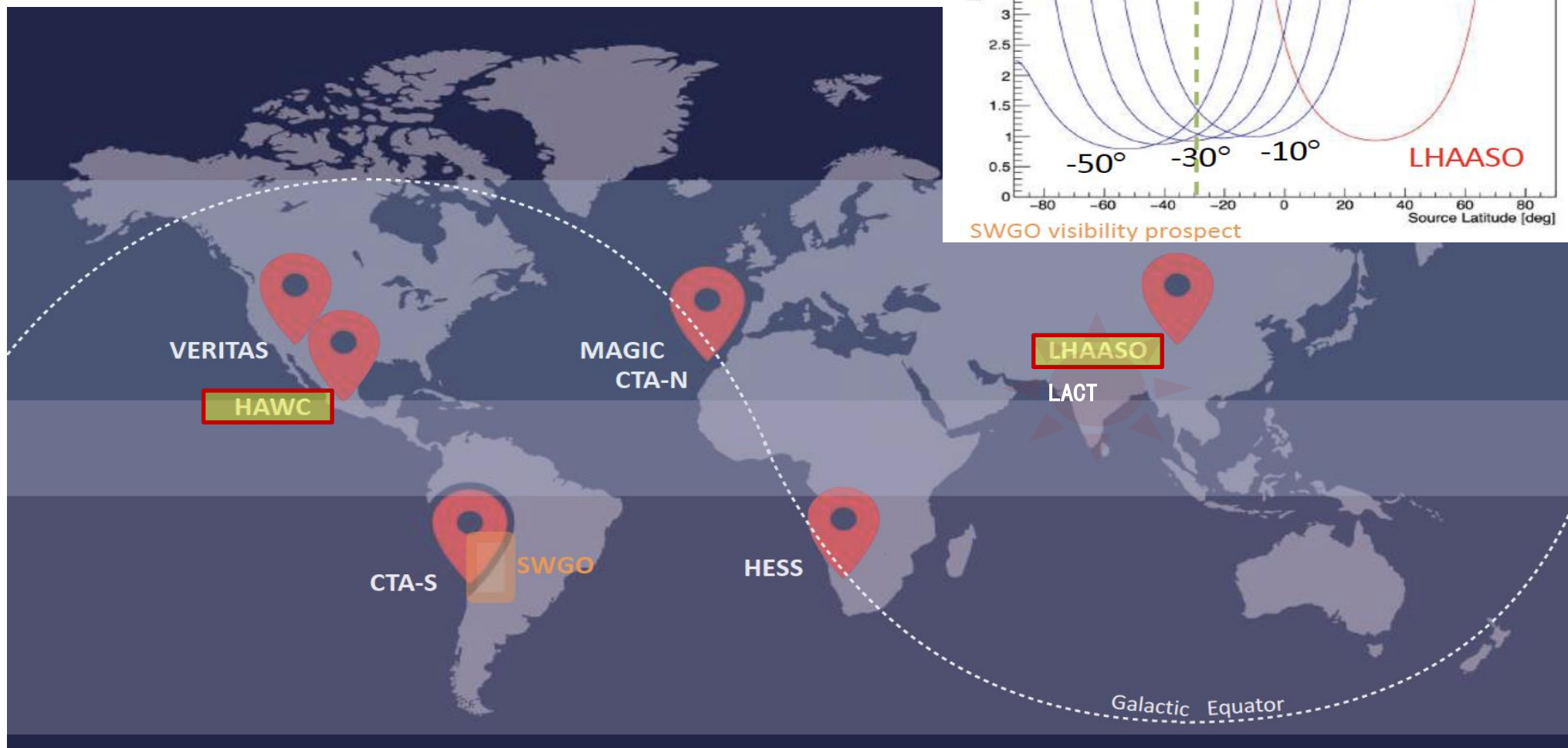
## ◆ Ground-level detector arrays :

(AS $\gamma$ , ARGO-YBJ, Milagro, HAWC, LHAASO...)

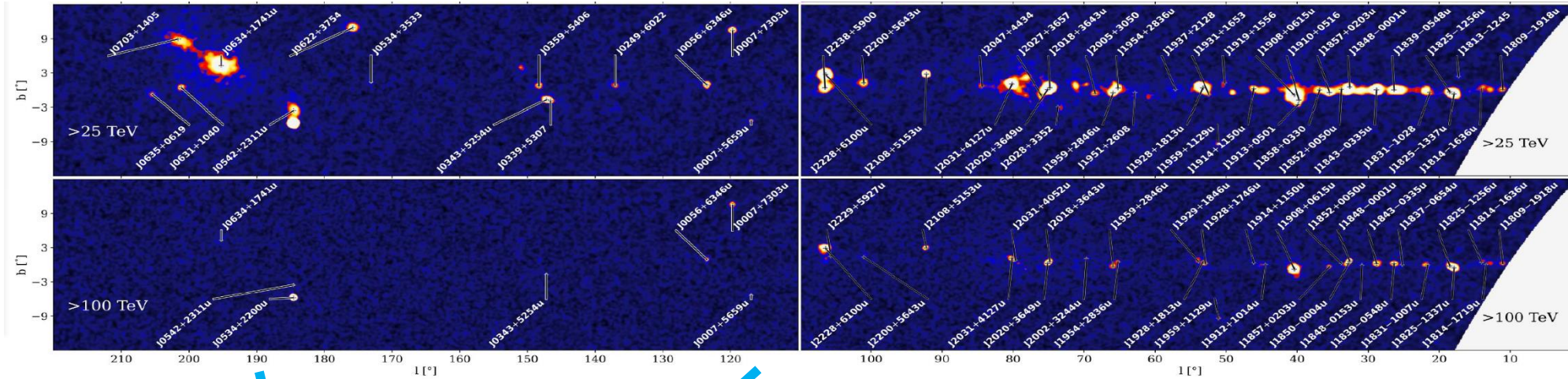
- ◆ Angular resolution:  $0.2 \sim 0.5^\circ$ ;
- ◆ Duty cycle:  $> 95\%$ ;
- ◆ field of view:  $> 2/3\pi$ ;
- ◆ Energy threshold :  $> 1 \text{ TeV}$ ;

➔ Good at sky survey, extended sources.

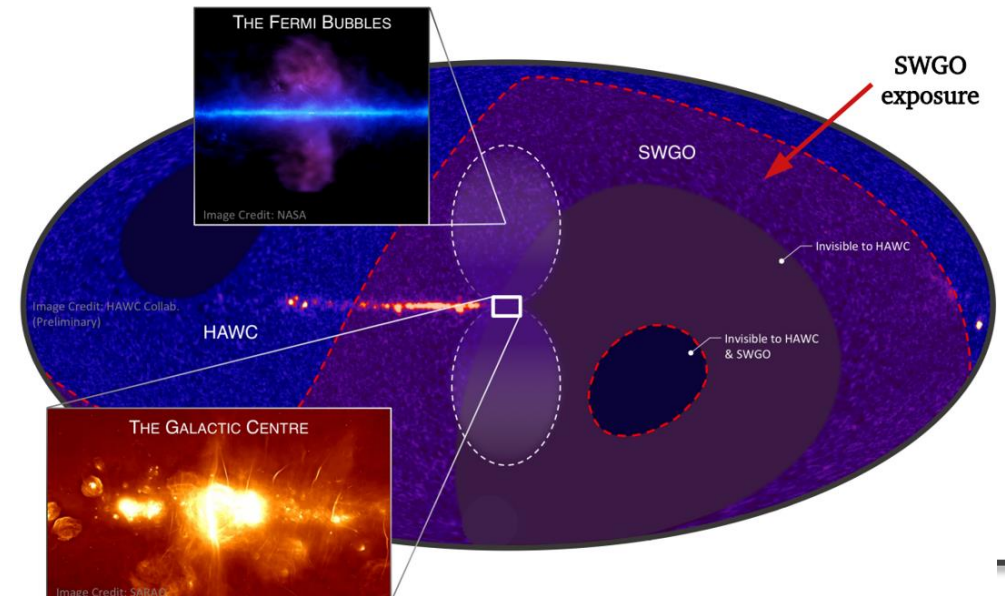
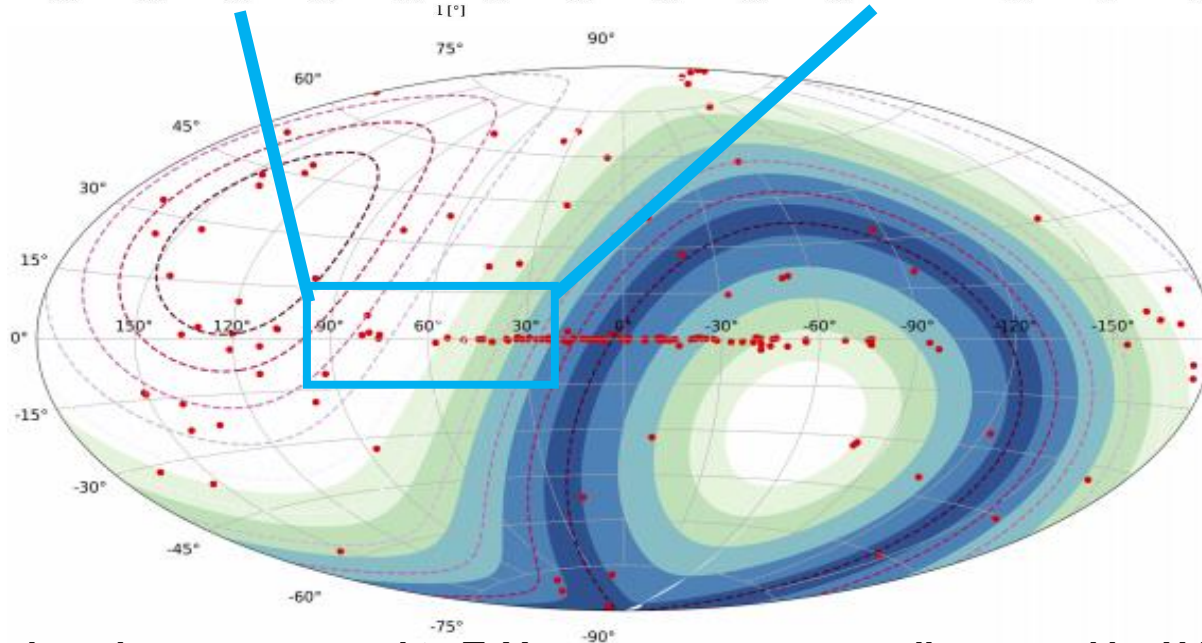
# Global Gamma-Ray Observatories Location



# SWGO: First Wide-field Instrument to the South Sky



TEV gamma-ray sources detected by LHAASO



The red markers correspond to TeV gamma-ray sources discovered by H.E.S.S., MAGIC and VERITAS

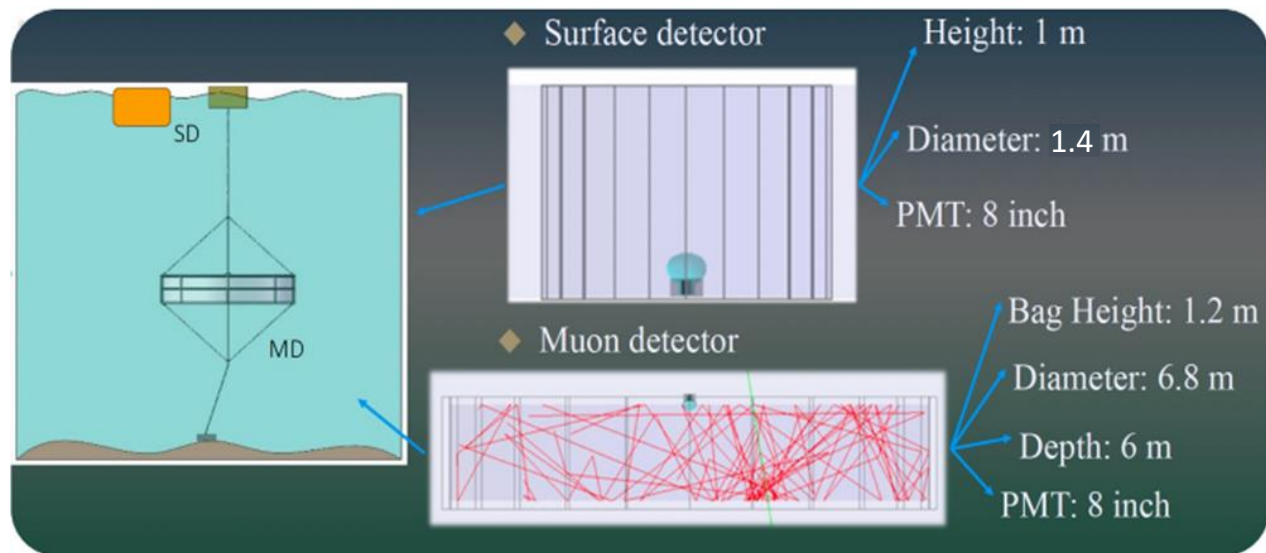
# SWGO Lake Concept



LHAASO MD: nearly 600,000 cubic meters construction volume.

## Project Goals:

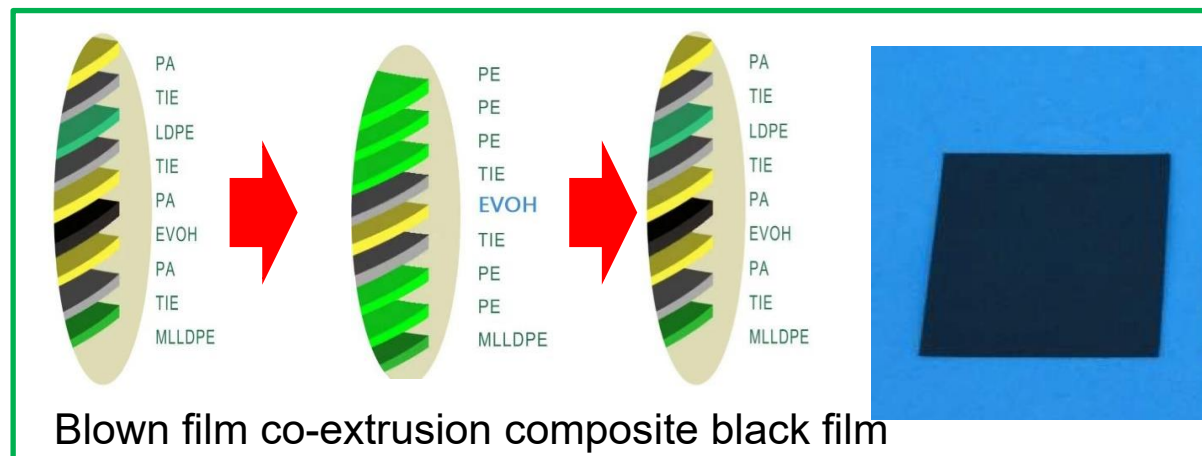
- Sensitive at UHE
- significantly larger area than LHAASO
- Design to suit conditions in South America.
- Lake concept



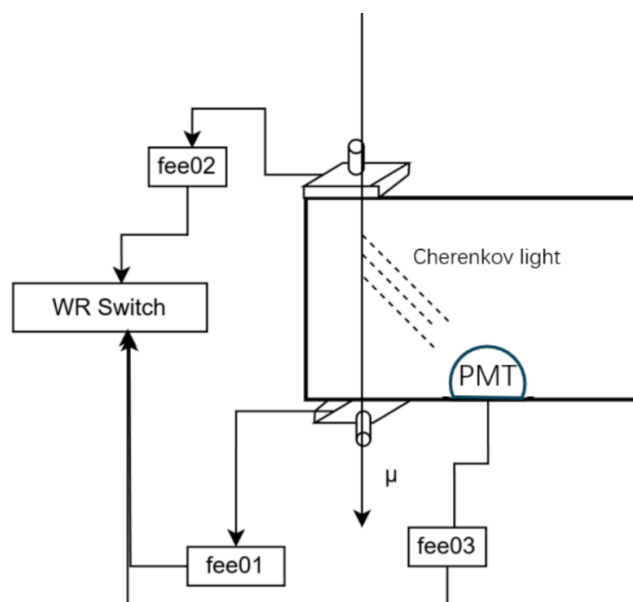
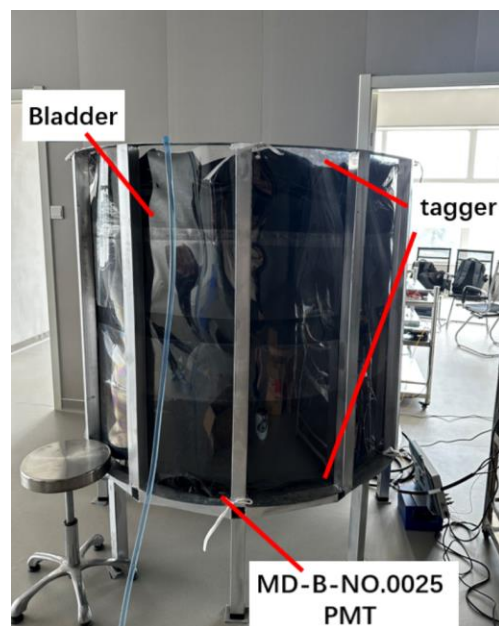
## Advantage:

- low cost
- modular design
- easy to install
- Artificial/Natural Lake

# Prototype Surface Detector test in lab

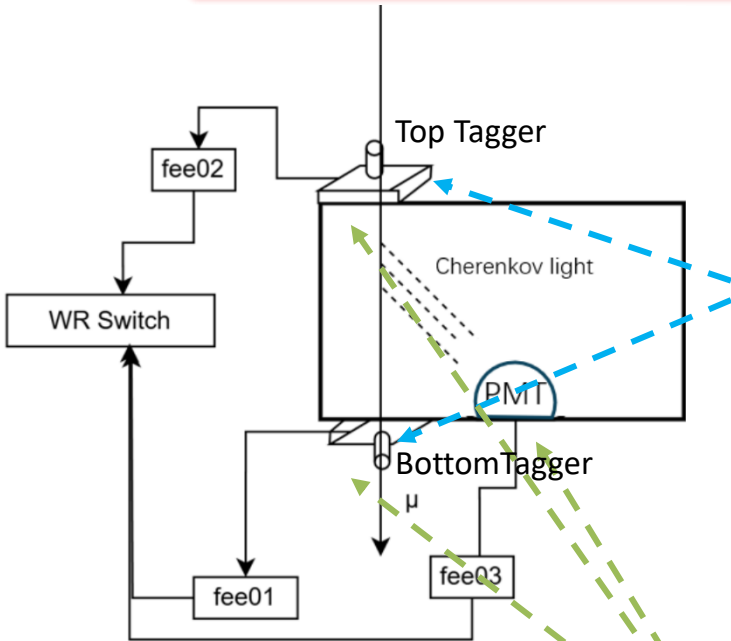


- Black film for bladder
  - multi-layer co-extrusion technology
  - PE + EVOH + MLLDPE
  - light-tight and durable



- Geometry: Diameter 1.4 m Height 80/106 cm
- 8-inch PMT(R5912) gain:  $6.7e5$
- threshold: 7mV( $\sim 5$  SPE)
- window: 100ns

# Event Rate and Detection Efficiency

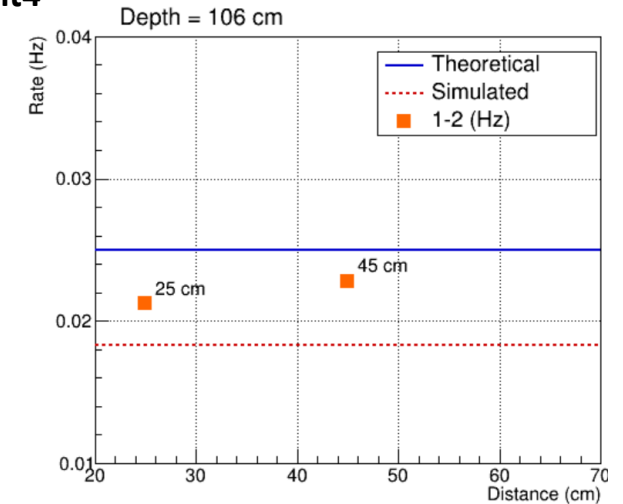
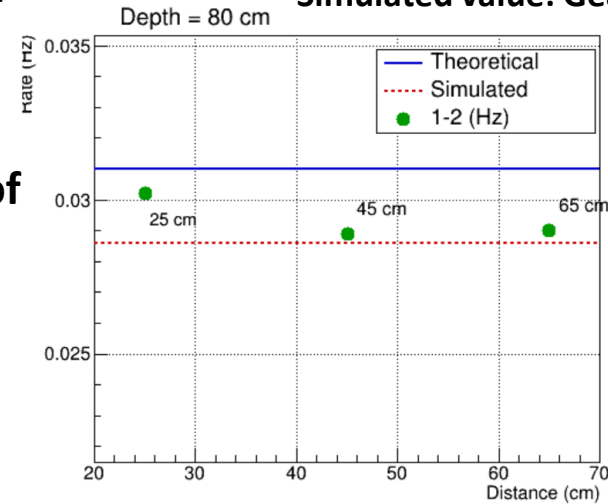


Depth: 80cm/ 106cm

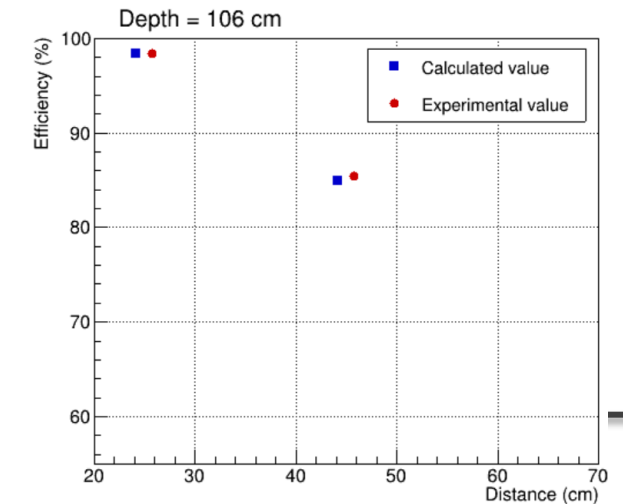
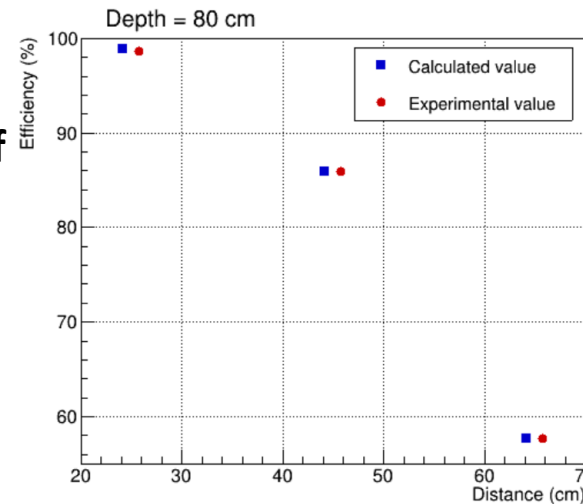
Event rate:  
2-coincidence rate of  
tagger 1,2

Detection efficiency:  
3-coincidence rate of  
tagger 1,2 and PMT

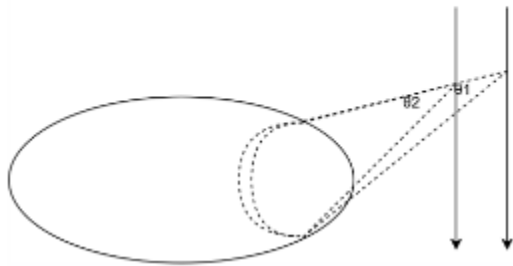
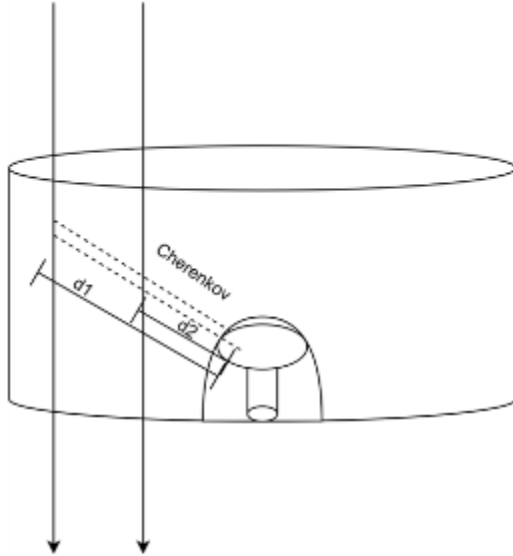
Theoretical value: calculated by dark count rate  
Simulated value: Geant4



change the position of taggers



# Light Yield Analysis



In the experiment, the measured results are as follows: The total number of photons generated between wavelengths  $\lambda_1$  and  $\lambda_2$  in a medium of length  $L$  is given by:

$$N = \int_{\lambda_1}^{\lambda_2} \frac{2\pi L \alpha z^2}{\lambda_1^2} \sin^2 \theta_c d\lambda$$

If only the visible light region is considered, i.e.,  $\lambda = 400 - 700 \text{ nm}$ , the number of photons per unit length is:

$$N/L = 490 \times \sin^2 \theta_c$$

To approximate the theoretical number of photons, consider the following approximations:

1.  $\theta_c \approx 41^\circ$
2. The light-sensitive surface is approximately circular with a diameter of  $\sim 20 \text{ cm}$ .
3. Let  $r$  be the distance from the probe to MD-B-NO.0025, and  $d$  be the distance from the Cherenkov light emission point to the receiving surface. Then:

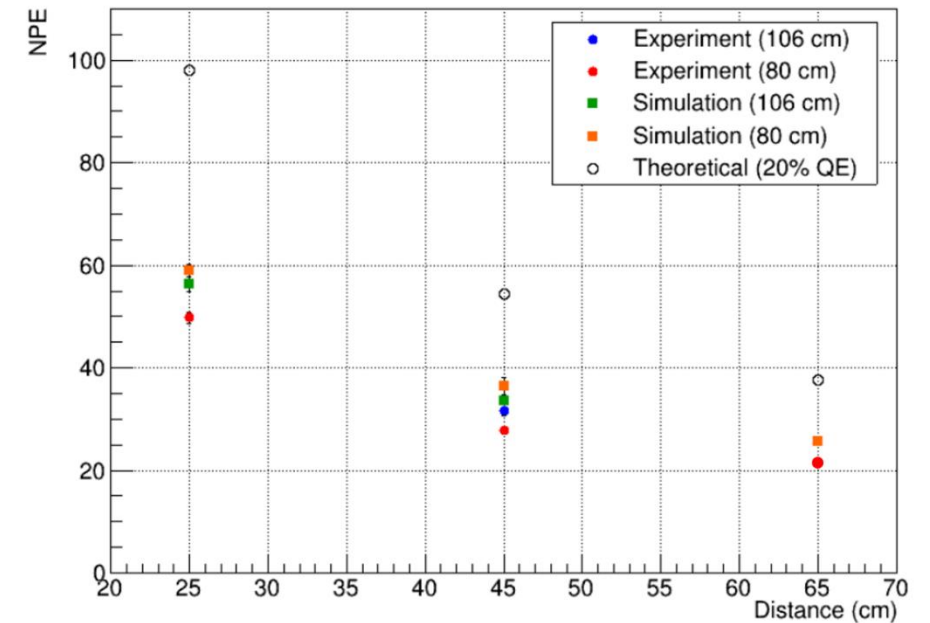
$$d = r / \tan \theta_c$$

The theoretical number of photons is:

$$N_{theory} \approx 490 \text{ cm}^{-1} \times \frac{1}{2} \times 20 \text{ cm} \times \frac{\pi \times 10^2}{\pi ((r+10)^2 - (r-10)^2)} = \frac{12250}{r}$$

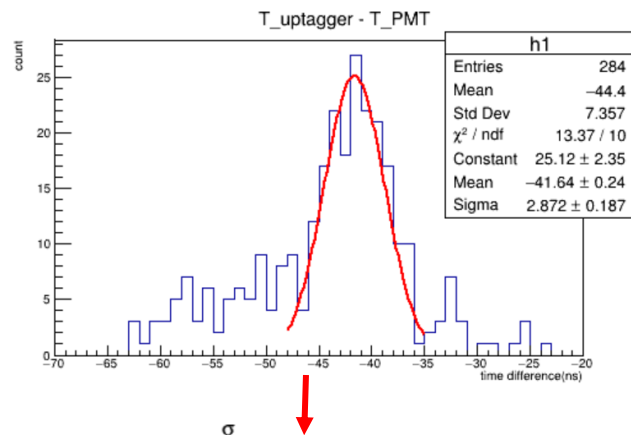
Water Depth (cm)	Distance from Center (cm)	NPE (p.e.)	product
106	25	$56.66 \pm 0.87$	1416.5
106	45	$31.74 \pm 1.16$	1428.3
80	25	$49.75 \pm 1.16$	1243.5
80	45	$27.82 \pm 0.41$	1251.9
80	65	$21.02 \pm 0.78$	1366.3

NPE vs Distance



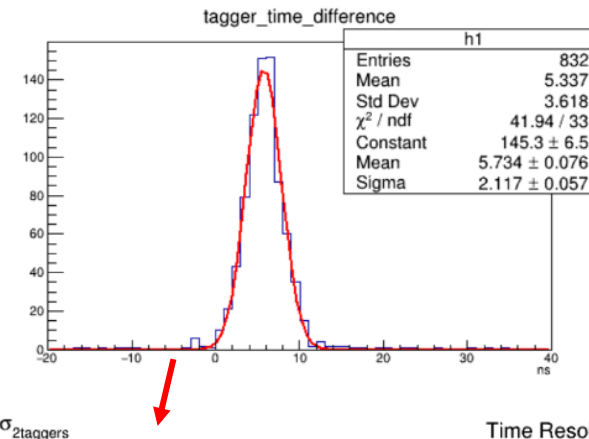
# Time Resolution

Time\_uptagger – Time\_PMT

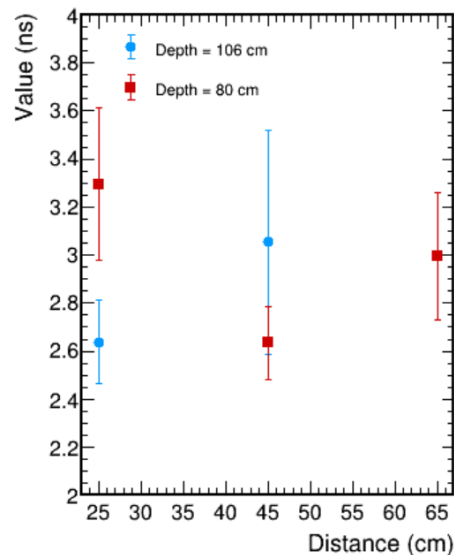


Fit with gaussian curve to get  $\sigma$  as time jitter

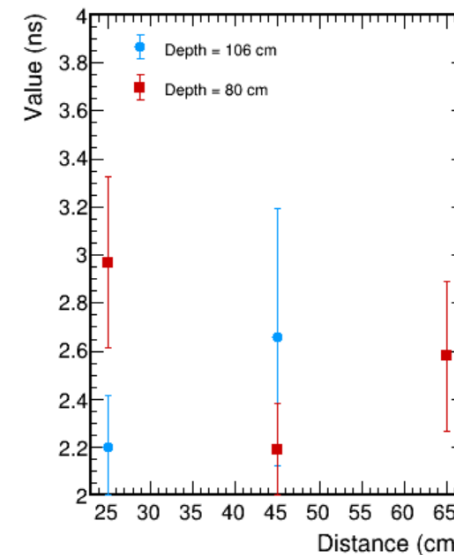
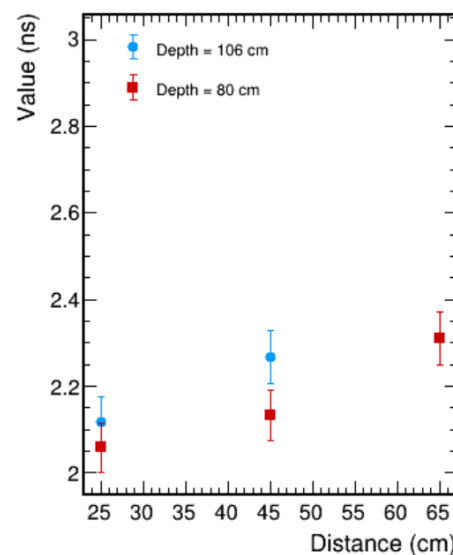
Time difference between taggers



$$\sigma_{\text{tagger}} = \frac{\sigma_{2\text{taggers}}}{\sqrt{2}}$$



Remove the influence of the tagger's intrinsic time jitter



$$\sigma_{\text{detector}} = \sqrt{\sigma_{\text{all}}^2 - \sigma_{\text{tagger}}^2}$$

# Installation at LHAASO site



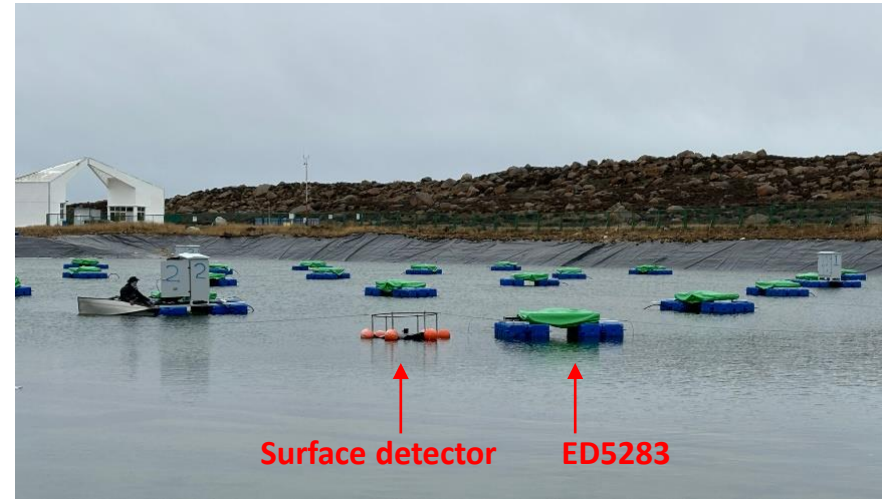
a small on-water  
array and  
experimental  
platform in the  
lake at the  
LHAASO Site



flange-mounted PMT



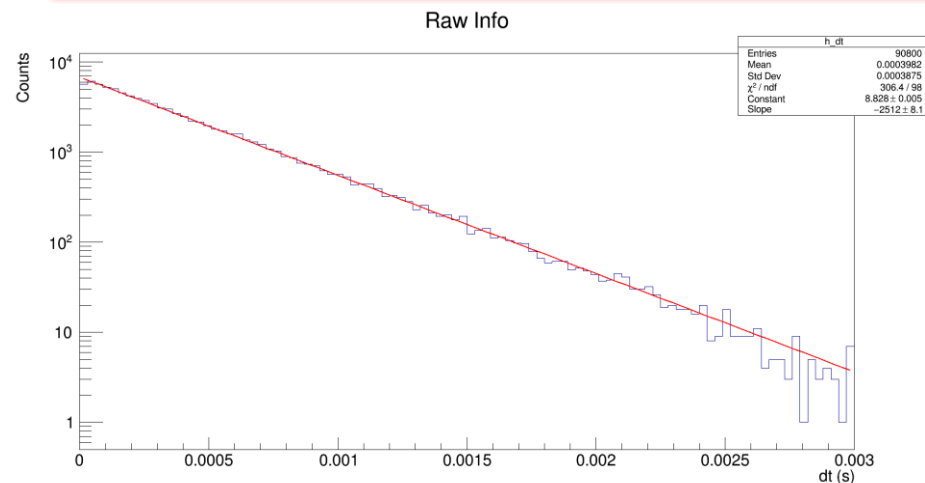
Detector Prototype Deployment



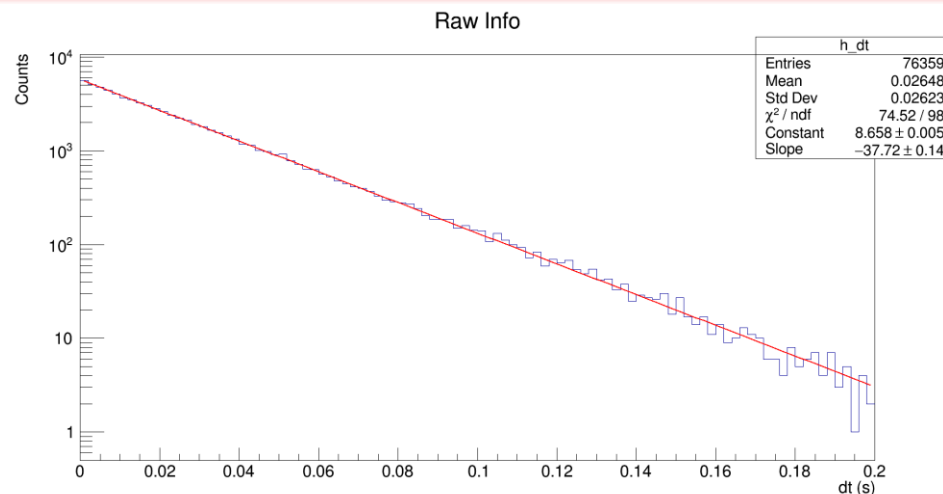
Surface detector

ED5283

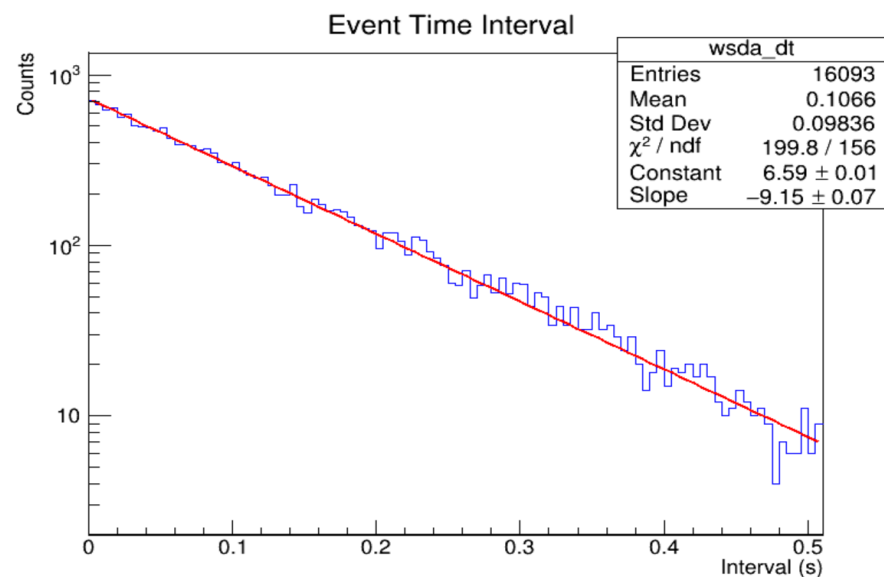
# Coincidence with KM2A



**LHAASO-KM2A event rate:  $\sim 2500\text{Hz}$**



**WSDA (water surface detector array) event rate:  $\sim 37\text{Hz}$**



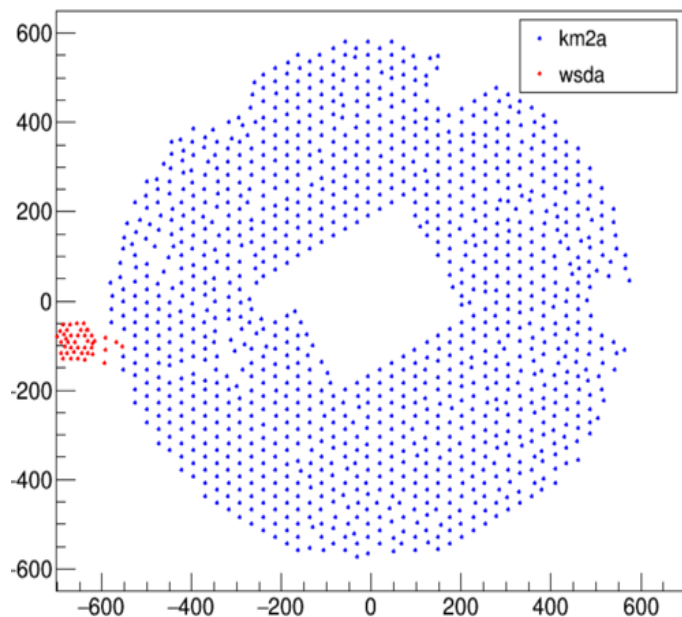
matching timestamps with a 1000ns window

Coincidence rate of WSDA and KM2A:  $\sim 9.15\text{Hz}$ .

Accident coincidence rate :  $\sim 0.19\text{Hz} \ll 9.15\text{Hz}$

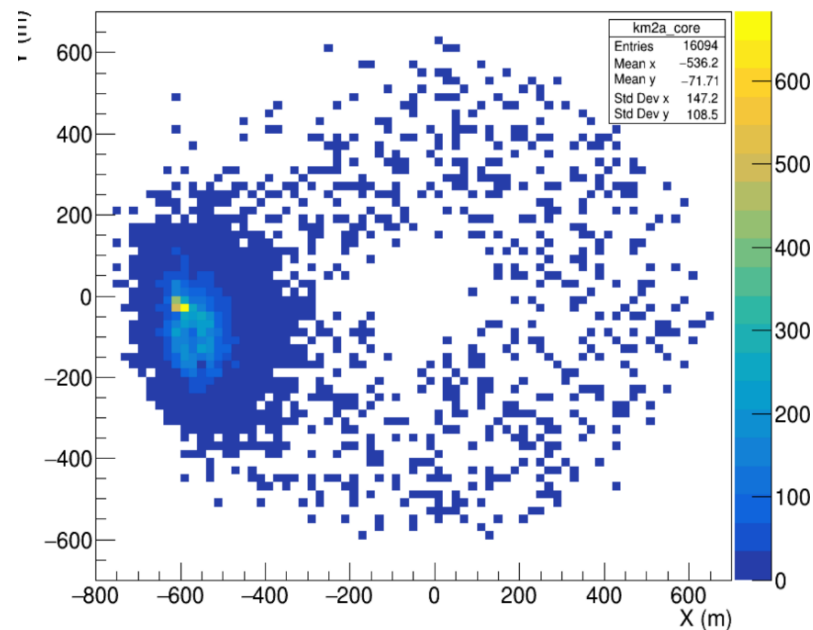
# Shower Reconstruction

42 EDs + 1188 MDs+WCD



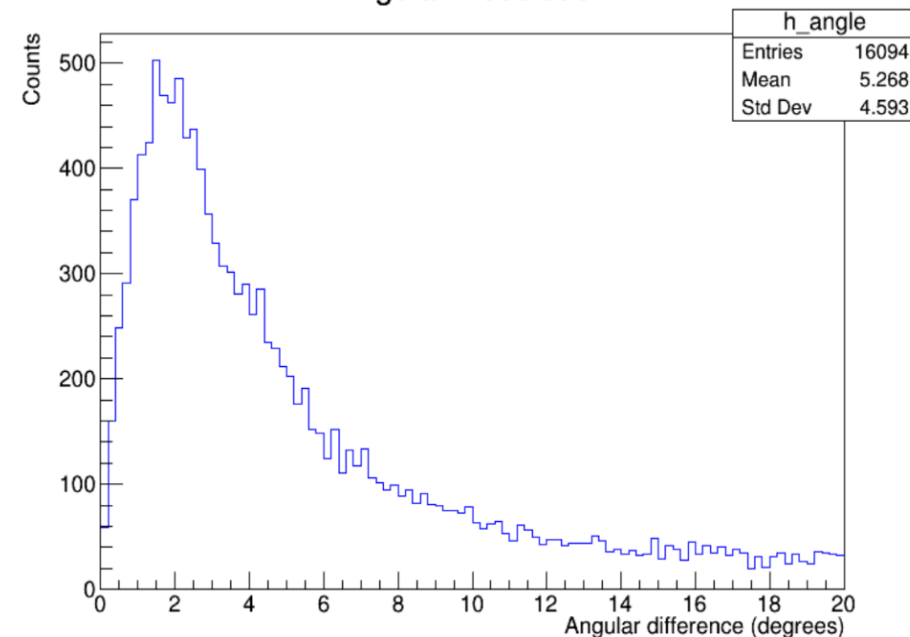
Position of WSDA and KM2A

Core Position



Core position reconstructed by KM2A

Angular Resolution



Space angular between WSDA and KM2A

# occupancy

Detector	Active area	Coincidence event	Hit number	Triggered by shower
ED5283	$\sim 1\text{m}^2$	16094	7057	6588
WCD	$\sim 1.5\text{m}^2$ (1m height)	16904	7528	7161

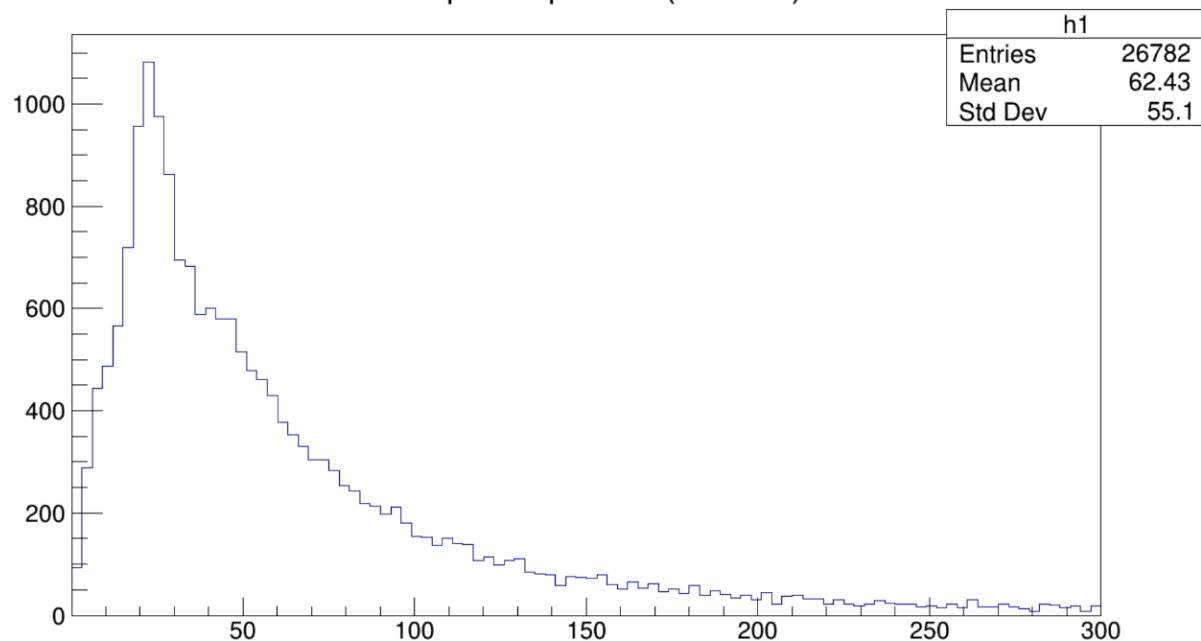
Significantly larger area, why not significantly higher occupancy?

- Cherenkov emission mechanism: charged particle and energy restrict
- The uniformity of detector: lower detection efficiency

✗ low energy  $\gamma$

# Single particle peak

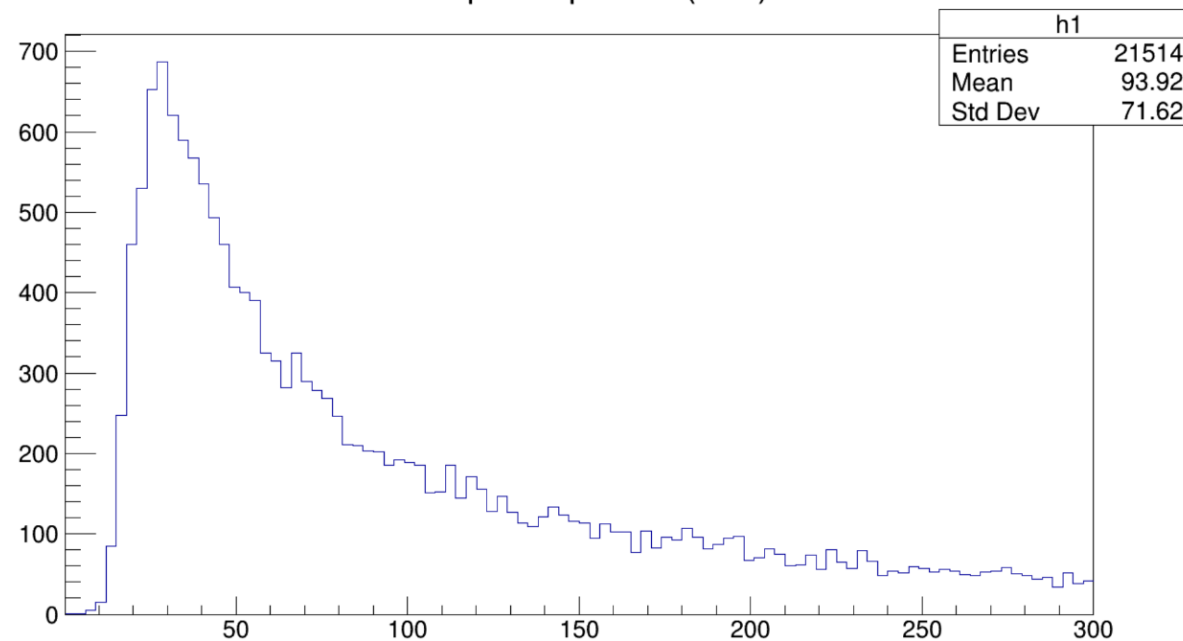
hits.qa-hits.peda/8 (ed5283)



ED: a counter for the number of particles

Charge: mostly around several tens of pCs.

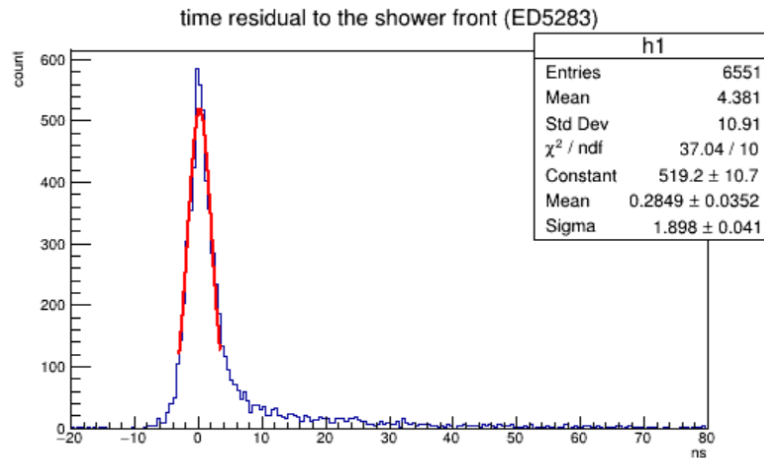
hits.qa-hits.peda/8 (wcd)



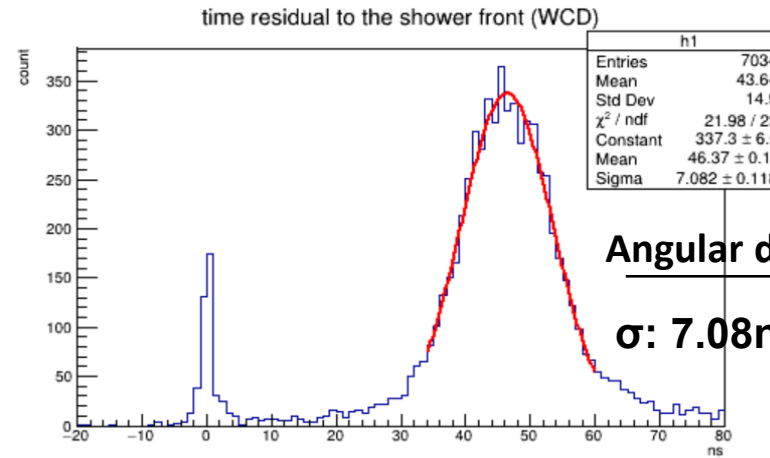
WCD: functions more like a calorimeter.

Charge: extend to several hundred pCs and even higher.

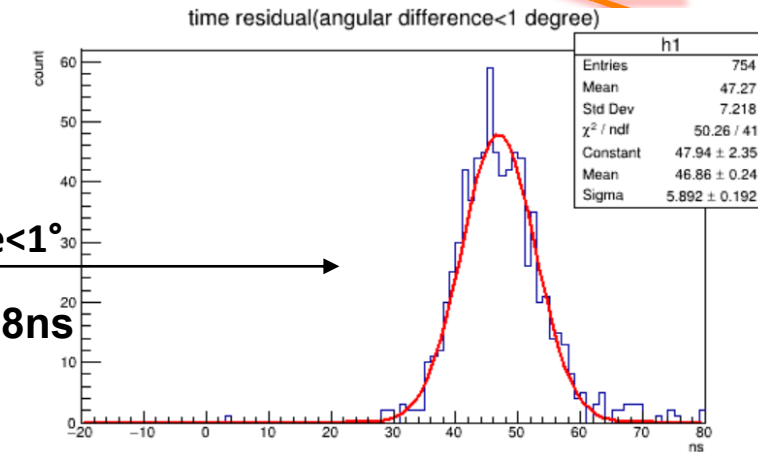
# Time residual to shower front



ED5283 time residual

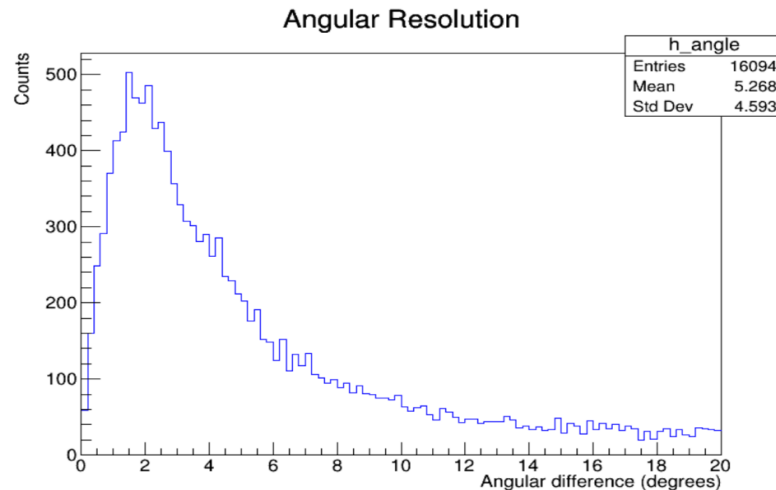


WCD time residual



WCD time residual

Angular difference < 1°  
 $\sigma: 7.08\text{ns} \rightarrow 5.98\text{ns}$



Space angular between WSDA and KM2A

## Influence factor

- Signal of electromagnetic particle / muon
- Lake environment
- Reconstruction accuracy of WSDA

# Summary

- ◎ **Muon & EM Tests Completed**
- ◎ **Room for Performance Improvement**
- ◎ **Valuable Experience for Surface Detector Construction**
- ◎ **Further plans:**

- **Adjusting the thresholds to test the changes in performance.**
- **Exploring optimizations for the detector structure.**
- **Conducting simulations for the full array.**

**Thank you!**

# backup

# Prototype Muon Detector and Underwater test

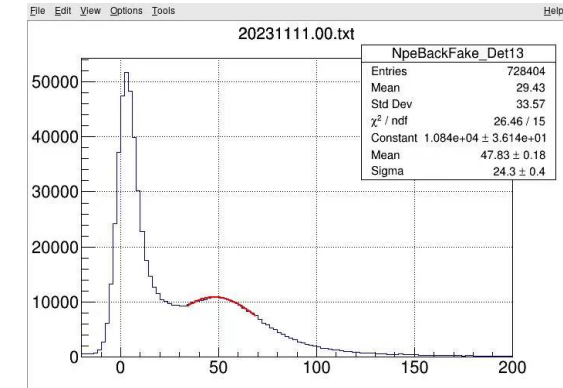
- ① Fill the bladder with the water
- ② Verify the installation on the water surface
- ③ The bladder is installed @Qingdao and LHAASO site respectively
- ④ Some data is collected @LHAASO site, the single muon signal and water attenuation length are similar as those of LHAASO Muon detector.



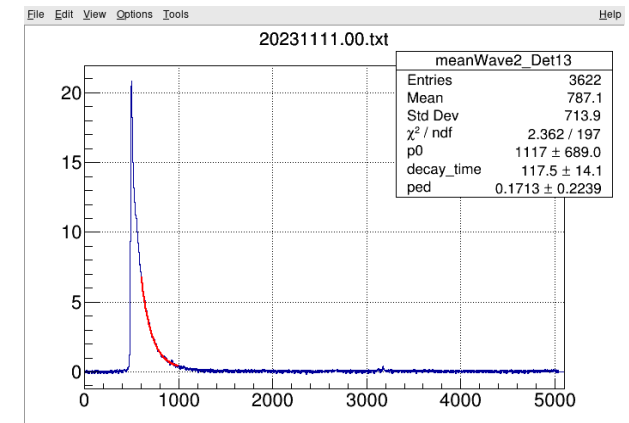
installation in the lake @Qingdao



Installation @ LHAASO

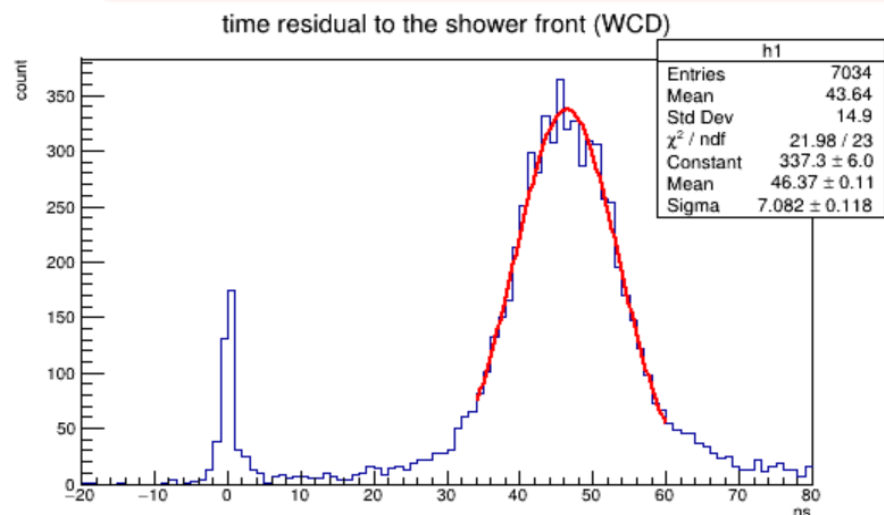


Charge resolution ~ 50%;  
Mixed with Punch-through electromagnetic signals

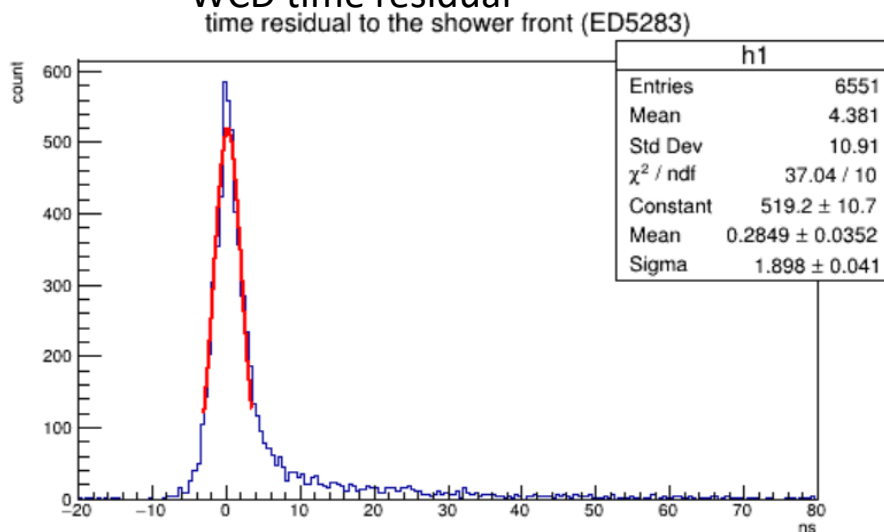


Decay time: 117.5ns

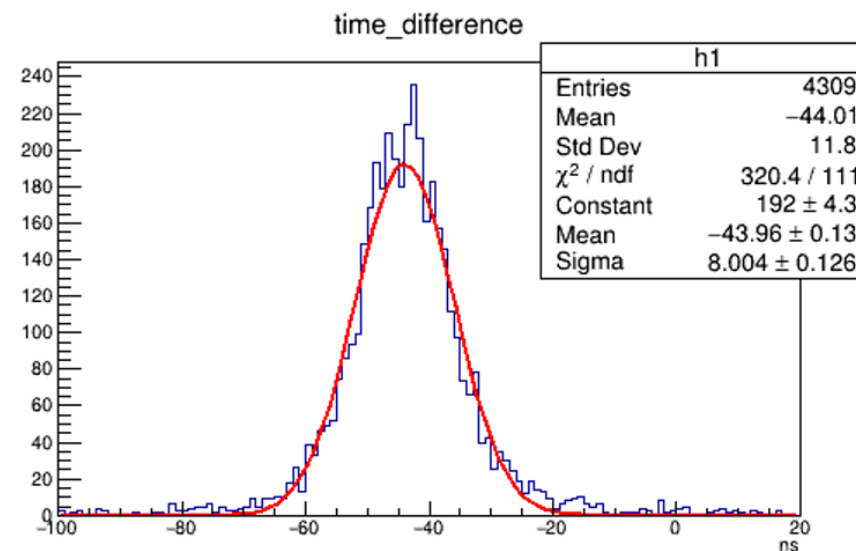
# Time residual to shower front



WCD time residual



ED5283 time residual



time difference between ED5283 and WCD