



Gamma-Ray and AntiMatter survey(GRAMS) experiment

August, 2025, TAUP 2025

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Northeastern University

On behalf of GRAMS collaboration

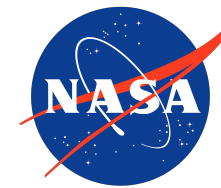
What is GRAMS?



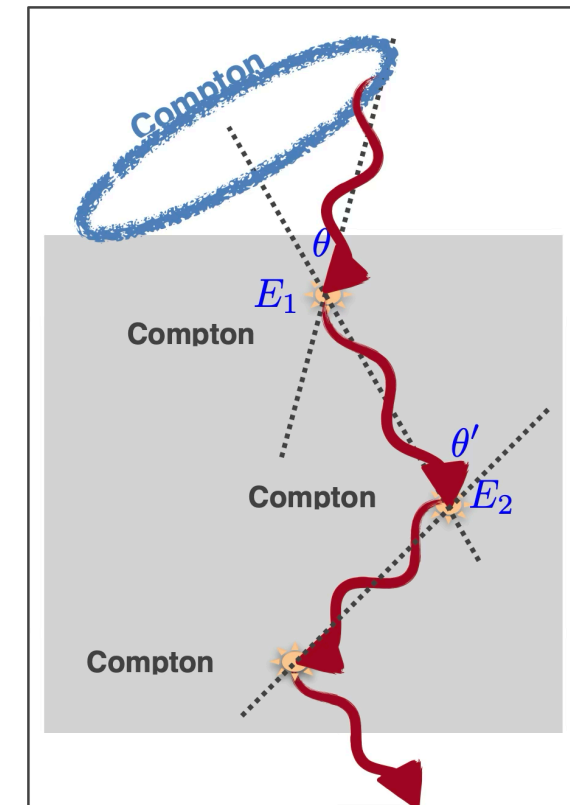
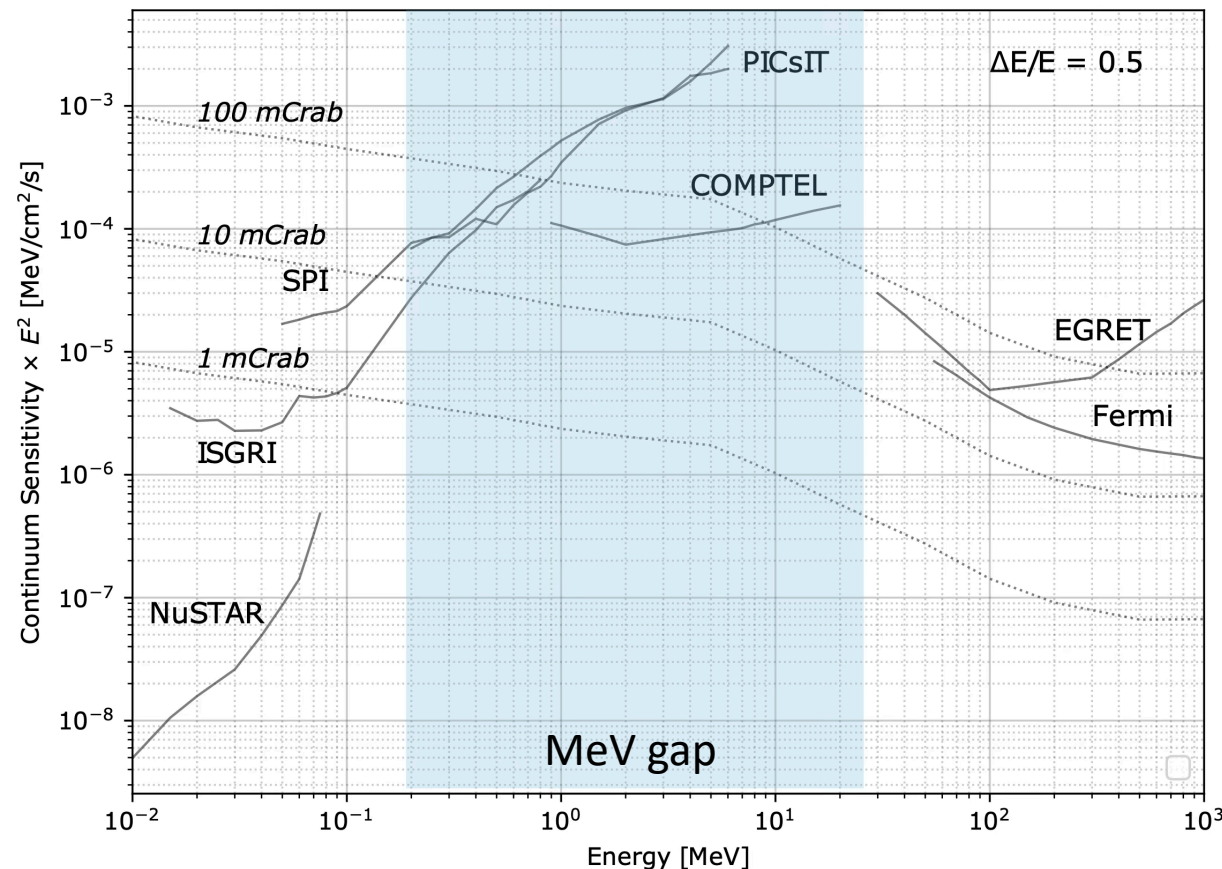
GRAMS = Gamma-Ray and AntiMatter Survey

- ❑ First balloon/satellite experiment to target both
 - ❑ Astrophysical observations with **MeV gamma rays**
 - ❑ Indirect dark matter searches with **cosmic antinuclei**
- ❑ First balloon/satellite mission with a low-cost, large-scale **LArTPC** (liquid argon time projection chamber) detector
- ❑ Funded by **NASA** as a **Physics of the Cosmos** suborbital experiment

MeV Gamma-Ray Observations



- MeV region is poorly measured due to domination of Compton scattering. Detectors require good spatial and energy resolution



MeV Gamma-ray Science



☐ Physics processes/nucleosynthesis

☐ Transition of physics processes in MeV

☐ Particle acceleration in Pulsar Wind Nebula

☐ Relativistic flows in stellar/super-massive BHs, pulsars, magnetars

☐ Nuclear lines from radioactive isotopes in astrophysical environments

☐ Galactic center, classic novae, SNe, etc

☐ r-process for the origin of heavy elements

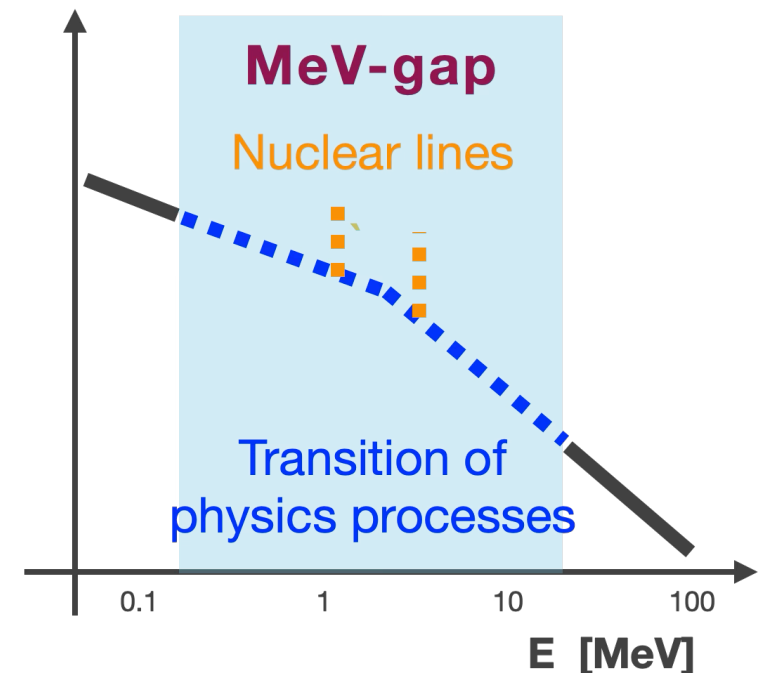
☐ Multi-messenger astrophysics

☐ MeV gamma rays + GWs

☐ MeV gamma rays + high-energy neutrinos

☐ NS/BH mergers, GRBs, SNe, AGNs

☐ Indirect Dark Matter and PBH searches



GRAMS detector design

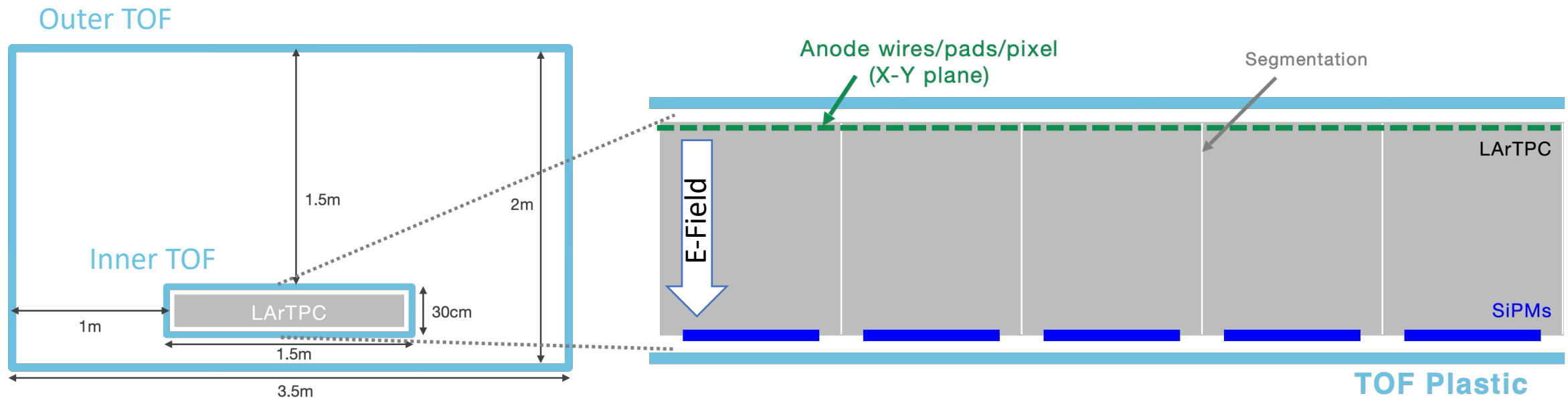


- ❑ Large-scale, low-energy threshold LArTPC has been **well-studied** and **widely-used** in underground **dark matter** and **neutrino** experiments

[Charged Particle](#)

[Gamma Ray: Compton Scattering](#)

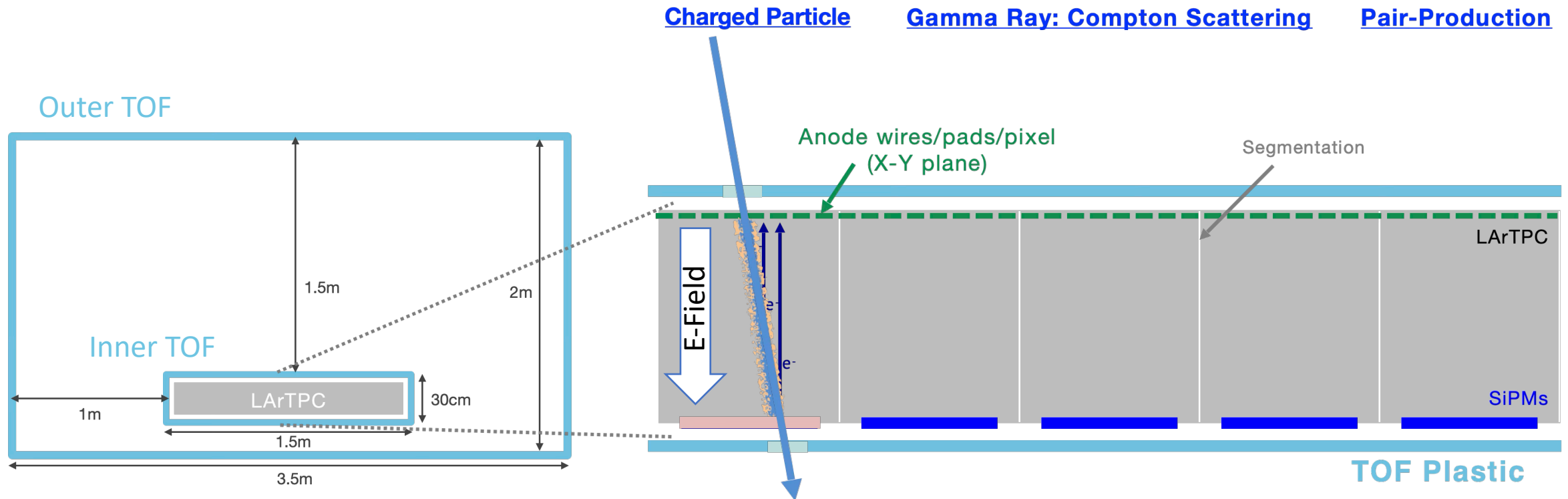
[Pair-Production](#)



GRAMS detector design



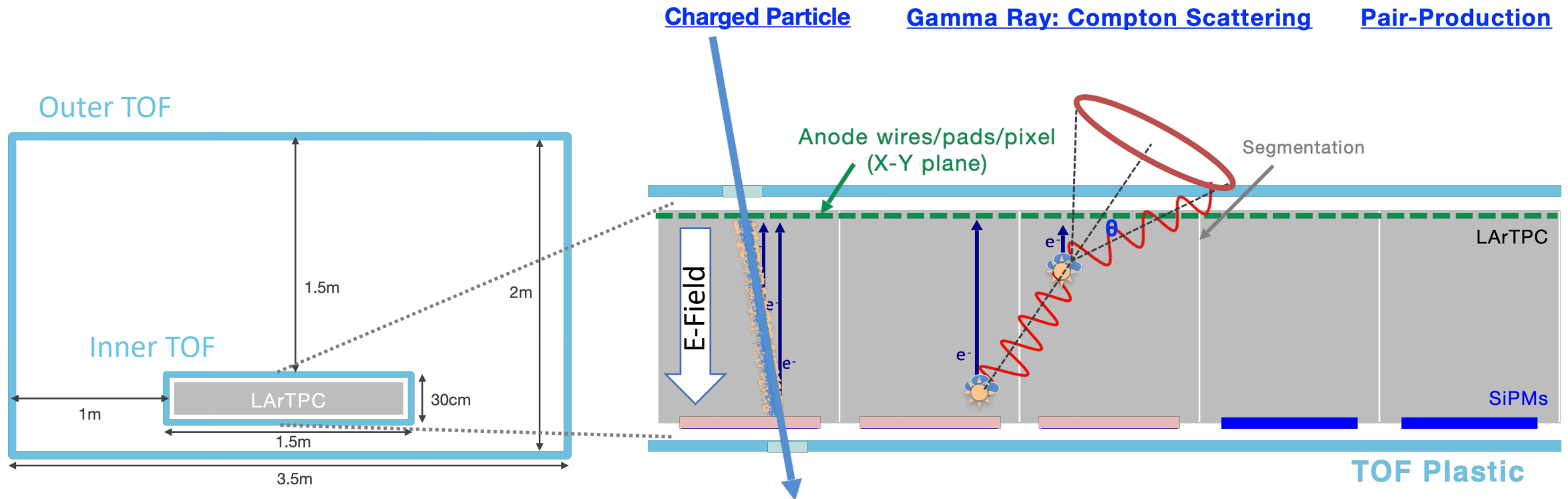
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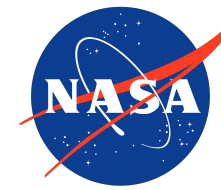
GRAMS detector design



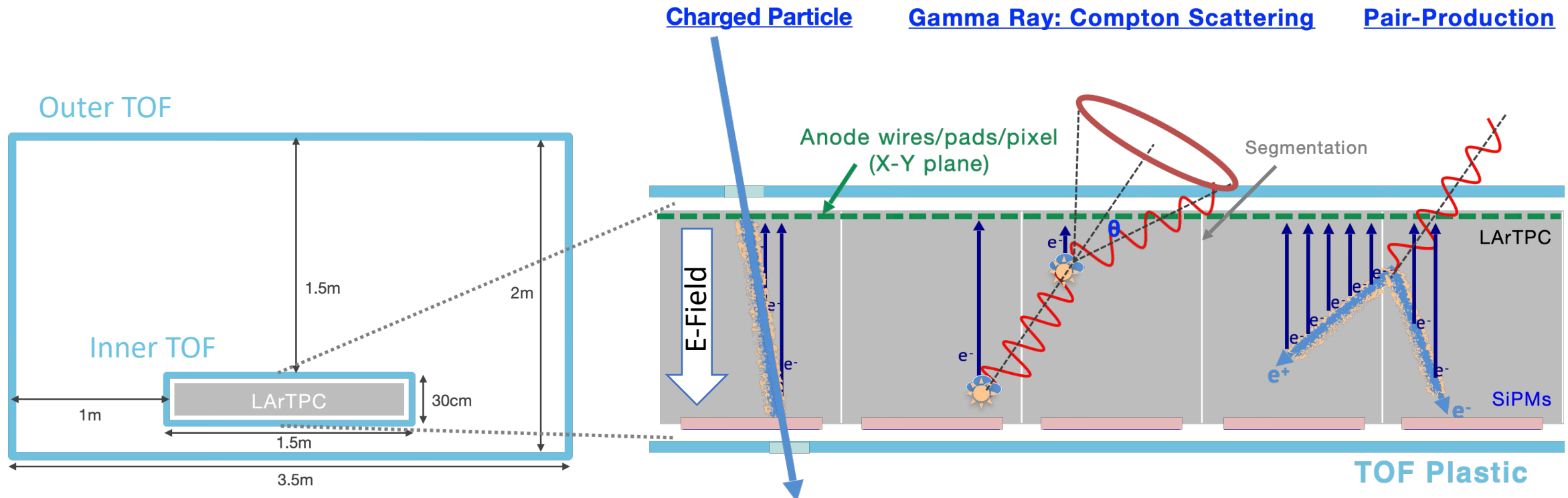
- ❑ Large-scale, low-energy threshold LArTPC has been **well-studied** and **widely-used** in underground **dark matter** and **neutrino** experiments



GRAMS detector design



- ❑ Large-scale, low-energy threshold LArTPC has been **well-studied** and **widely-used** in underground **dark matter** and **neutrino** experiments

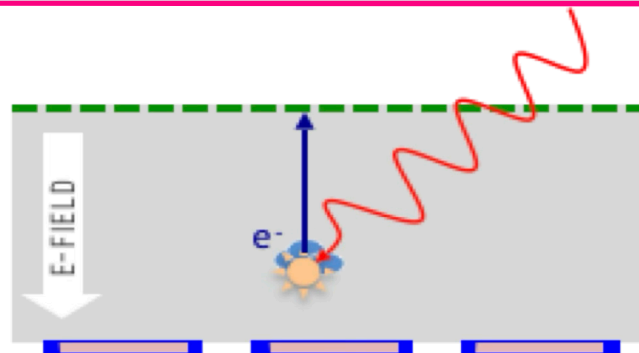


Why LArTPC?

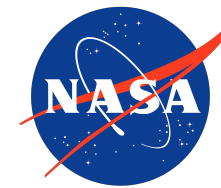


❑ LArTPC is **cost-effective** and easily expandable to a **larger scale**, Almost **no dead volume** and **high detection efficiency**

	LArTPC	Semiconductor/Scintillator
ρ (g/cm ³)	1.4	2.3/5.3 (Ge/Si)
T _{operation}	~80K	~240K/~80K
Cost	\$	\$\$\$
Signals	Scintillation light + ionization electrons	Electrons, holes
X, Y positions	Wires/pads on anode plane (X-Y)	Double-sided strips
Z position	From drift time	From layer #
# of layers	Single layer	Multi-layers
# of electronics	#	###
Dead volume	Almost no dead volume	Detector frame, preamps
Neutron bkg	Identified with pulse shape	No rejection capability



MeV Gamma-ray Observations

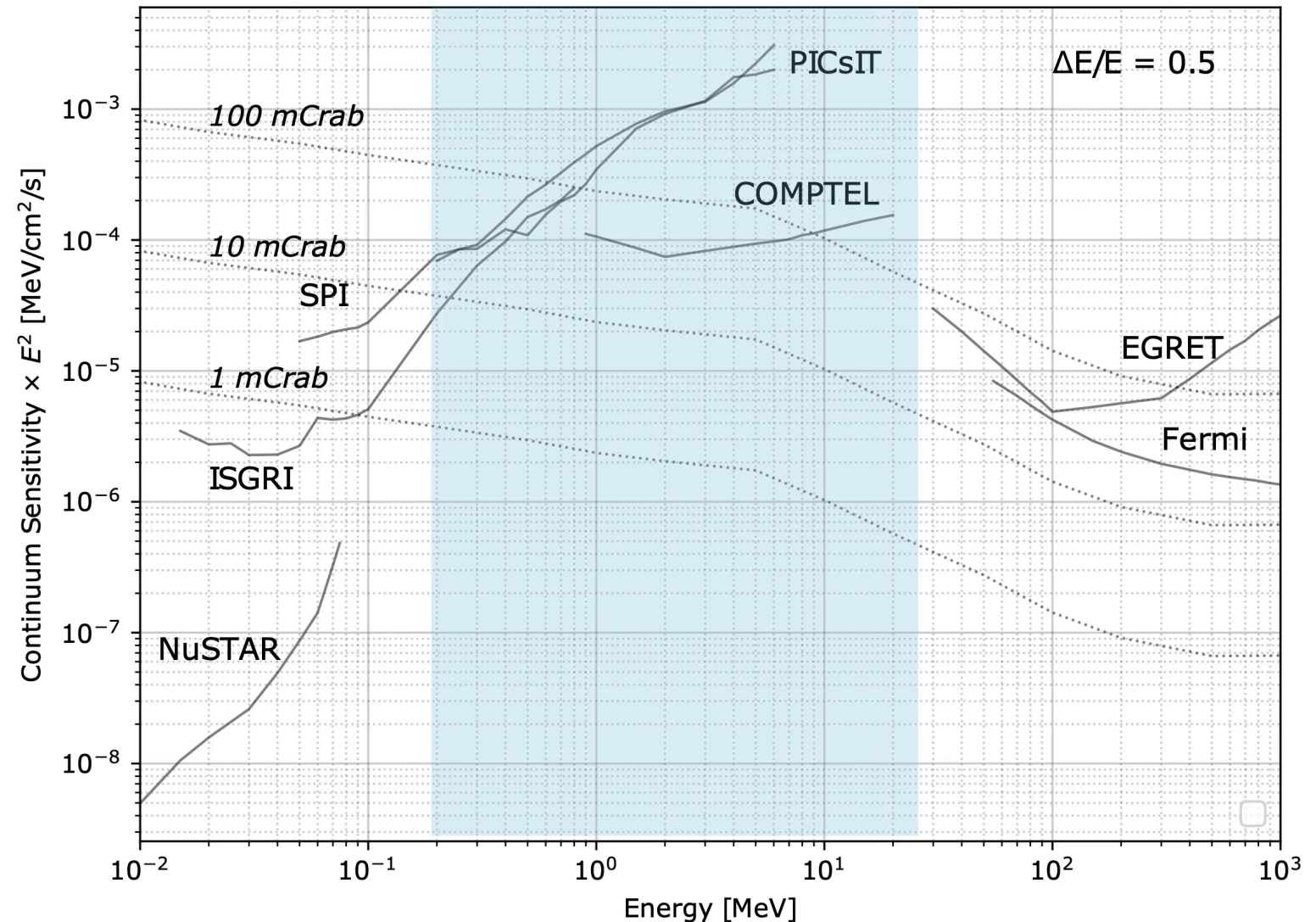


❑ **Single balloon flight:**

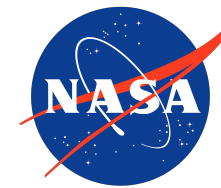
❑ Order of magnitude improved

❑ **Satellite mission:**

❑ Comparable to future missions



MeV Gamma-ray Observations

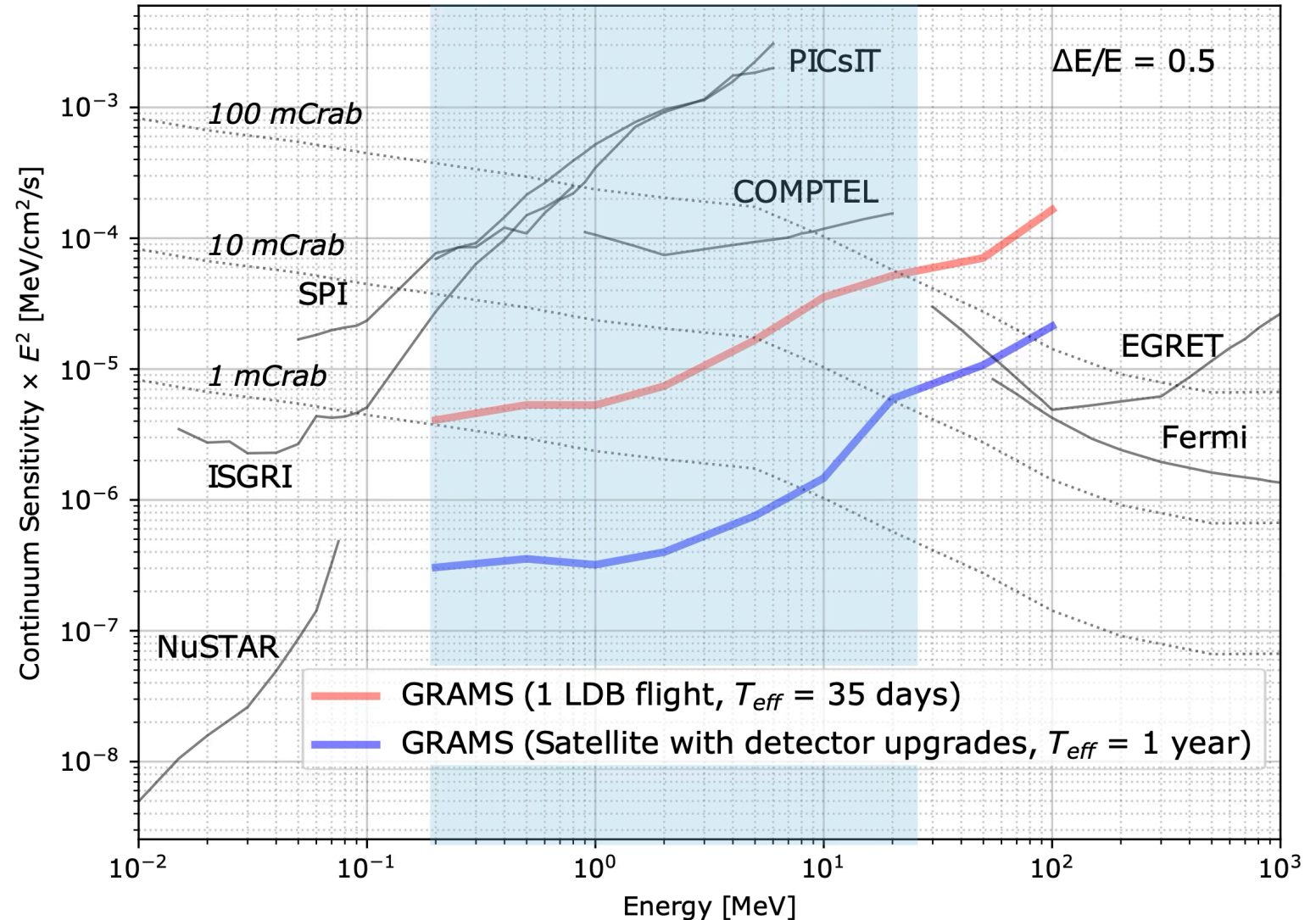


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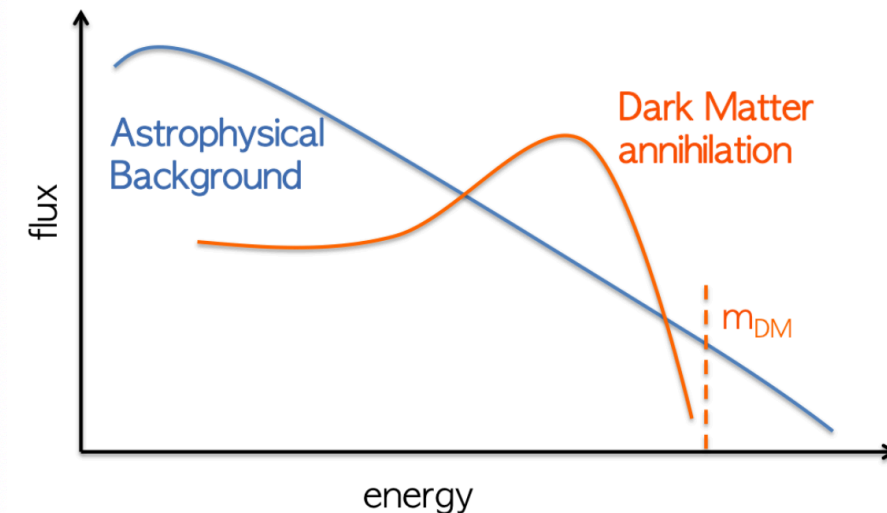
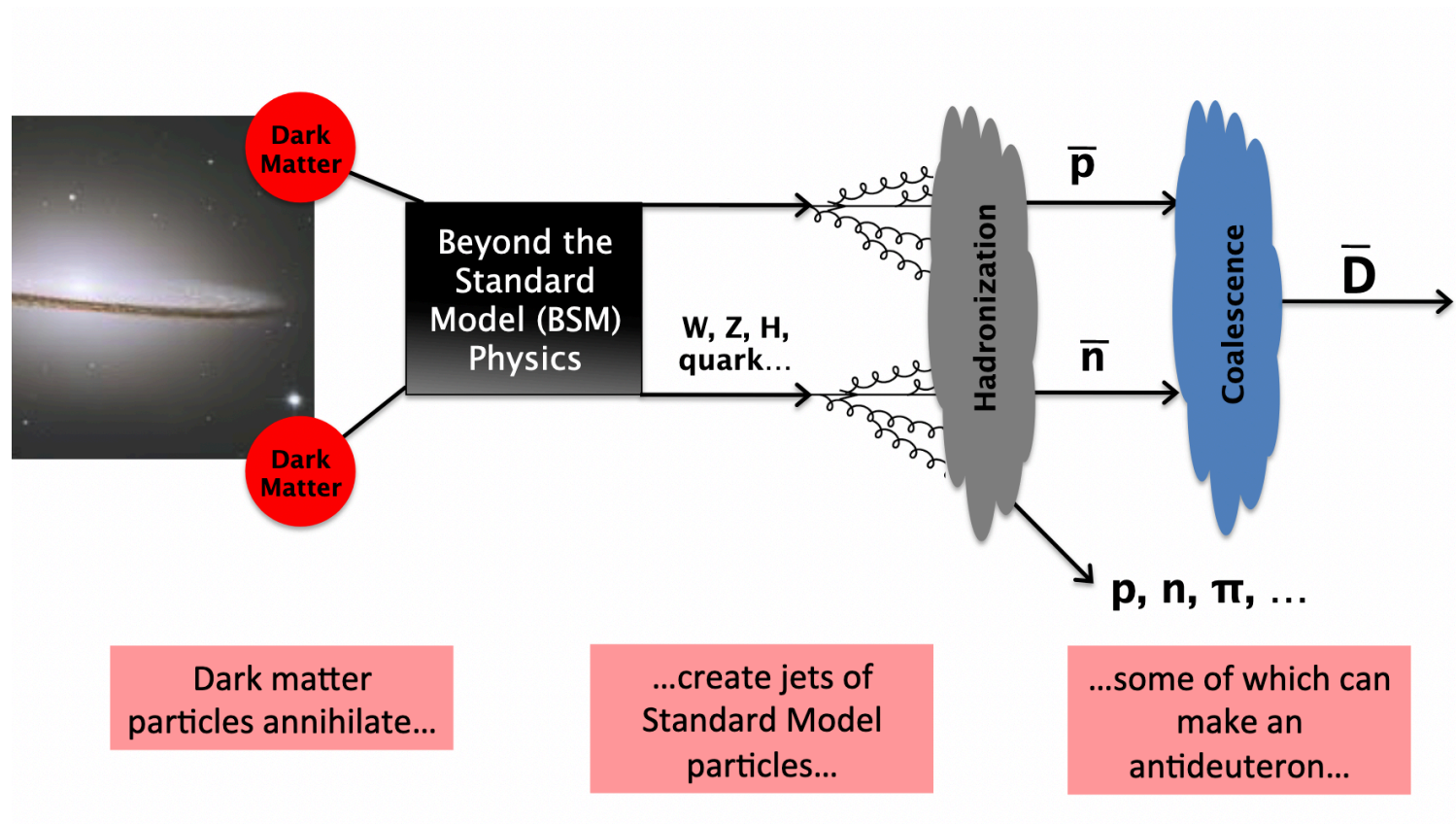
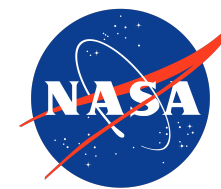
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Indirect DM Searches

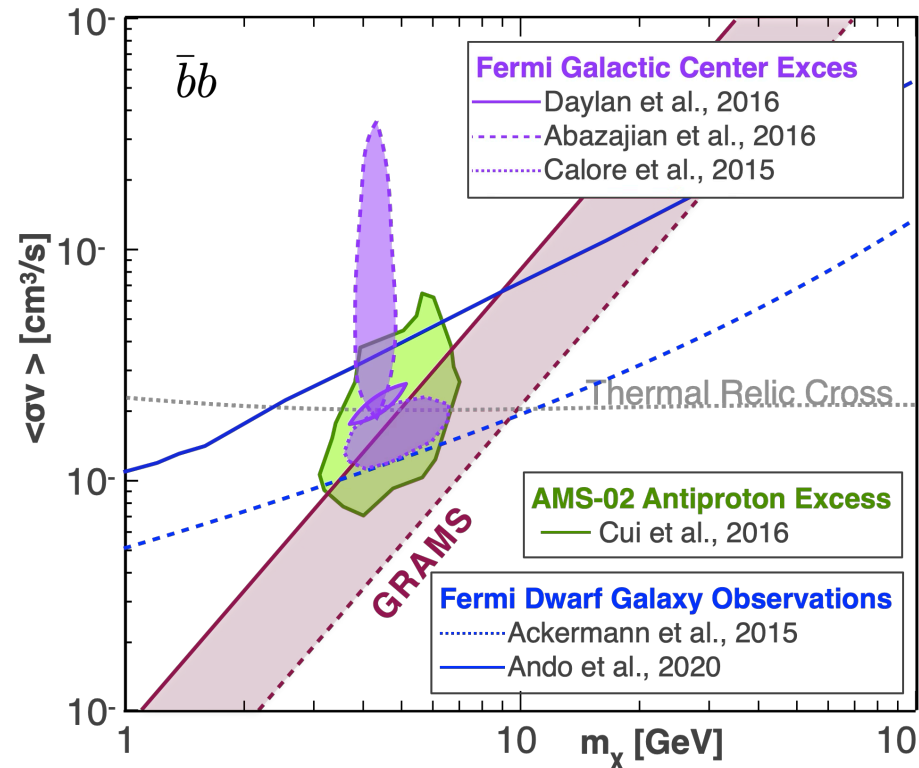
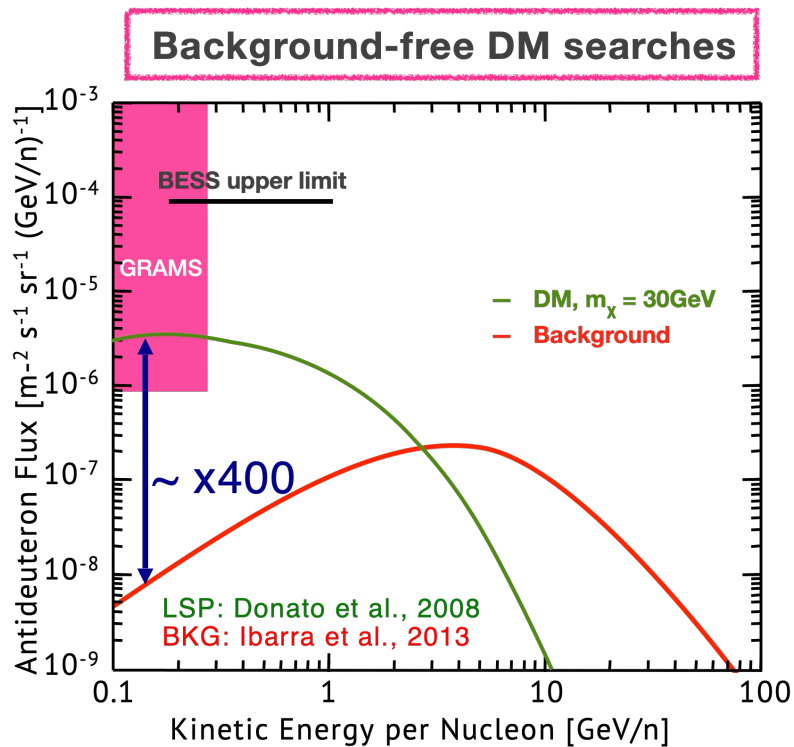


Finding products generated by DM self interaction

GRAMS Antideuteron Sensitivity

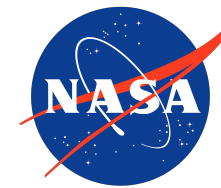


- ❑ Low energy \bar{D} gives an essentially background-free signature
- ❑ Extensively explore DM parameter space and validate Fermi/AMS-02 results

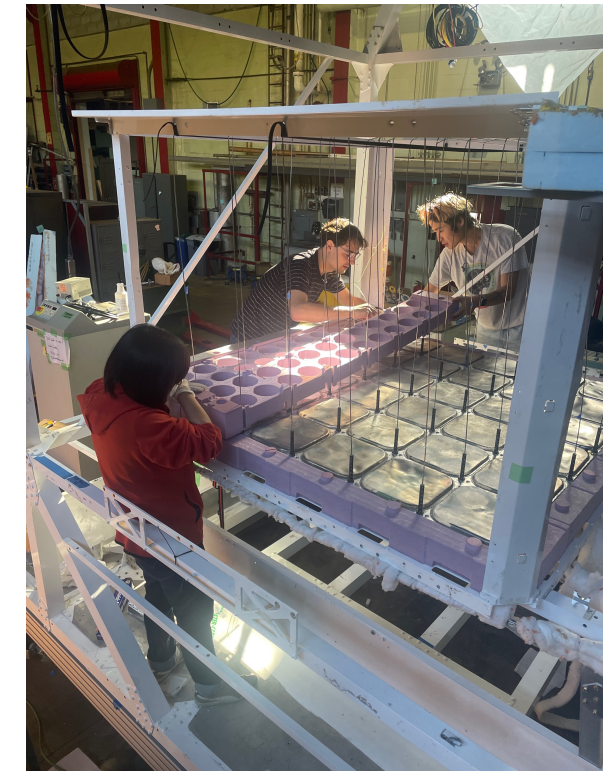
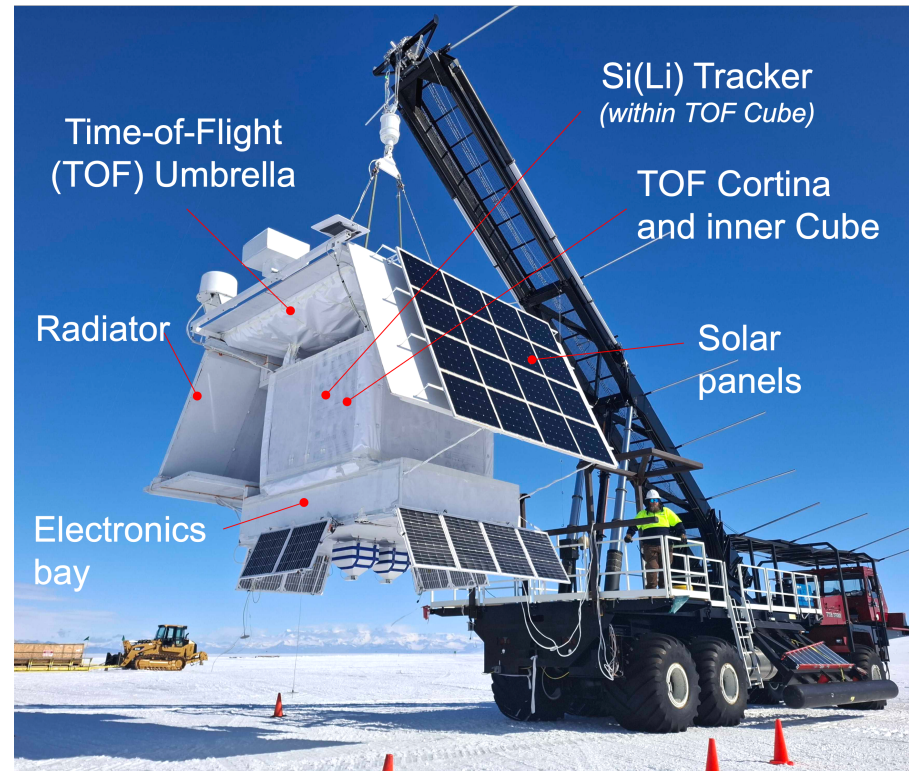
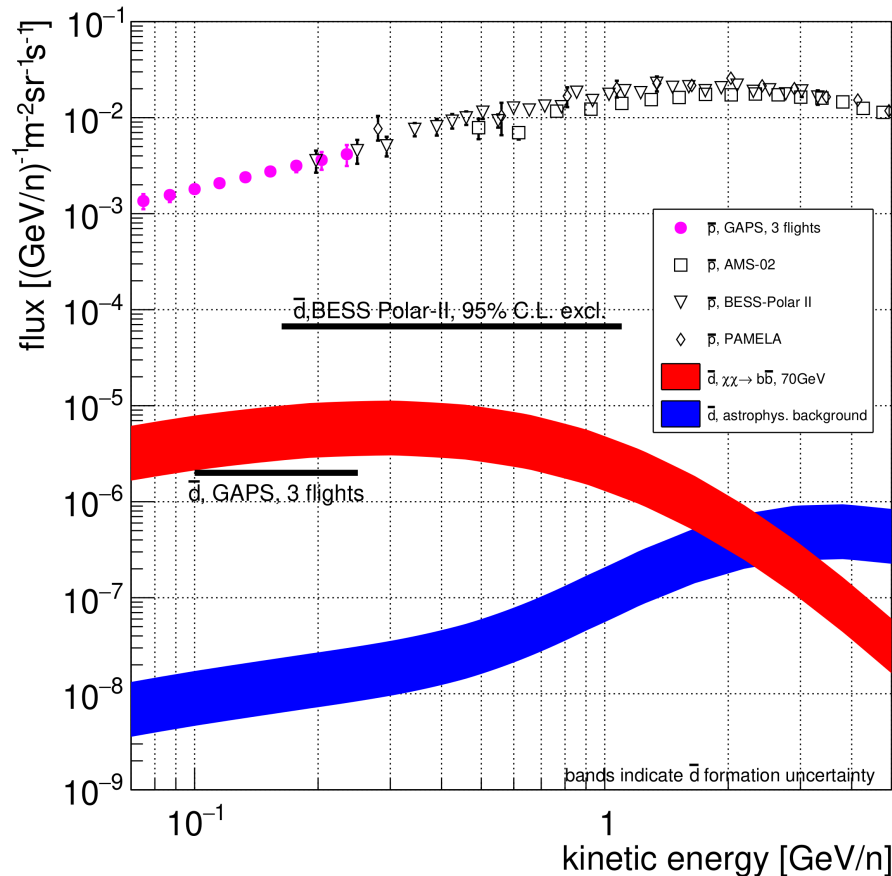


Next generation measurement beyond GAPS

General AntiParticle Spectrometer



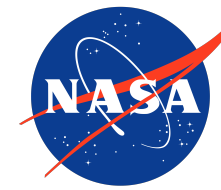
Current generation low energy antiparticles experiment



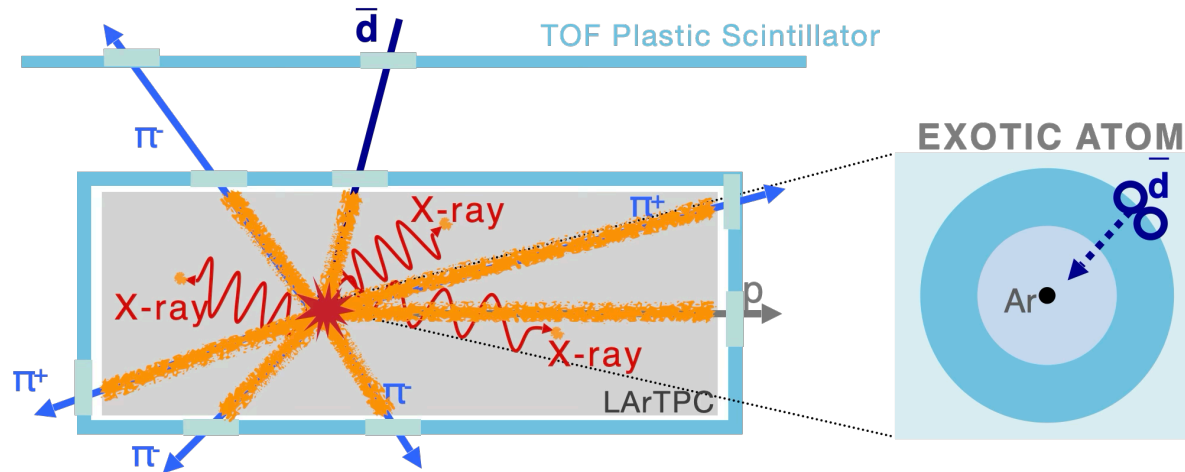
Mengjiao Xiao's talk [Search for Cosmic-Ray Antinuclei from Dark Matter with the GAPS Antarctic Balloon Mission](#) yesterday

GAPS instrument is READY in Antarctica planed for flight in late 2025

GRAMS Antimatter detection

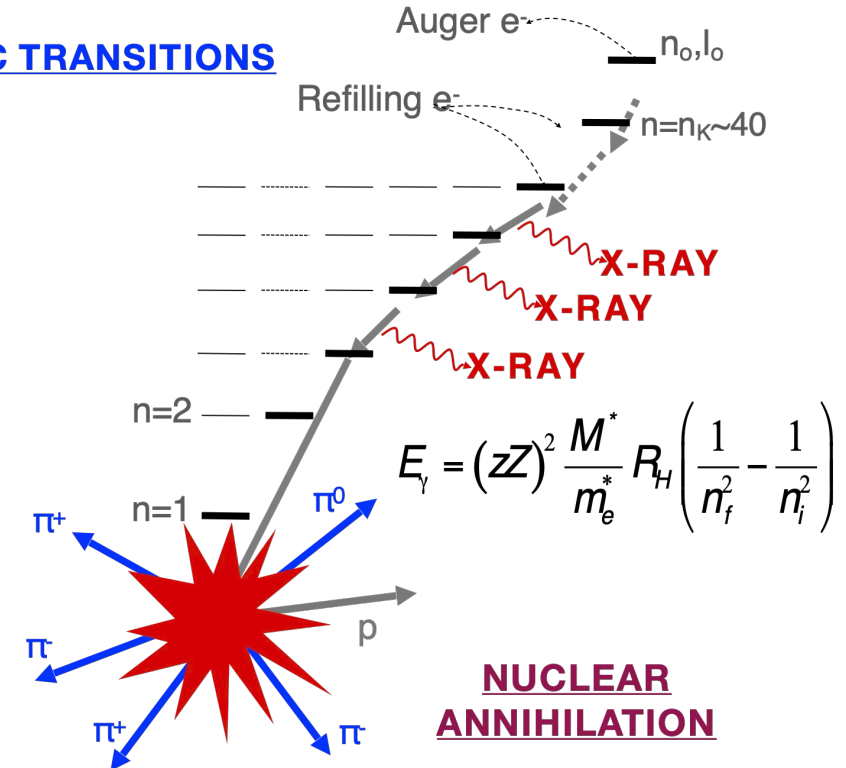


Measure atomic **X-rays** and **annihilation** products



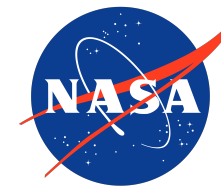
- A time of flight (TOF) system tags candidate events and records velocity
- The antiparticle slows down & stops, forming an excited exotic atom
- De-excitation X-rays provide signature
- Annihilation products provide additional background suppression

ATOMIC TRANSITIONS

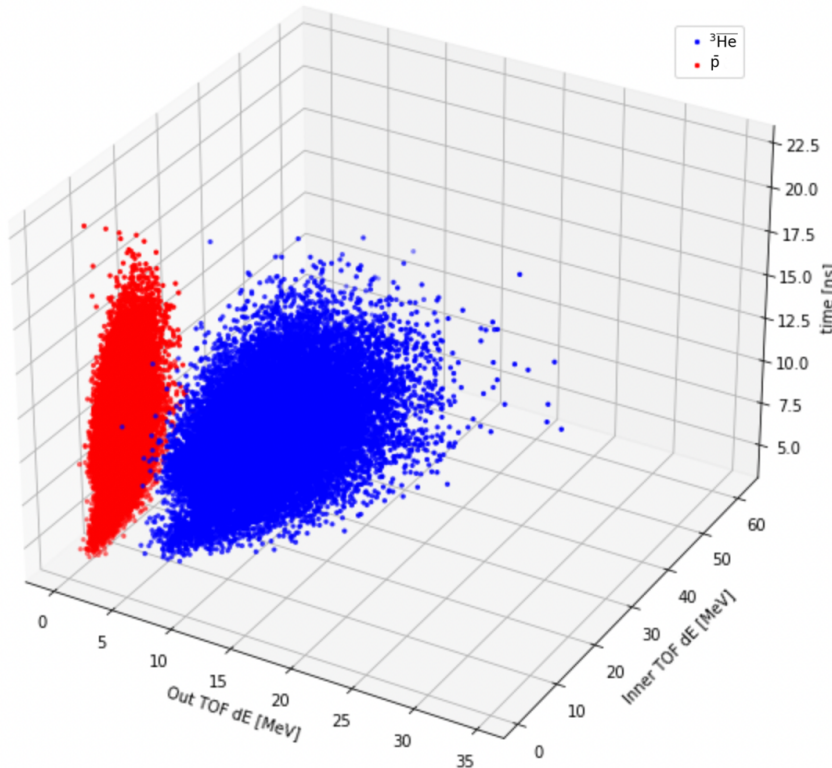


Particle identification with the combination of
TOF timing, stopping range, pion/proton multiplicities, energy depositions in plastic/LArTPC, atomic X-rays

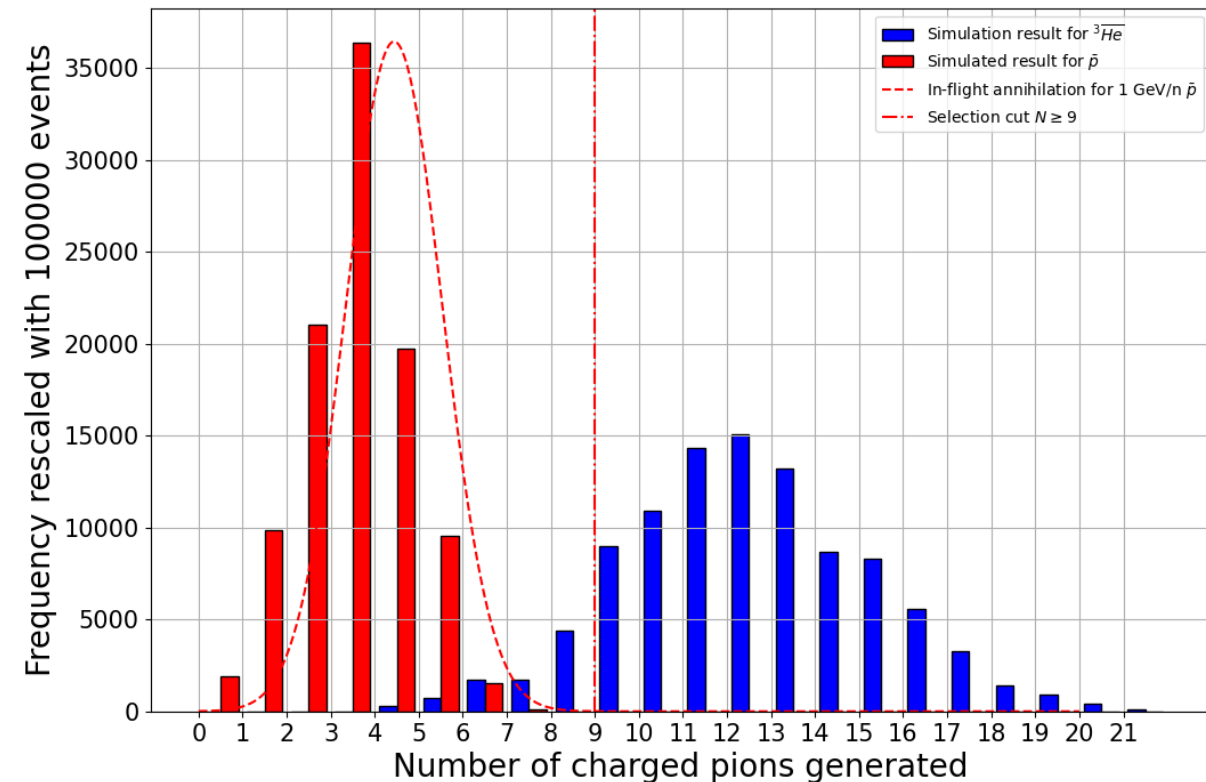
GRAMS Antimatter detection



- ❑ Aside from unique X-Ray peak, Antihelium-3 or Antideuteron has unique behavior in LArTPC detector



Charge-2 CRs particles tend to deposit more energy in TOF detector

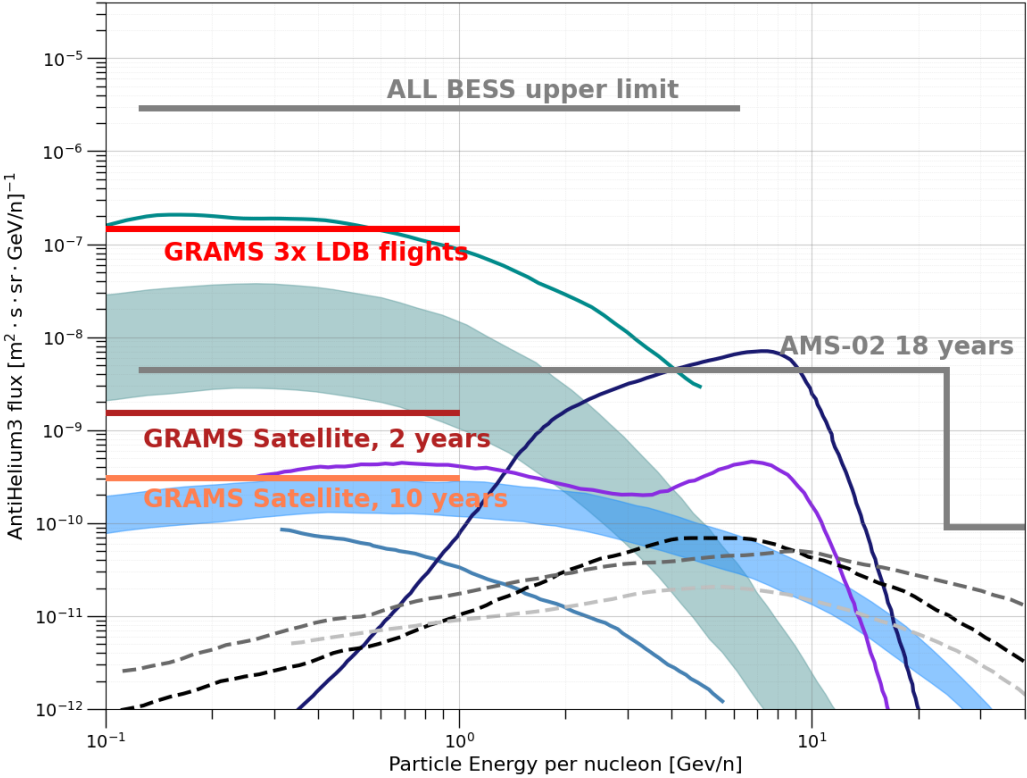


GRAMS antihelium-3



GRAMS could investigate a variety of DM models, complementary to AMS-02

GRAMS Antihelium-3 Sensitivity



Antihelium-3 sensitivity for the GRAMS experiment

J. Zeng^{1,*,}, T. Aramaki¹, K. Aoyama^{2,3}, S. Arai⁴, S. Arai², J. Asadi⁵, A. Bamba⁴, N. Cannady⁶, P. Coppi⁷, G. De Nolfo⁶, M. Errando⁸, L. Fabris⁹, T. Fujiwara¹⁰, Y. Fukazawa¹¹, P. Ghosh⁶, K. Hagino⁴, T. Hakamata¹⁰, N. Hiroshima¹², M. Ichihashi⁴, Y. Ichinohe¹³, Y. Inoue^{10,14,15}, K. Ishikawa², K. Ishiwata¹⁰, T. Iwata⁴, G. Karagiorgi¹⁶, T. Kato⁴, H. Kawamura¹⁰, D. Khangulyan^{17,18}, J. Krizmanic⁶, J. Leyva¹, A. Malige¹⁶, J.G. Mitchell⁶, J.W. Mitchell⁶, R. Mukherjee¹⁹, R. Nakajima², K. Nakazawa²¹, H. Odaka¹⁰, K. Okuma²¹, K. Perez¹⁶, N. Poudyal¹, I. Safa¹⁶, K. Sakai²⁰, M. Sasaki⁶, W. Seligman¹⁶, J. Sensenig¹⁶, K. Shirahama¹⁰, T. Shiraishi²², S. Smith²³, Y. Suda¹¹, A. Suraj¹, H. Takahashi¹¹, S. Takashima⁴, T. Tamba^{4,3}, M. Tanaka², S. Tandon¹⁶, R. Tatsumi¹⁰, J. Tomsick²⁴, N. Tsuji²², Y. Uchida²⁵, Y. Utsumi², S. Watanabe²³, Y. Yano², K. Yawata²⁶, H. Yoneda²⁷, K. Yorita², M. Yoshimoto¹⁰

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Antiproton

ABSTRACT

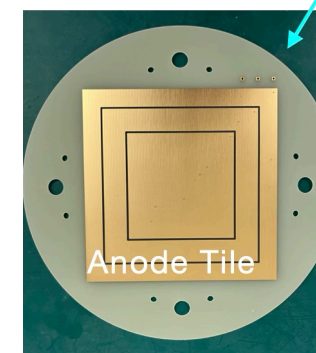
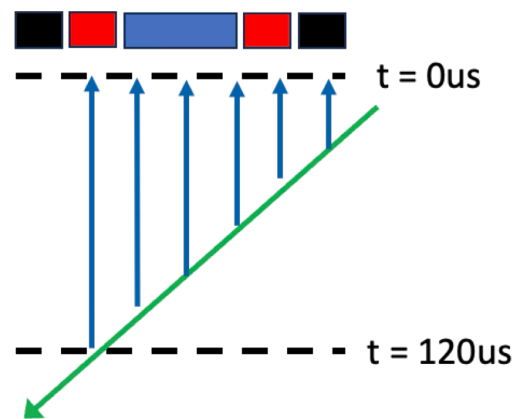
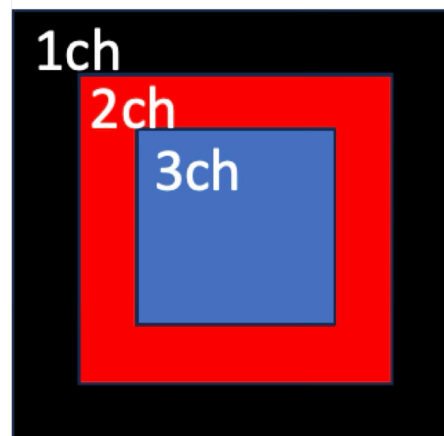
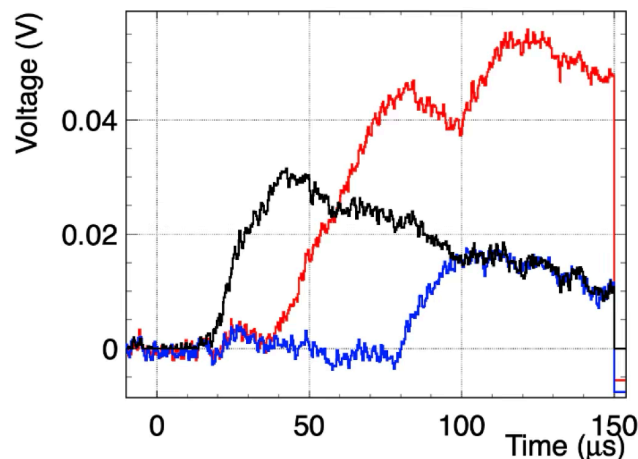
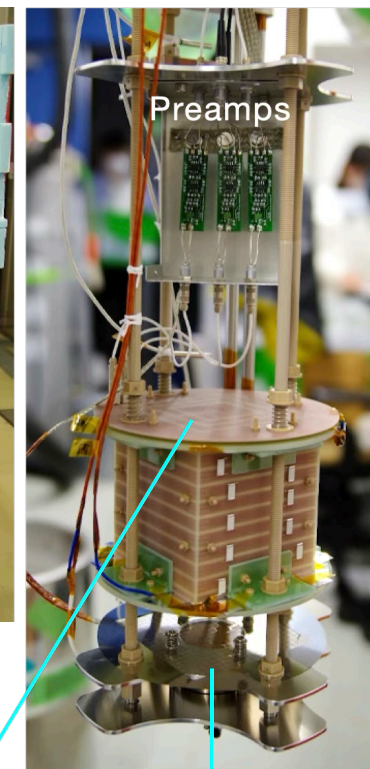
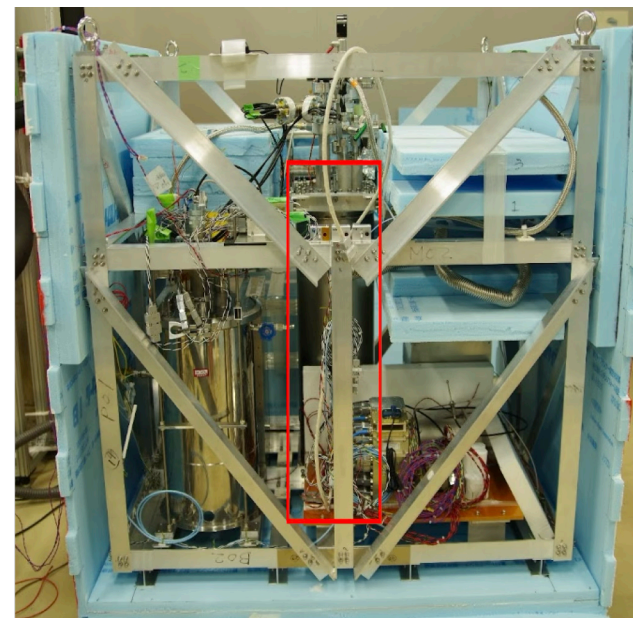
The Gamma-Ray and AntiMatter Survey (GRAMS) is a next-generation balloon/satellite mission utilizing a Liquid Argon Time Projection Chamber (LArTPC) detector to measure both MeV gamma rays and antinuclei produced by dark matter annihilation or decay. The GRAMS can identify antihelium-3 events based on the measurements of X-rays and charged pions from the decay of the exotic atoms, Time of Flight (TOF), energy

Engineer flight eGRAMS

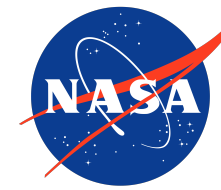


❑ Successful engineering flight @JAXA TARF

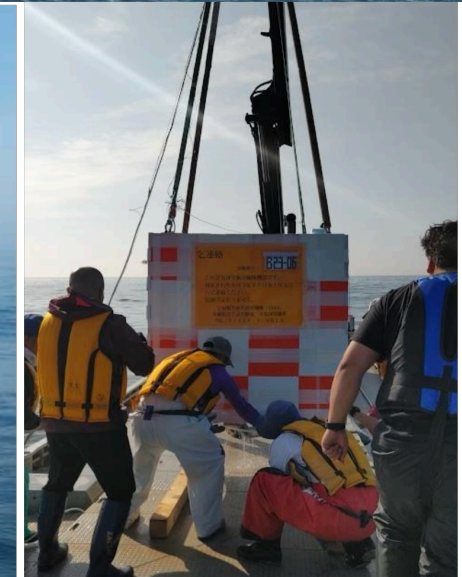
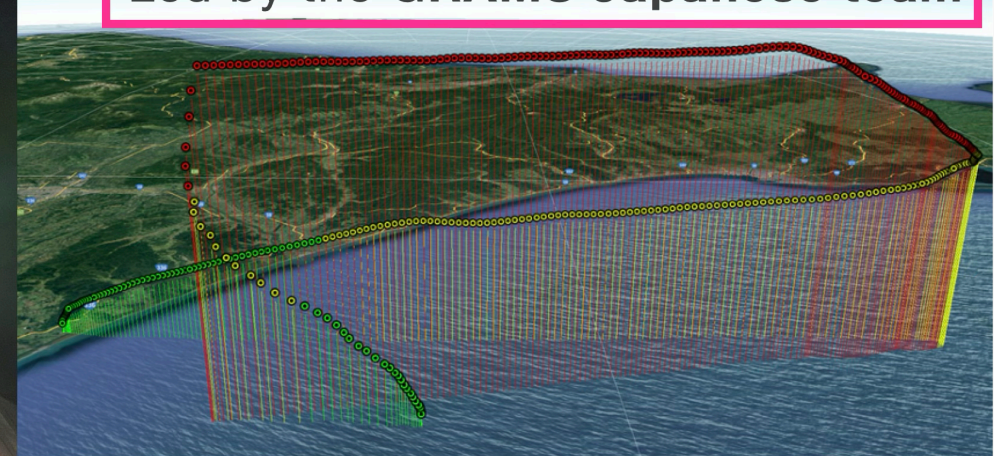
- ❑ First LArTPC operation at stratosphere
- ❑ TPC: $10 \times 10 \times 10 \text{ cm}^3$
 - ❑ 1 PMT (Hamamatsu R6041-06) at the bottom
 - ❑ 3 charge channels (pGRAMS preamps)
 - ❑ No cooling/circulation system
- ❑ Pressure vessel for RPi/DAQ
- ❑ Obtained ~400k stable events with light + charge



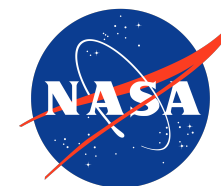
eGRAMS in July 2023



Led by the **GRAMS Japanese team**



Antiproton Beam Test at J-PARC, February 2025



❑ @J-PARC in Feb 2025

❑ Antiprotons with 0.7 GeV/c

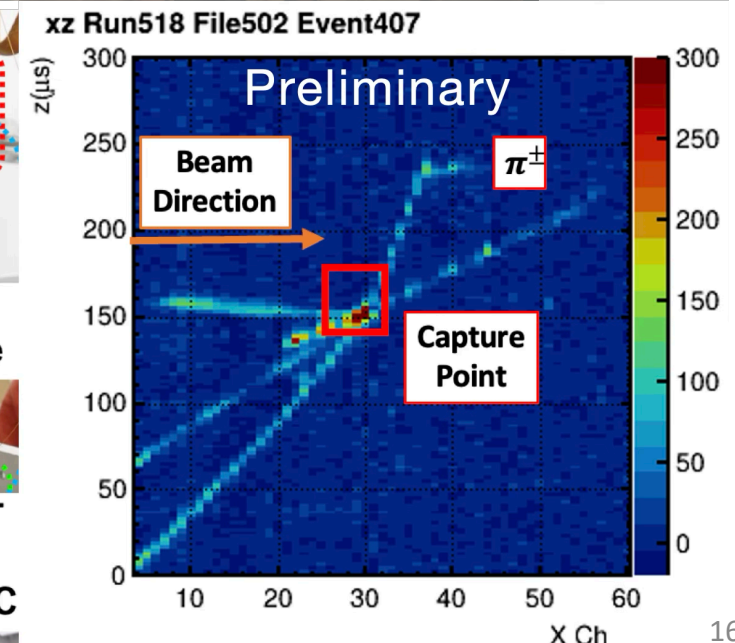
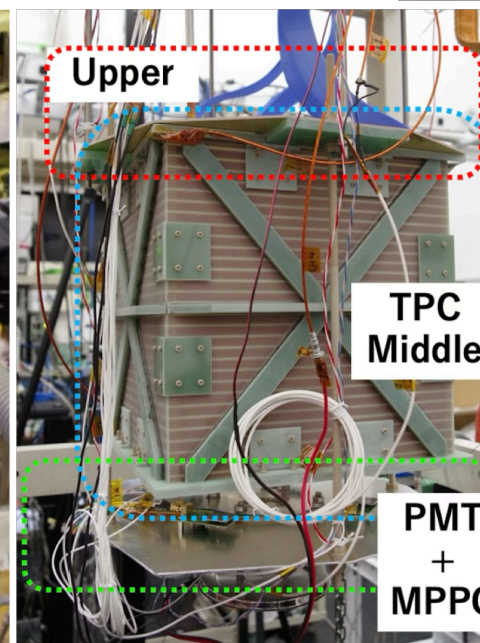
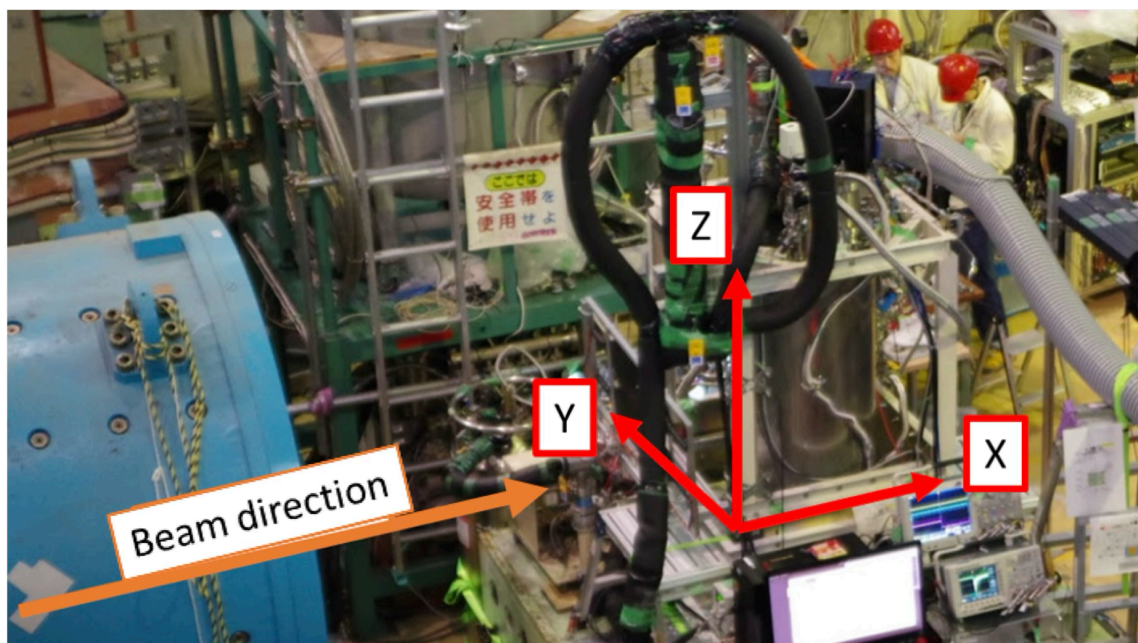
❑ Validated LArTPC performance as an antimatter detector

❑ TOF + Aerogel Cherenkov detector to reject pion events

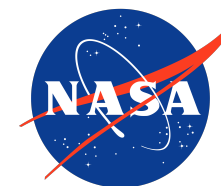
❑ Waseda TPC: 30 cm x 30 cm x 30 cm

❑ Measure annihilation products: pions (and protons)

Led by the **GRAMS Japanese team**

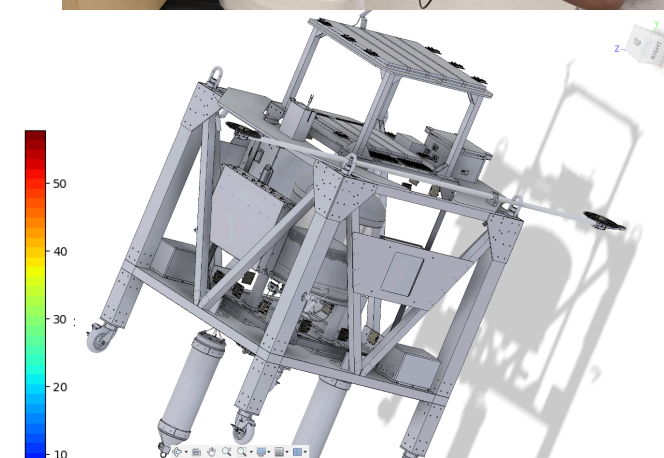
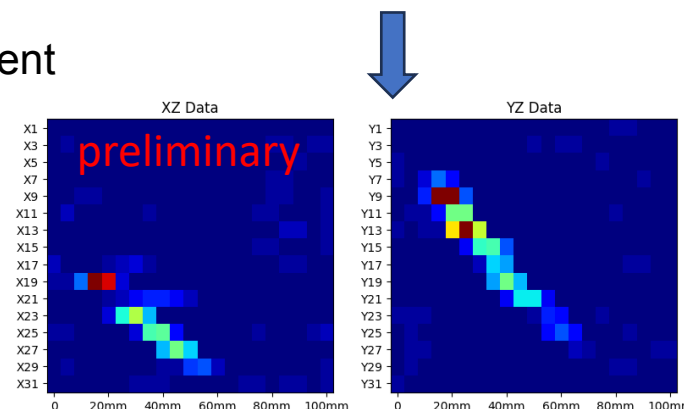
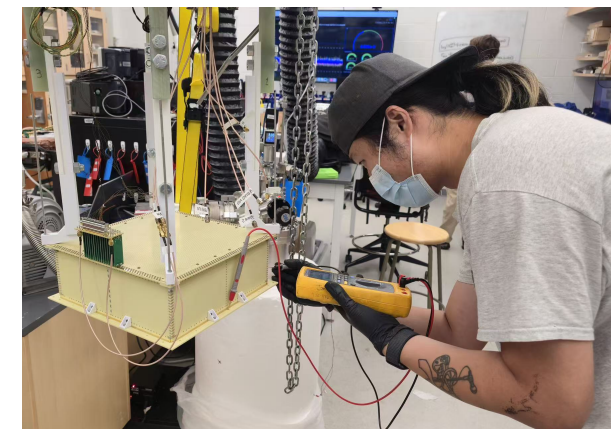
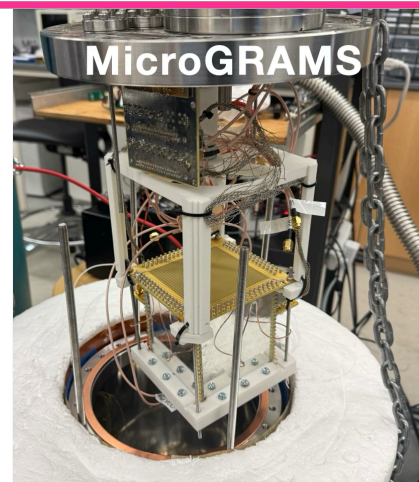


Prototype Flight (pGRAMS), Spring 2026



- ❑ Scheduled @Arizona in **Spring 2026**
 - ❑ Demonstrate LArTPC performance in flight
 - ❑ **Particle tracking** for charged particles
 - ❑ **Gamma-ray detection**
 - ❑ **MiniGRAMS**: 30 x 30 x 20 cm³ segmented into **9 cells**
 - ❑ Tile/pads (~3mm pitch) for x/y directions
 - ❑ 180 charge channels, 36 light channels
 - ❑ Currently testing **MicroGRAMS** @Northeastern
 - ❑ TPC size: 10 x 10 x 10 cm³
 - ❑ Demonstrate the particle tracking and event reconstruction
- ❑ Science flights with **MiniGRAMS**
 - ❑ One of the **largest** Compton cameras
 - ❑ Cooling/circulation system onboard

Funded by **NASA APRA2022**



Gondola and cryostat designed by NASA GSFC group

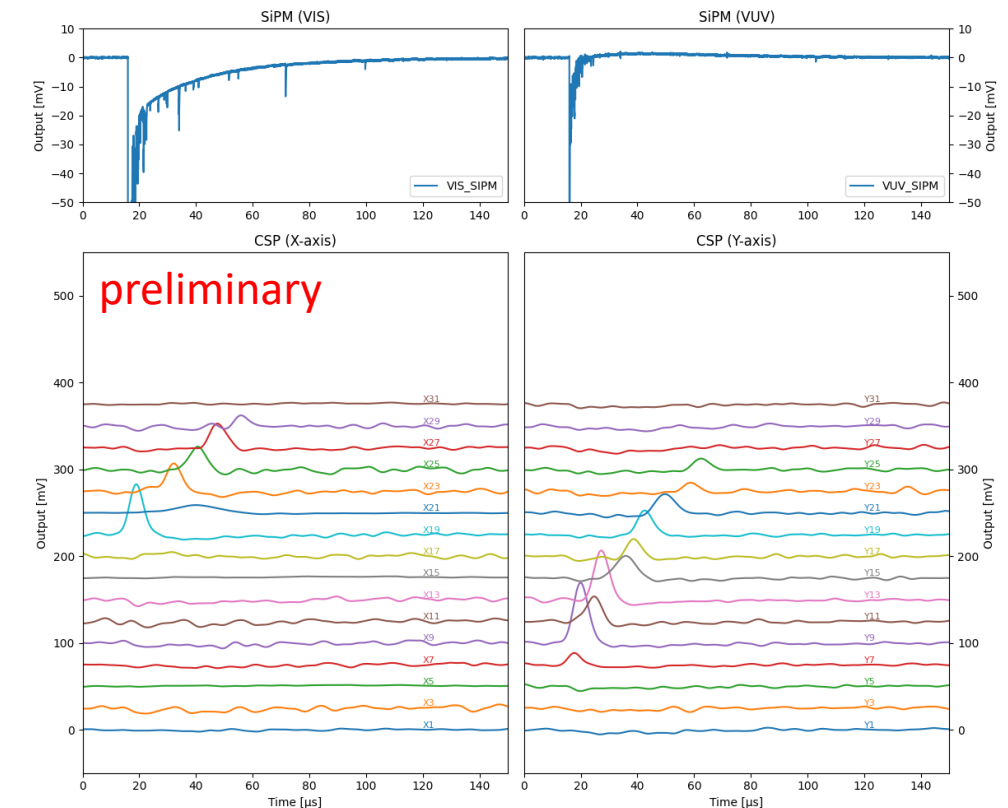
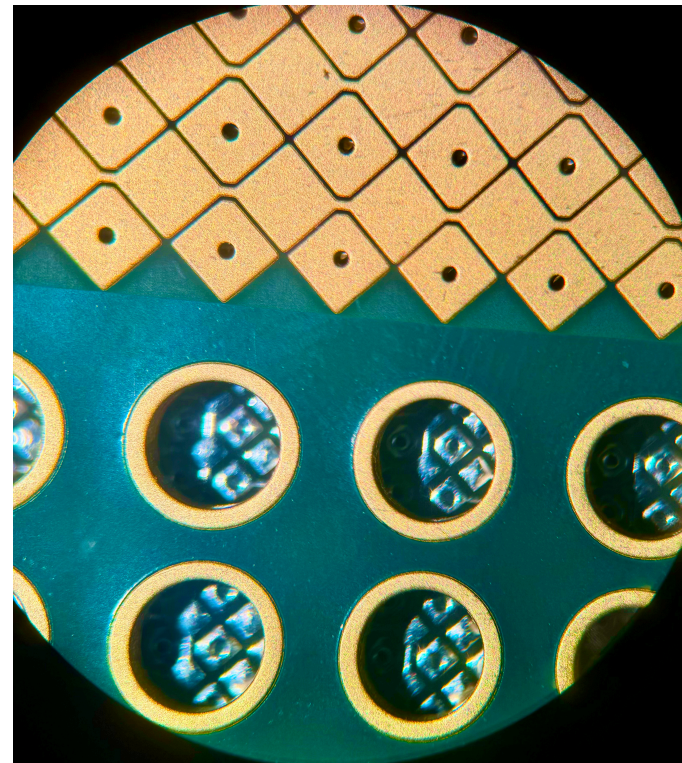
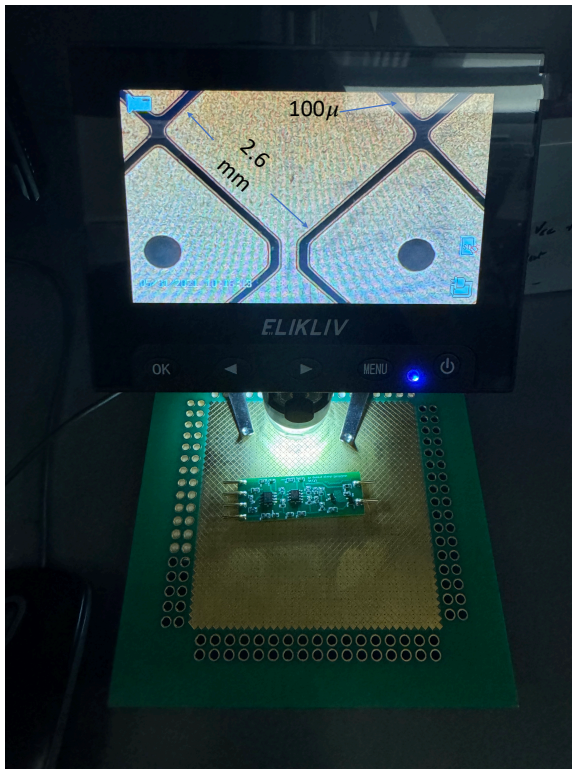
First science flight focus on MeV gamma-ray and particle tracking

Charge readout tile



❑ 2D readout tile

- ❑ More stable than wire-mesh readout for balloon payload
- ❑ Power saving comparing to pixel readout

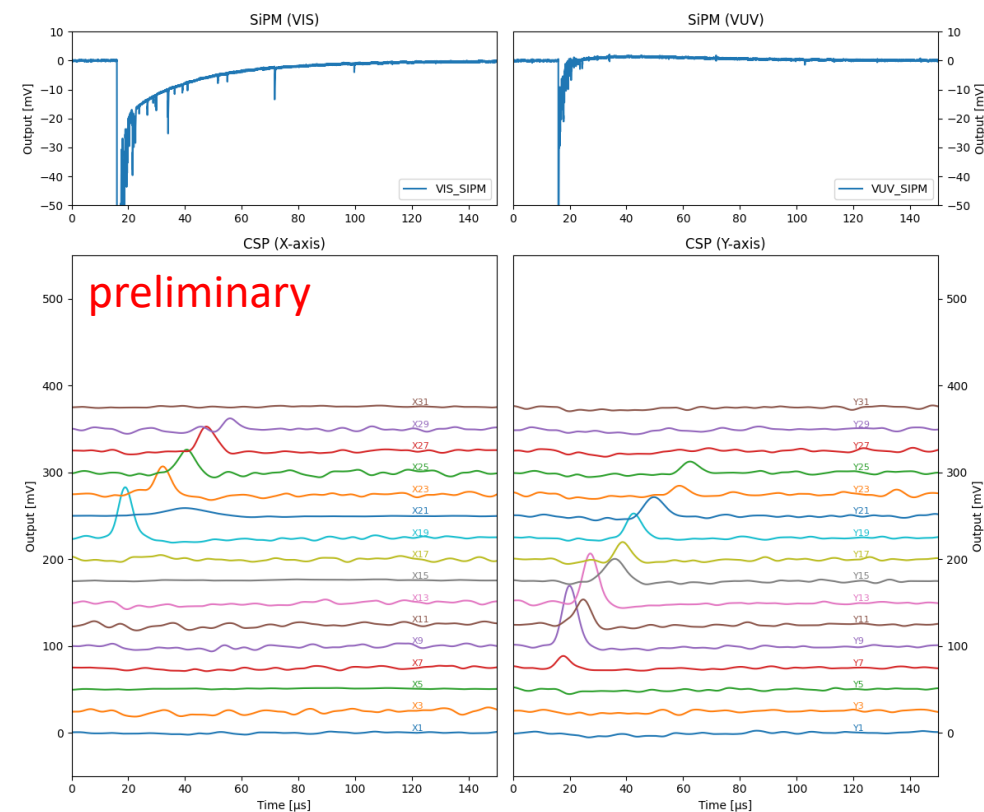
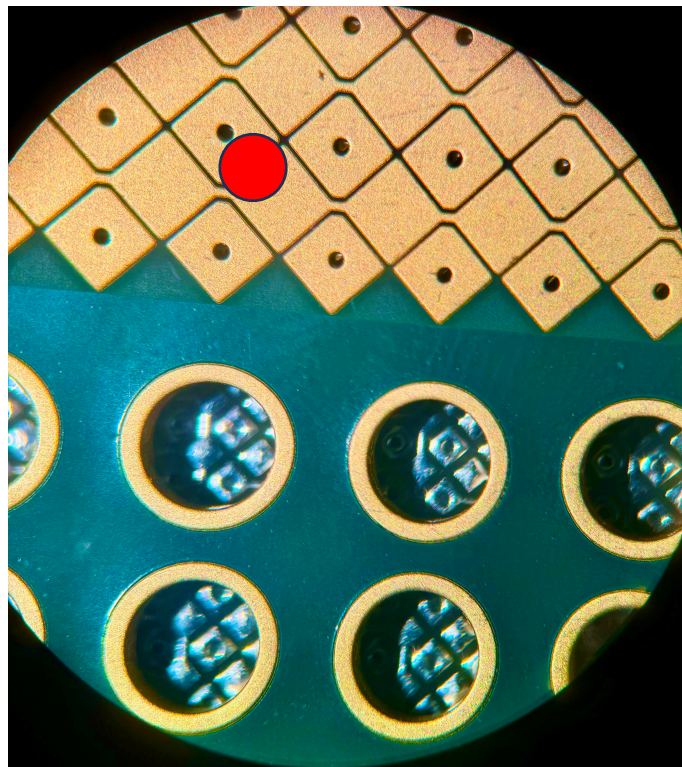
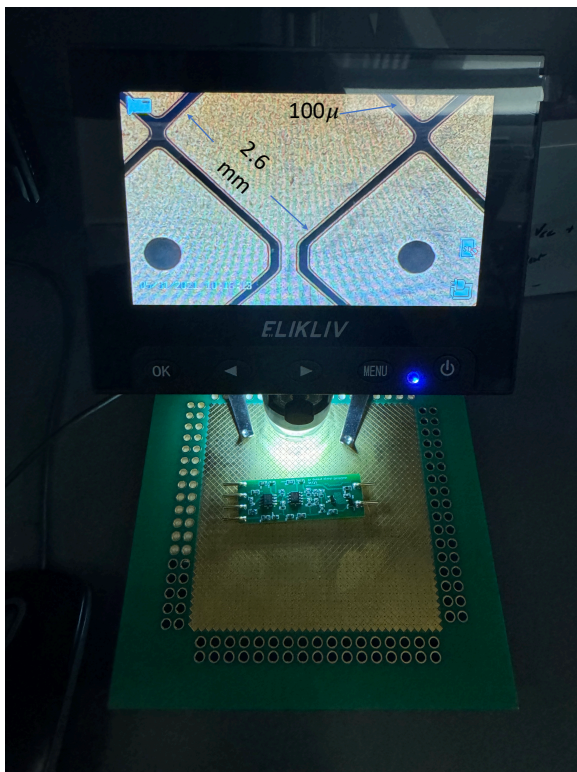


Charge readout tile



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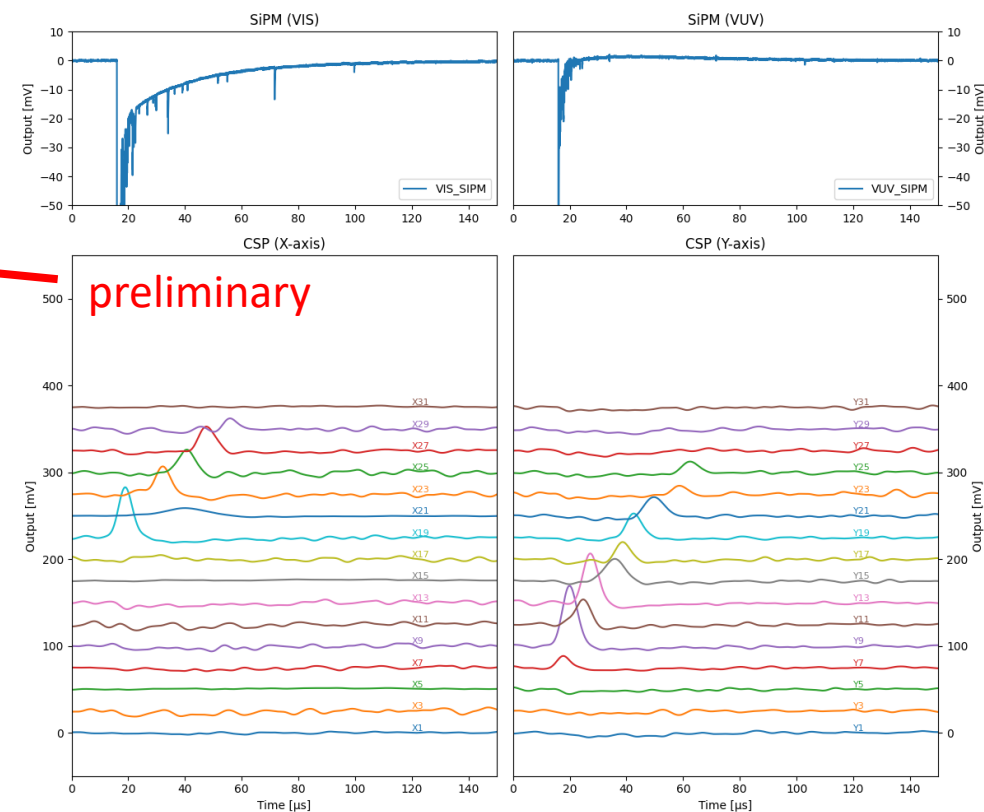
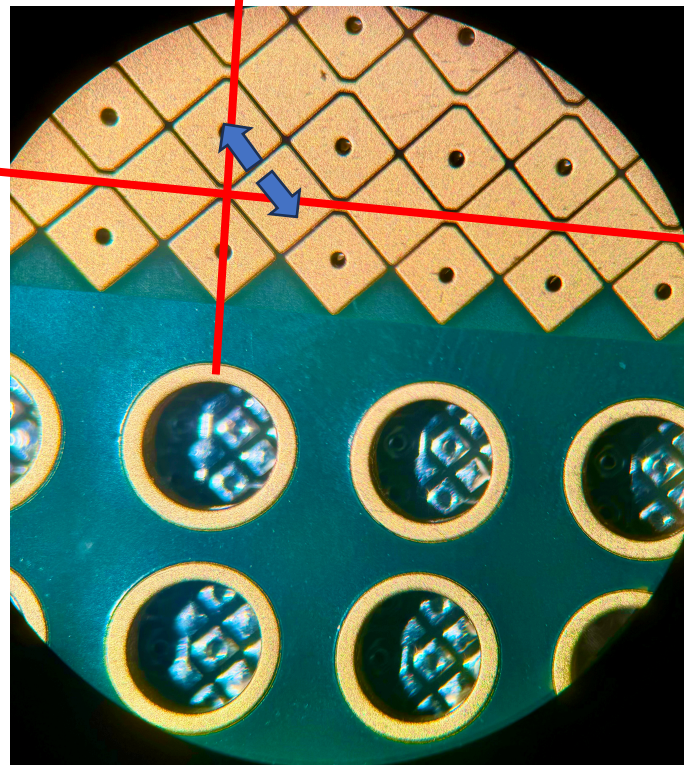
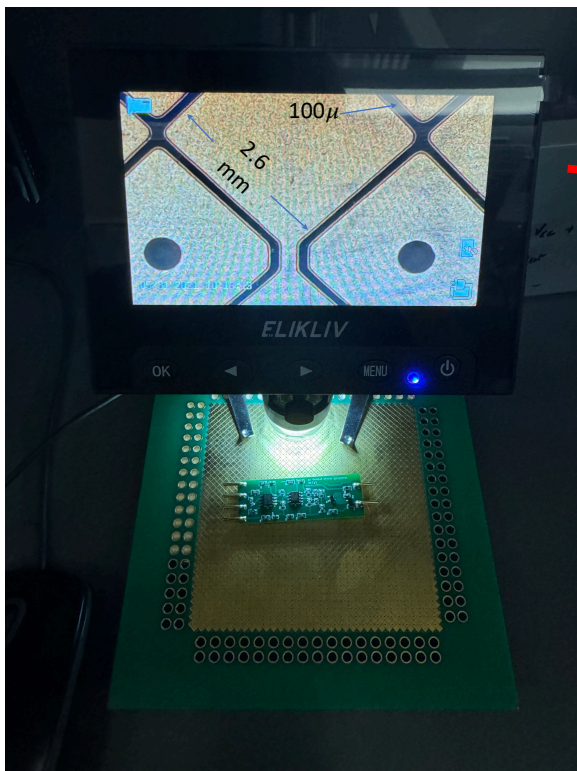


Charge readout tile

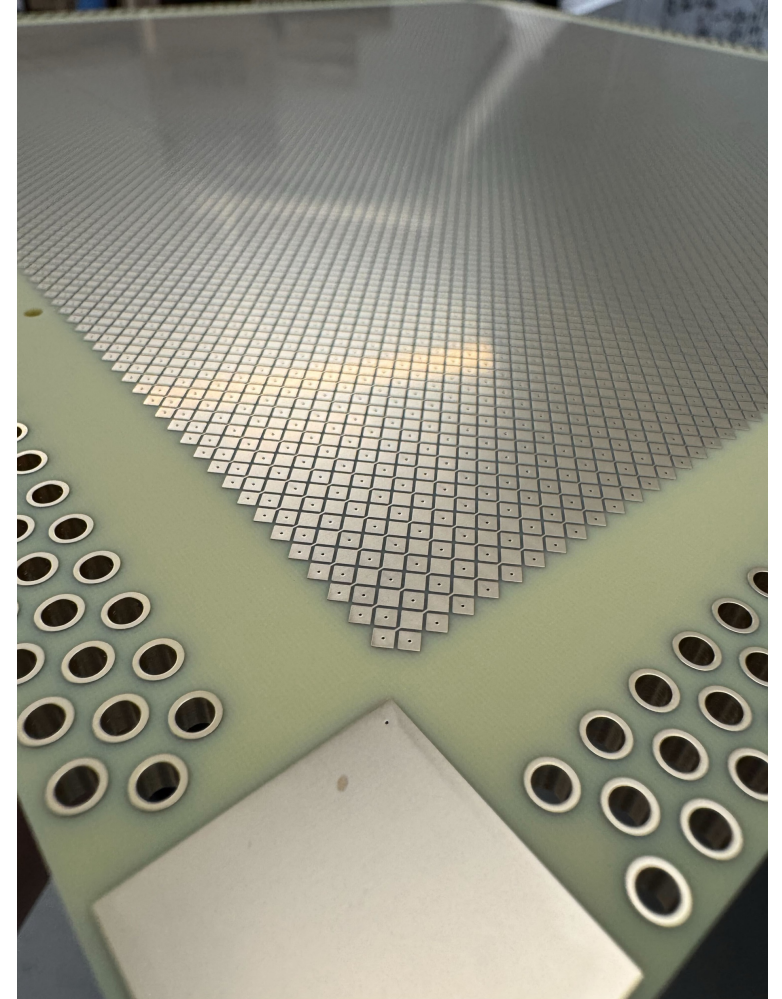
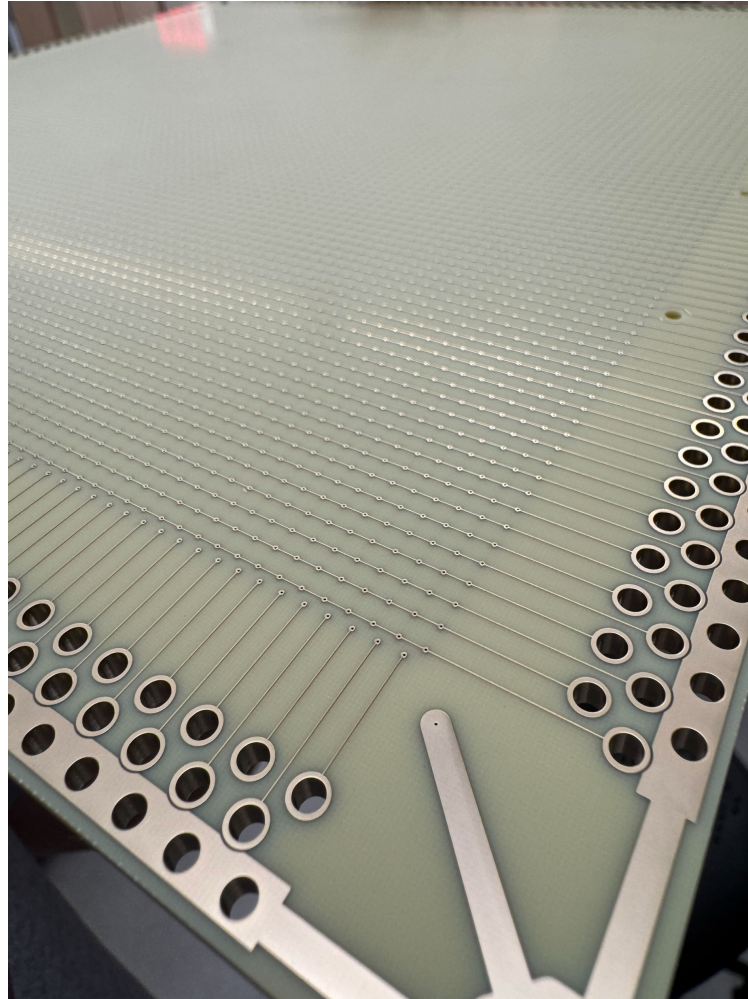
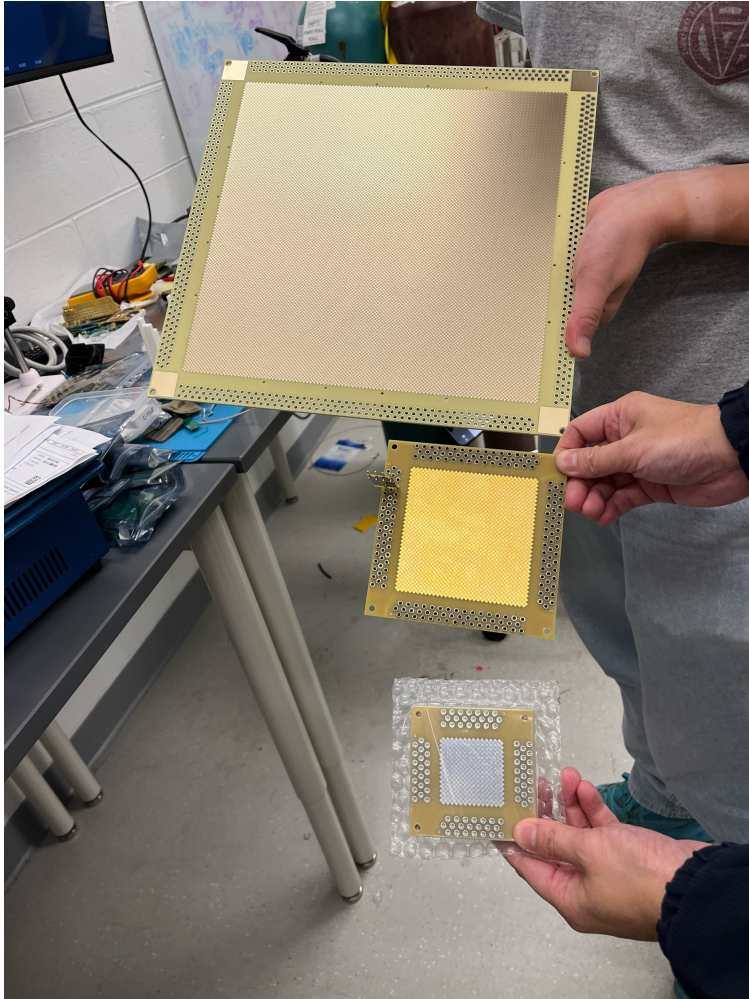


❑ 2D readout tile

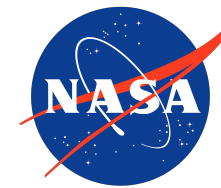
- ❑ More stable than wire-mesh readout for balloon payload
- ❑ Power saving comparing to pixel readout



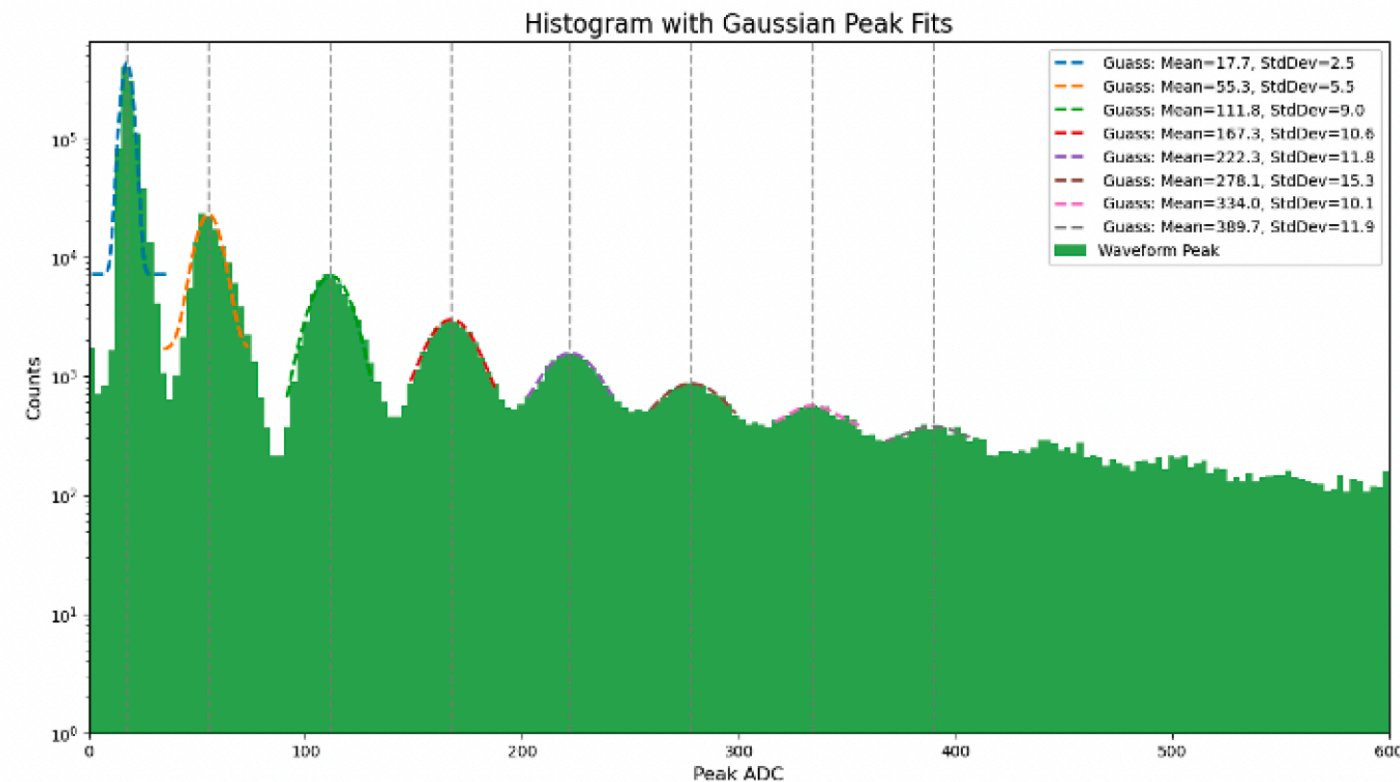
MiniGRAMS tile



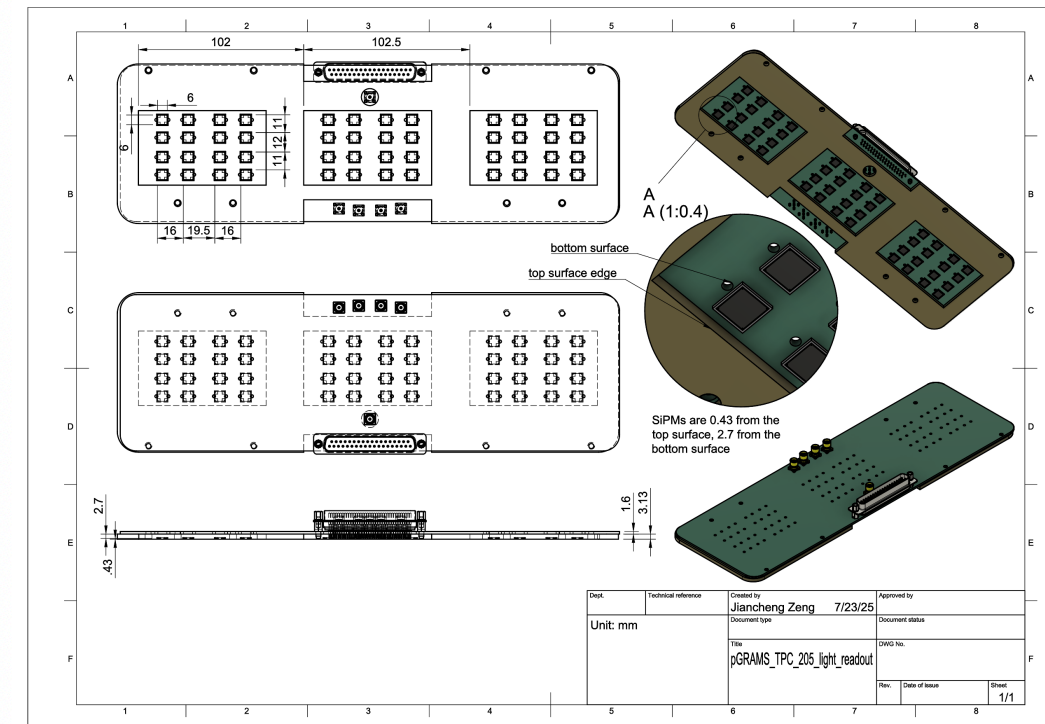
MiniGRAMS light readout



- 9x SiPM cell array
- more compact than PMT

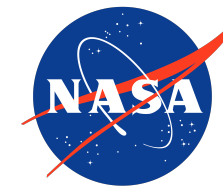


One VIS SiPM channel P.E spectrum, cr: Jon Sensenig



SiPM board designed by Jonathan LeyVa

GRAMS collaboration and timeline



International collaboration with different backgrounds/expertise

Gamma-rays, X-rays, Cosmic-rays, Neutrinos, Direct/Indirect DM searches

International

- ☐ Hiroshima University
- ☐ IPMU
- ☐ JAXA
- ☐ Kanagawa University
- ☐ Kyoto University
- ☐ Nagoya University
- ☐ National Defense Medical College
- ☐ Osaka University
- ☐ RIKEN
- ☐ Rikkyo University
- ☐ Tokyo University of Science
- ☐ Universität Würzburg
- ☐ University of Tokyo
- ☐ Waseda University
- ☐ Yokohama National University

USA

- ☐ Barnard College
- ☐ Columbia University
- ☐ Howard University
- ☐ NASA GSFC
- ☐ Northeastern University
- ☐ Oak Ridge National Lab
- ☐ University of California, Berkeley
- ☐ University of Chicago
- ☐ UT Arlington
- ☐ Washington University
- ☐ Yale University



2020

2025

2030

◆ eGRAMS flight

◆ pGRAMS flight

◇ Science flights

◇ Satellite

◆ Antiproton beam test

Summary



- ❑ GRAMS aims for both **gamma-ray** observations in the **poorly explored MeV** range and **indirect dark matter searches** with **antimatter**. The project started with a **balloon** experiment and will be expanded to a **satellite** mission.
- ❑ With a cost-effective, large-scale LArTPC detector, the sensitivity to MeV gamma rays can be **an order (two orders)** of magnitude improved with a **single balloon flight (Satellite)** compared with the previous missions.
- ❑ GRAMS low-energy **antinuclei** measurements can be essentially **background-free** dark matter searches while investigating and validating the possible dark matter signatures indicated in **Fermi GCE** (Galactic Center Excess) and **AMS-02 antiproton excess**.
- ❑ We successfully demonstrated the LArTPC detector performance during the **engineering flight** from the JAXA balloon facility and the **antiproton beam test** at J-PARC.
- ❑ As a step forward for future science flights, we will have a **prototype flight** scheduled for Spring 2026, supported by the NASA APRA program.



Thanks!

zeng.jia@northeastern.edu

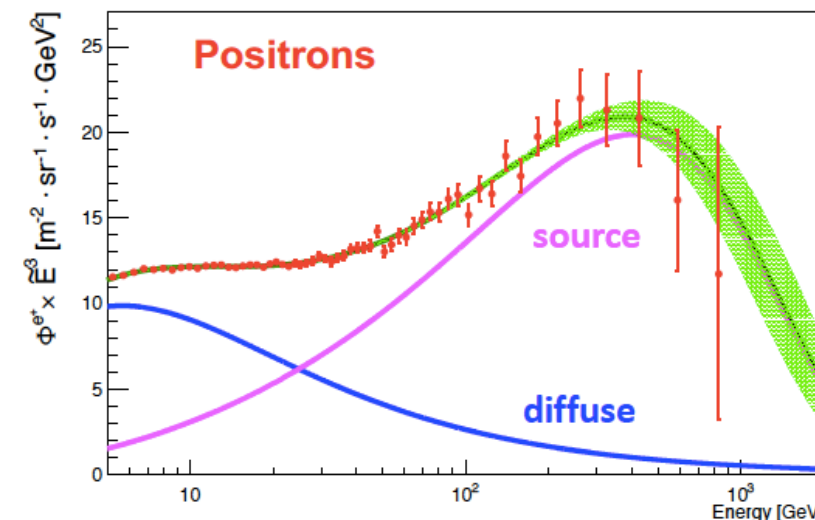
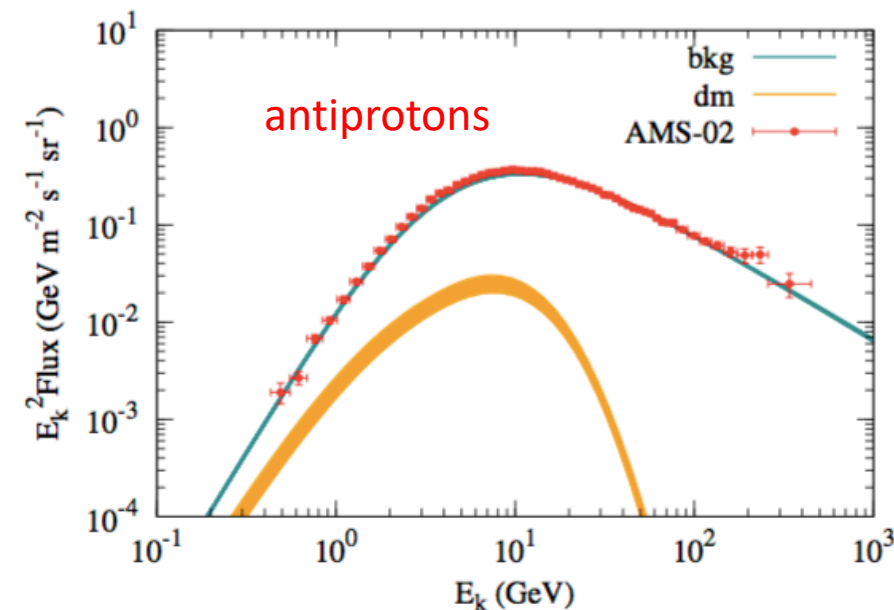
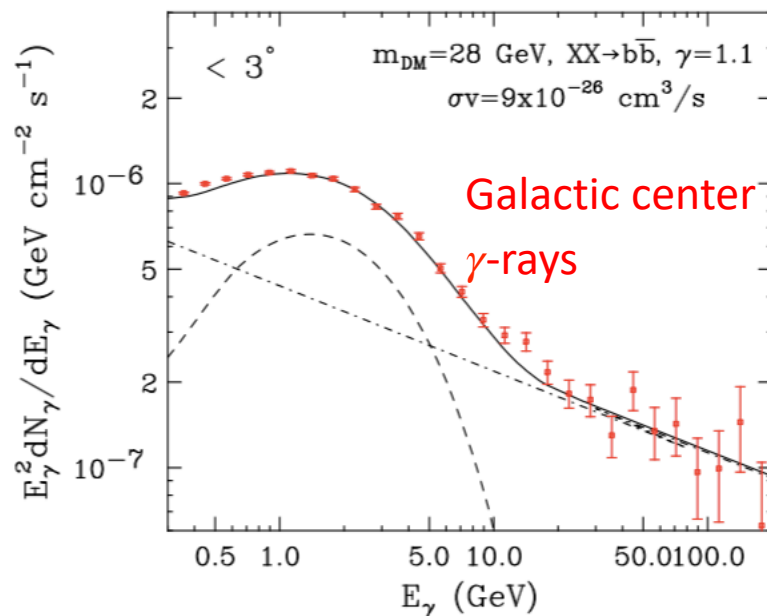
JCZeng1412@gmail.com

Backup slides

Indirect DM Searches



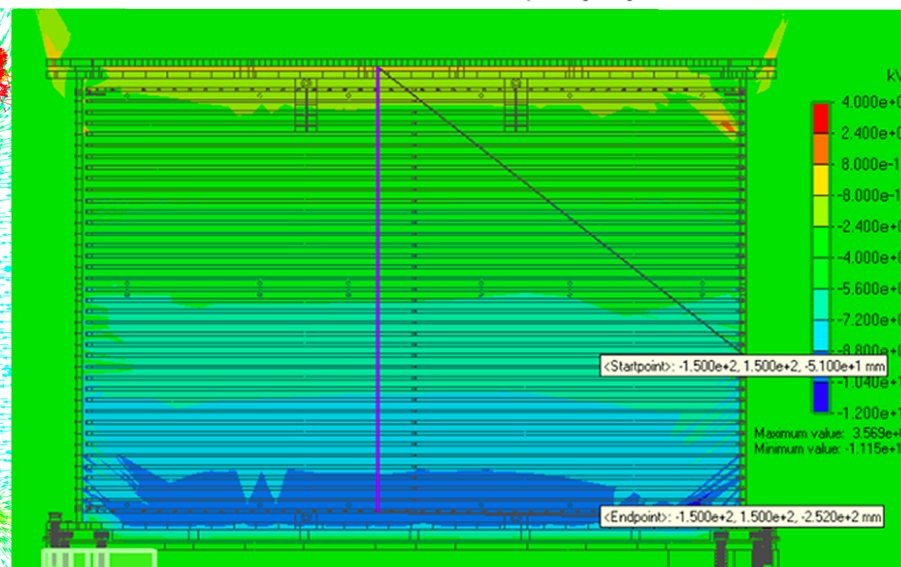
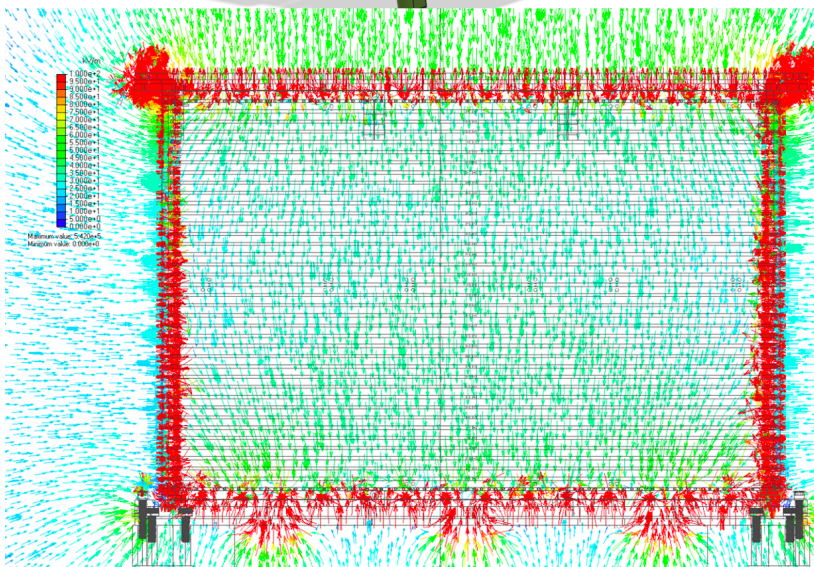
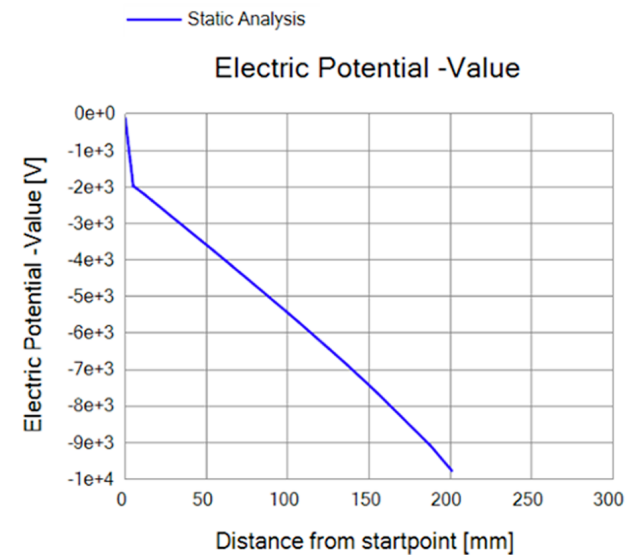
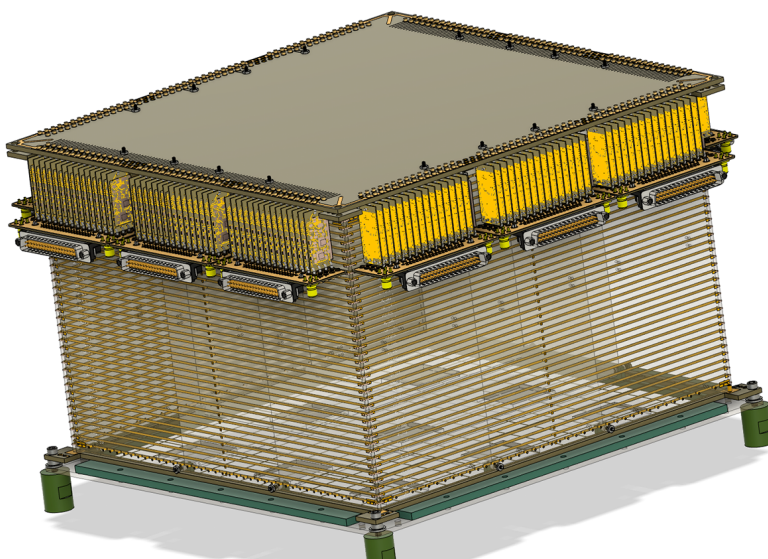
❑ Uncertain astrophysical backgrounds make indirect searches harder



We need background-free searches!



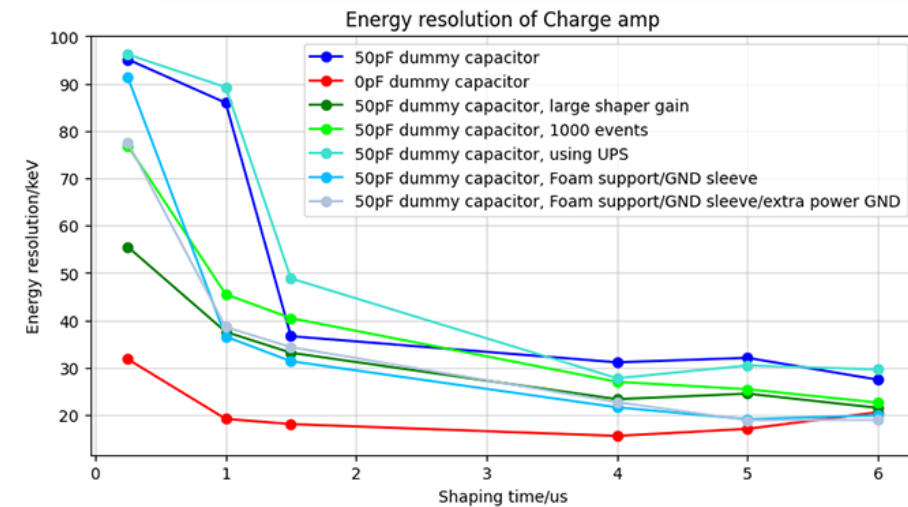
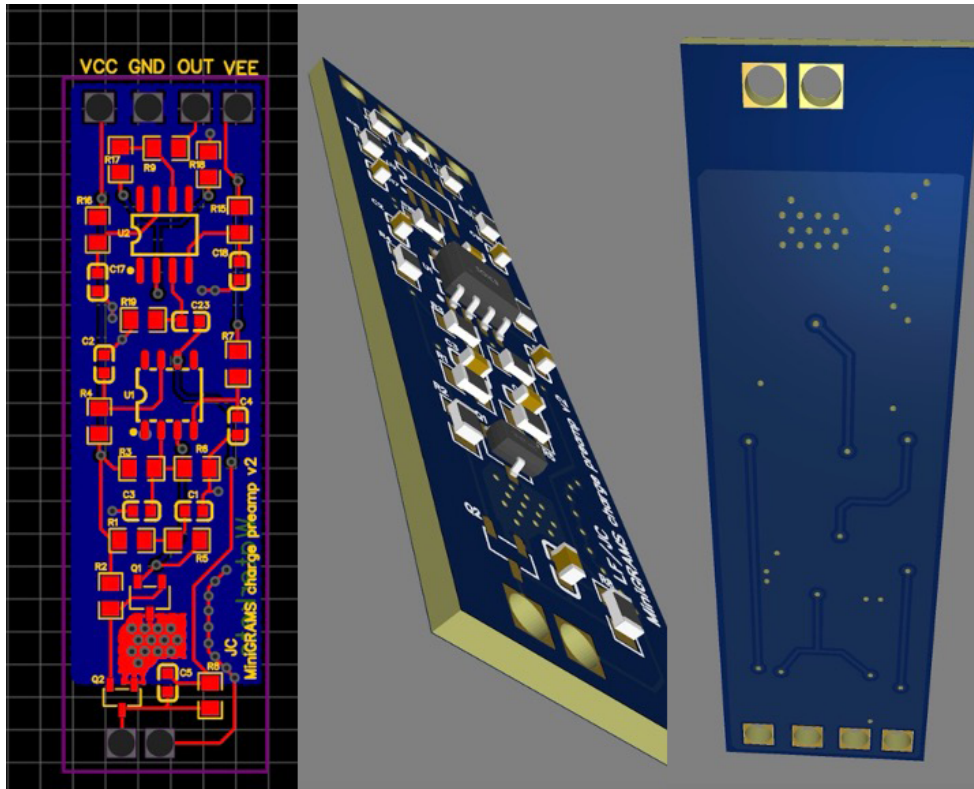
pGRAMS E-field



pGRAMS electronics

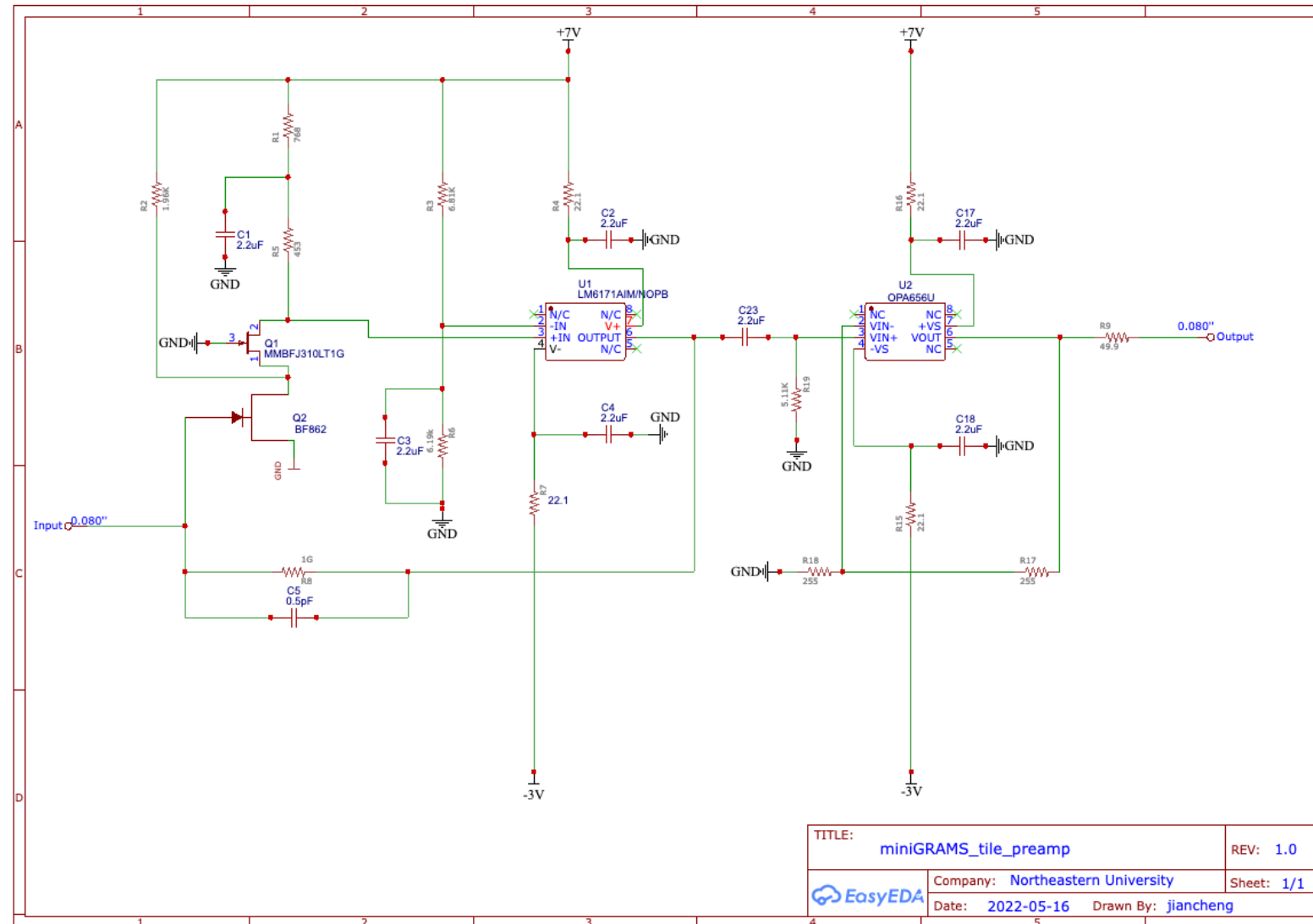
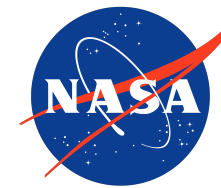


- ❑ Charge Sensitive Pre-amp(CSP) that works in cryo front end, minimize the pickup noise after readout tile

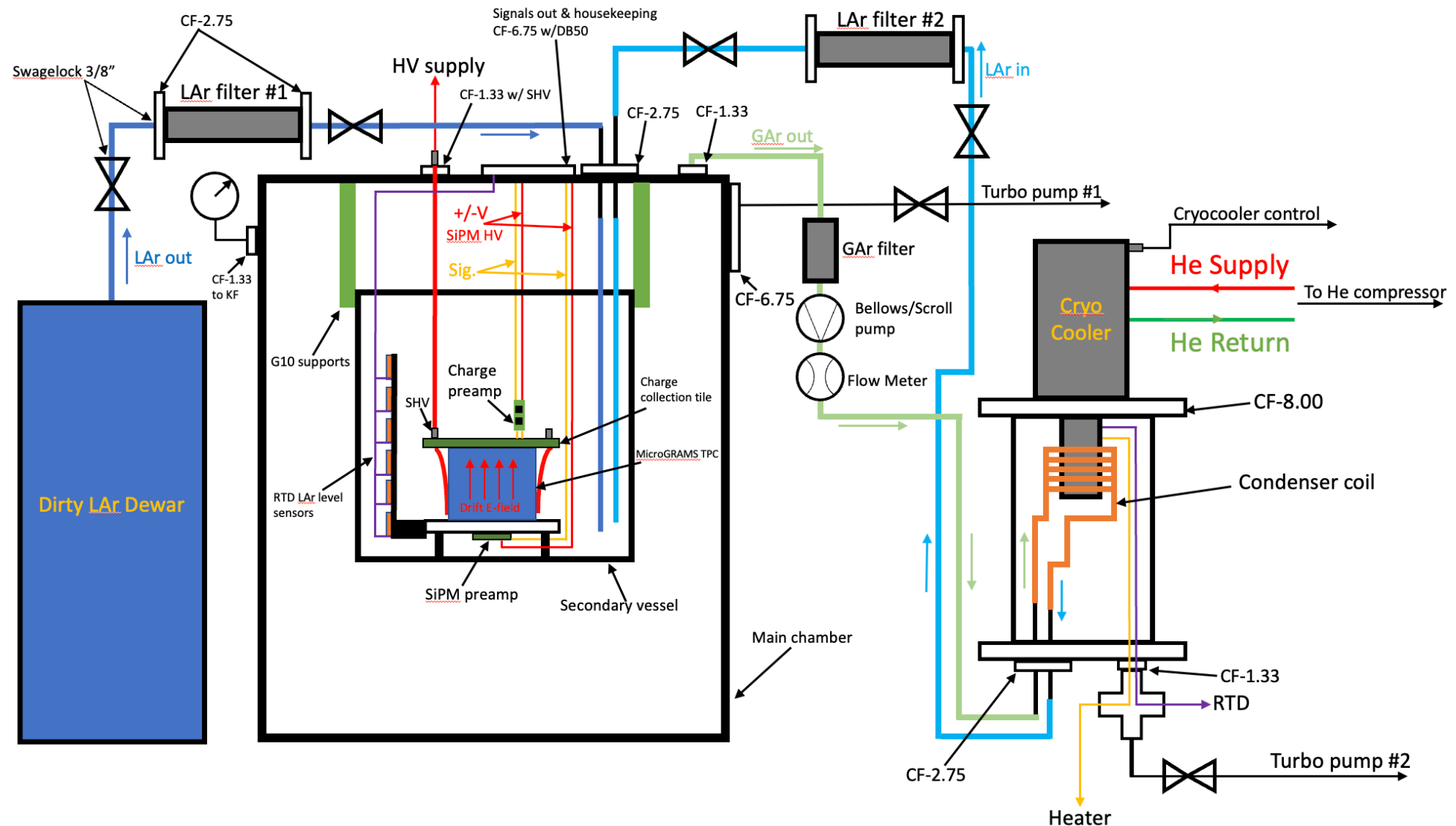
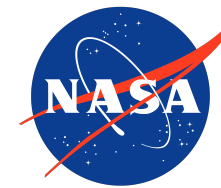


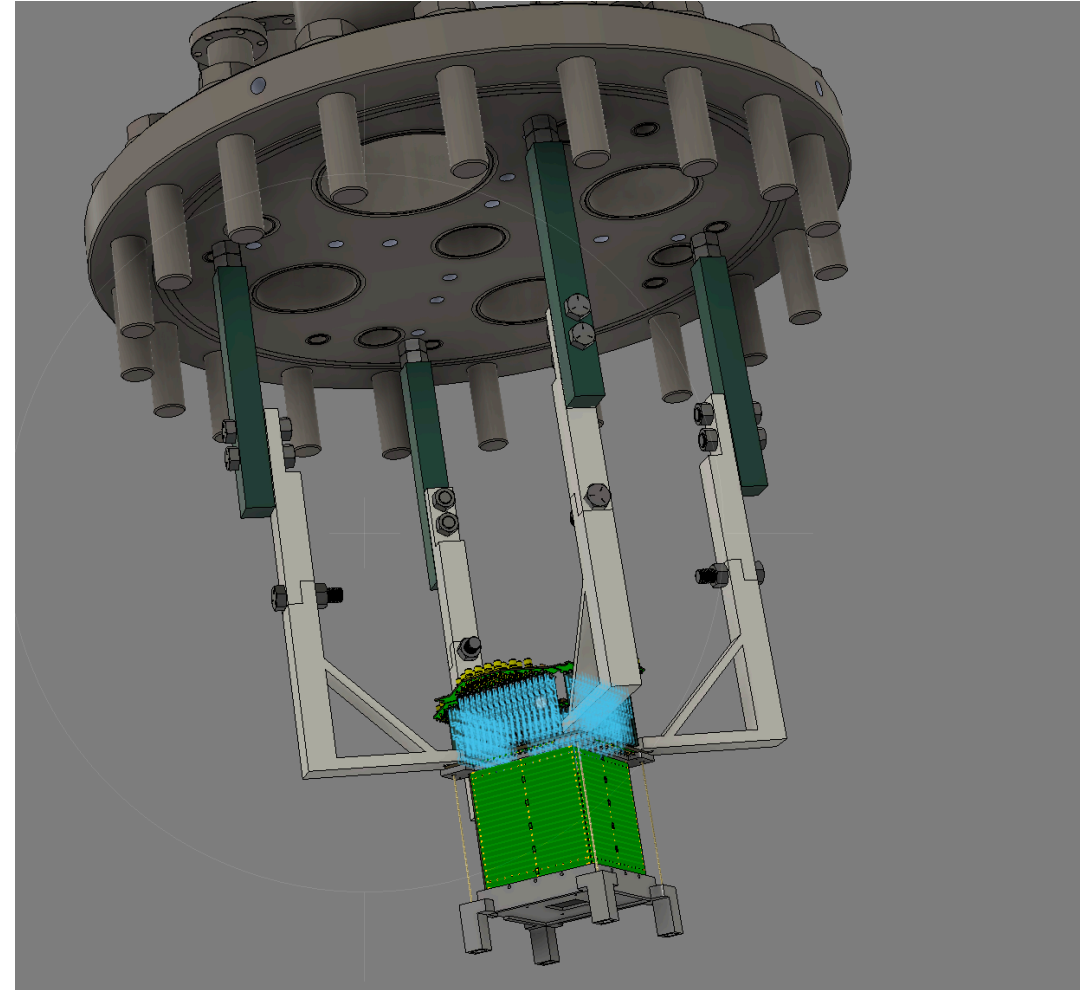
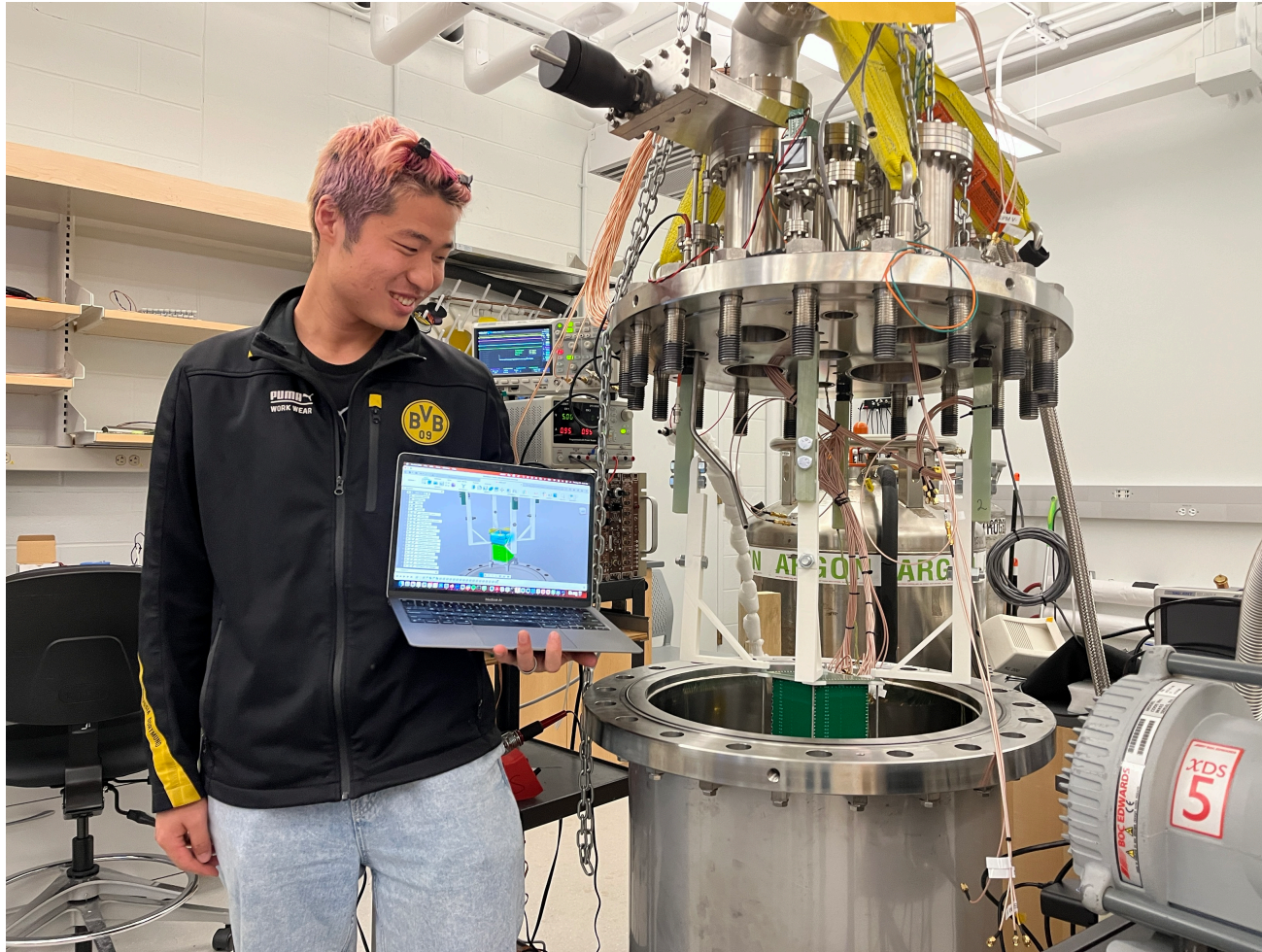
	0.25us/keV	1us/keV	1.5us/keV	4us/keV	5us/keV	6us/keV
0pF dummy	31.8	19.2	18.1	15.6	17.1	20.6
50pF dummy	55.517	37.58	33.165	23.363	24.528	21.523
50pF 1000 events	76.872	45.527	40.48	27.033	25.46	22.642
50pF no pump	185.83	43.64	47.49	25.89	21.3	21.35
50pF UPS	96.17	89.26	48.86	27.78	30.48	29.64
50pF ceiling GND	66.66	37.06	NaN	22.53	24.7	22.71
50pF foam/sleeve/ceiling GND	91.32	36.58	31.38	21.64	19.13	19.95
50pF extra power GND	77.54	38.6	34.35	22.72	18.92	18.98

S. Arathi preparing performance analysis paper



miniGRAMS



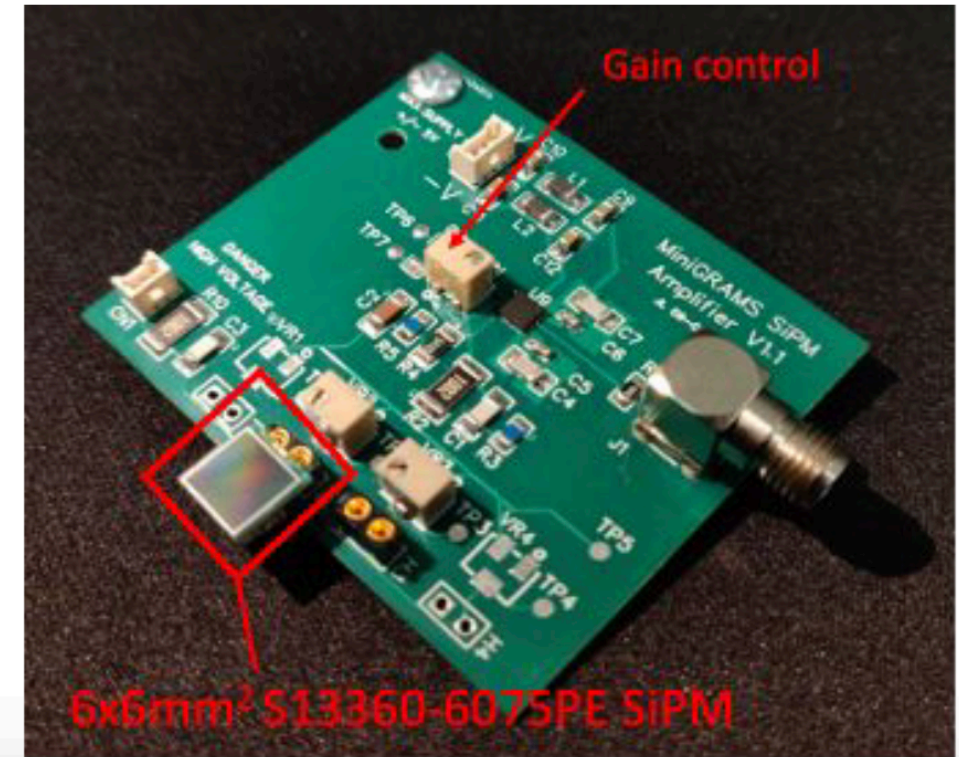
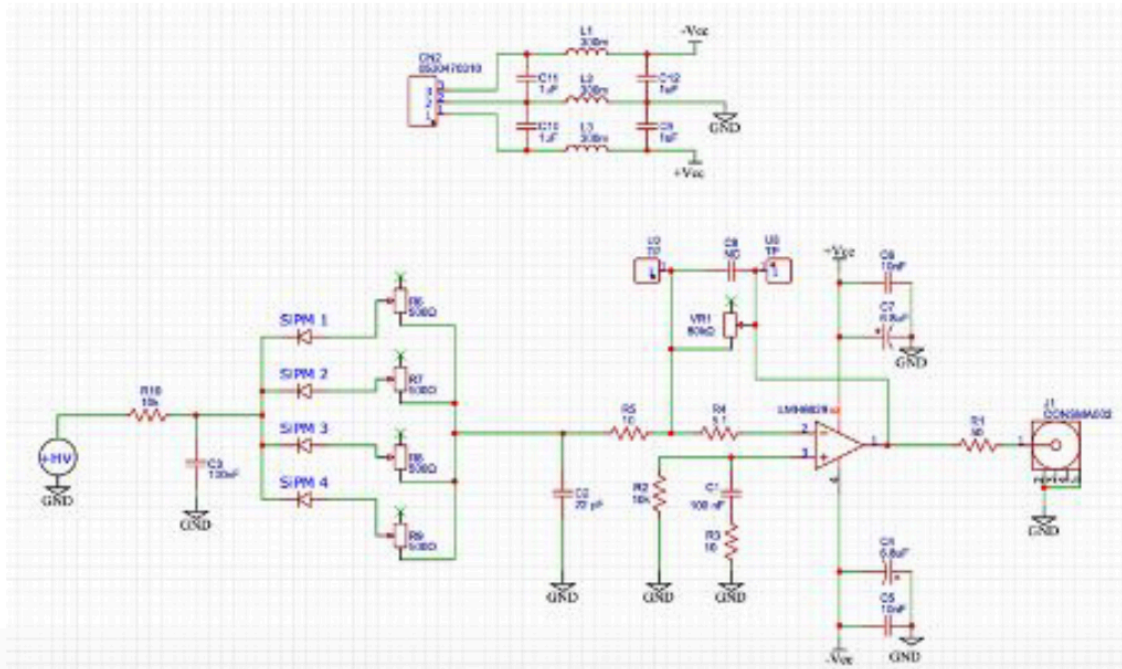


SiPM board

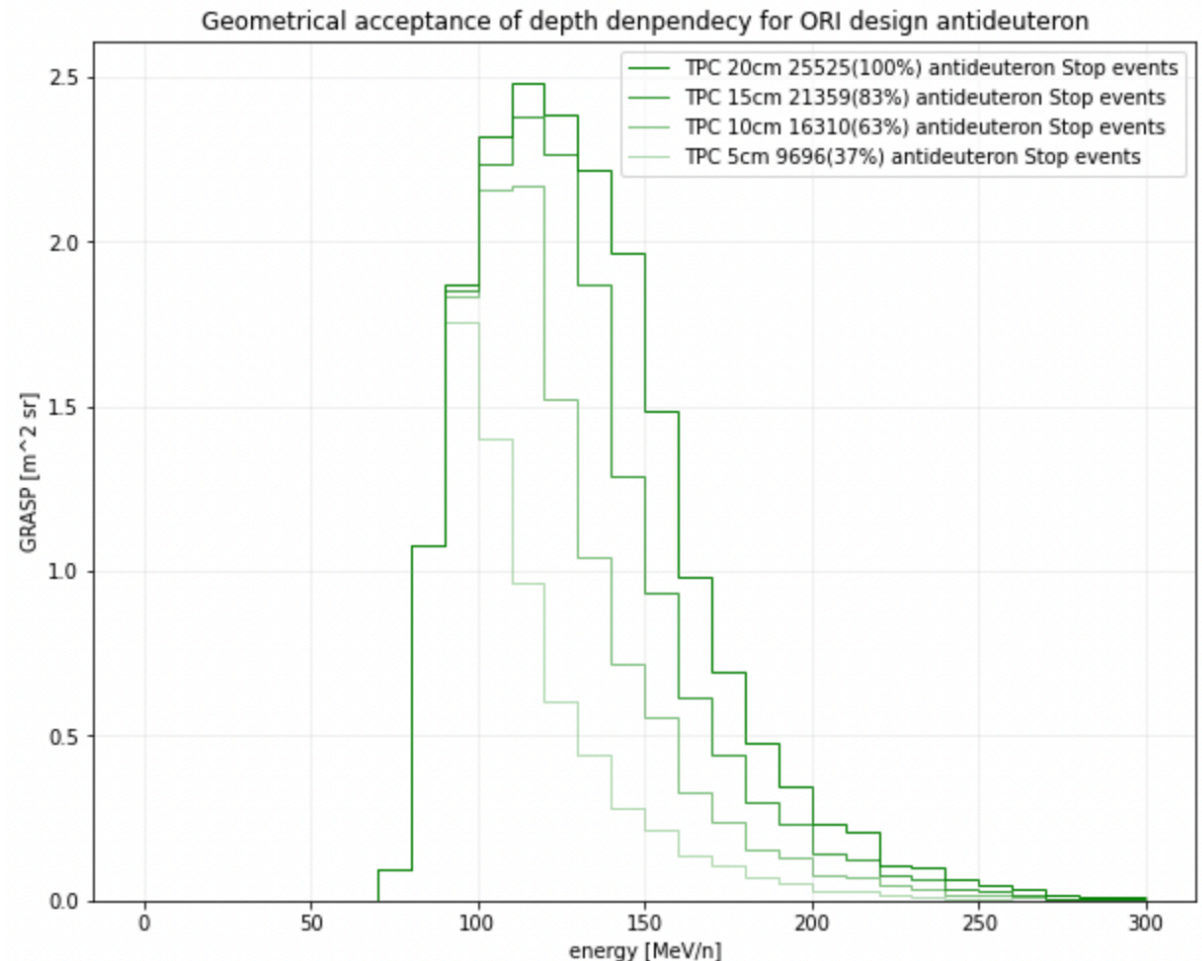


- Texas Instruments LMH6629

- Input noise: $0.69 \text{ nV}/\sqrt{\text{Hz}}$ and $2.6 \text{ pA}/\sqrt{\text{Hz}}$
- 900 MHz bandwidth (includes controllable compensation feature that sacrifices bandwidth for improved stability at gains as low as 4V/V)
- $1600 \text{ V}/\mu\text{s}$ slew rate
- Hetero-junction BJT, good for low temp stability
- Typical power consumption with 3.4V (+/- 1.7V) of dynamic range $\rightarrow \sim 30 \text{ mW}$ @ $T=87\text{K}$



-



Detector horizontally expand

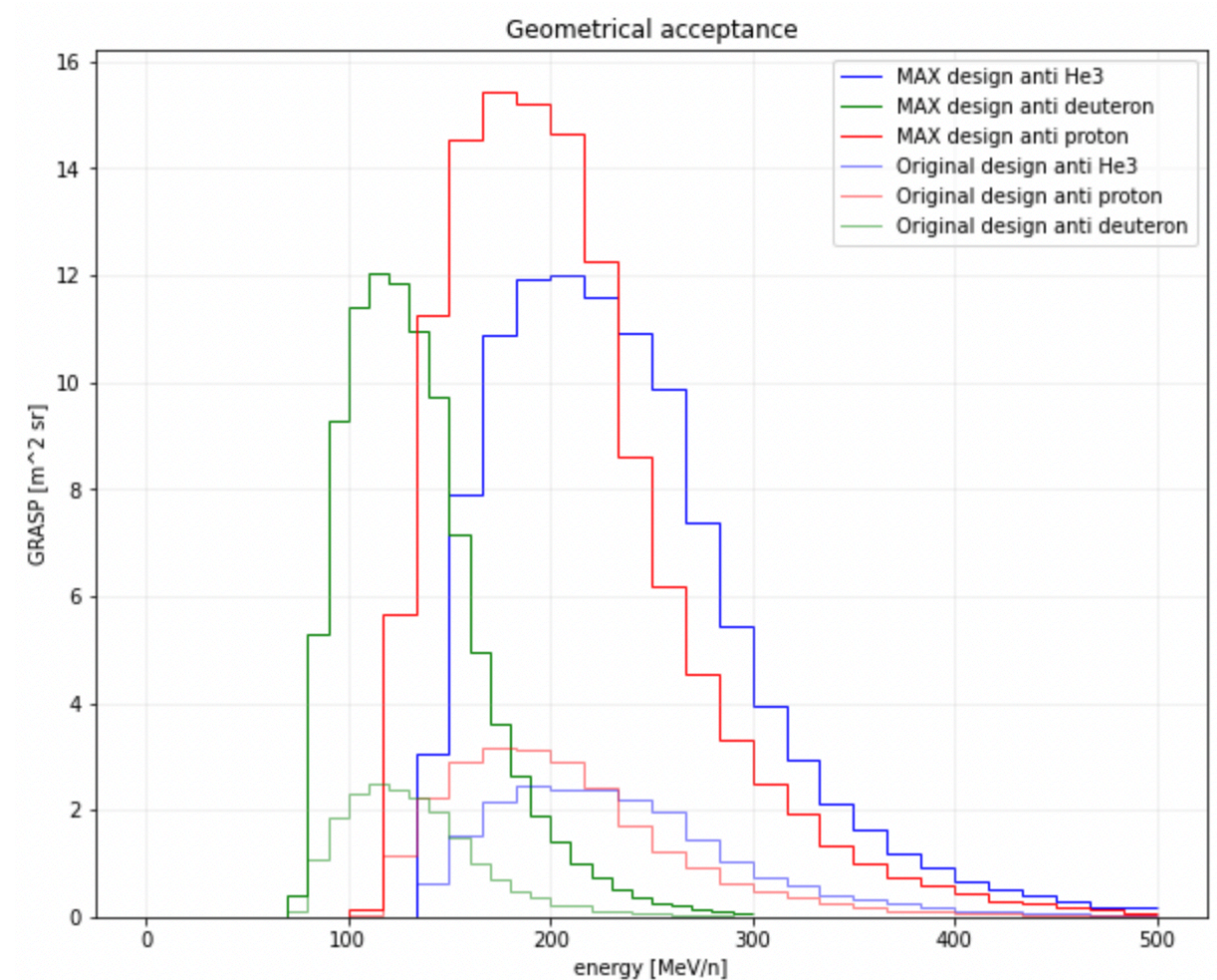
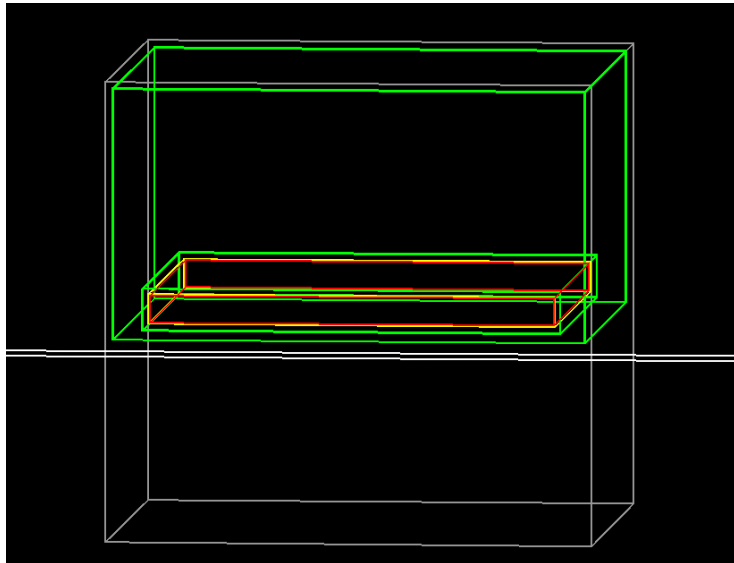


❑ Original design:

$140cm \times 140cm \times 20cm$

❑ MAX design:

$320cm \times 320cm \times 20cm$



TPC shape



Tracker + TOF 900kg TOF CONF 1424	
TPC length [cm]	TPC thickness [cm]
50	88.2
60	62.8
70	46.2
80	34.8
90	26.6
100	20.7
110	16.2
120	12.7
140	7.8
160	4.6
180	2.3
200	0.8
300	-2.9
600	-4.8

Tracker + TOF 1100kg TOF CONF 1424	
TPC length [cm]	TPC thickness [cm]
50	130
60	93.7
70	69.9
80	53.5
90	41.8
100	33.1
110	26.6
120	21.6
140	14.4
160	9.7
180	6.5
200	4.1
300	-1.4
600	-4.4

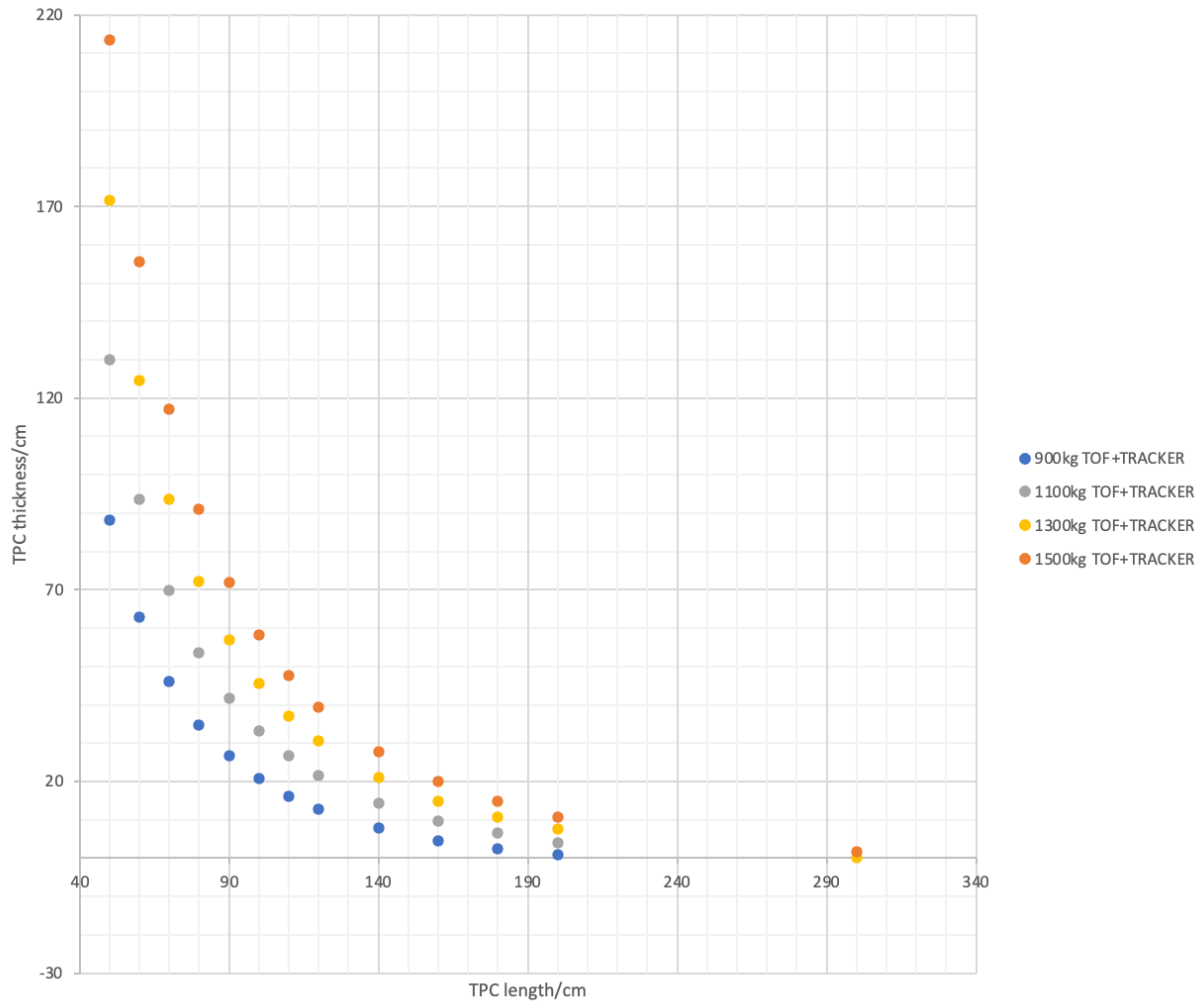
Tracker + TOF 1300kg TOF CONF 1424	
TPC length [cm]	TPC thickness [cm]
50	171.7
60	124.6
70	93.6
80	72.2
90	56.9
100	45.6
110	37.1
120	30.5
140	21.1
160	14.9
180	10.6
200	7.5
300	0.2
600	-4

Tracker + TOF 1500kg TOF CONF 1424	
TPC length [cm]	TPC thickness [cm]
50	213.5
60	155.5
70	117.2
80	90.9
90	72
100	58.1
110	47.5
120	39.4
140	27.7
160	20
180	14.7
200	10.8
300	1.7
600	-3.7

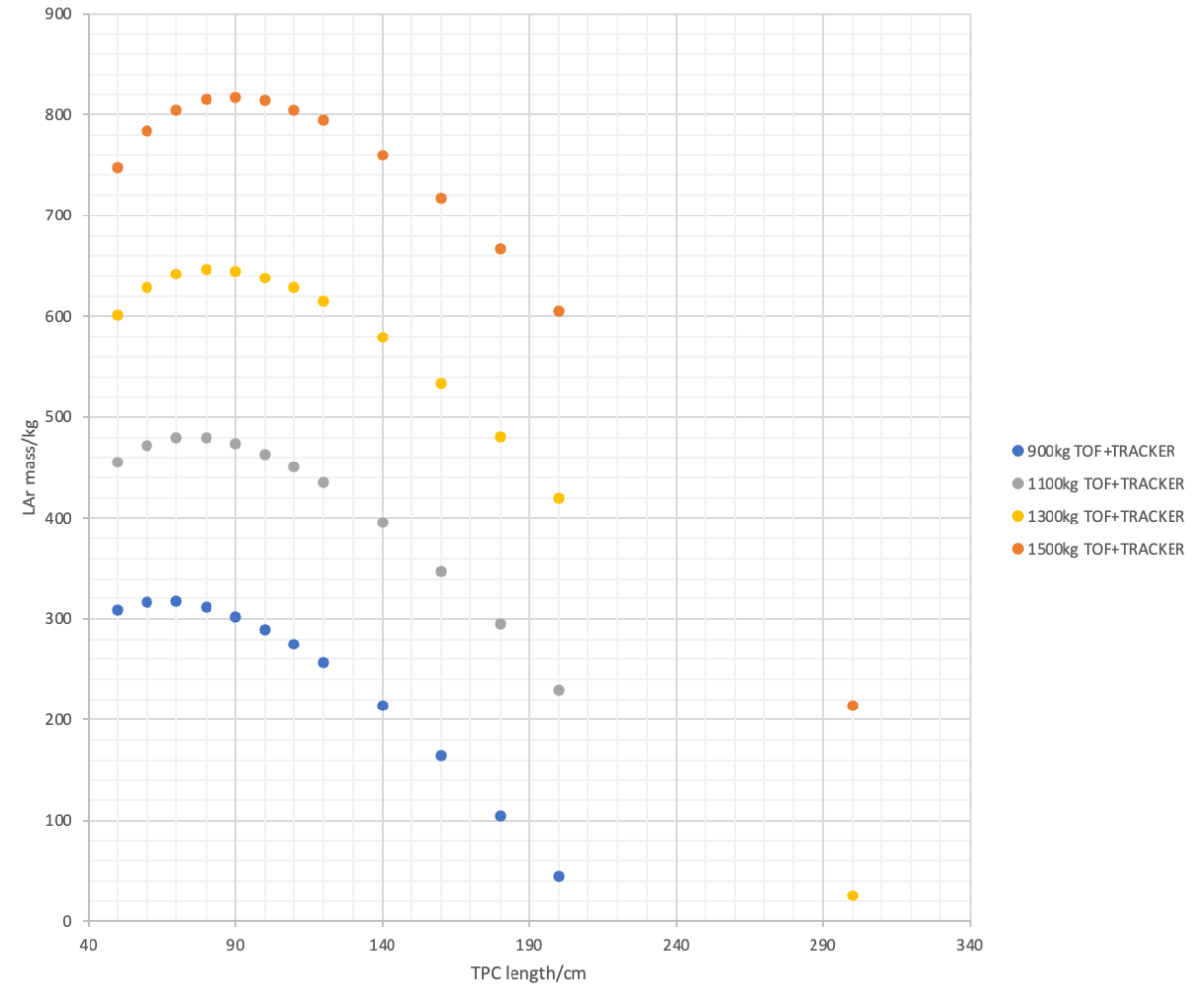
TPC Shape



Realistic TPC shape based on fixed TOF+TRACKER mass budget



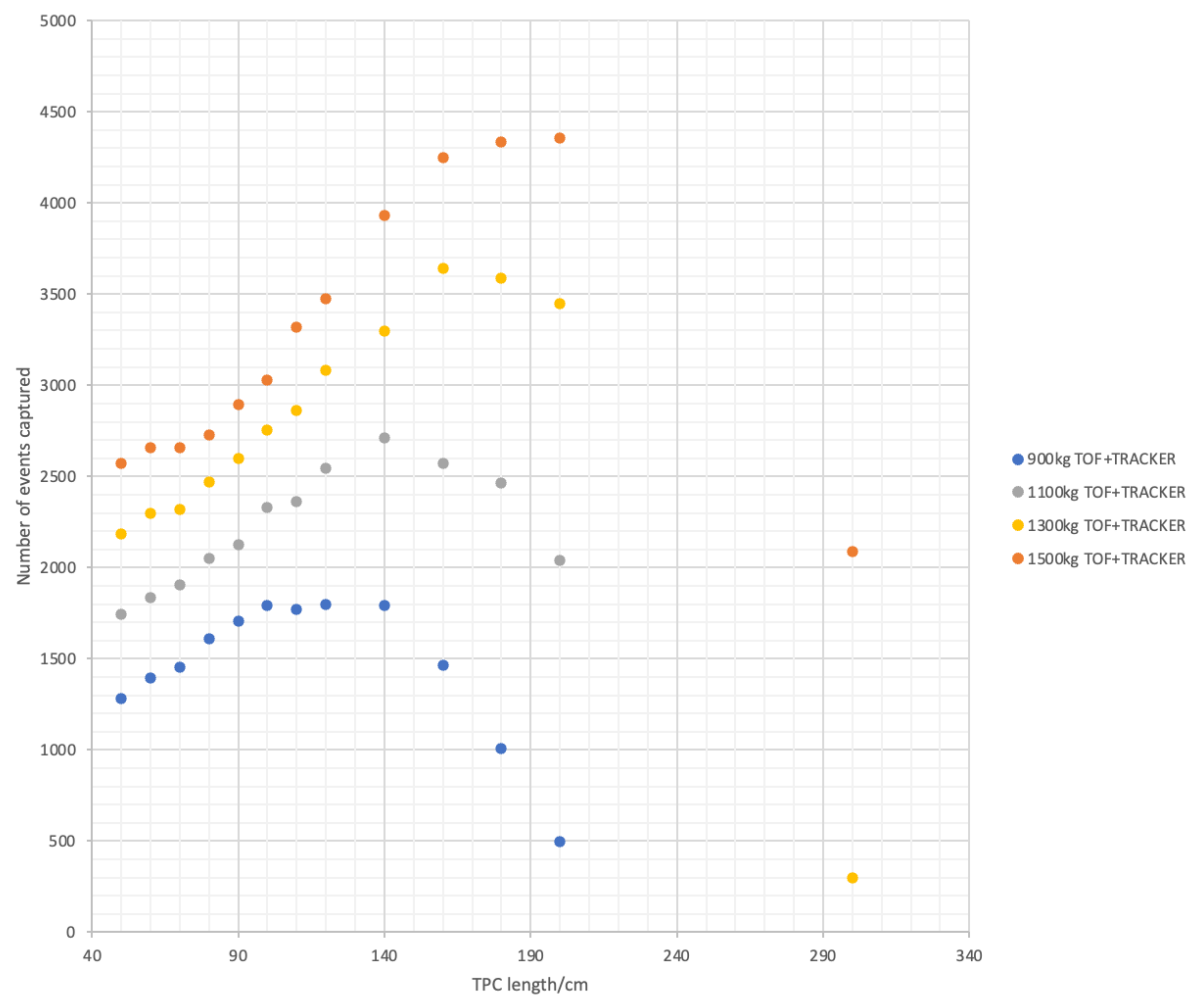
Realistic TPC shape based on fixed TOF+TRACKER mass budget



Mass budget comparison



Stop Events capture efficiency based on realistic TPC shape



Stop+In-Flight capture efficiency based on realistic TPC shape

