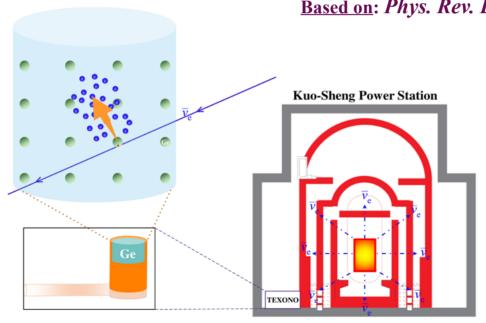
# New Limits on the Coherent Neutrino-Nucleus Elastic Scattering Cross Section at the Kuo-Sheng Reactor-Neutrino Laboratory

# S. Karmakar, Manoj. K. Singh, H.T. Wong

Institute of Physics, Academia Sinica

[On behalf of the TEXONO Collaboration]

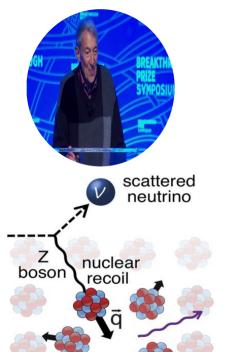
Based on: Phys. Rev. Lett. 134, 121802 (2025)







#### Introduction to CvA<sub>el</sub>



scintillation



PHYSICAL REVIEW D

VOLUME 9, NUMBER 5

Coherent effects of a weak neutral current

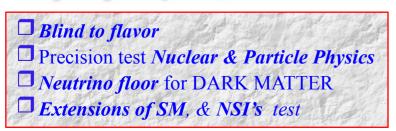
Daniel Z. Freedmant

National Accelerator Laboratory, Batavia, Illinois 60510 and Institute for Theoretical Physics, State University of New York, Stony Be (Received 15 October 1973: revised manuscript received 19 November

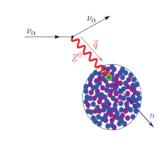
- Proposed right after the discovery of Neutral Current D. Z. Freedman [1974]
- ightharpoonup Signature  $\rightarrow$  Nuclear Recoil < O(1) (keV)
- $\triangleright v + N(A,Z) \rightarrow v + N(A,Z)$
- $\triangleright$  Coherent: Outgoing nucleon wave-functions are in phase  $[E_{ij} < O(10)MeV]$
- Elastic: Target remains in the same energy state

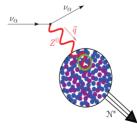
$$\frac{d\sigma_{C\nu A_{el}}}{dT} = \frac{G_F^2}{\pi} m_A \mathcal{Q}_W^2 \left( 1 - \frac{m_A T_A}{2E_\nu^2} \right) \mathcal{F}^2(T_A)$$

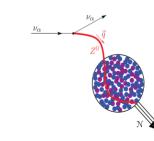
 $\left| \frac{d\sigma_{C\nu A_{el}}}{dT} = \frac{G_F^2}{\pi} m_A \mathcal{Q}_W^2 \left( 1 - \frac{m_A T_A}{2E_\nu^2} \right) \mathcal{F}^2(T_A) \right| \left| \mathcal{Q}_W = g_V^p Z + g_V^n N = \left( \frac{1}{2} - 2 \sin^2 \theta_W \right) Z - \frac{1}{2} N \right|$ 



recoils







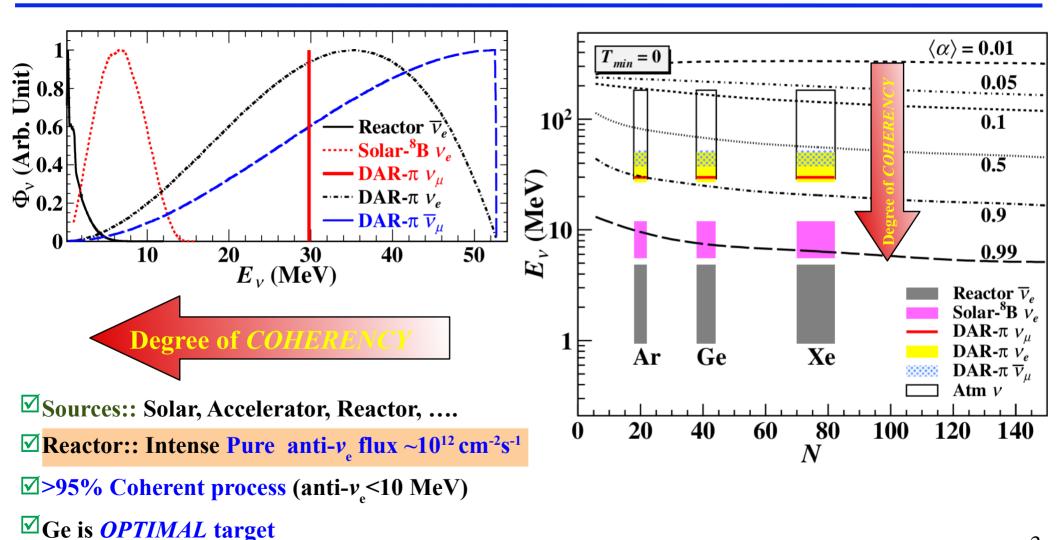
Inelastic incoherent  $\lambda_{Z^0} \ll 2R$ 

Elastic incoherent  $\lambda_{Z^0} \lesssim 2R$ 

Elastic coherent ( $CE\nu NS$ )  $\lambda_{Z^0} \gtrsim 2R$ 

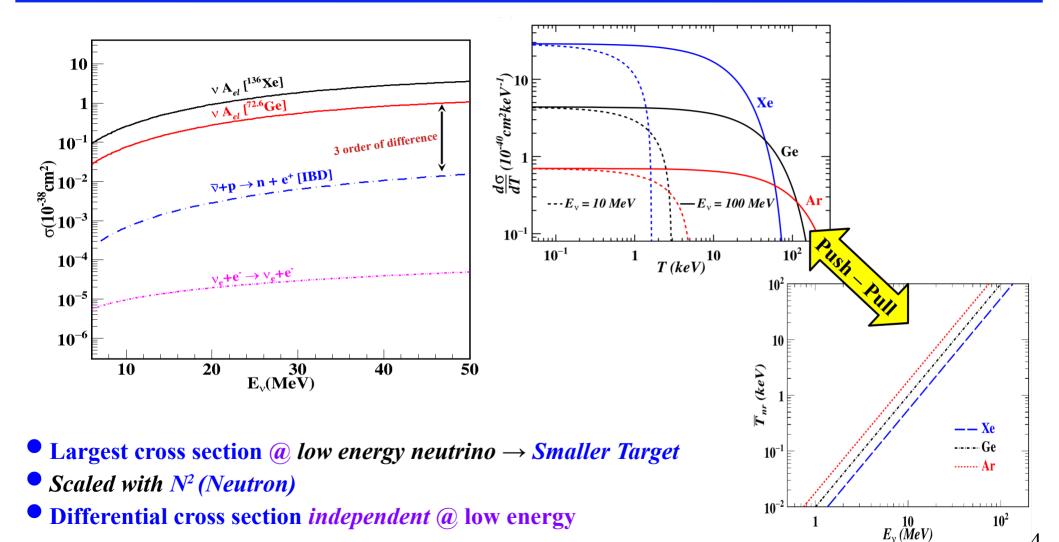
MARCH 1974

# Reactor and Ge for CvA<sub>el</sub> study

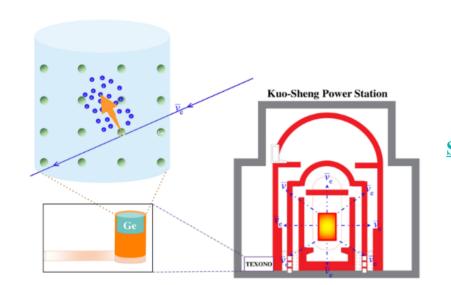


3

## Reactor and Ge for CvA<sub>el</sub> study



# Quenching Factor for CvA<sub>el</sub> study



**The** recoil spectrum → Convoluted with QUENCHING FACTOR

 $\square$  Lindhard k = 0.157 (Ge)

**☑**No Binding effect

**☑** Ionized electrons do not produce recoil atoms of appreciable energy

**☑** *Migdal effect* 

#### **Quenching factor**

$$\frac{dR}{dE_I} = \frac{dR}{dE_R} \left( \frac{1}{Q} - \frac{E_I}{Q^2} \frac{dQ}{dE_I} \right)$$

"q" can be (+/-) in Sign

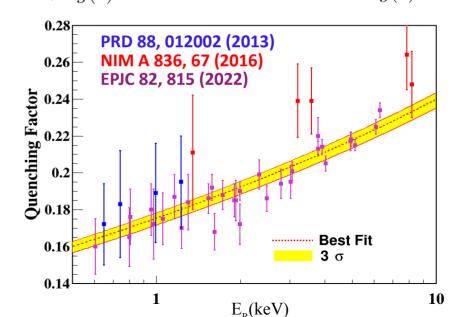
☑ (+) Sharp cutoff to energy transfer

☑ (-) Enhancement to energy transfer

**☑** (0) For the Current Work

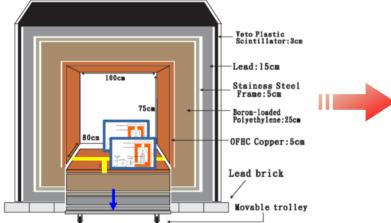
#### **Standard Lindhard QF definition**

$$Q(E_R) = \frac{kg(\epsilon)}{1 + kg(\epsilon)} \qquad \qquad Q(E_R) = \frac{kg(\epsilon)}{1 + kg(\epsilon)} - \frac{q}{\epsilon}$$



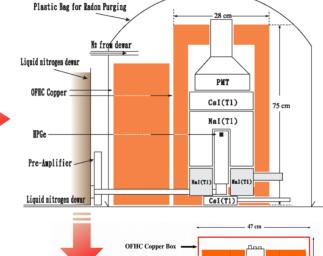
# **TEXONO** [Taiwan EXperiment On NeutrinO]

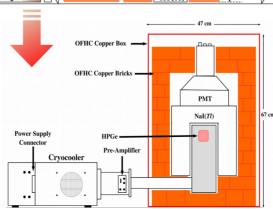
- **□** Location: Kuo-Sheng Nuclear Power Plant -II on northern shore of <u>Taiwan</u>
- **Theme:** Low Energy Neutrino Physics and Dark Matter Searches
- **☐** Collaboration: India, China [CDEX] & Turkey
- Flux: Reactor Power of 2.9 GW gives 6.35×10<sup>12</sup> cm<sup>-2</sup>s<sup>-1</sup> @ distance of 28 m
- **Shielding:** 30 m.w.e. overburden



## Eelectrocool Upgrade

- **Extra Space**
- **☐** Two detector configuration in working
- Custom Cold-tip temp & Real-time monitoring
- **☑** No LN<sub>2</sub> required & Less human exposure
- Less micro-phonic noise

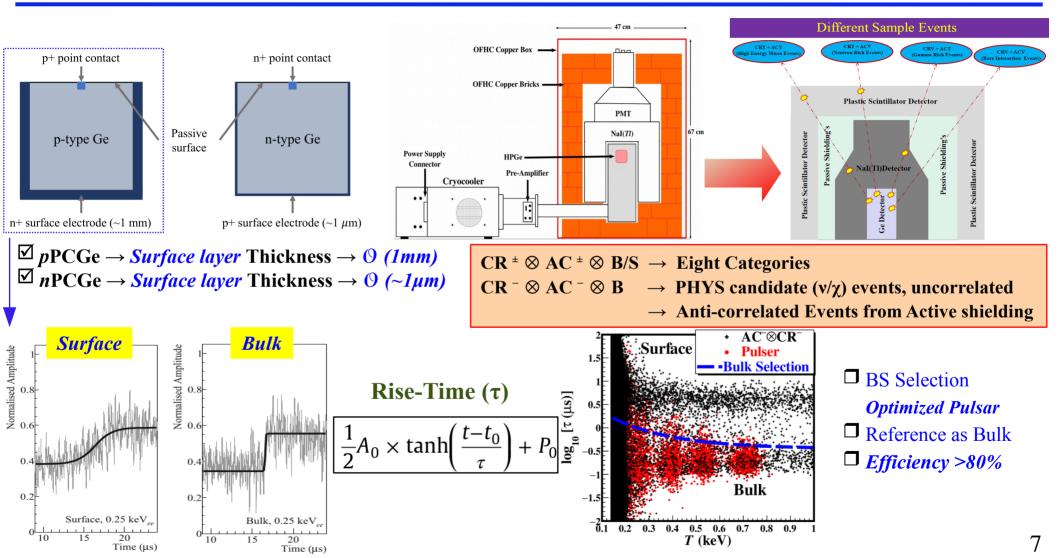




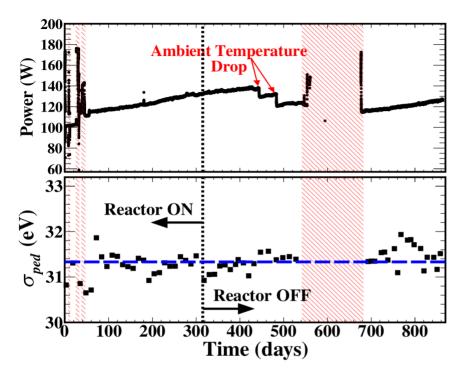




# Shielding & Bkg @ TEXONO



# Generations of HPGe Detector @ TEXONO & Stability

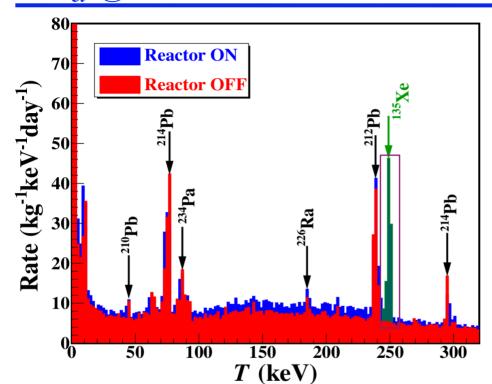


		Generation	Mass (g)	Pulsar FWHM (eV)	Threshold
$LN_2$	Y	G1	500	130	500
		G2	900	100	300
Electro-Cooling	V	G3	500	70	200
			900	70	~230
		$G3^{+}$	1383	70	200
		G4	900	<50	<140



- □ 200 eV Threshold
- □ Pulsar FWHM 70 eV
- ☐ Controlled Background
- ☐ Stable: >3.5 years

- → Pedestal fluctuation < 0.5 eV @ <3.5 eV
- **→** Power consumption
  - Ambient temperature
  - Outgasing



PHYSICAL REVIEW D 75, 012001 (2007)									
TABLE IV. Summary of $\gamma$ -lines intensity measured in Period-III.									
Energy (keV)	Isotopes	Source/ Decay Series	$ au_{1/2}$	Intensity (kg <sup>-1</sup> day <sup>-1</sup> )					
66.7	<sup>73</sup> <i>m</i> Ge	cosmic	0.5 s	$15.4 \pm 0.4$					
92.6	<sup>234</sup> Th	$^{238}\mathrm{U}$	24.1 d	$11.9 \pm 0.5$					
143.8	$^{235}U$	$^{235}U$	$7.0 \times 10^{8} \text{ y}$	$5.1 \pm 0.8$					
185.7	$^{235}U$	$^{235}U$	$7.0 \times 10^8 \text{ y}$	$17.2 \pm 0.4^{a}$					
186.2	<sup>226</sup> Ra	$^{238}\mathrm{U}$	1600 y	$17.2 \pm 0.4^{a}$					
238.6	<sup>212</sup> Pb	<sup>232</sup> Th	10.6 h	$18.8 \pm 0.5$					
249.8		unidentified		$11.6 \pm 0.5$					
295.2	<sup>214</sup> Pb	$^{238}\mathrm{U}$	26.8 m	$6.3 \pm 0.3$					
338.3	$^{228}Ac$	<sup>232</sup> Th	6.2 h	$3.7 \pm 0.5$					
351.9	<sup>214</sup> Pb	$^{238}\mathrm{U}$	26.8 m	$17.1 \pm 0.4$					
463.0	$^{228}Ac$	<sup>232</sup> Th	6.2 h	$1.6 \pm 0.3$					
583.2	$^{208}{ m Tl}$	<sup>232</sup> Th	3.1 m	$14.4 \pm 0.3$					

#### **To Achieve Better Sensitivity:**

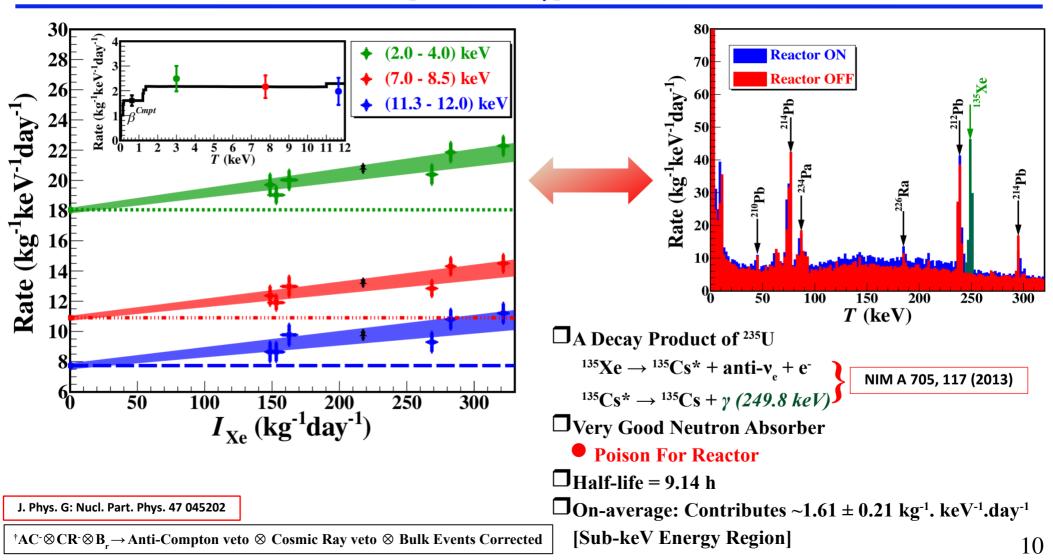
- □ Reactor OFF data collection
- R&D to achieve Low Energy Threshold & Less Background

# Background

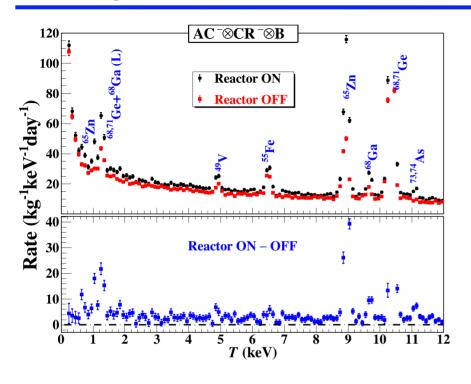
In sub-keV region: ~50 counts (kg⁻¹ keV⁻¹ day⁻¹)

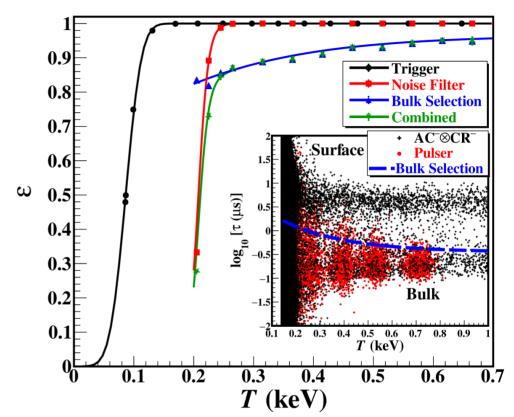
135Xe @ REACTOR ON

# **TEXONO:** <sup>135</sup>Xe Subtraction [250-keV γ]



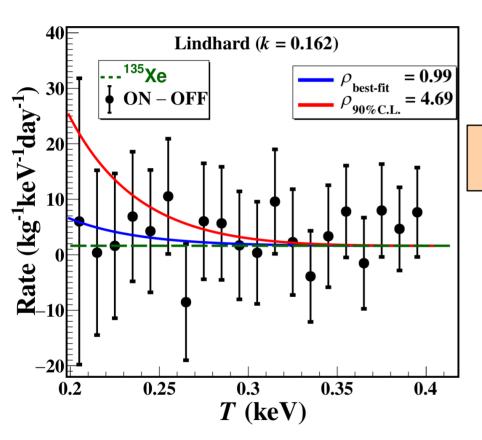
# CvA<sub>el</sub> @ KSNL





- The cosmogenic x-ray lines are identified
- Reactor ON-OFF → Finite <sup>135</sup>Xe Compton excess</sup>
- **Cosmogenic peaks** are observed → *Isotopes with half-lives comparable to exposure*

## **Limits (a) TEXONO**



- $\supseteq \rho$  estimate the excess over SM prediction
- $\supset \beta_{Cmpt}$  <sup>135</sup>Xe excess calibrated @ Sub-keV

Classical Statistic 
$$\chi^{2}(\rho, \beta; k) = \sum_{i} \left[ \frac{N_{i} - \rho \nu_{i}^{\text{SM}}(k) - \beta}{\Delta_{i}} \right]^{2} + \left[ \frac{\beta - \beta^{\text{Cmpt}}}{\Delta^{\text{Cmpt}}} \right]^{2}$$

- ⇒ < 280 eV contribute 90% of vA<sub>al</sub> signal
- **⊃** Spectral uncertainty 4.26 cpkkd < 280 eV

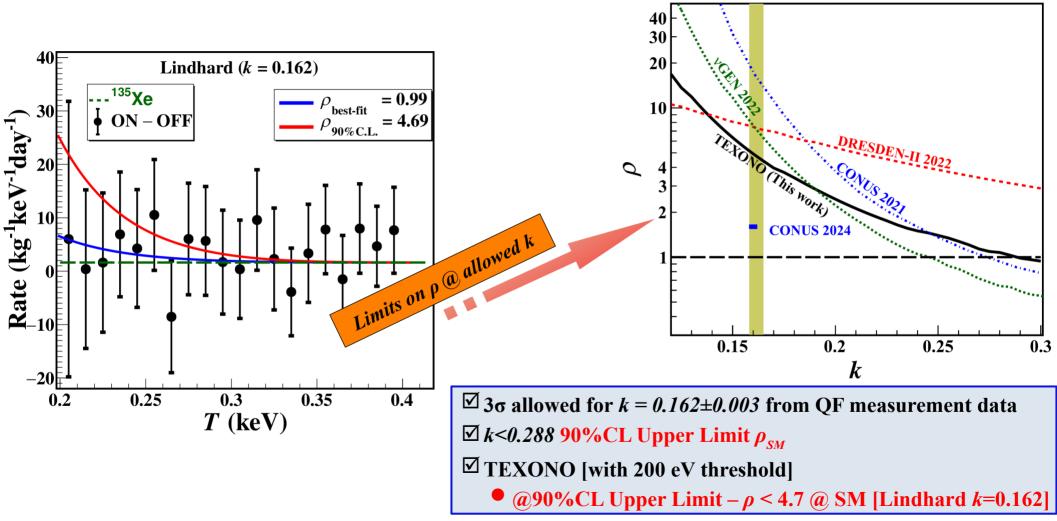
$$\beta = 1.62 \pm 0.22$$
 cpkkd

- ⇒β < Spectral Uncertainty @ RoI
- Prising Uncertainty → COMBINED

  Efficiency < 250 eV

\*cpkkd → counts.kg<sup>-1</sup>. keV<sup>-1</sup>.day<sup>-1</sup>

# Limits @ TEXONO



### **Result & Conclusion**

- ✓ Achieved [with 200 eV threshold]
  - $\rho = 0.99 \pm 2.23 \text{ (Stat.)} \pm 0.05 \text{ (Sys.)}$  @ SM [Lindhard k=0.162]
  - $\bigcirc$  90%CL  $\rightarrow \rho$  < 4.7 @ SM [Lindhard k=0.162]
  - **OReactor ON[OFF]**  $\rightarrow$  242[357] kg-days
- Have not observed any excess above SM prediction
- Set the benchmark for precision testing of SM & Beyond







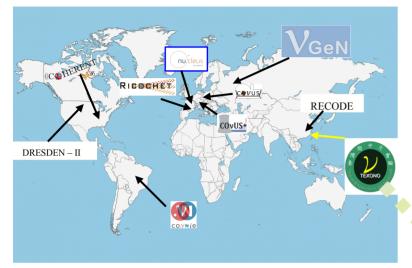
[RECODE program]

See the Talk of "Yufeng Wang"

ID: 408 (Wednesday, August 27, 17:20)

□ KS Reactor decommissioned 2023 → Permission of data taking till end of 2028
 □ R&D continues to achieve → Lower (~150 eV<sub>ee</sub>) threshold Cross-Correlation, Optimized Pulsar, etc.

- □ New [G4] Detector → Characterization and Commissioning for Dark Matter studies
- □ New Reactor site → RECODE [Sanmen Reactor @ Zhejiang]



# Thank You! 謝謝

