Dark Matter Annual Modulation Analysis with Combined Nuclear & Electron Recoil Channels

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[On behalf of the TEXONO Collaboration]

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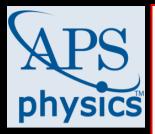
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(TEXONO Collaboration)

Overview

- ☐ Introduction and Motivation
- □Input Data
- **□**Expected Spectra
- **□**Cross Section
- **□**Analysis
- □ Results and Interpretations
- **□**Summary and Conclusions

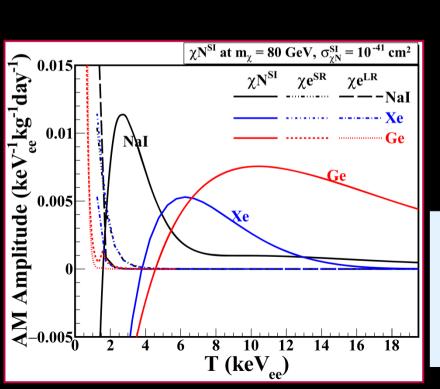
Introduction & Motivation

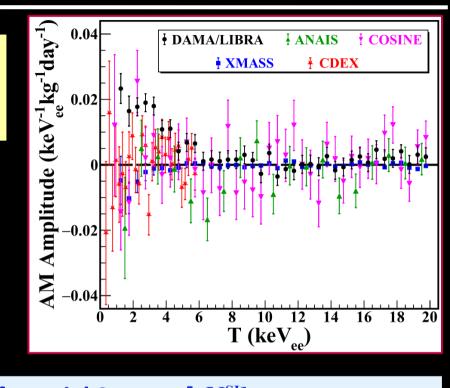
- Compelling Experimental Evidence (Cosmological Observations) ☐ One-quarter of Energy density of Universe == Composed of Dark Matter ☐ Exact Nature & Properties remain Unknown **Decades of Experimental Efforts** ■ Non-luminous & Non-baryonic ☐ Inferred Only from Gravitational Effects Only Result:: Consistent with Positive WIMP Signatures is from AM analysis on xN from DAMA/LIBRA (DL) [NaI(TI)] Search Efforts **►►► Challenged & Rejected::** Numerous Experiments + Variety of **□**Numerous Directions **Targets □**Diverse Techniques Attempts to Explain:: Scenarios other than xN detection [e.g. ☐ Intense areas of Fundamental Research complications in Analysis procedures] □ Favored Candidate | ■ WIMPs (χ) **Present Study::** Analysis including both χe & χN interactions Three Interactions:: SI χN^{SI}, Long & Short-range χe^{LR} & χe^{SR} Direct Experimental Search \square Assume Finite Interactions:: WIMPs with Electrons (χ_{α}) & Nuclei (χ_{α})
 - **☑** Positive Signatures::
 - **Excess Events** over known BKG [Measured Time-Integrated (TI) Energy Spectra] Sensitive to Uncertainties of BKG Modeling
 - Annual Modulation (AM) [Changes of Relative velocity between Earth & WIMPs in Galactic Halo]
 Only requires stable BKG with time & Independent of other details

Input

1. Published AM Amplitudes Data!

- ☐ [NaI(TI)]:: [DL, ANAIS, COSINE], [Xe] XMASS, [Ge] CDEX
- ☐ Complementary in Strength:: Probe Different Parameter Space

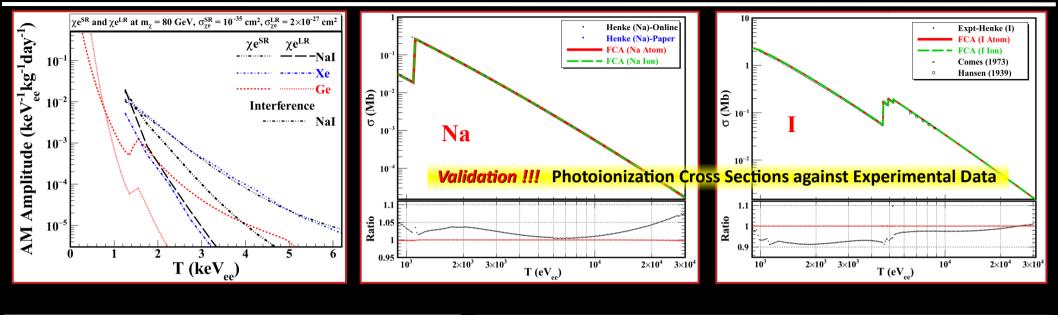




2. Expected <u>Differential Spectra</u> $[\chi N^{SI}]$

- Origin:: Differences in χ-velocity relative to Earth
- ☐ Maximum (minimum) Amplitudes :: June 1 (December 2)
- ☐ Feature :: Drop from Enhancement to Deficit at LE
- \square Turning Point $\longrightarrow m_{\gamma}$ -dependent

Input



4. Cross-Section



- I. Rapidly Rising Spectra
- II. Only Data < 4 keVee Contribute to Analysis
- **Σ** χε^{LR} Interactions has Additional 1/q² term
- I. Rise More Steeper @ Low Recoil Energy
- II. Favor experiments with Lower Detection Threshold
- \square Very Different Response at LE from χN^{SI}

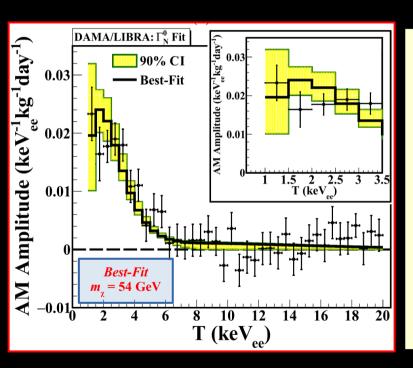
- [Talk: Prof. Cheng-Pang Liu → ID 359 !]
- **The Detection Channel** $\chi + A \rightarrow \chi + e^- + A^+$
- Frozen Core Approximation (FCA) PRD 102,123025 (2020) I. Experience on Ge & Xe Extend same approach to Na & I
 - II. Consistency Within ~ 5% [Across Energy Range of 1-30 keVee]
 - **III.** Identical Results [Whether Targets treated as **Atoms** or **Ions**]
 - **IV. Indicates::** Reliable Modeling to Interactions of χ with Atoms

Analysis

Best-fit Estimate Cross Sections @ Given m_{y}

$$\chi^2 = \sum_{i} \frac{1}{\Delta_i^2} \{ n_i - [\sigma_{\chi N}^{SI} \phi_{\chi N}^{SI}(T_i) + \sigma_{\chi e}^{LR} \phi_{\chi e}^{LR}(T_i) + \sigma_{\chi e}^{SR} \phi_{\chi e}^{SR}(T_i) + 2\sqrt{\sigma_{\chi e}^{LR} \sigma_{\chi e}^{SR}} \phi_{\chi e}^{int}(T_i)] \}^2$$

 $n_i & \Delta_i$ [AM Amplitudes & Uncertainties], ϕ [Normalized Spectral Functions]



DL Data::

- I. Positive AM Signatures & Reject Null Hypothesis [@ Large Significance]
- II. Best-fit Spectra only with χN -channel $[\Gamma_N^0: \sigma^{LR}] = \sigma^{SR}] = 0$
- III. Analysis Expands :: All Three Channels as free fitting variables

Two DM Scenarios::

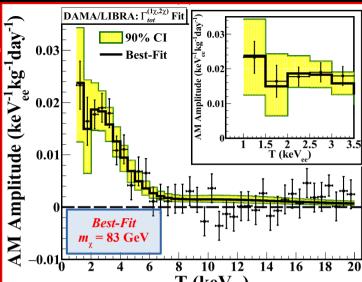
- 1. Parametrized by f_{γ} as DM Relic Density Fraction from χ interacting via χe
- 2. $\Gamma^{1\chi}_{tot} \rightarrow \text{Both } \chi \text{N \& } \chi \text{e Interactions are due to a single } \chi (f_{\chi} = 1)$ [Same constraints on m_{χ} apply to All Channels]
- 3. $\Gamma^{2\chi}_{tot}$ \Rightarrow Case of Independent constraints [Two different χ with fractional density f_{χ} & $(1-f_{\chi})$ interact separately via χe & χN]
- 4. Limiting Case $(f_{\gamma} = 0) \rightarrow$ Corresponds to Baseline Γ^0_N

Analysis

- Combined <u>best-fit</u> in $\Gamma^{1\chi}_{tot}$ == Spectra for $\Gamma^{2\chi}_{tot}$ Shifted m_{ν} from 54 GeV in Γ^{0}_{ν} to 83 GeV in Γ^{1}_{ν}
- Addition of xe-Channels Better Description data < 4 keVee
- <u>Interpretation</u> of DL data incorporating $\Gamma^{(1\chi,2\chi)}$...:
- I. Statistical Significance higher than Γ^0_N alone
- II. LE data (1-4 keVee) gives p-values of 0.52 for $\Gamma^{(1\chi,2\chi)}$ but only 0.07 for Γ^0_N
- III. Differences in $\chi^2/d.o.f.$ between Γ^0_N & $\Gamma^{(1\chi,2\chi)}_{tot}$ p-values of 0.02 & 0.008
- IV. For Complete 1-20 keV dataset :: $\chi^2/\text{d.o.f.}$ in $\Gamma^{(1\chi,2\chi)}$

Implies Additional physical processes ($\chi e^{LR} \& \chi e^{SR}$) to Explain AM spectrum

- Suggests a Scenario Published uncertainties are overestimated
- V. TEST Case :: Uncertainties of DL data Uniformly reduced by 20%
- Resulting in p-value = 0.5
- Γ Tension against Γ^0_N as a Valid Hypothesis is Stronger
- $\Gamma^{(1\chi,2\chi)}$ shows a Perfect Agreement with Data



aplita			
AM Amplitu	Best-Fit $m_{\chi} = 83 \text{ GeV}$		
~ 333 <u>- 1</u> 0	2 4 6	8 10 12 14 T (keV _{ee})	16 18
	Γ^0_{N} :: χN^{SI}	$\Gamma^{(1\chi,2\chi)}_{tot} :: \chi \mathbf{N}^{\mathrm{SI}} + \chi \mathbf{e}^{\mathrm{LR}} + \chi \mathbf{e}^{\mathrm{SR}}$	Compare
Data		χ²/d.o.f.	Δχ²/d.o.f.

₹ -0.01	2 4 6	8 10 12 14 T (keV _{ee})	16 18 2
	Γ^0_{N} :: χN^{SI}	$\Gamma^{(1\chi,2\chi)}_{tot} :: \chi \mathbf{N}^{SI} + \chi \mathbf{e}^{LR} + \chi \mathbf{e}^{SR}$	Compare
Data (keVee)	χ²/d.o.f. (p-value)		Δχ²/d.o.f. (p-value)
	Pi	ublished Data	

	$\chi_{\mathbf{N}^{\mathrm{SI}}}$	$ \begin{array}{c} \Gamma^{(1\chi,2\chi)} & :: \\ \chi N^{SI} + \chi e^{LR} + \chi e^{SR} \end{array} $	Compare
Data (keVee)	(Δχ²/d.o.f. (p-value)	
	Pı	ıblished Data	
1-20	32.06/36	22.40/34	9.66/2

		· ee-	
	Γ^0_{N} :: χN^{SI}	$\Gamma^{(1\chi,2\chi)}_{tot}::$ $\chi N^{SI} + \chi e^{LR} + \chi e^{SR}$	Compare
Data (keVee)	χ²/d.o.f. (p-value)		Δχ²/d.o.f. (p-value)
	Pı	ublished Data	
1-20	32.06/36 (0.66)	22.40/34 (0.94)	9.66/2 (0.008)

\ ee'				
	Γ^0_{N} :: χN^{SI}	$\Gamma^{(1\chi,2\chi)}_{tot} :: \chi \mathbf{N}^{SI} + \chi \mathbf{e}^{LR} + \chi \mathbf{e}^{SR}$	Compare	
Data (keVee)	χ²/d.o.f. (p-value)		Δχ²/d.o.f. (p-value)	
	Pi	ublished Data		
1-20	32.06/36 (0.66)	22.40/34 (0.94)	9.66/2 (0.008)	
1-4	8.6/4	1.3/2	7.26/2	

	Γ^0_{N} :: χN^{SI}	$\Gamma^{(1\chi,2\chi)}_{tot} :: \chi N^{SI} + \chi e^{LR} + \chi e^{SR}$	Compare	
Data	χ²/d.o.f.		Δχ²/d.o.f.	
(keVee)	(p-value)		(p-value)	
Published Data				
1-20	32.06/36	22.40/34	9.66/2	
	(0.66)	(0.94)	(0.008)	
1-4	8.6/4	1.3/2	7.26/2	
	(0.07)	(0.52)	(0.02)	

Test Case (Effects or reduced uncertainties)*

33.33/34

(0.50)*

1.94/2

(0.38)

48.1/36

(0.086)

12.8/4

(0.012)

1-20

1-4

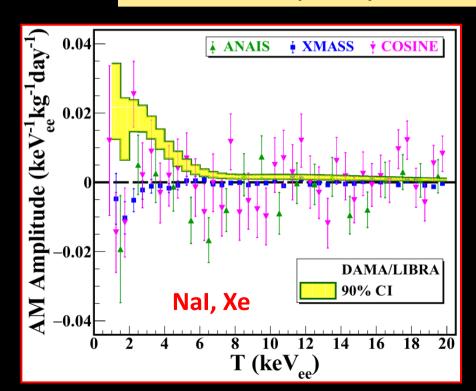
14.8/2

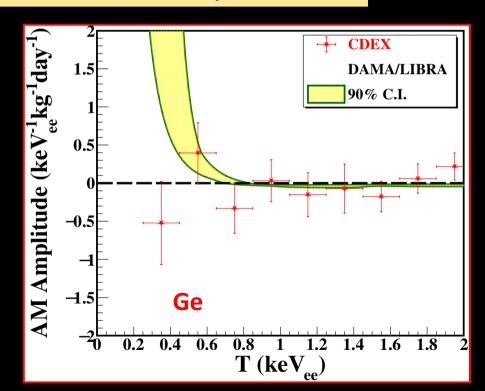
(0.0006)

10.86/2_ (0.0044)

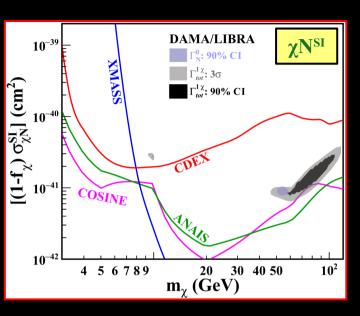
Analysis

- No AM Signatures Observed:: COSINE, ANAIS, XMASS, CDEX
 - **☑** Data consistent with Null Hypothesis
- Predicted Spectra for Null AM Experiments
 - Solution Due to $\Gamma^{(1\chi,2\chi)}_{tot}$ best-fit values of (χN^{SI}, χe^{LR}, χe^{SR}) derived from DL AM data
 - Shows:: Incompatibility of DL best-fit values with Null AM Experiments





Results & Interpretations:: Case of $\Gamma^{1\chi}_{tot}$



Two Allowed Regions in Γ⁰_N::

 @ Low & High m_χ
 Na & I-recoils

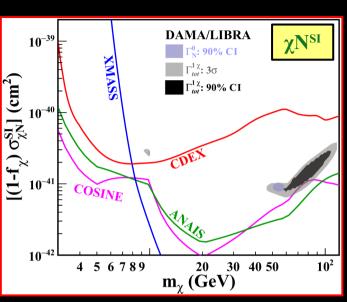
 With χe^{LR} & χe^{SR} Added in Γ^{1χ}_{tot}::

 Only high m_χ Region [I-recoils]
 Allowed
 @ Same Significance

 Best-fit Solution of m_χ::

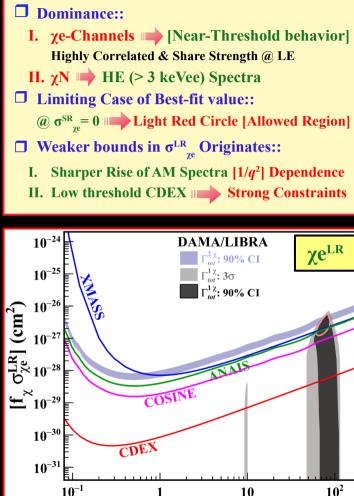
 Shifted from 54 GeV in Γ⁰_N to 83 GeV in Γ^{1χ}_{tot}

Results & Interpretations:: Case of $\Gamma^{1\chi}_{tot}$



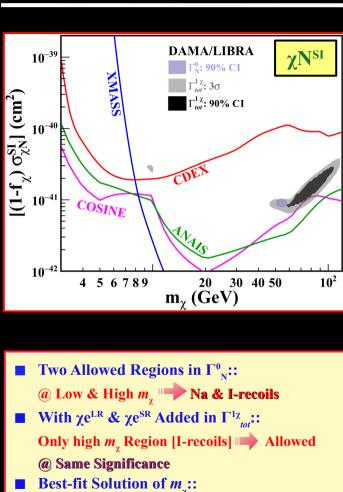
- Two Allowed Regions in Γ^0_N ::
 - @ Low & High m, Na & I-recoils
- With χe^{LR} & χe^{SR} Added in $\Gamma^{1\chi}_{tot}$::
- Only high m_{χ} Region [I-recoils] \Longrightarrow Allowed
 - @ Same Significance
- Best-fit Solution of m_{χ} ::

Shifted from 54 GeV in $\Gamma^0_{\ N}$ to 83 GeV in $\Gamma^{1\chi}_{\ tot}$

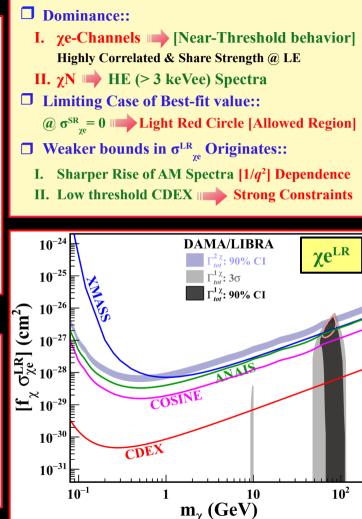


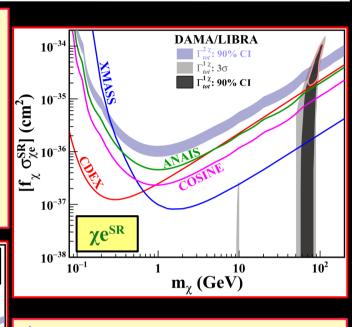
 m_{γ} (GeV)

Results & Interpretations:: Case of $\Gamma^{1\chi}_{to}$



Shifted from 54 GeV in Γ^{0}_{N} to 83 GeV in $\Gamma^{1\chi}_{tot}$



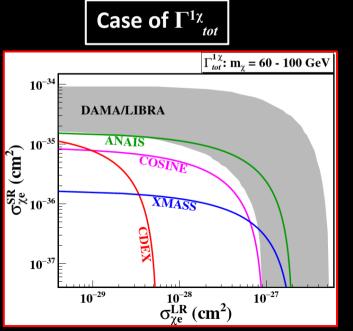


- Limiting Case of Best-fit value::

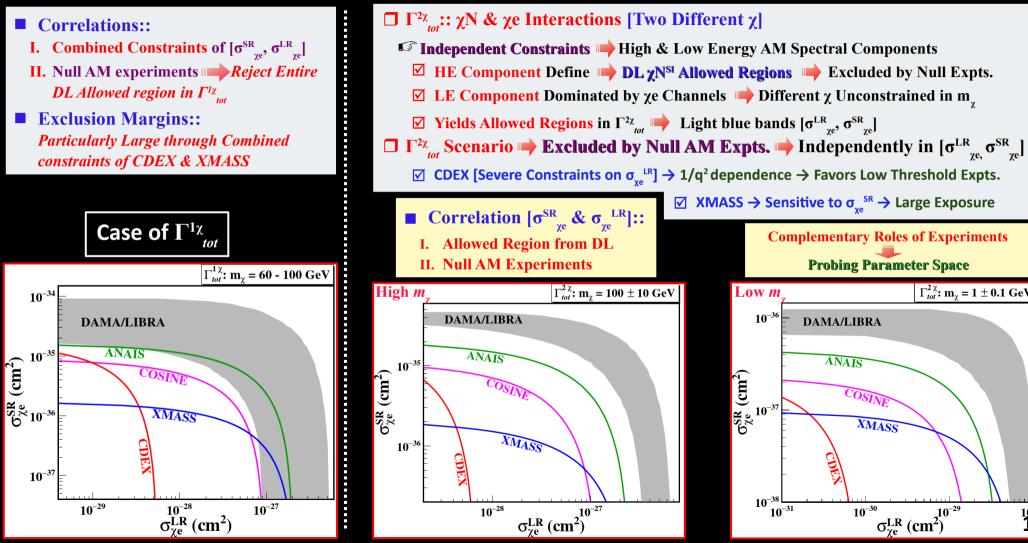
 (a) σ^{LR}_{χe} = 0 || Light Red Circle [Allowed Region]
- ▲ Interdependent Allowed Regions::
- I. Low Cross Section Portion of $[\sigma^{LR}_{\chi e}, \sigma^{SR}_{\chi e}]$ Allowed Regions remains Unprobed
- II. Correlation Allowed Space::

Correlated with High Cross Section Region of its Counterparts Rejected by Null AM Experiments

Results & Interpretations:: Case of $\Gamma^{1\chi}_{tot}$



Results & Interpretations:: Case of $\Gamma^{2\chi}$



 $\Gamma_{tot}^{2\chi}$: $m_{\gamma} = 1 \pm 0.1 \text{ GeV}$

Summary and Conclusions

"First analysis":: Explores How χe AM can Play a Role in DL AM

Present Study:: Expand Investigations
 Addition of χe^{LR} and χe^{SR} Interactions to χN
 Using Frozen Core Approximation
 Considered Two scenarios
 χN & χe Processes:: Single χ (Γ¹χ tot) or Two Different χ (Γ²χ tot)
 Combined fits [χN & χe] Provides
 Stronger significance to DL AM data

☑ Compatible with Presence of Additional Physical Effects

All DL AM Allowed Parameter Spaces::

ΣχΝ & χε Channels under both Γ¹χ tot

Ruled out @ 90% CL by Combined null AM results

☑ Beyond xN Alone



Acknowledgments

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All Members of the TEXONO & TDMC Collaborations



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Dr. Mukesh Kumar Pandey



Dr. Shuvadeep Karmakar



Prof. Henry Tsz-King Wong

Thank You! 谢谢