



BULLKID-DM

searching for light WIMP with monolithic arrays of detectors

Matteo Folcarelli on behalf of the collaboration Università La Sapienza di Roma, INFN - Roma 1





BULLKID-DM: A WIMP-like Dark Matter experiment



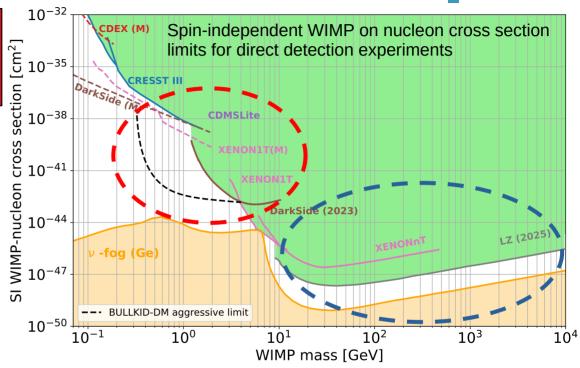












Kg-scale solid state phonon detectors

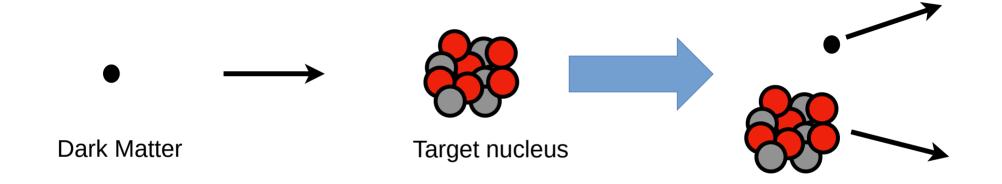
Multi-ton liquid scintillators

Detection principle: nuclear recoils

Target: Silicon (Si) + ongoing effort to port technology on Germanium

Si: A. Cruciani et al, Appl. Phys. Lett. 121, 213504 (2022)

Ge: D. Delicato et al, Appl. Phys. Lett. 126, 153502 (2025)

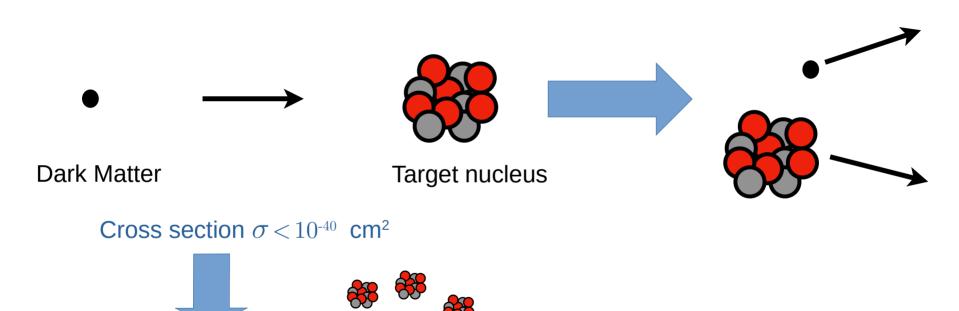


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Large number of targets O(kg)

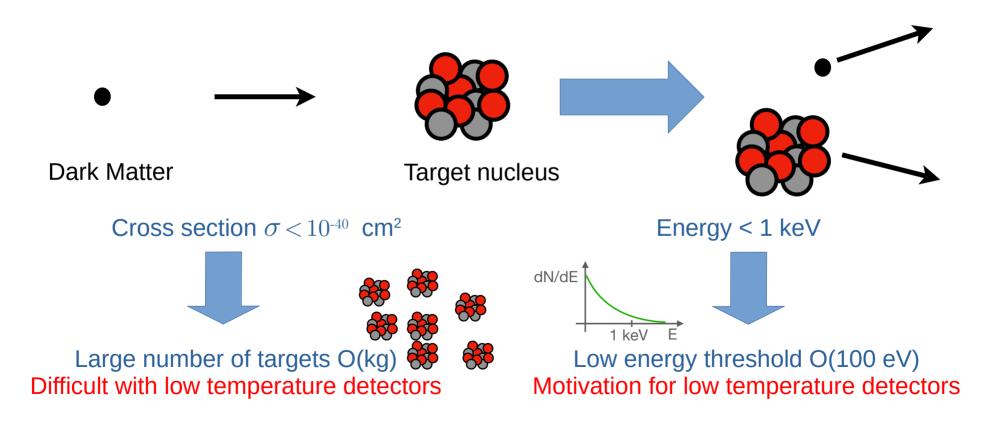
Difficult with low temperature detectors

Detection principle: nuclear recoils

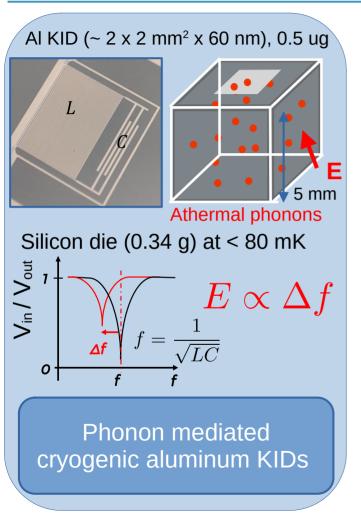
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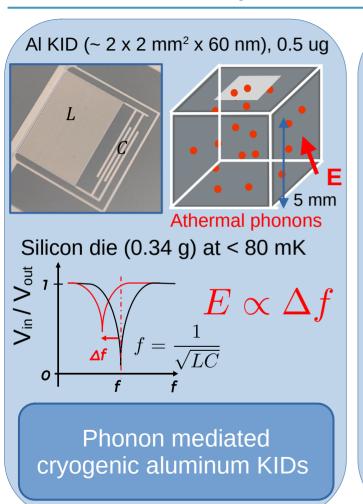
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Detector: Fully active structure of cryogenic calorimeters

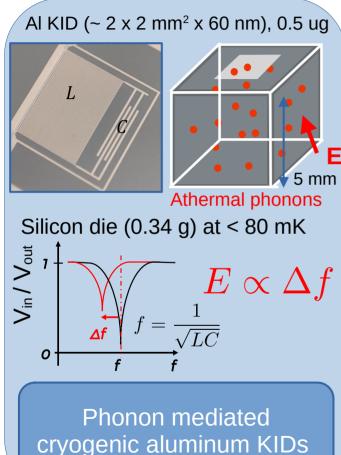


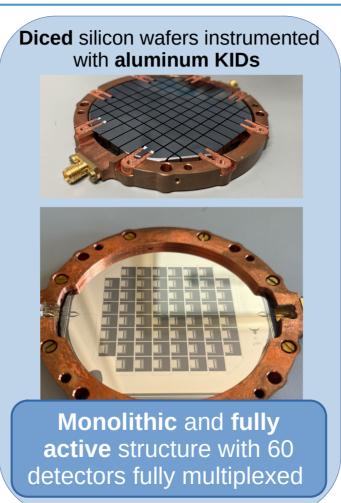
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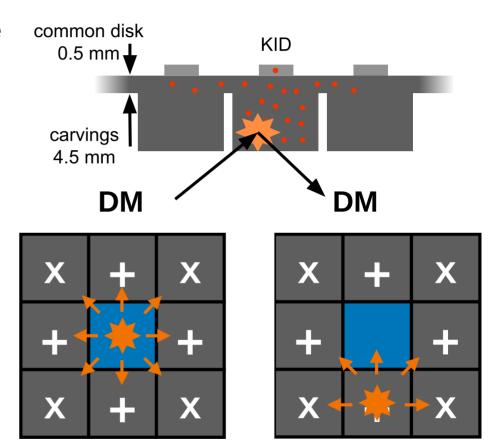




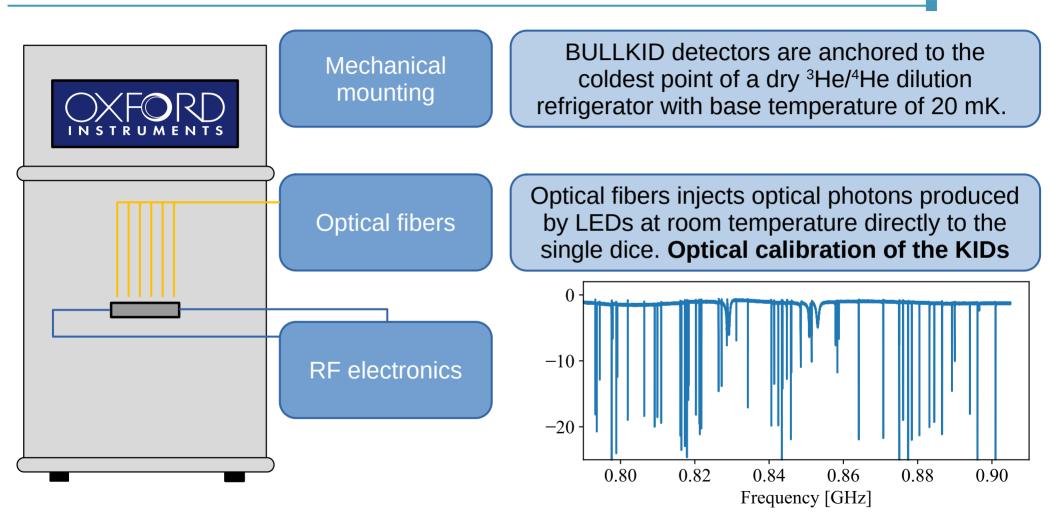
Phonon leakage and mapping

- Only part of the phonons are absorbed by the KID
- The rest leaks in nearby dice:
 - ~20% in each "+" die
 - ~10% in each "x" die
 - the rest in outer dice

This effect reduces the phonon focusing on the KID but is exploited to identify the interaction voxel

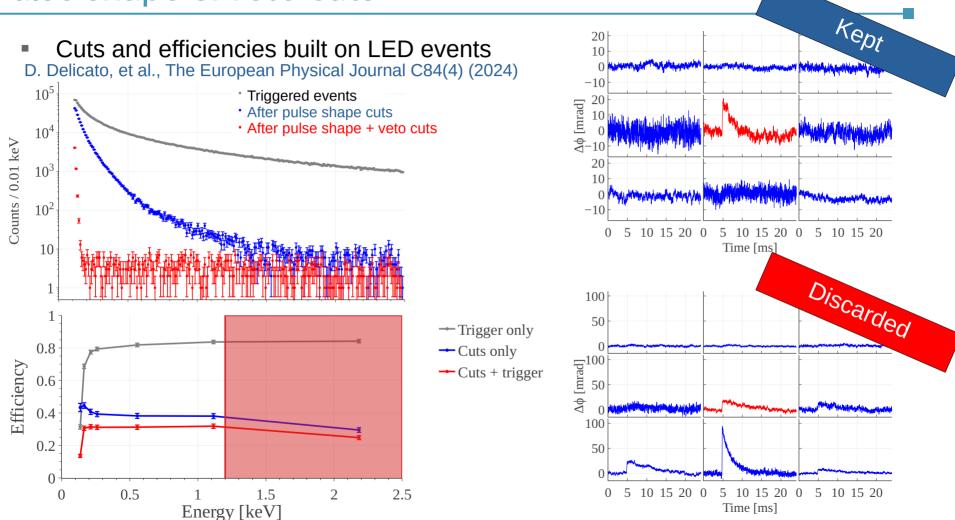


Experimental setup of BULLKID detectors



Matteo Folcarelli

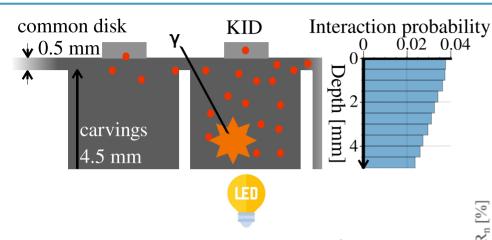
Pulse shape & veto cuts



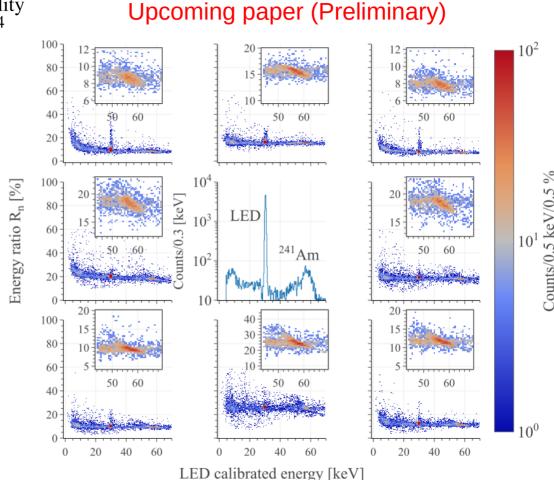
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Energy calibration of bulk events (241Am)

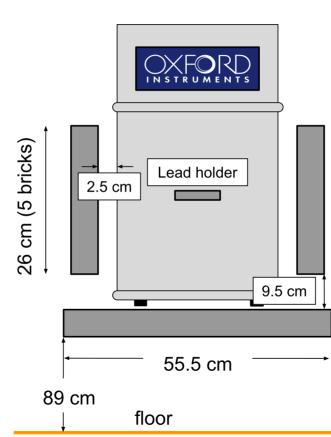


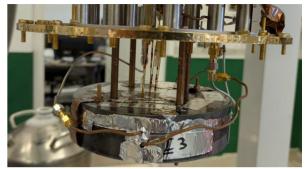
- Dedicated cooldown with ²⁴¹Am source (59.5 keV γ ray)
- Less than 10% deficit between LED (surface) and Americium (bulk) calibration
- Same phonon leakage between bulk and LEDs



Background measurement with mild shield

- Location: Surface laboratory (Sapienza, 1st floor)
- Midly shielded environment with lead (internal and external)
- Optical fibers for detector calibration and monitoring
- Run time:
 19th Apr 22:00 → 5th May 08:00
 (290 live hours)
- Acquisition of 15 KIDs: 2 triggering + 13 as veto





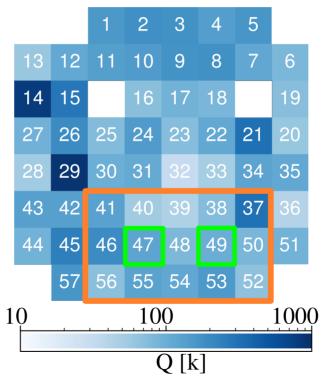


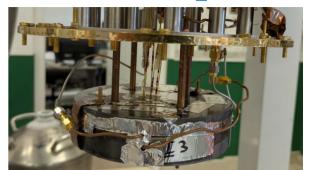
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Triggering KIDs KIDs acquired in

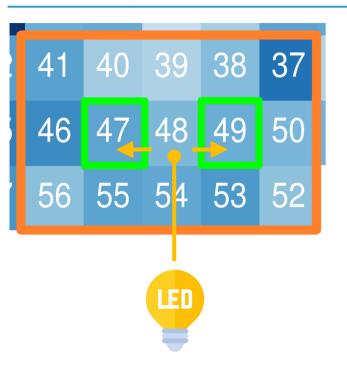




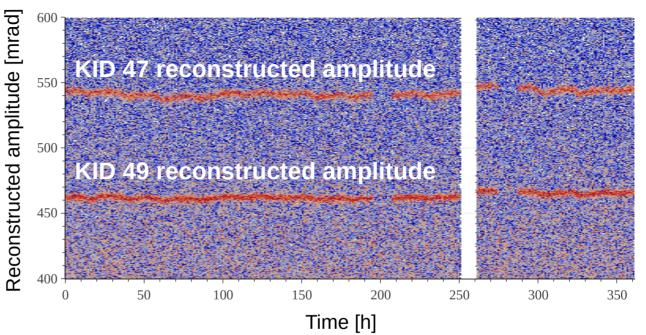




Stability evaluation with controlled LED pulses



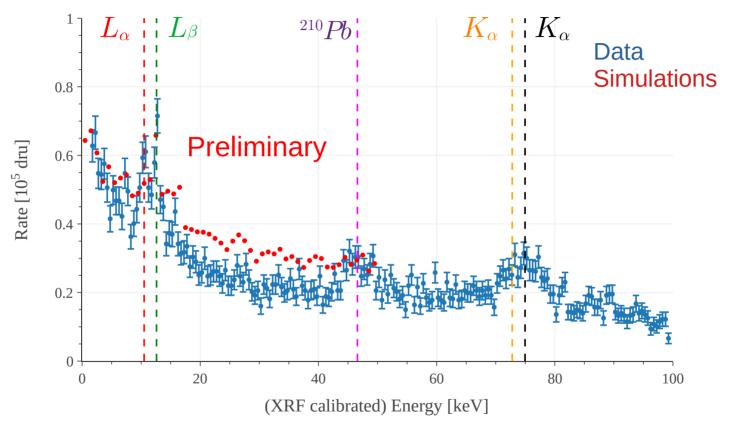
KID 47: Noise RMS 35 eV KID 49: Noise RMS 30 eV



- Stability of the triggering KIDs evaluated with leakage signal from KID 48 via controlled LED pulses
- Detector stable over the entire data acquisition

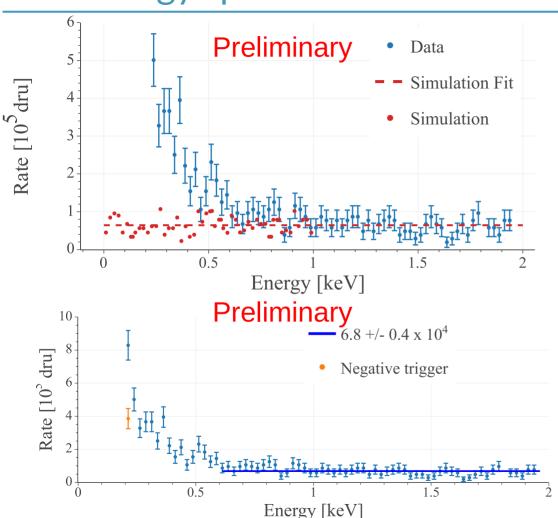
Counts/1 h/0.5 mrad

High energy spectrum and simulation comparison



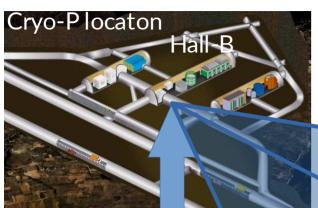
- Visible Pb X-rays well reconstructed (under 10% discrepancy) with LED calibration
- Overall fairly well matching with simulations but Pb rays not well reproduced

Low energy spectrum



- Flat background: $(68\pm4)\cdot10^3$ d.r.u (simulated $(61\pm3)\cdot10^3$ d.r.u)
- In < 600 eV measured a rise up to a factor ~ 5
- Spectra compatible between the two triggering KIDs
- Rise not compatible with negative triggers (noise false positives)

Background measurement underground (LNGS)



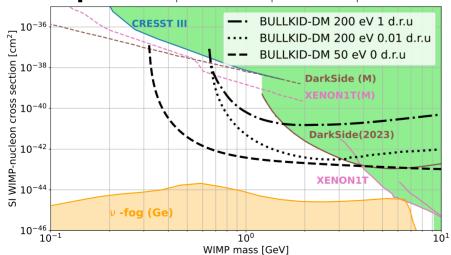
The cryo-platform is a cryogenic underground hub for mK applications:

- Oxford Proteox Cryostat successfully commissioned in July (base temperature < 10 mK)
- Overall facility ready in next few months



BULLKID-DM roadmap

	BULLKID prototype	BULLKID-DM Demonstrator		BULLKID-DM
mass	20 g	60 g		800 g
# of sensors	60	180		2300
Threshold	160 eV	200 eV		< 200 eV
Bkg (c / keV Kg d)	2 x 10 ⁶	< 10 ⁵		1 - 0.01
laboratory	Sapienza U.	Sapienza U.	LNGS	LNGS
installation	2023	2025	2026	2027



We are now in the **Demonstrator phase!**

Status of the demonstrator:

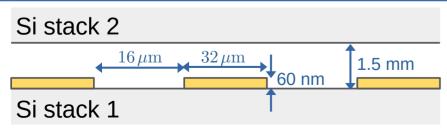
- stack of 3 wafers
- RF electronics

Working in view of the experiment:

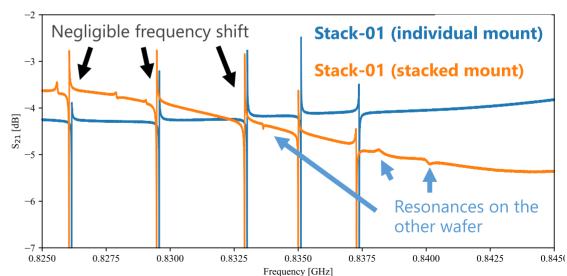
- 100 mm arrays: tests, assembly and serial production
- Shielding and veto

Status of the 3-wafer stack





- Holding structure: thermalization and mounting
- Reproducibility of electrical coupling
- Reproduction of the results of the unstacked wafers



RF electronics

Current electronics:

Ettus Research USRP X310:30 KIDs / board

New electronics:

ZCU215 Evaluation Board with 16 lines):

Goal >= 150 KIDs / line 2300 KIDs / board

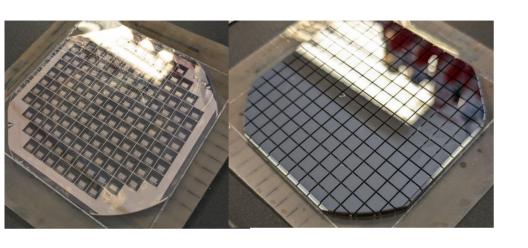
- Custom analog Front-End
- Control firmware by KIT group
- Status: First tests on BULLKID prototype ongoing

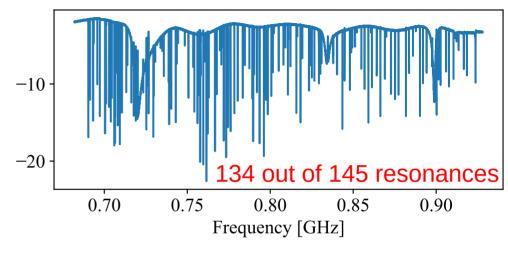
Upcoming paper

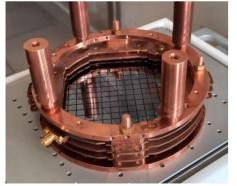




Scalability for the 100 mm mask







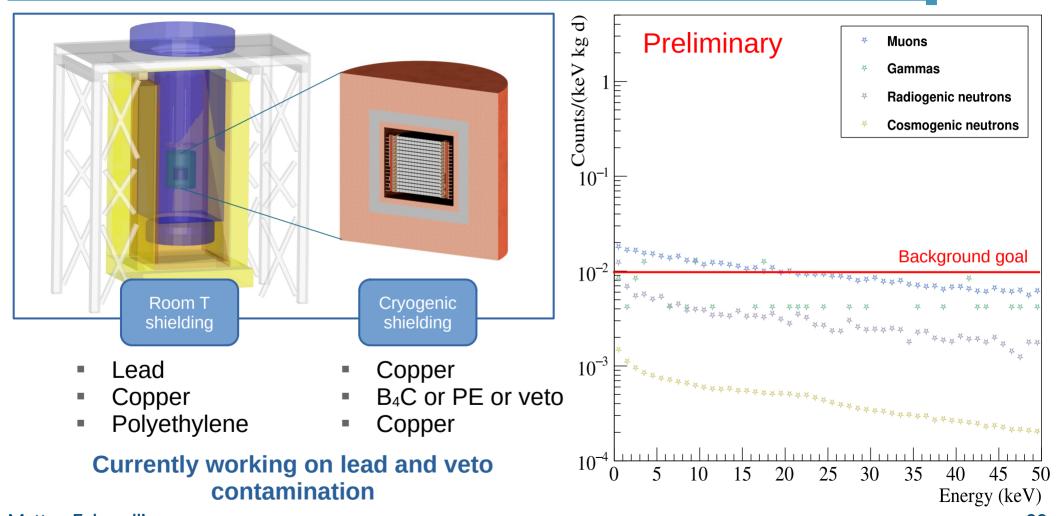


Next steps:

- Series wafer production
- Improvement of the lithography

22

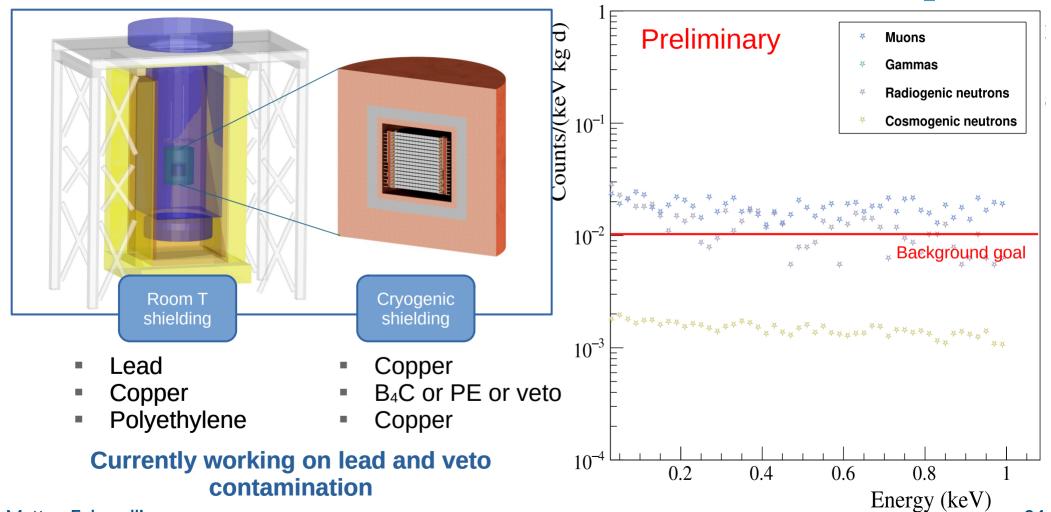
Simulations: shields and veto



Matteo Folcarelli

BULLKID-DM: searching for light WIMP with monolithic arrays of detectors

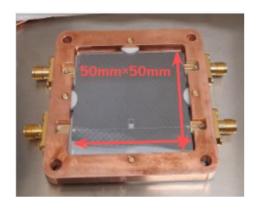
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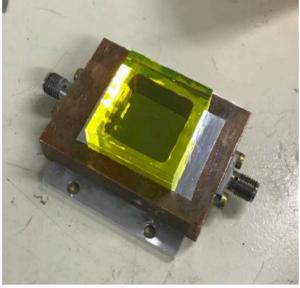
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Cryogenic veto







Idea: Use scintillating crystals (BGO, GSO or GAGG) read by KID light detectors as cryogenic veto

Pros:

- Dense materials
- Same readout of the central experiment
- Mass produced in industry

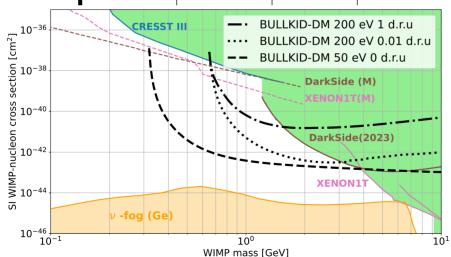
Cons:

- Low radio-purity
- Materials not well tested at cryogenic temperatures
- KID readout to be optimized

Goal: energy threshold < 50 keV

BULLKID-DM roadmap

	BULLKID prototype	BULLK Demons		BULLKID-DM
mass	20 g	60 g		800 g
# of sensors	60	180		2300
Threshold	160 eV	200 eV		< 200 eV
Bkg (c / keV Kg d)	2 x 10 ⁶	< 10 ⁵		1 - 0.01
laboratory	Sapienza U.	Sapienza U.	LNGS	LNGS
installation	2023	2025	2026	2027



- 800 g of silicon target
- 2300 detector units (dice)
- Unique features for background suppression:
 - No inert material in the detector volume
 - Fully active
 - Fiducialization (600 g)
- Will it help with the unknown backgrounds?



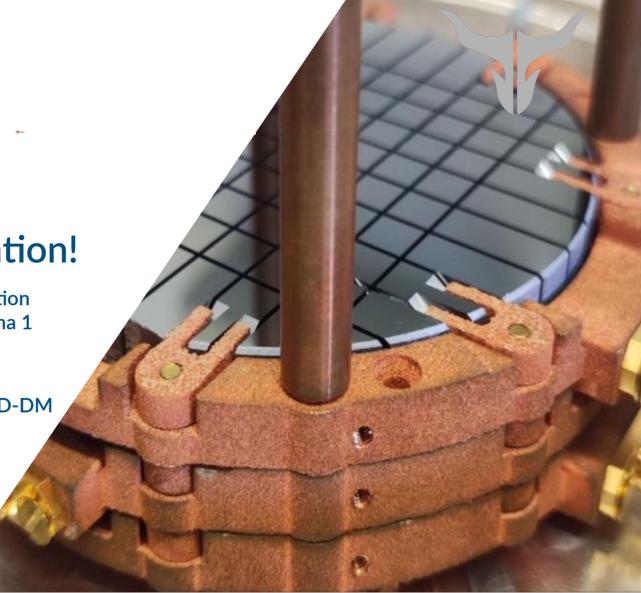
Thank you for the attention!

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Other contribution to TAUP2025:

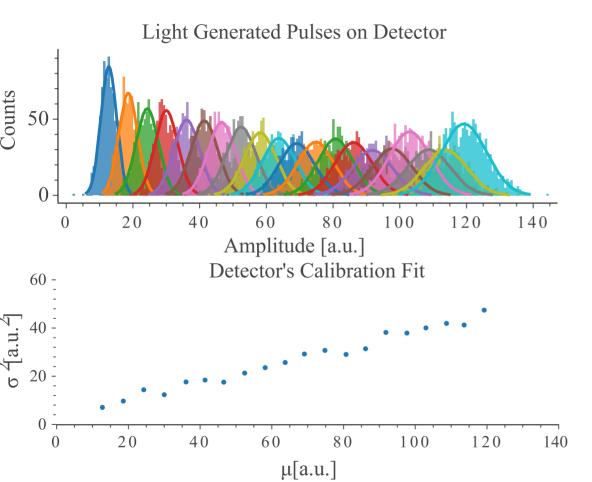
POSTER: Energy calibration of the BULLIKID-DM experiment (speaker Matteo Folcarelli)

This work was partially supported through the European Research Council through the Consolidator Grant DANAE number 101087663



Backup Slides

Optical calibration



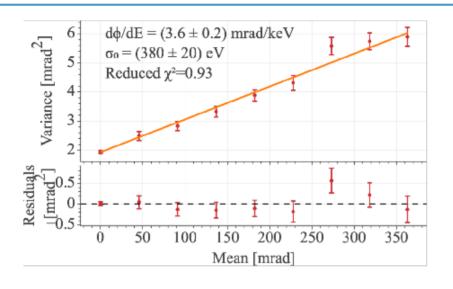
The amplitude of the LED induced pulse is a Poisson distributed variable.

Then the mean and the variance of the distributions can be expressed as follows:

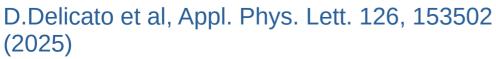
$$\sigma^{2}_{LED} = \sigma^{2}_{0} + r \cdot \epsilon \cdot \mu$$

Where ϵ is the photon energy, r is the responsivity of the KID

KIDs on Germanium fro neutrino detection



	$d\phi/dE \ [{ m mrad/keV}]$	σ_0 [eV]	η [%]	$rac{A_{ m KID}}{A_{ m TOT}}$
KID-1	3.6 ± 0.2	380 ± 20	2.0 ± 0.2	0.12
KID-2	2.5 ± 0.1	450 ± 22	2.0 ± 0.2	11
KID-3	2.2 ± 0.2	540 ± 31	1.6 ± 0.2	11
CALDER-17	5.8	115 ± 6	7.4 - 9.4	0.42

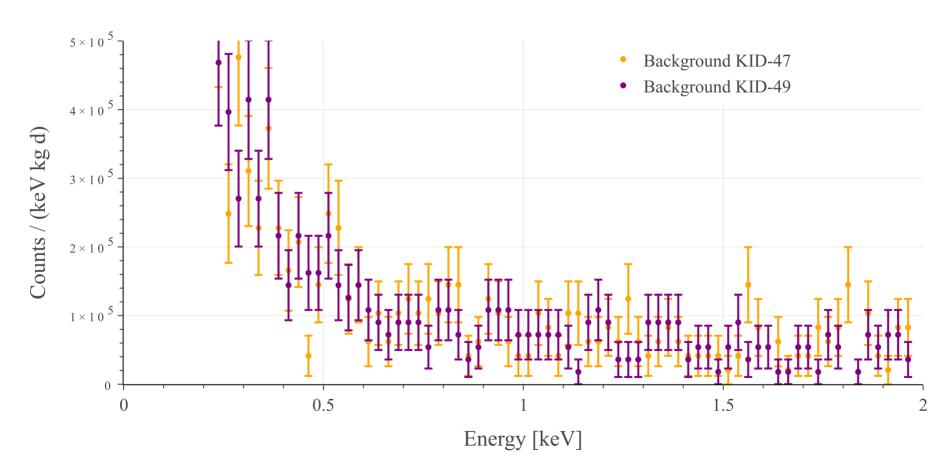




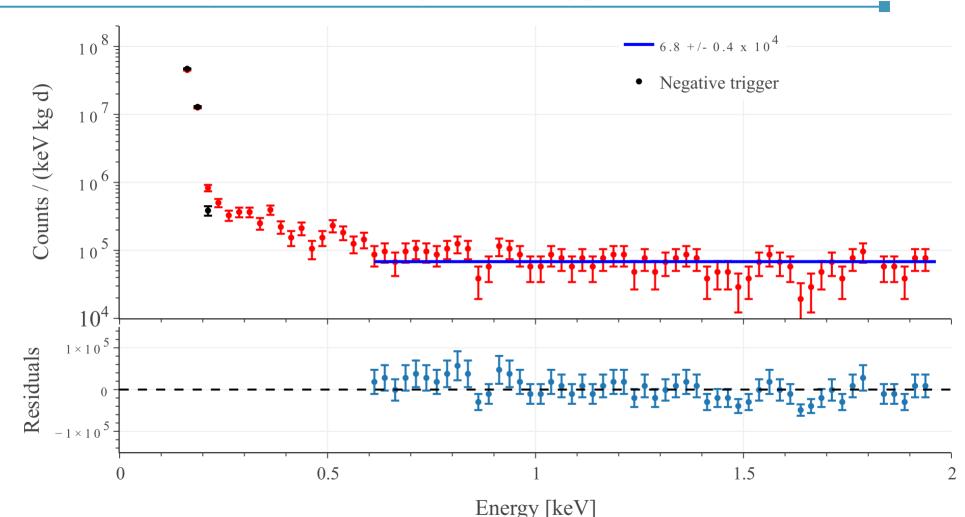


Comparison between px 47 and 49

Energy after cluster cuts



Spectrum in logscale for Negative Triggers

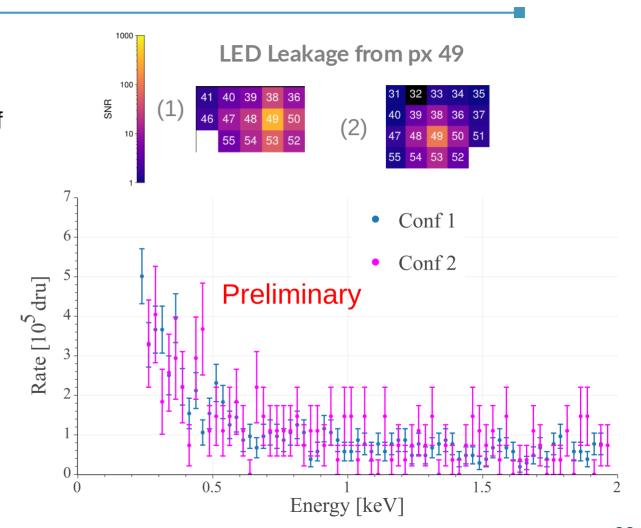


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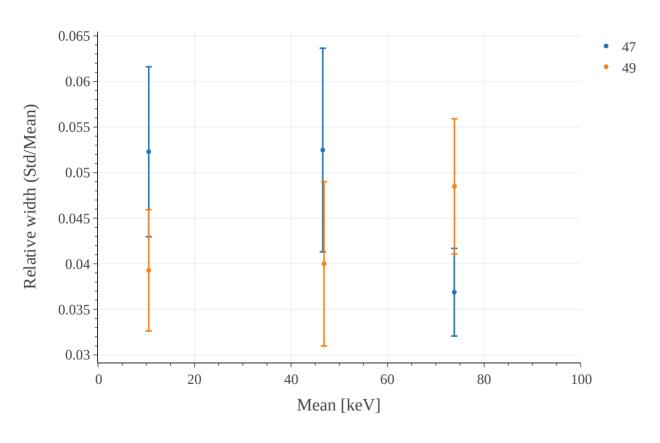
Repeated data taking

- Different working conditions tried:
 - 1) Optimizing the resolution of the triggering pixels
 - Uniform leakage SNR in between pixels
- Data taking period in the same cooldown:
 - 1) $19/04 \rightarrow 05/05 \ (\sim 15 \ days)$
 - 2) $19/05 \rightarrow 26/05 \ (\sim 7 \text{ days})$

Observed no difference in between the energy spectra in the two configurations



Resolution as function of the energy

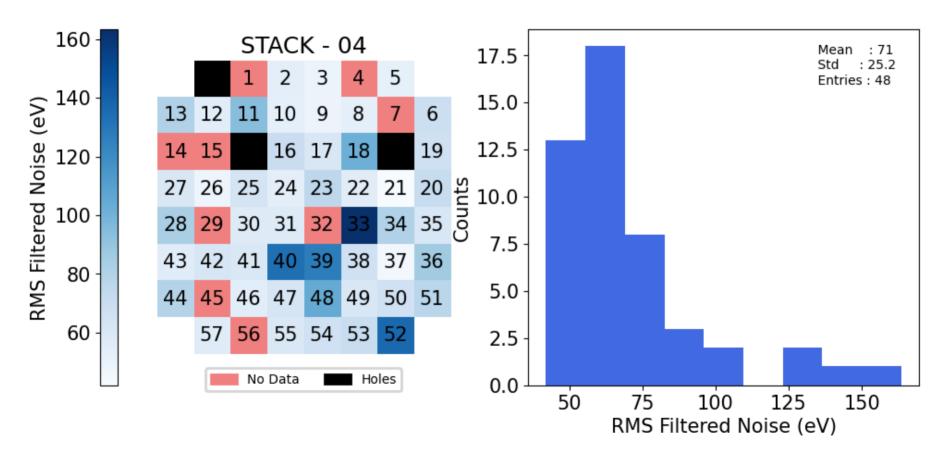


- Pb peaks allows the evaluation of the energy resolution of a BULLKID detector
- Each peak is fitted with a Gaussian over a linearly decreasing background

A resolution around 5% is found for energies above 10 keV

Uniformity of the wafer response

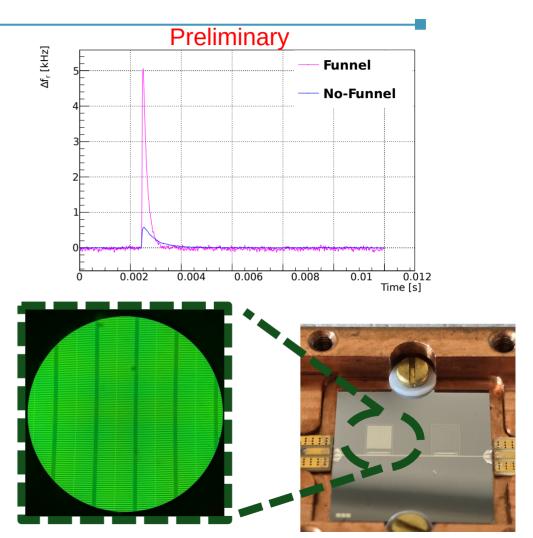
Noise after calibration (FS) @ low bias power



R&D efforts: phonon funnels

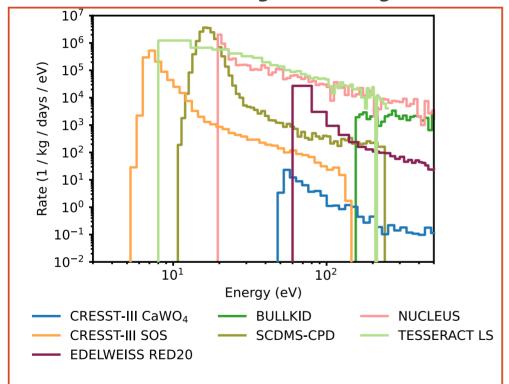
Upcoming paper

- Development of KIDs with dedicated phonon collection structures (funnels)
- Improvement in collection efficiency and resolution
- Trilayer KID (77 nm Al/Ti/Al) + funnel volume at higher gap (Al) for phonon funneling effect



Background issue in low-T experiments

Not understood excess background rising at low energies



- Phonon bursts (crystal-support friction) ?
- Lattice relaxations after cool down?
- Phonon leakage from interactions in the supports?



Excess workshop 2025
Santa Fe, 31 May
https://indico.cern.ch/event/1502420/

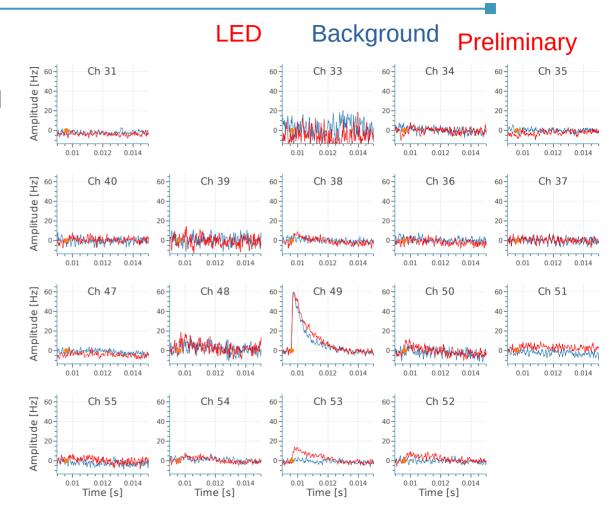
This background limits the sensitivity of present experiments

Singles observed in BULLKID

BULLKID architecture cannot distinguish Shared/Singles at threshold (small phonon leakage)

At higher energies:

- Singles visible at ~2 keV and above
- Could be direct KID hits due to high metalized area (Al absorbs well 100 eV photons)
- Could it be that the rise is due to Singles?



Electronic x-talk

