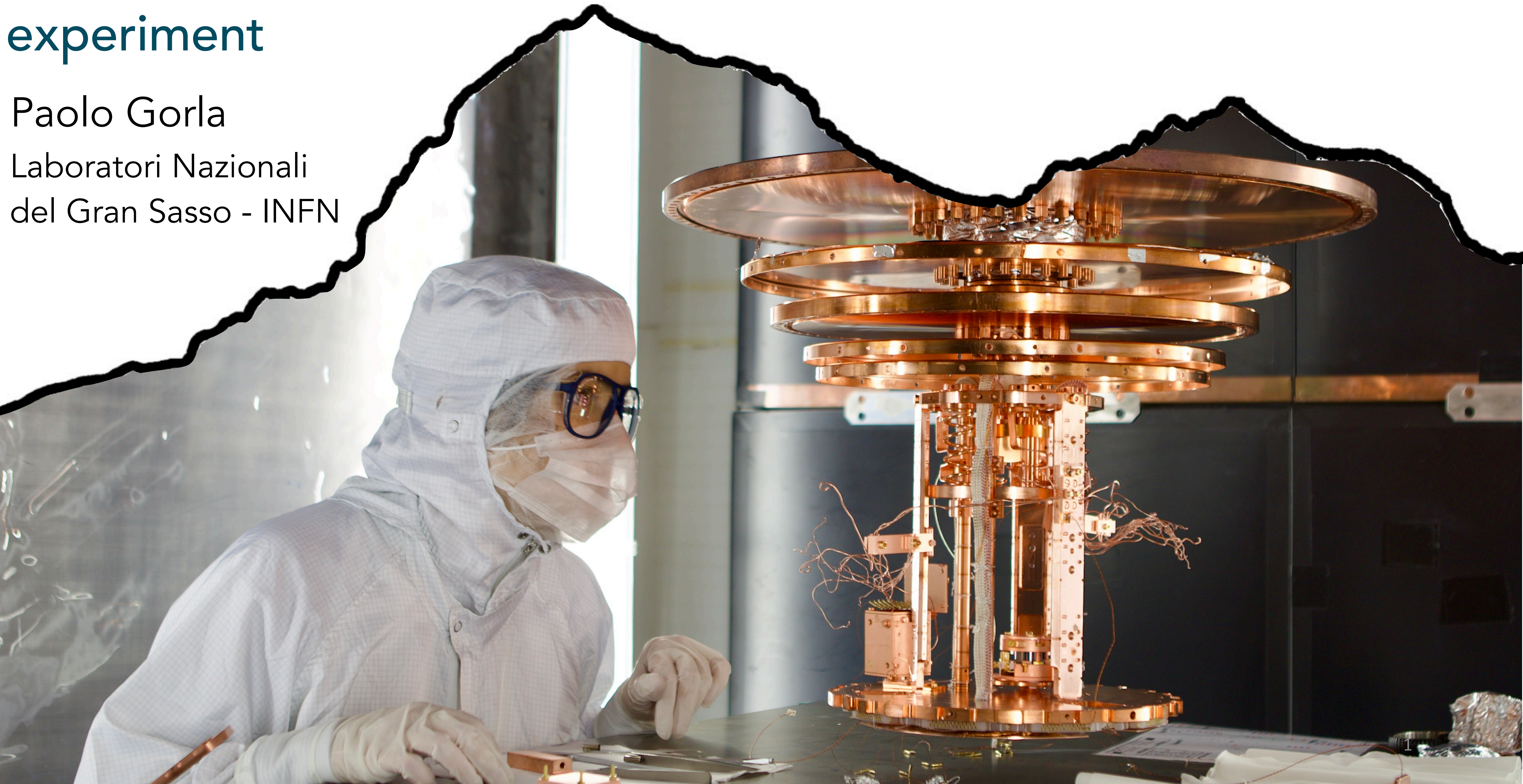


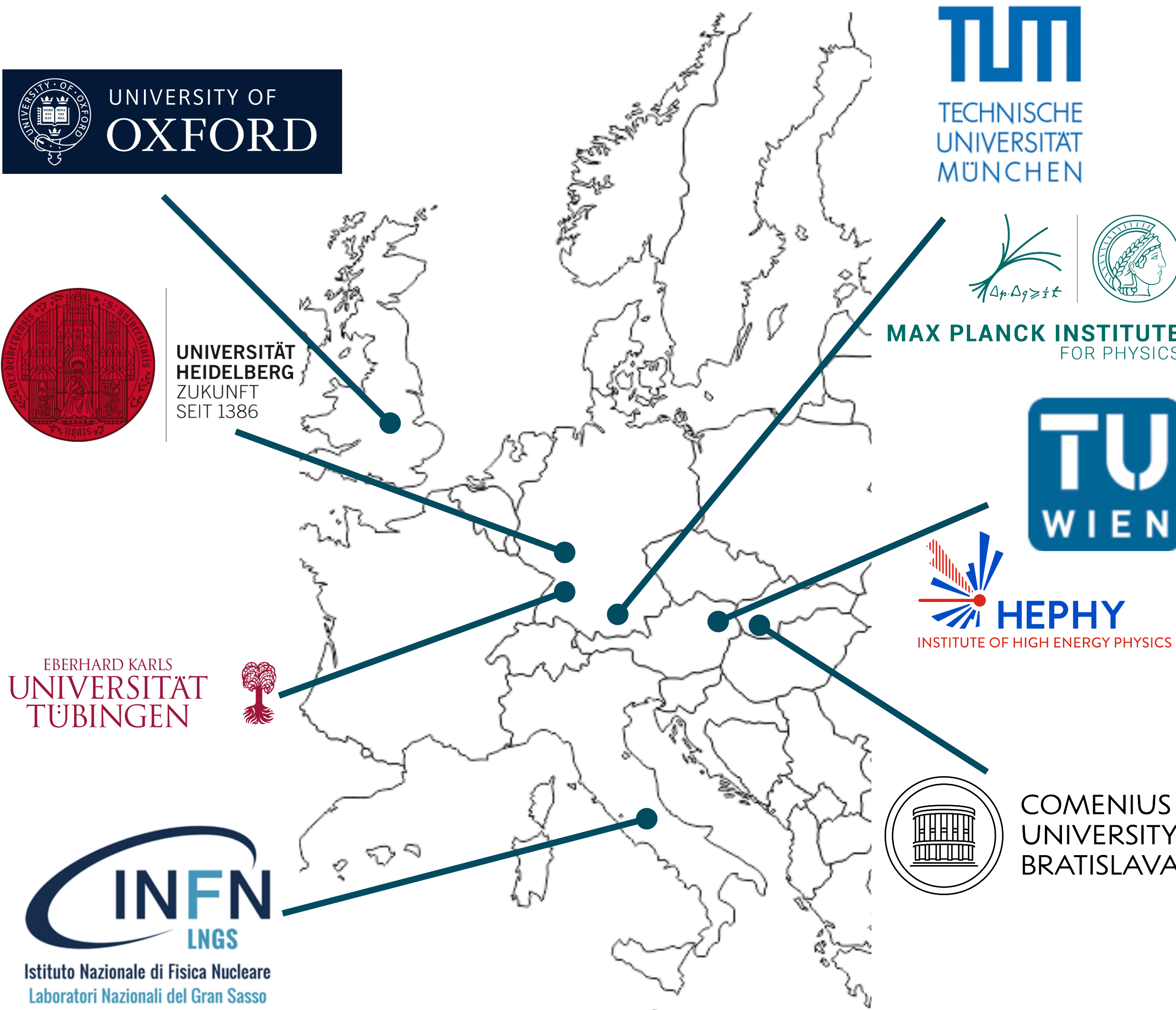
Direct Dark Matter search with the CRESST-III experiment

Paolo Gorla

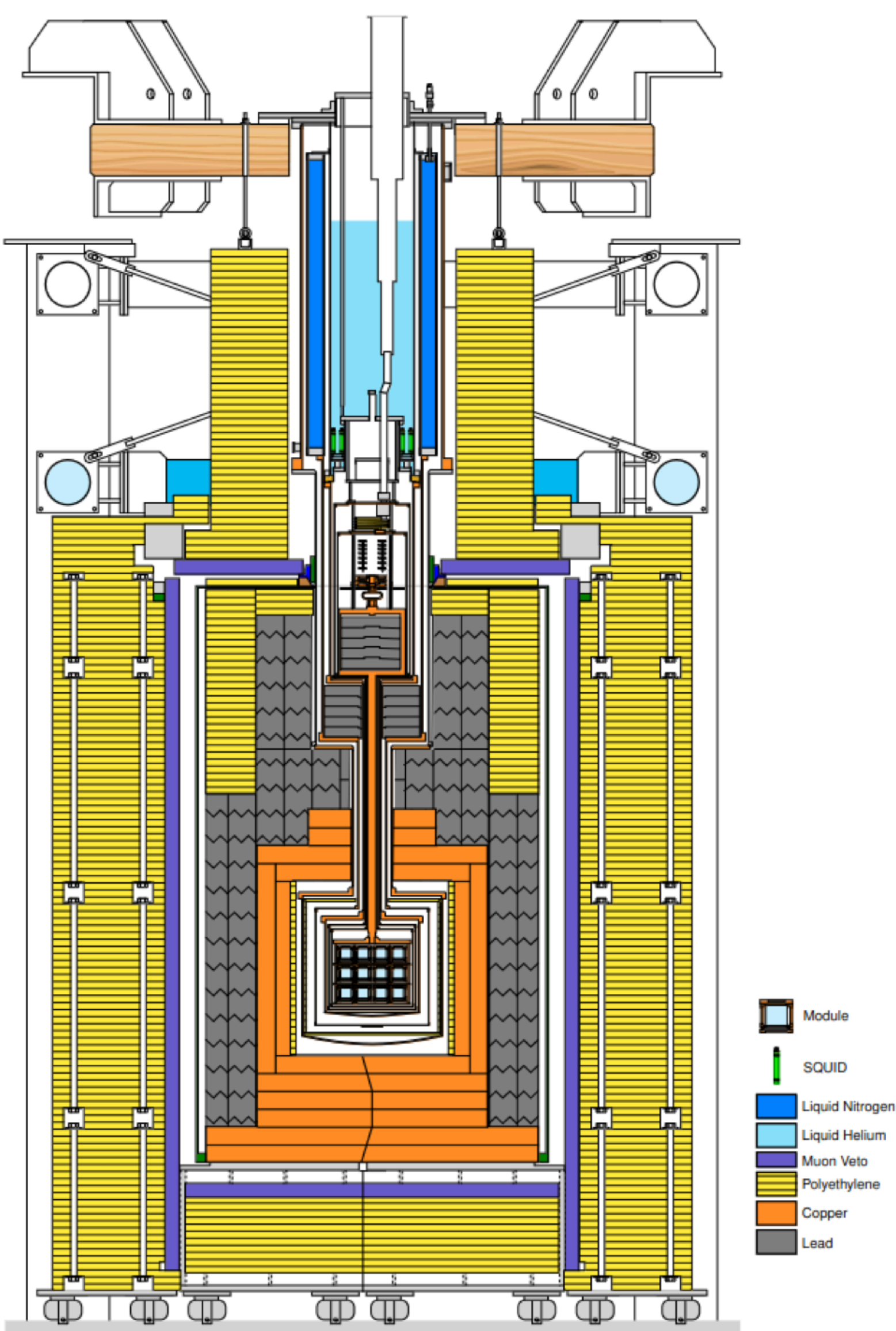
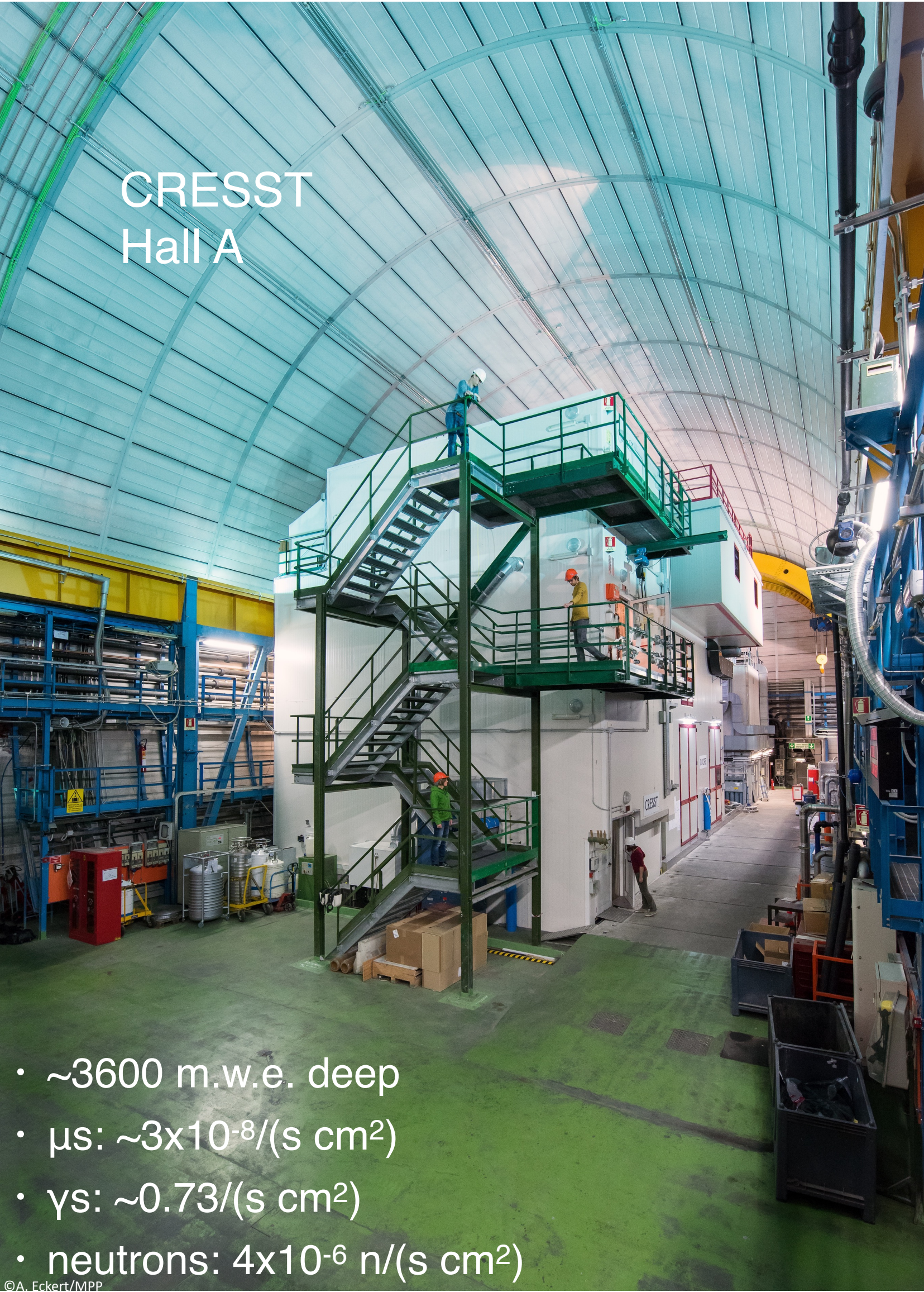
Laboratori Nazionali
del Gran Sasso - INFN



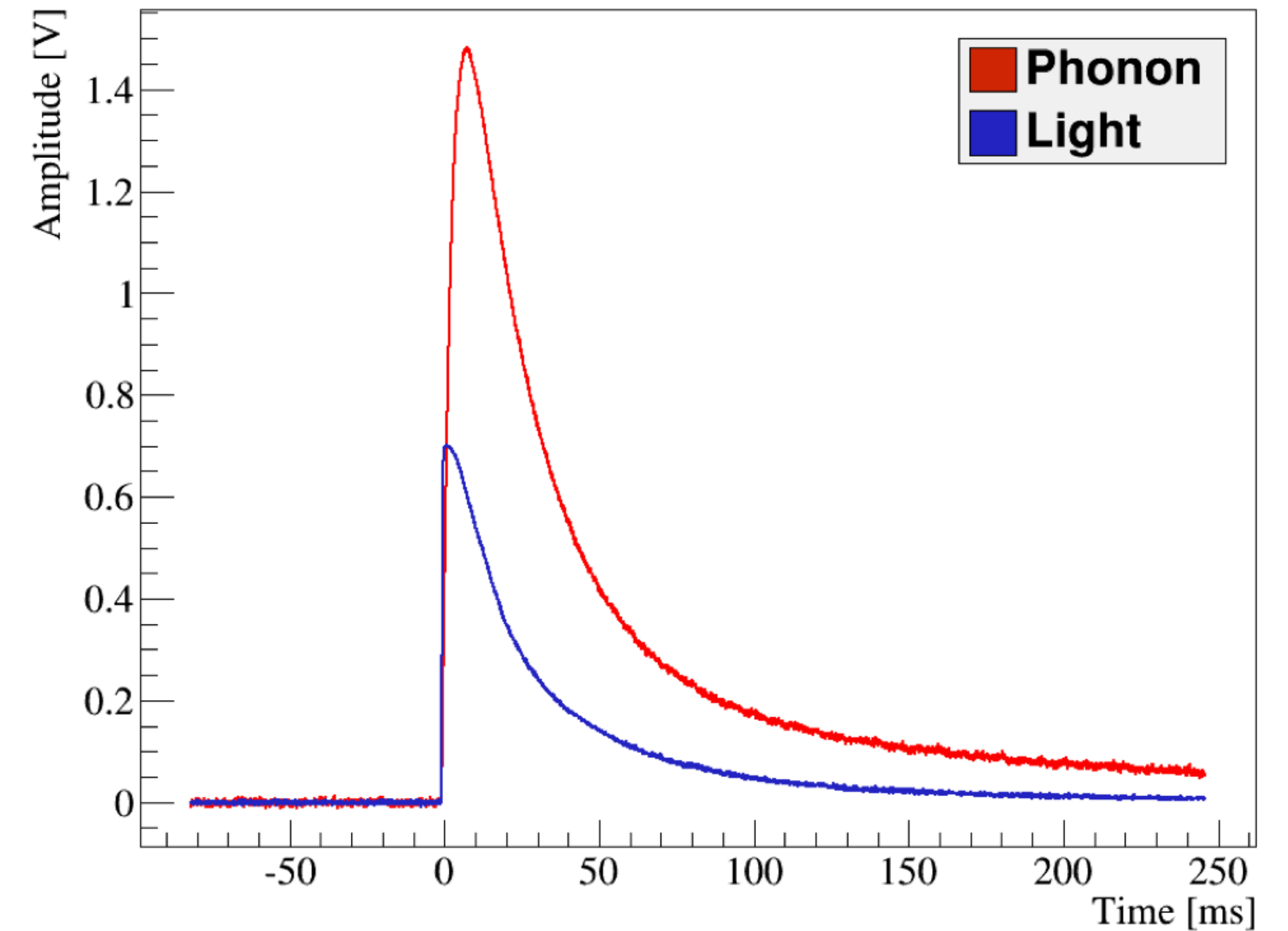
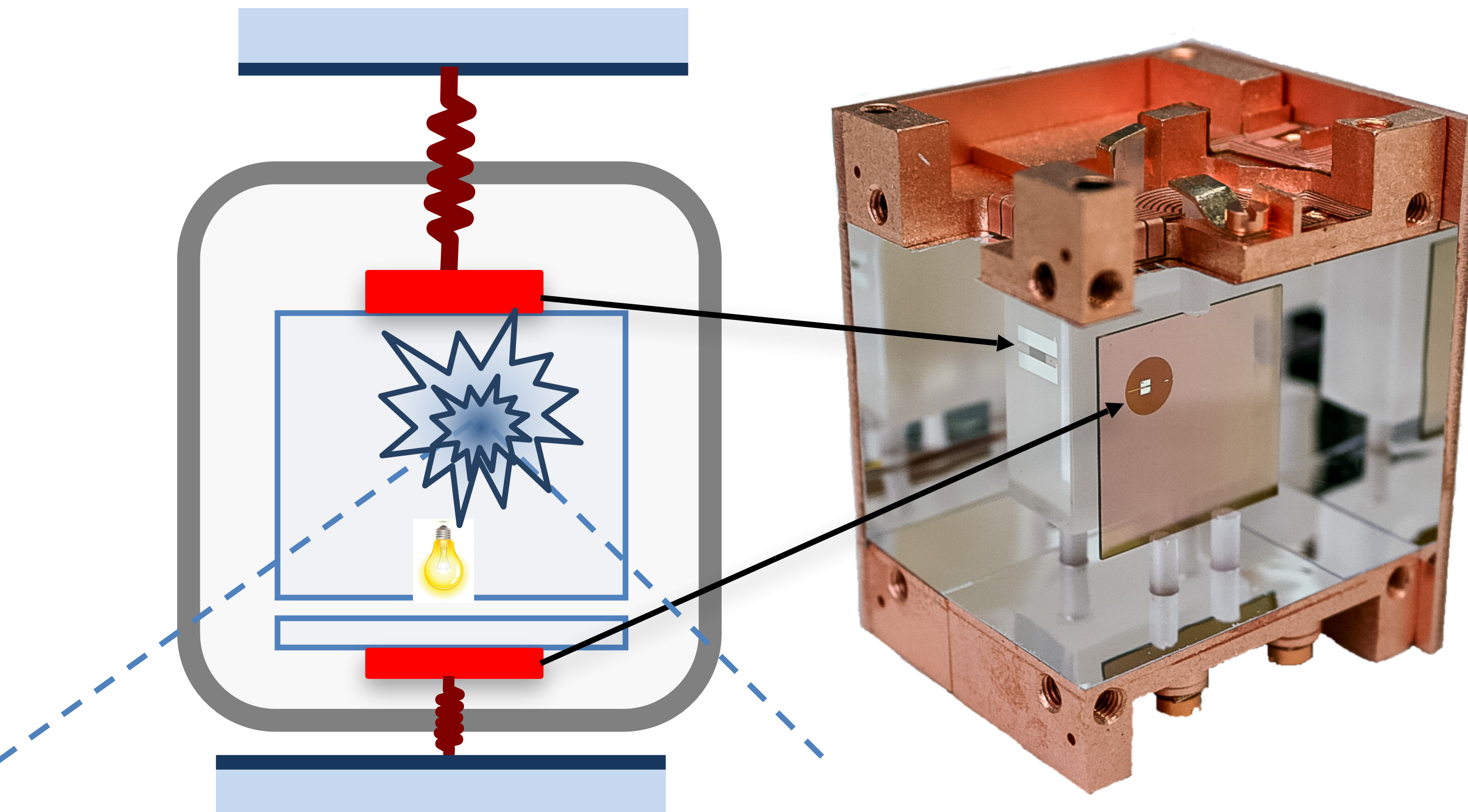
The CRESST collaboration



The CRESST experiment



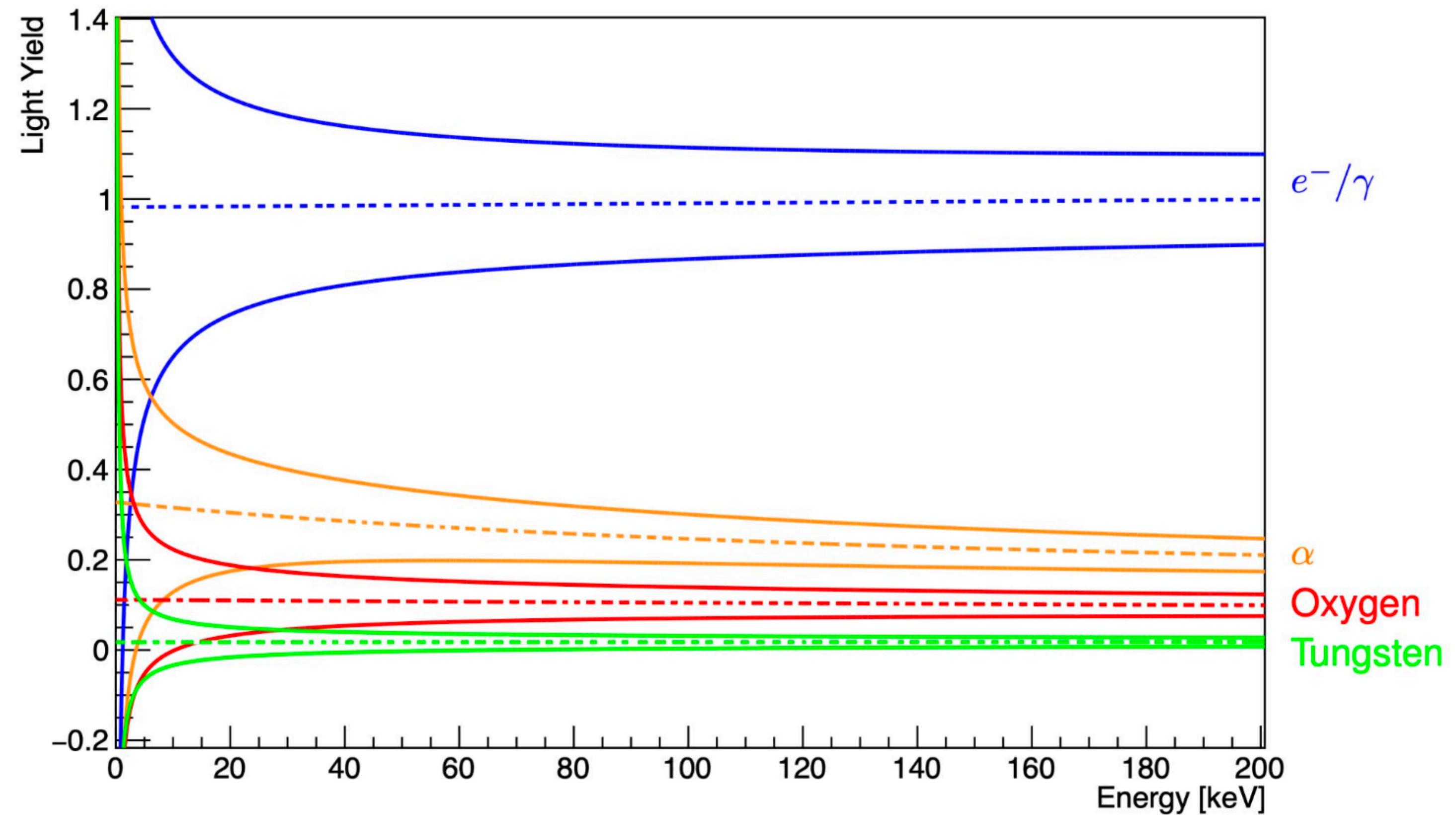
CRESST goal: direct detection of dark matter particles via their scattering off target nuclei in cryogenic detectors, operated at ~ 15 mK

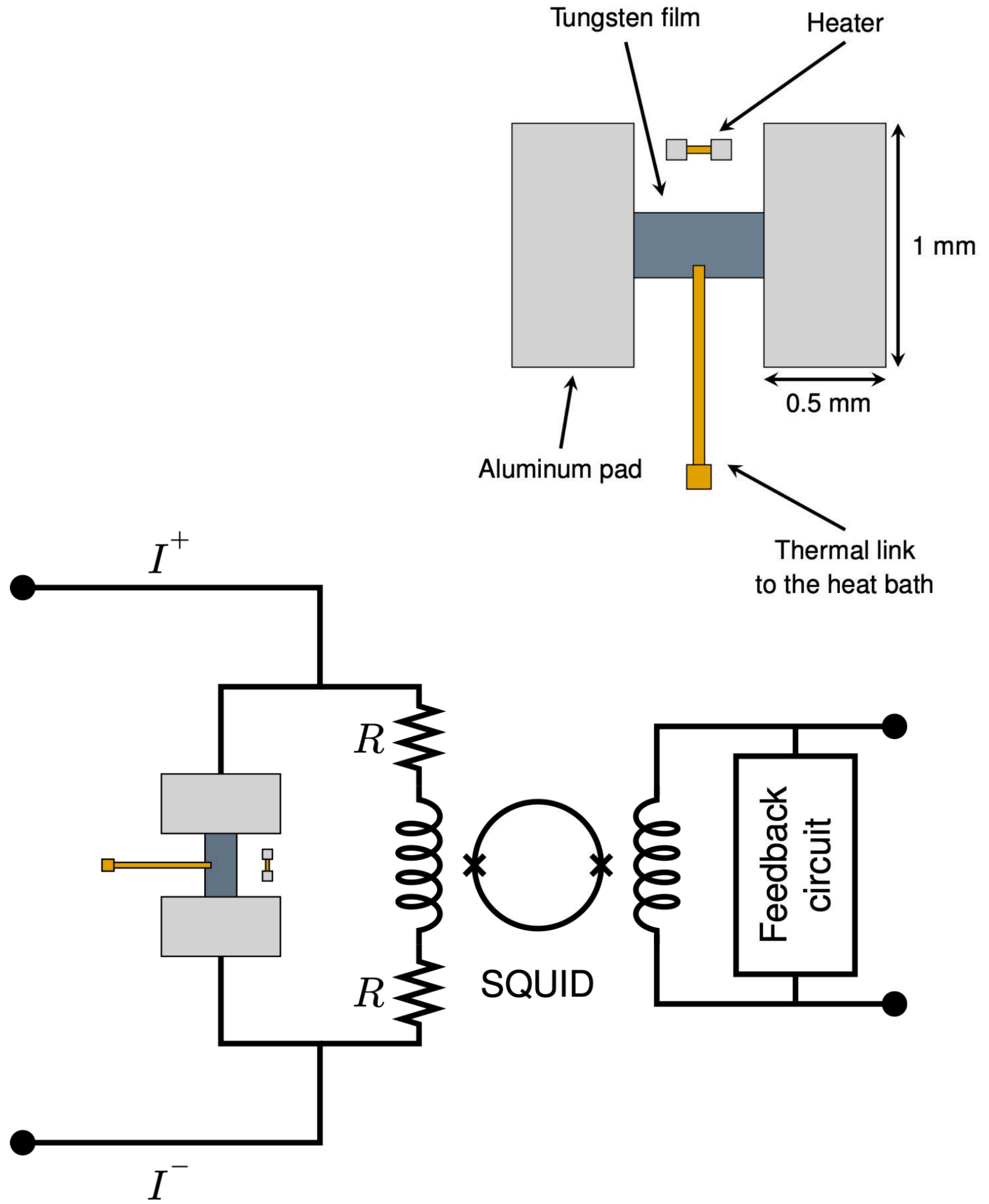
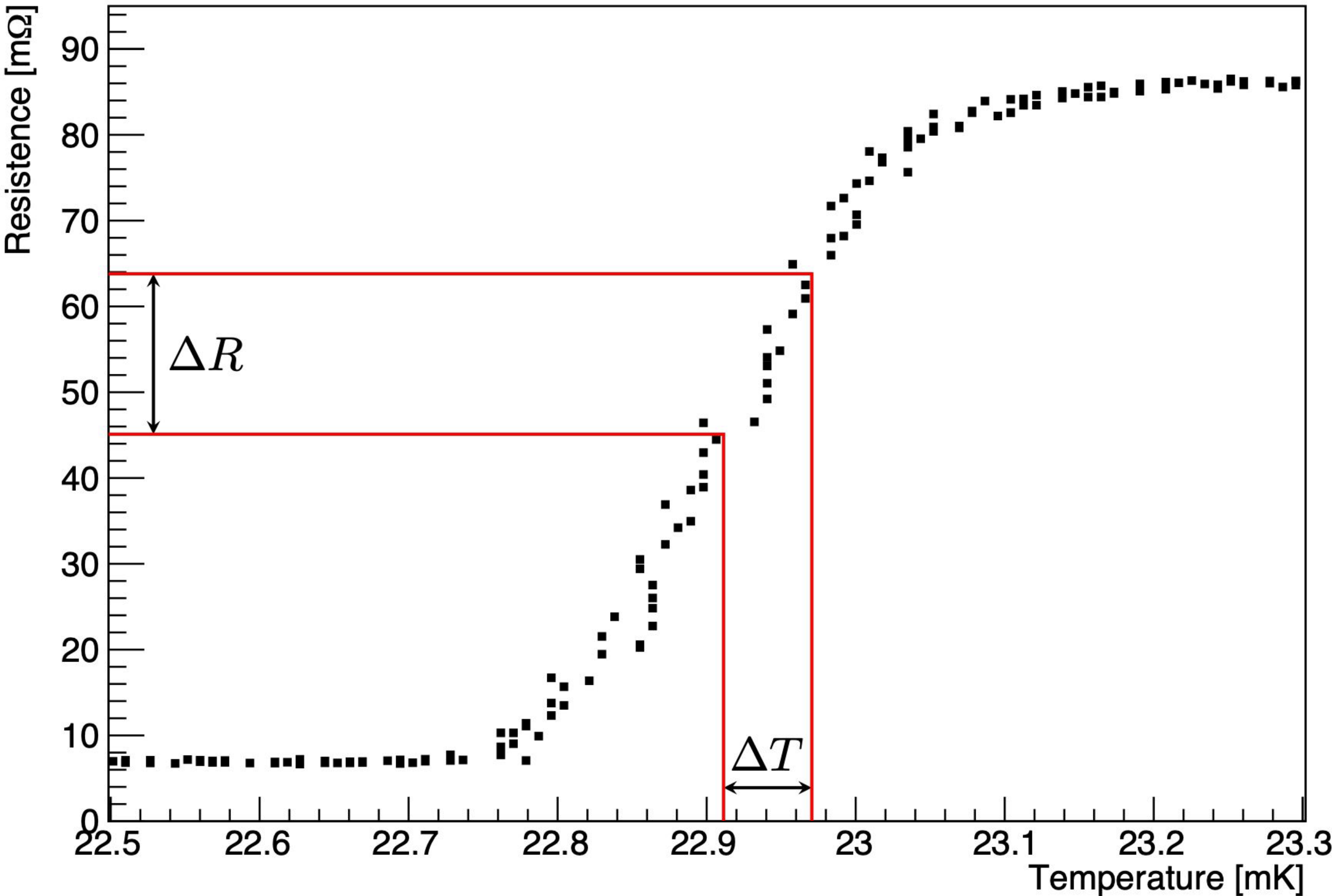


CRESST standard detectors are composed of a scintillating crystal as the main absorber, paired with a Silicon-on-Sapphire wafer as light detector.

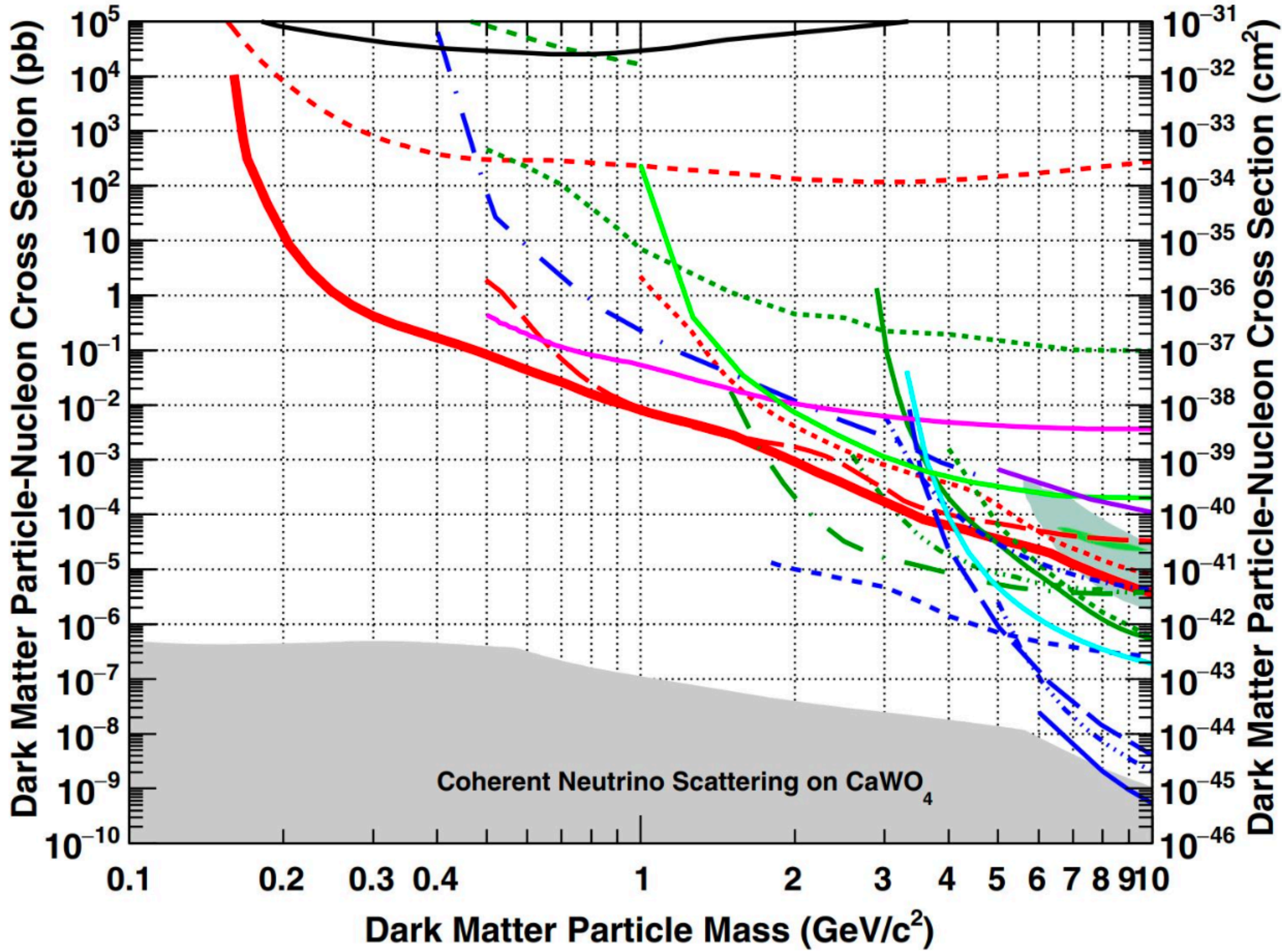
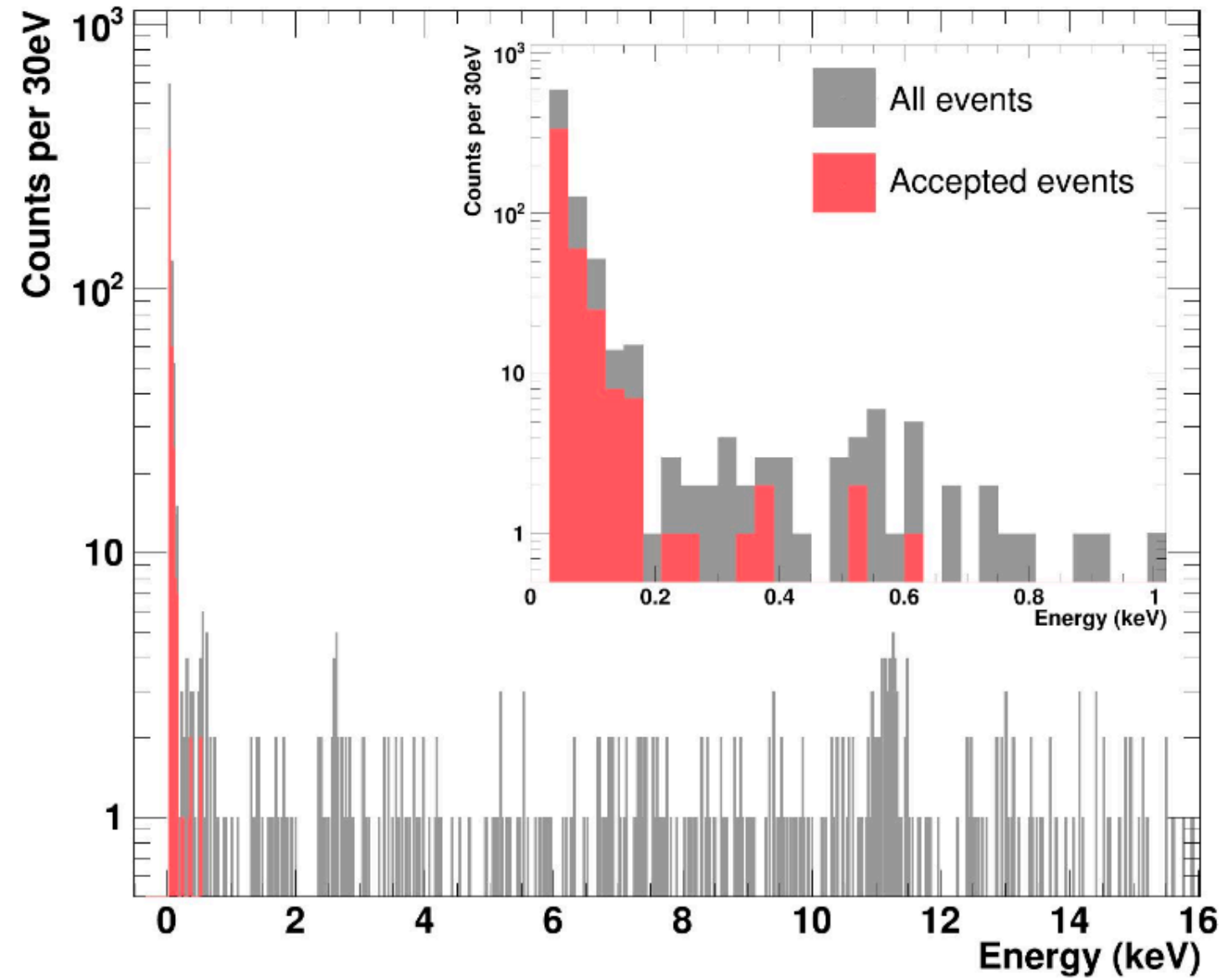
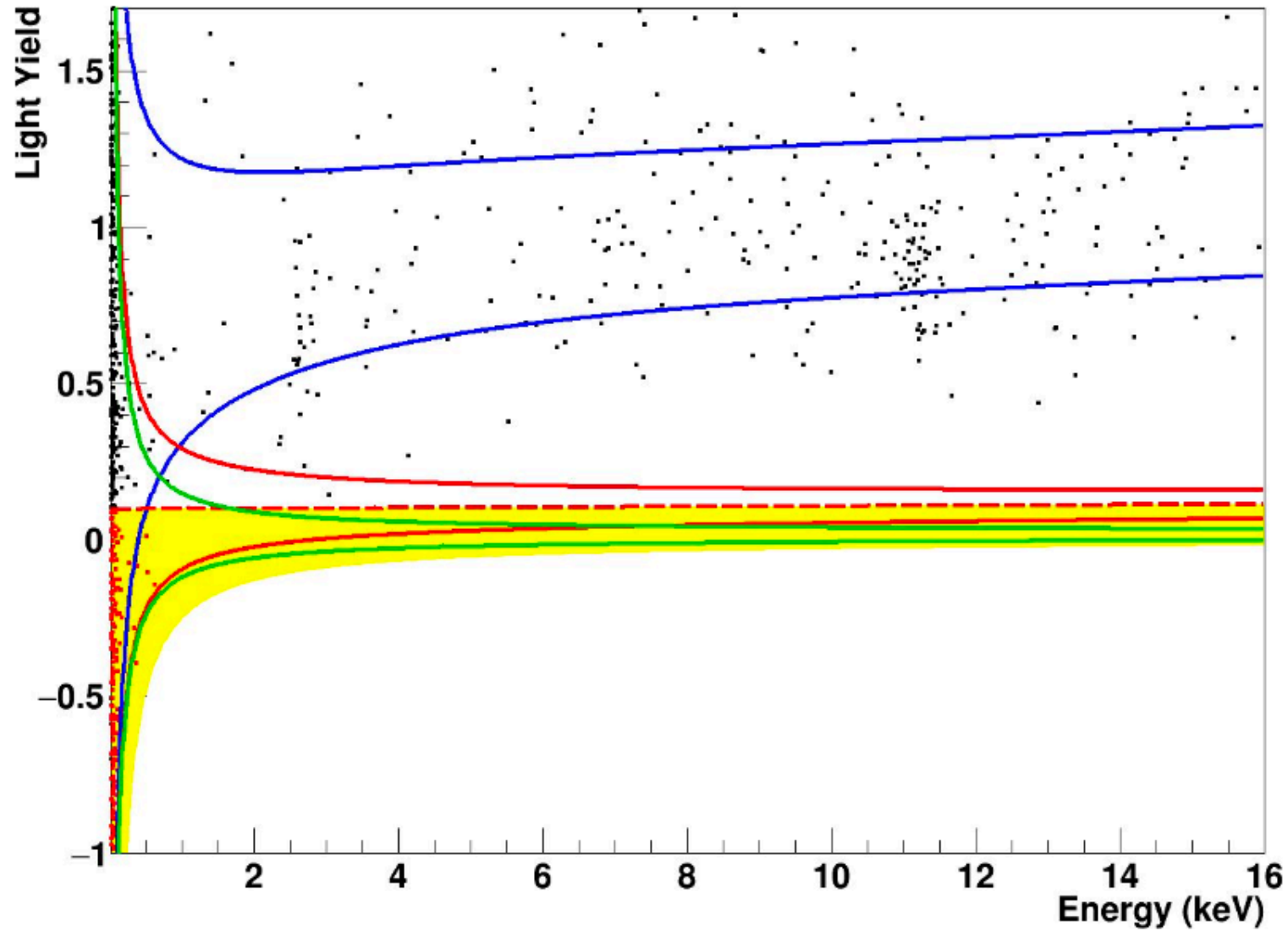
The double readout of light and phonon signals allows for separating potential signal events (nuclear recoils) and the dominant radioactive background (electron recoils) and defining different bands.

$$LY = \frac{\text{Light signal}}{\text{Phonon signal}}$$

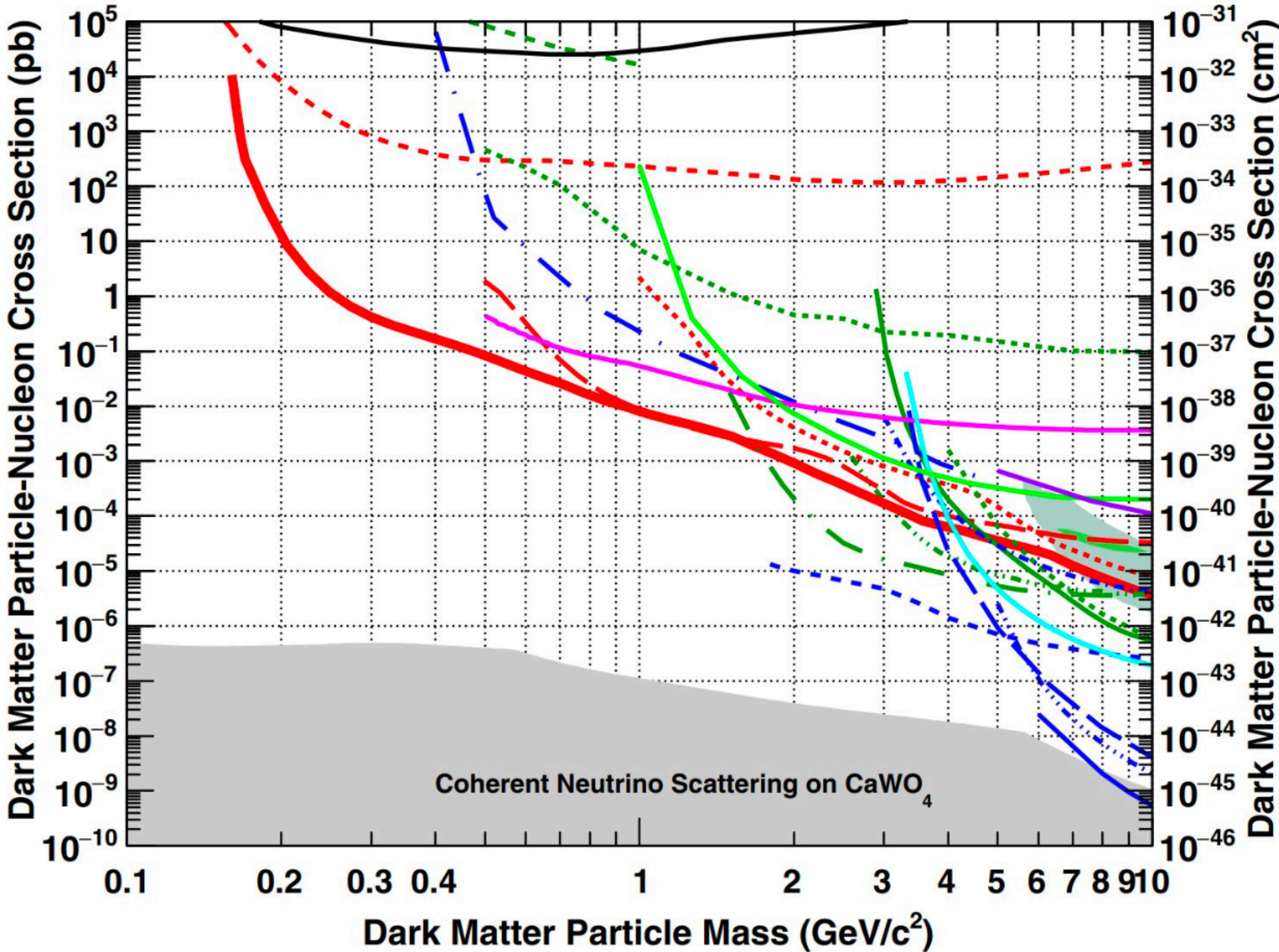
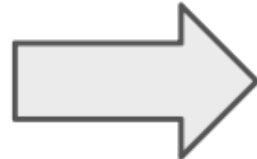
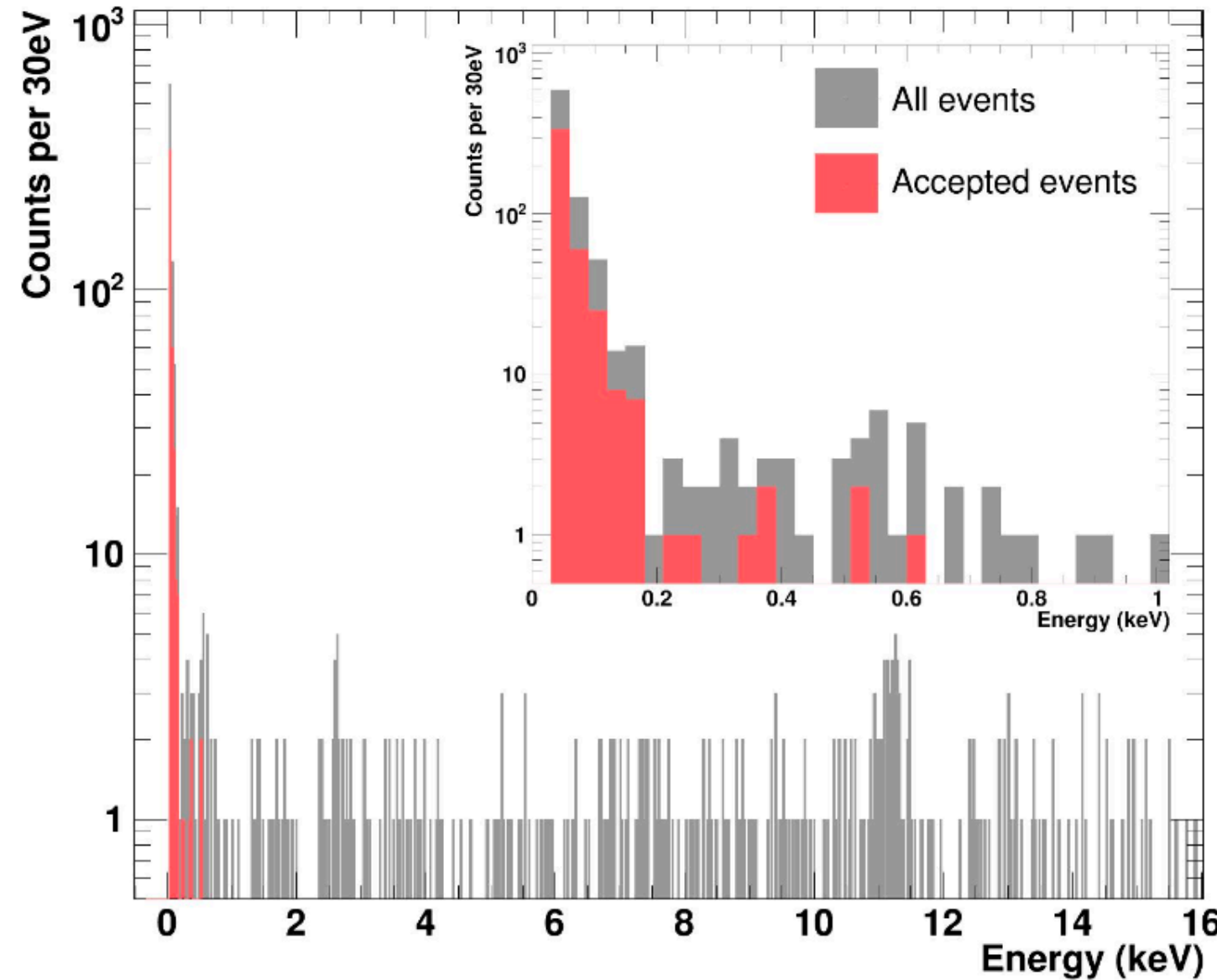
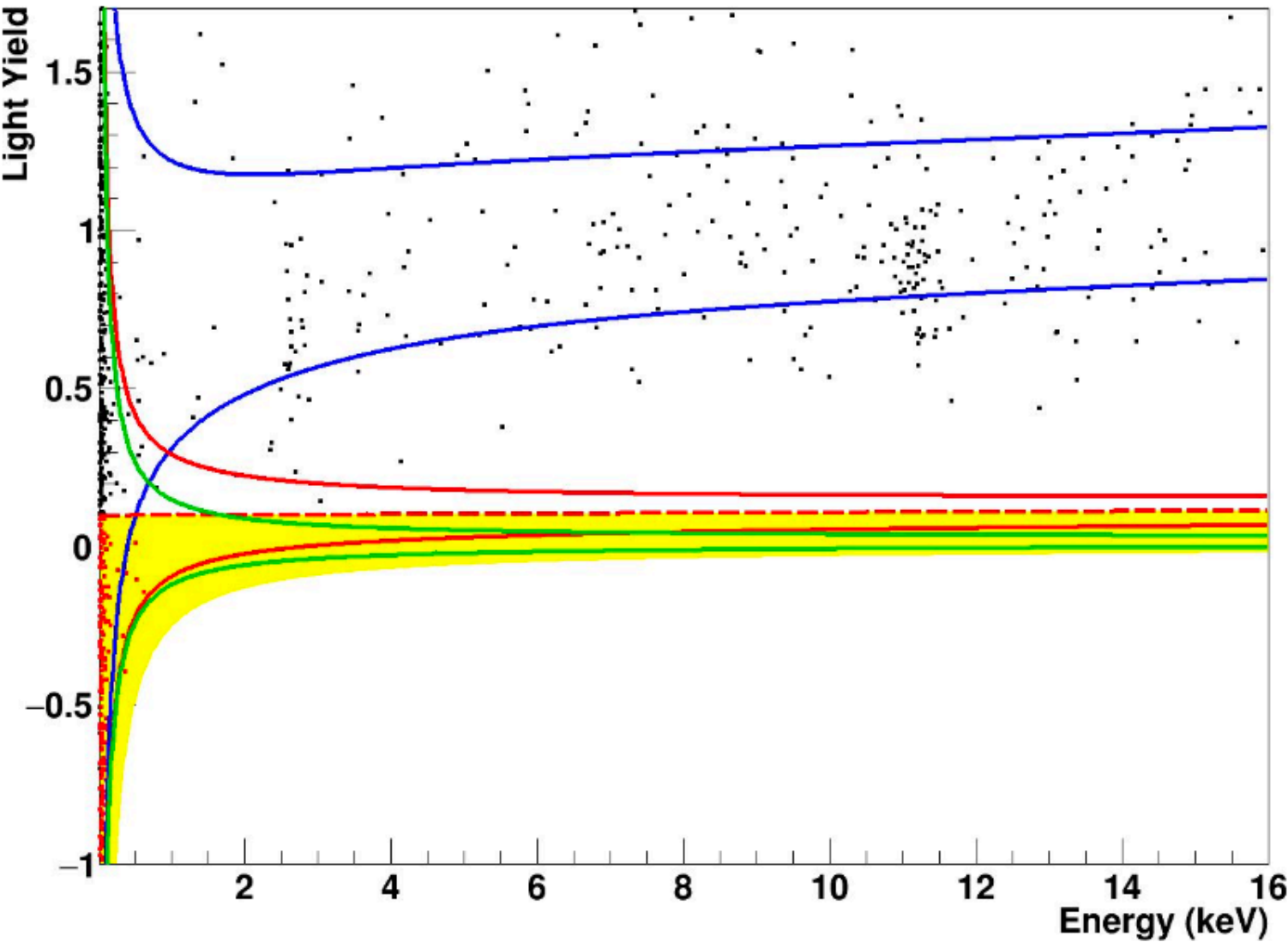




$E \sim \text{keV} \Rightarrow T \sim \mu\text{K} \Rightarrow R \sim \text{m}\Omega$



[Phys. Rev. D 100, 102002](#)

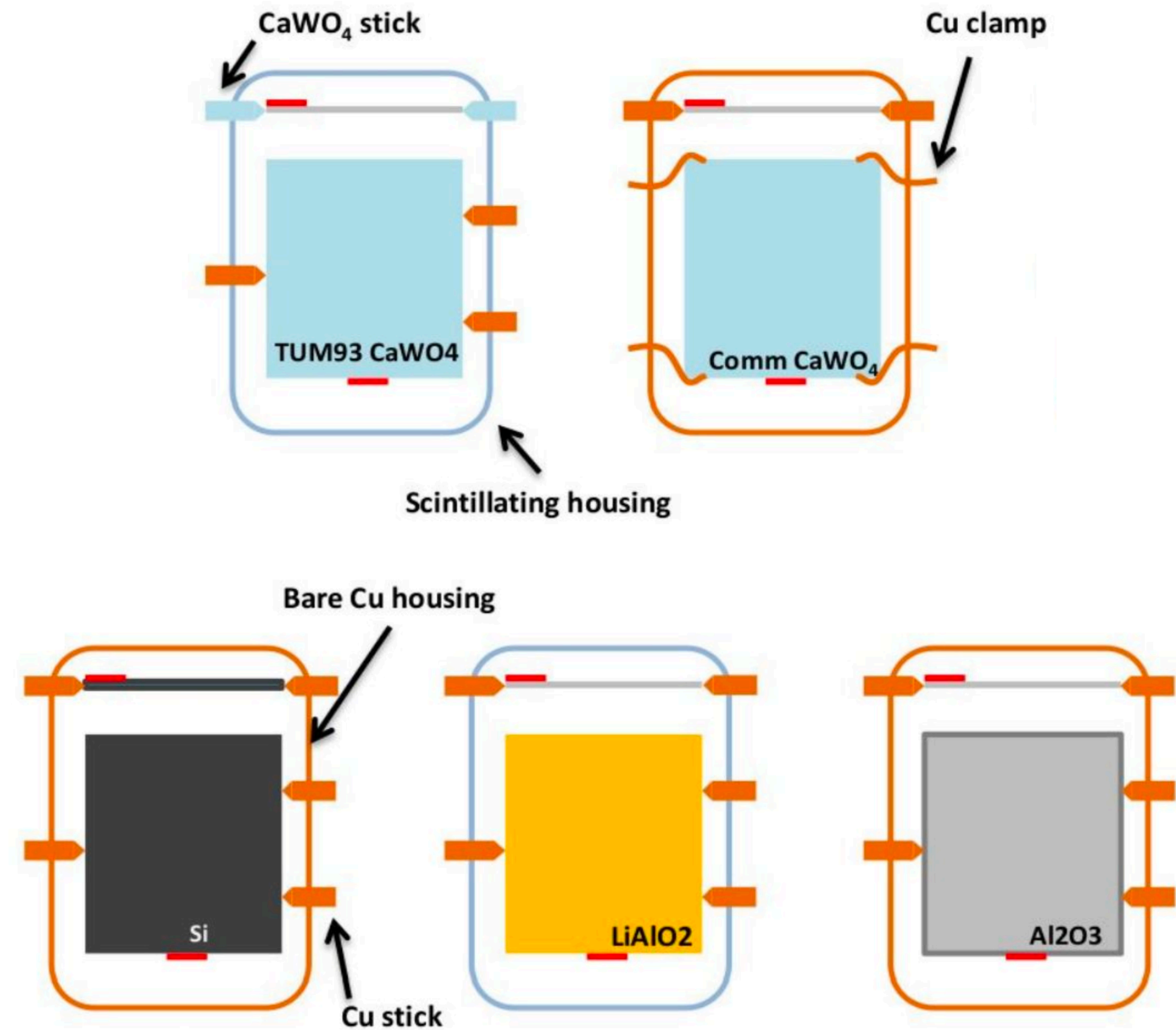


[Phys. Rev. D 100, 102002](#)

Breakthrough result + First observation of the unexpected rise of events below 200 eV, Low Energy Excess (LEE), limiting Dark Matter sensitivity

Physics campaign to pinpoint the LEE

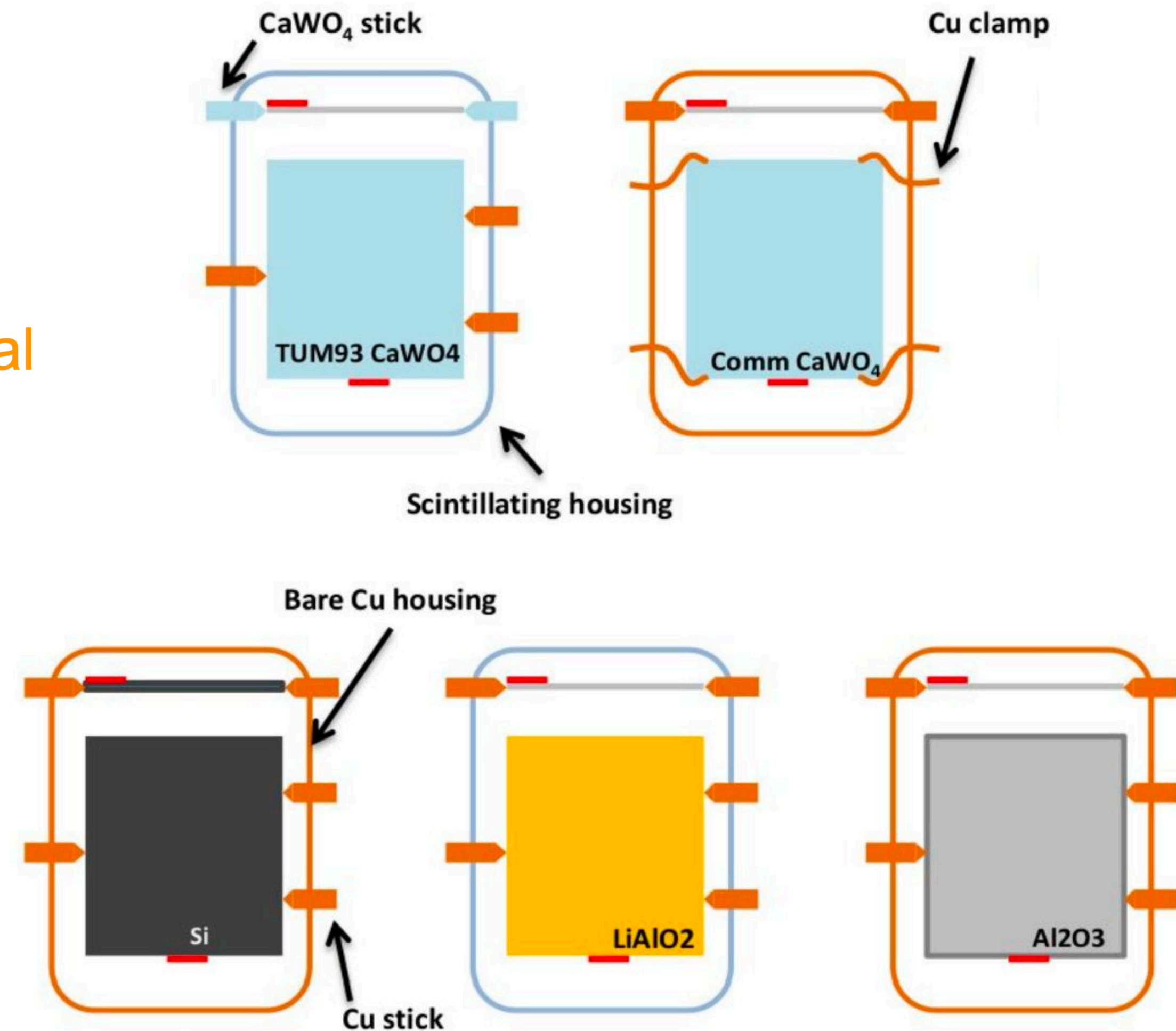
Target material?
Stress induced by holder?
Crystal growth parameters?
Scintillating material?
Geometry?

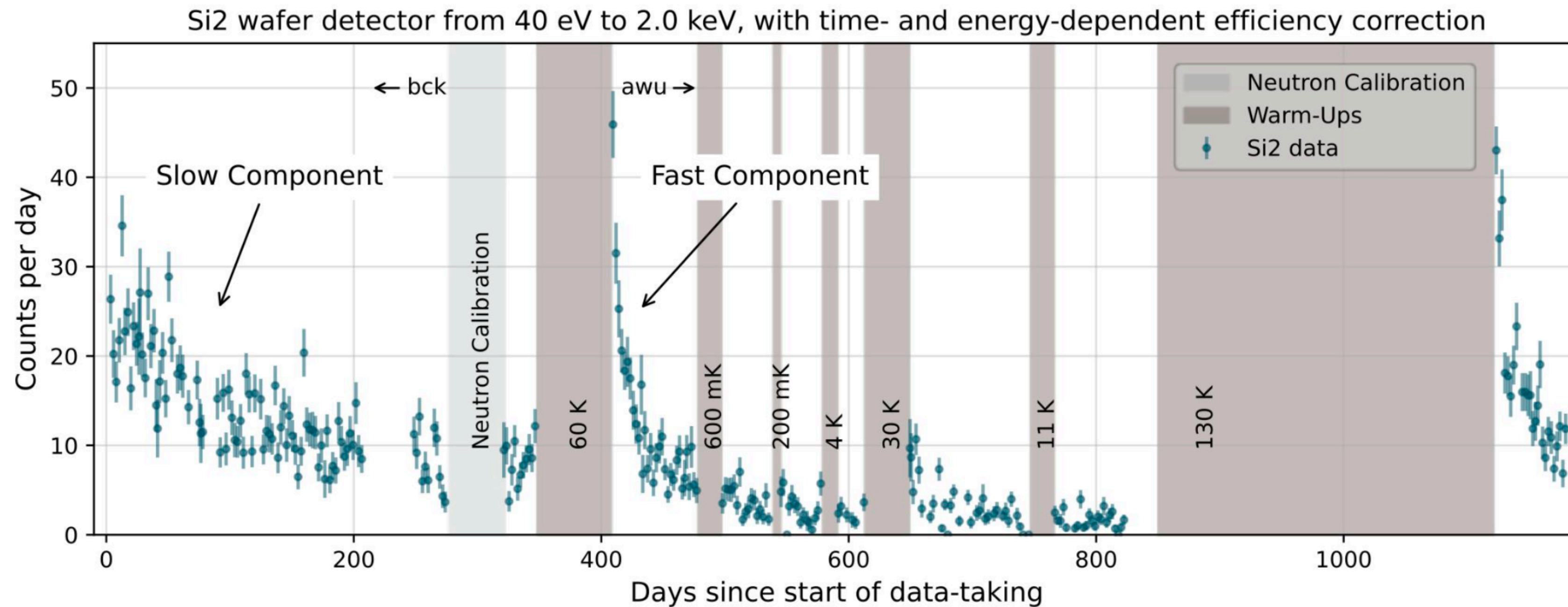


Physics campaign to pinpoint the LEE

Target material? Different materials
Stress induced by holder? Different holder
Crystal growth parameters? Tested on slow grown crystal
Scintillating material? Removed scintillating material
Geometry? LEE studies also on wafer

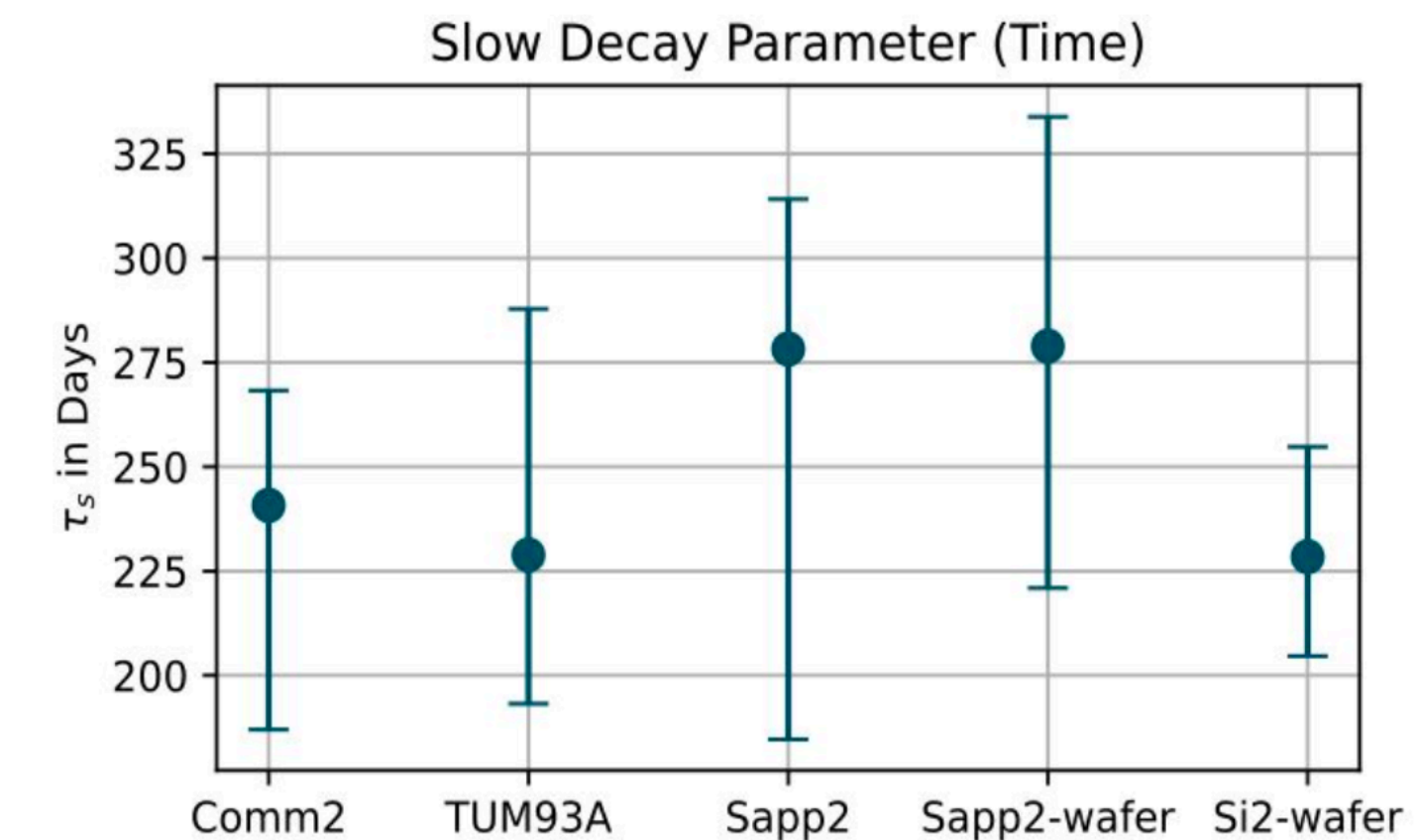
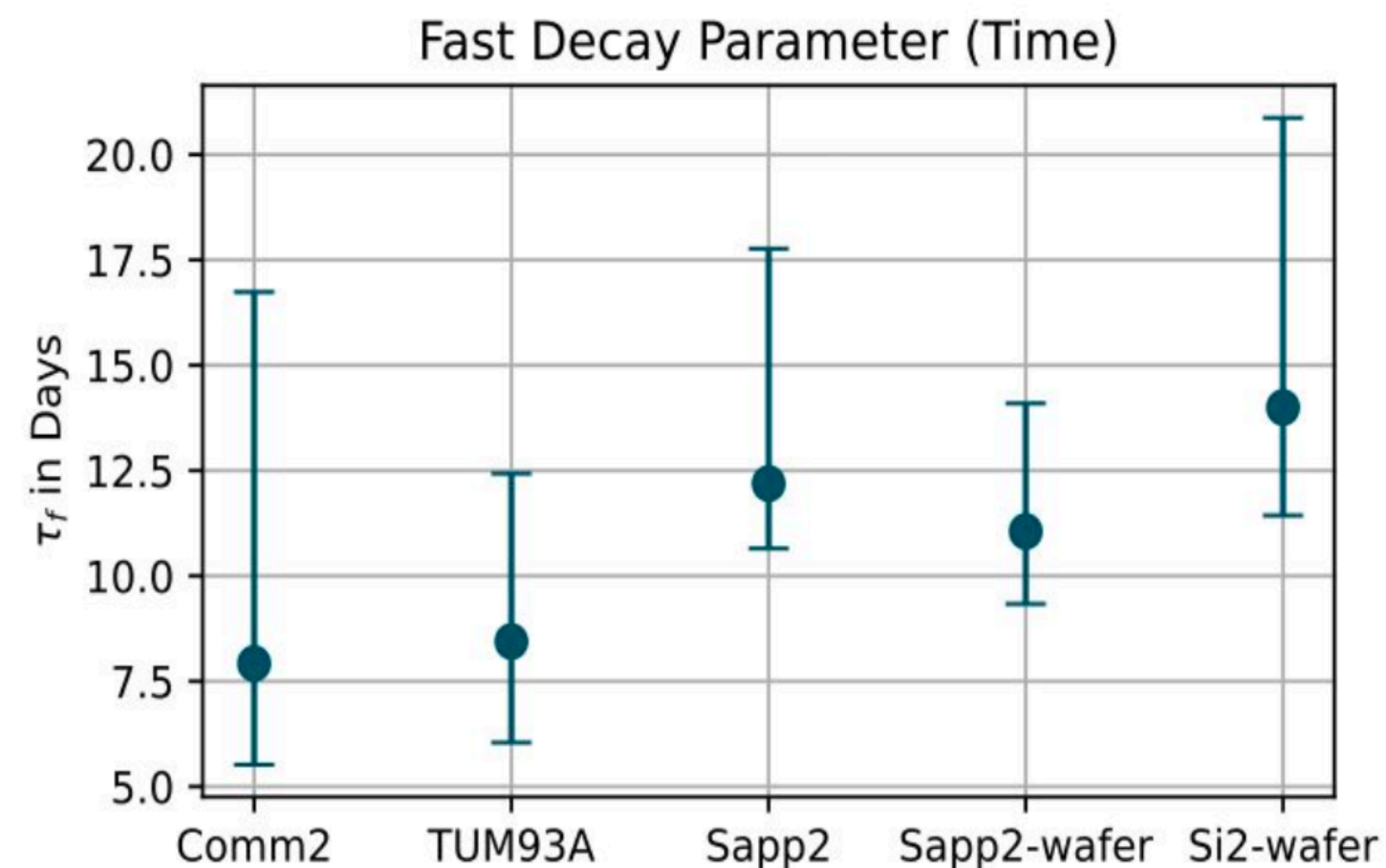
None of the modification had a significant impact on the LEE

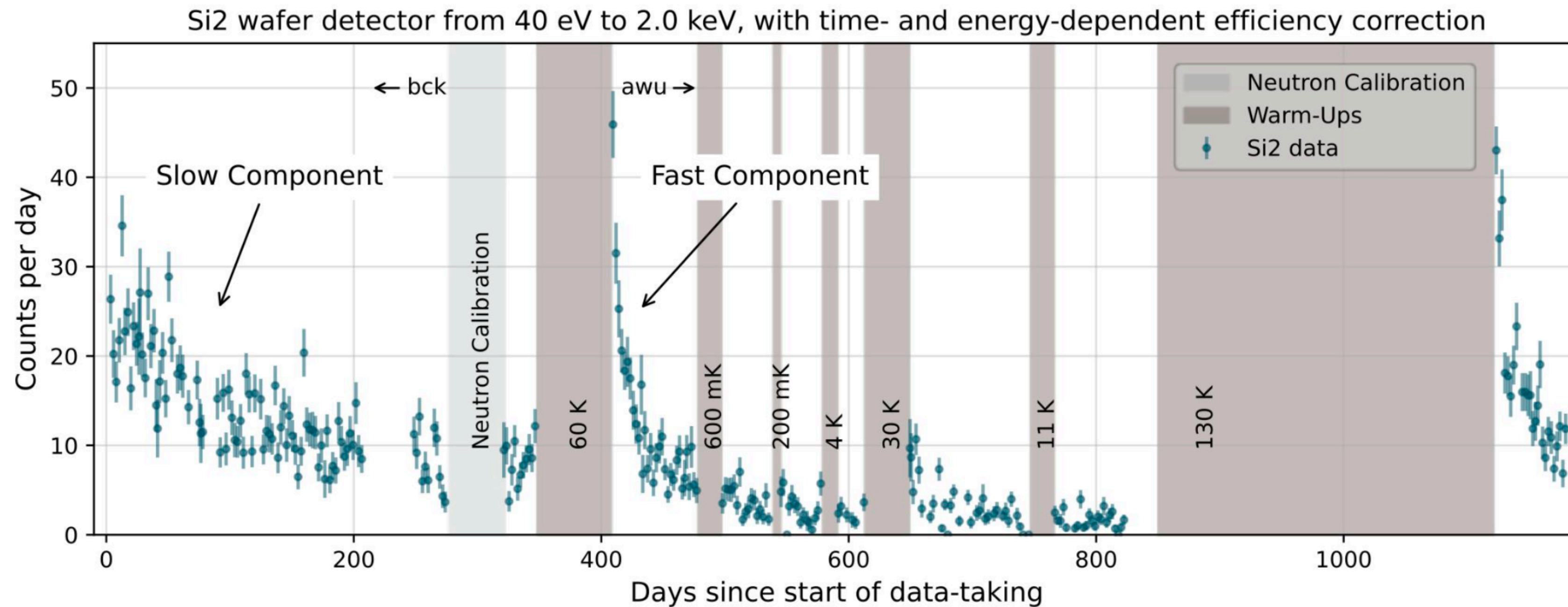




Rate monitoring over time:

- Rate decays exponentially in time
- Reset of the rate after warm-up cycles
- Two decay constants
 - Fast decay ~ 10 days
 - Slow decay ~ 250 days

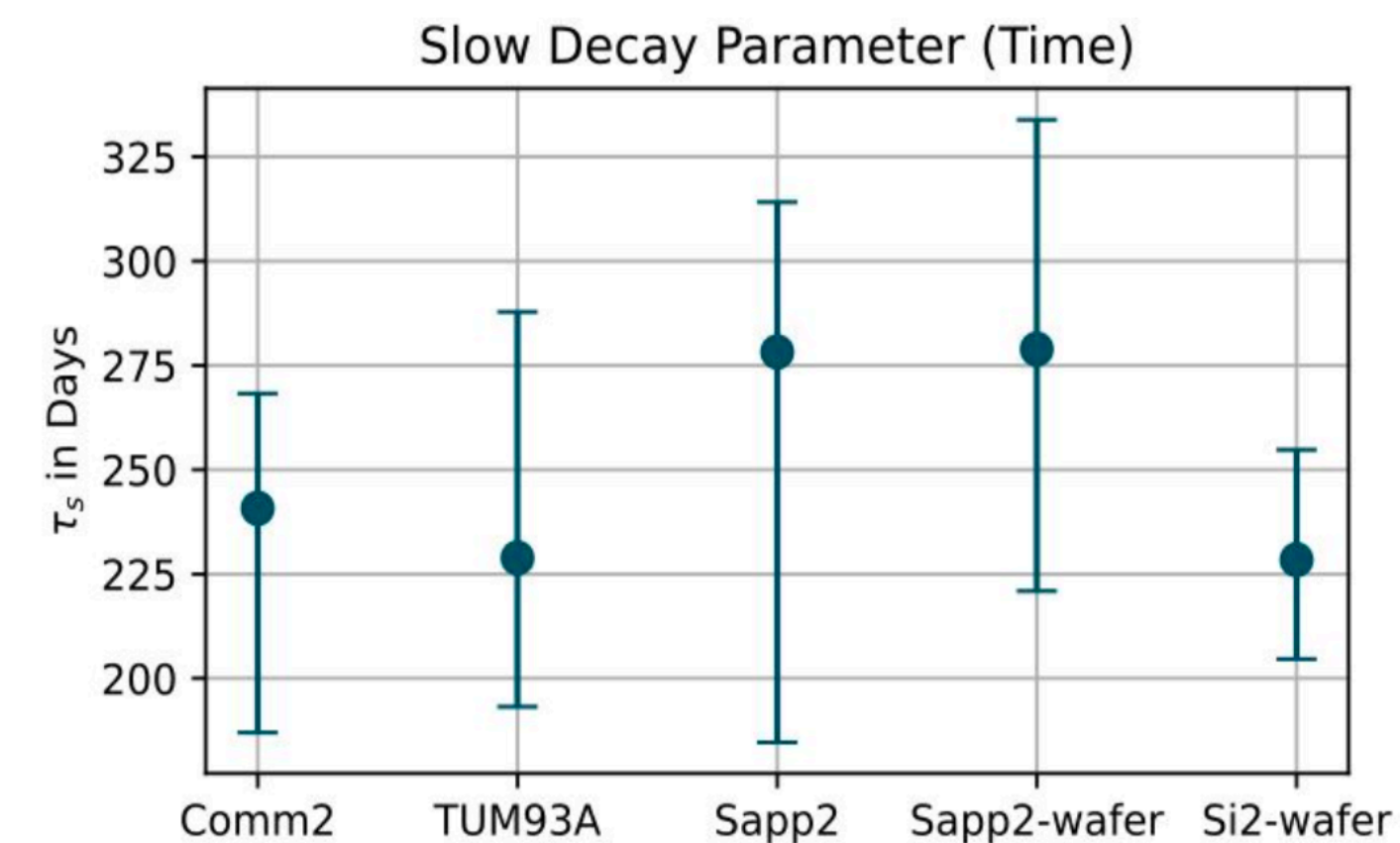
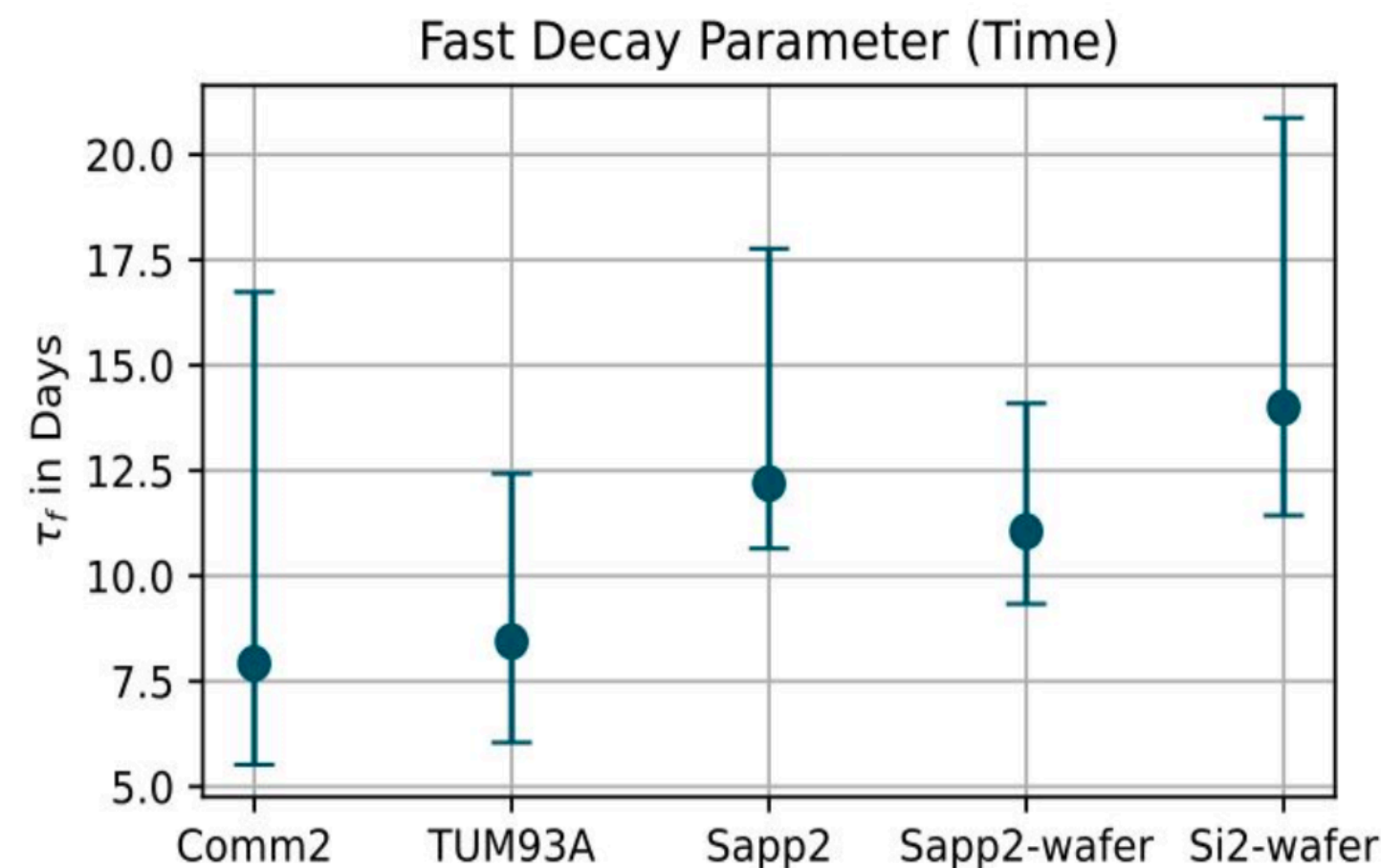




Not caused by radioactivity

Rate monitoring over time:

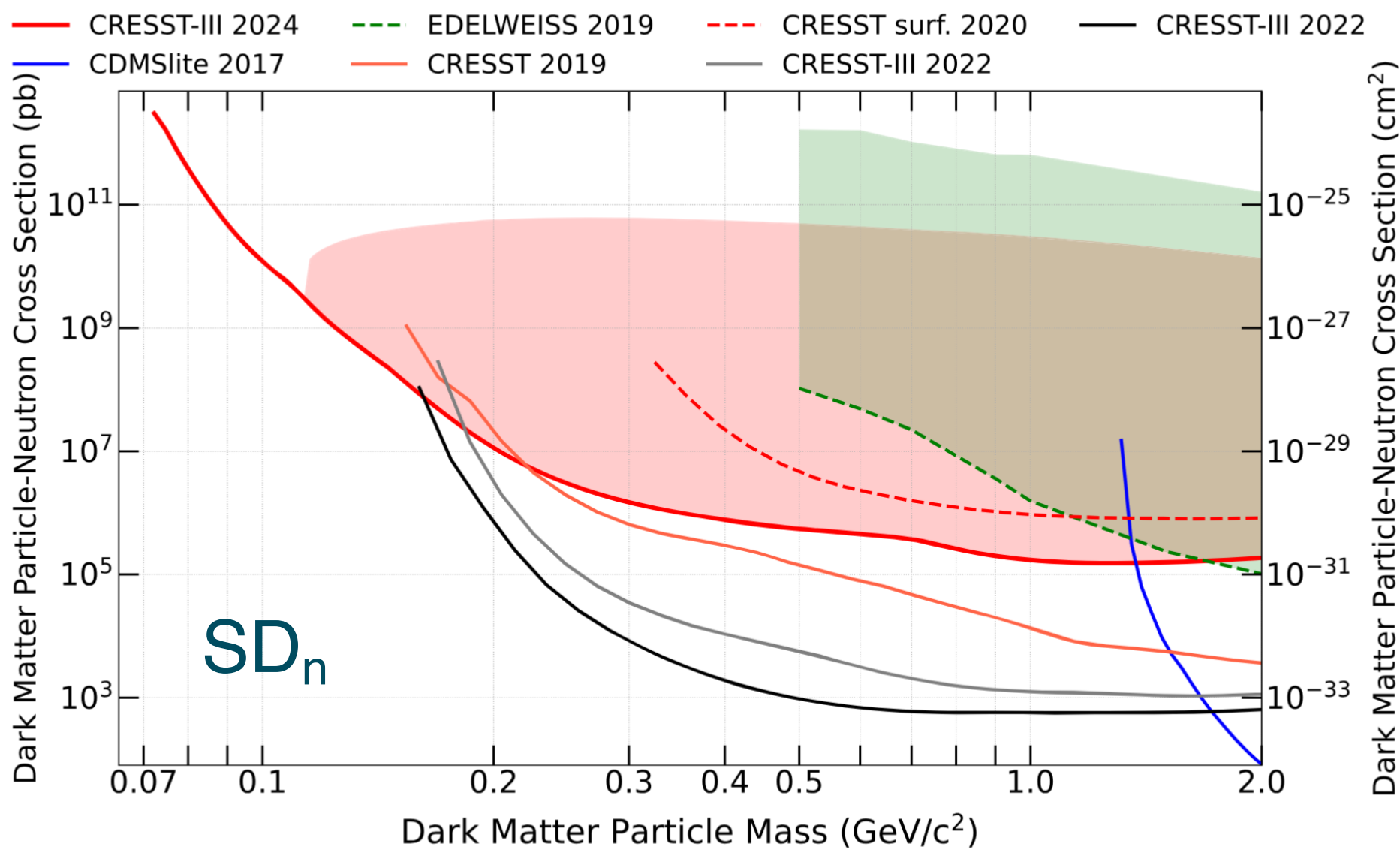
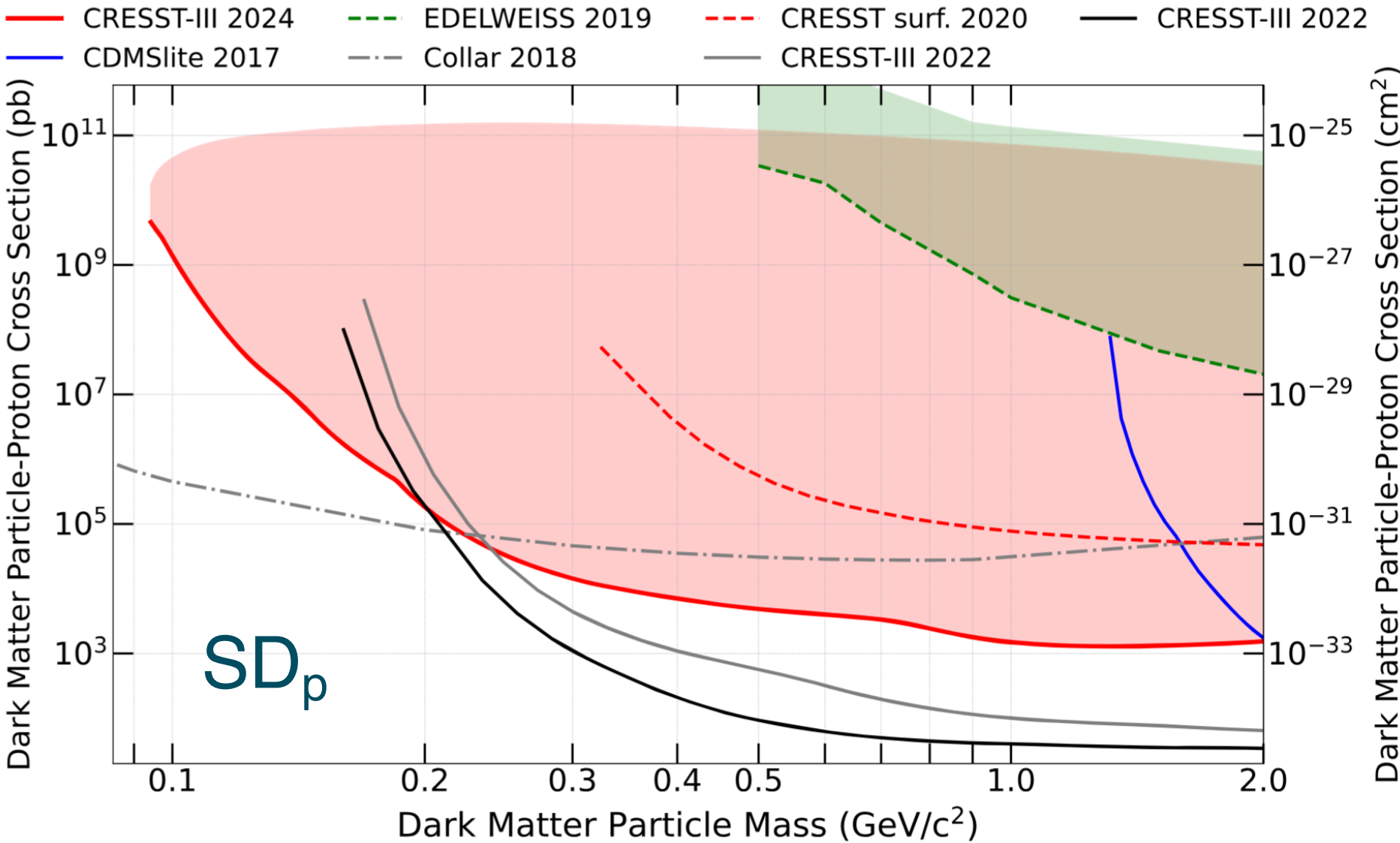
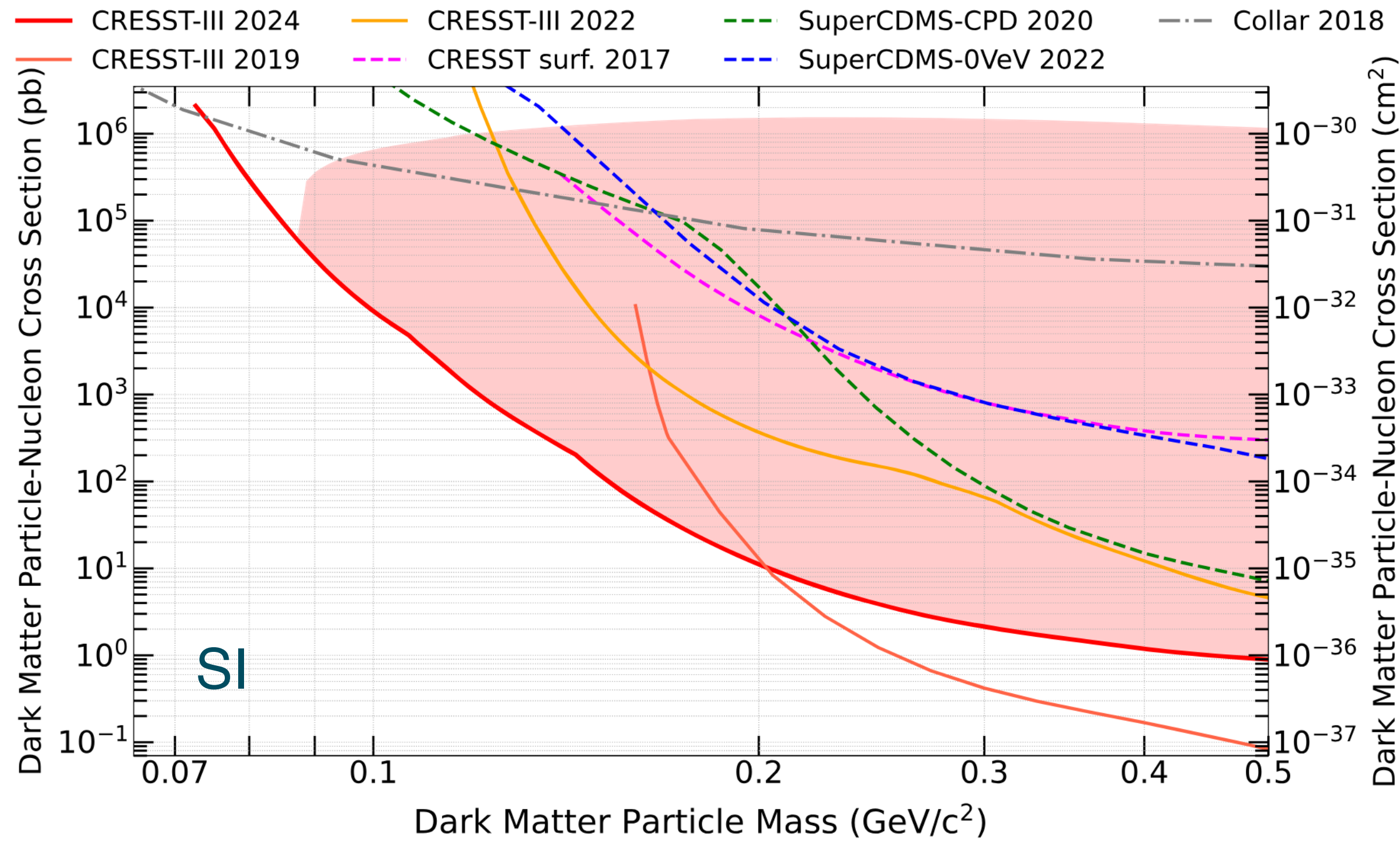
- Rate decays exponentially in time
- Reset of the rate after warm-up cycles
- Two decay constants
 - Fast decay ~ 10 days
 - Slow decay ~ 250 days



Low-threshold Dark Matter results

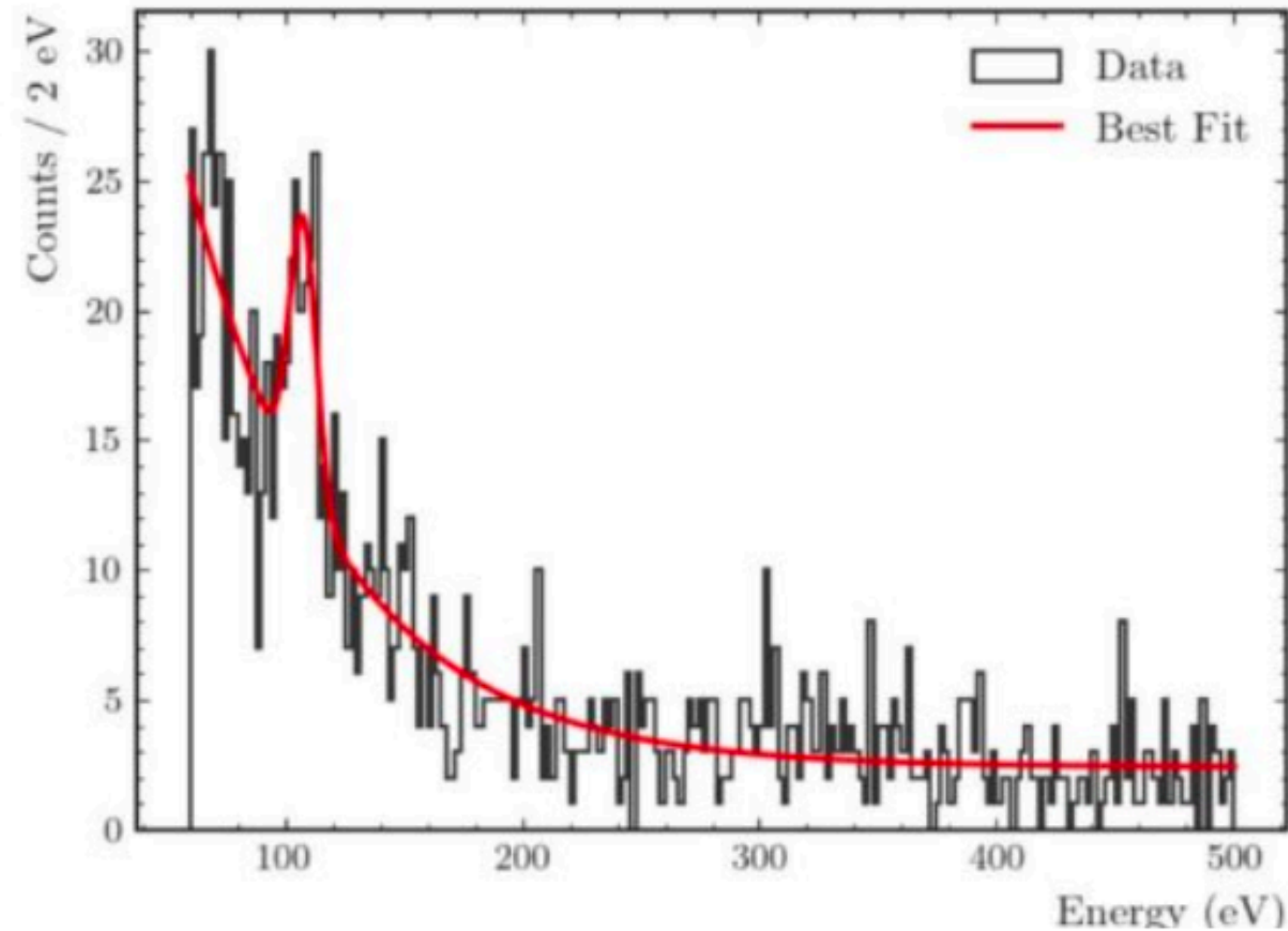
Light detector analyzed for dark matter limit:

Material: Al_2O_3
Mass: 0.6 g
Exposure: 138 g · days
Resolution: 1 eV
Threshold: 6.7 eV_{nr}



[10.1103/PhysRevD.110.083038](https://arxiv.org/abs/10.1103/PhysRevD.110.083038)

[Phys. Rev. D 108, 022005](#)

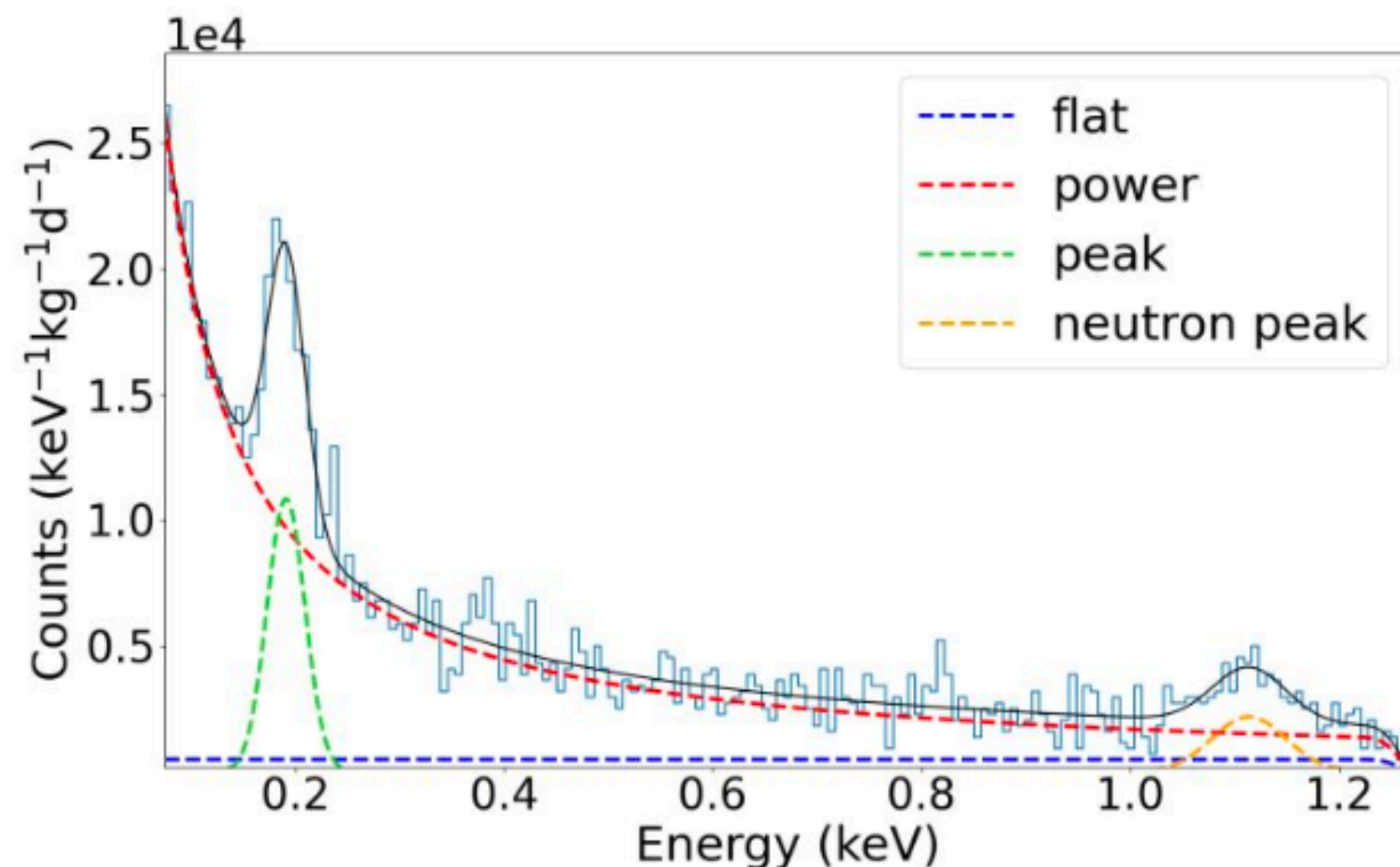


New technique for low nuclear recoil energy calibration for CaWO_4 and Al_2O_3 crystals through (n, γ) reactions



De-excitation gamma of 6.1 MeV and W nuclear recoil of 112eV

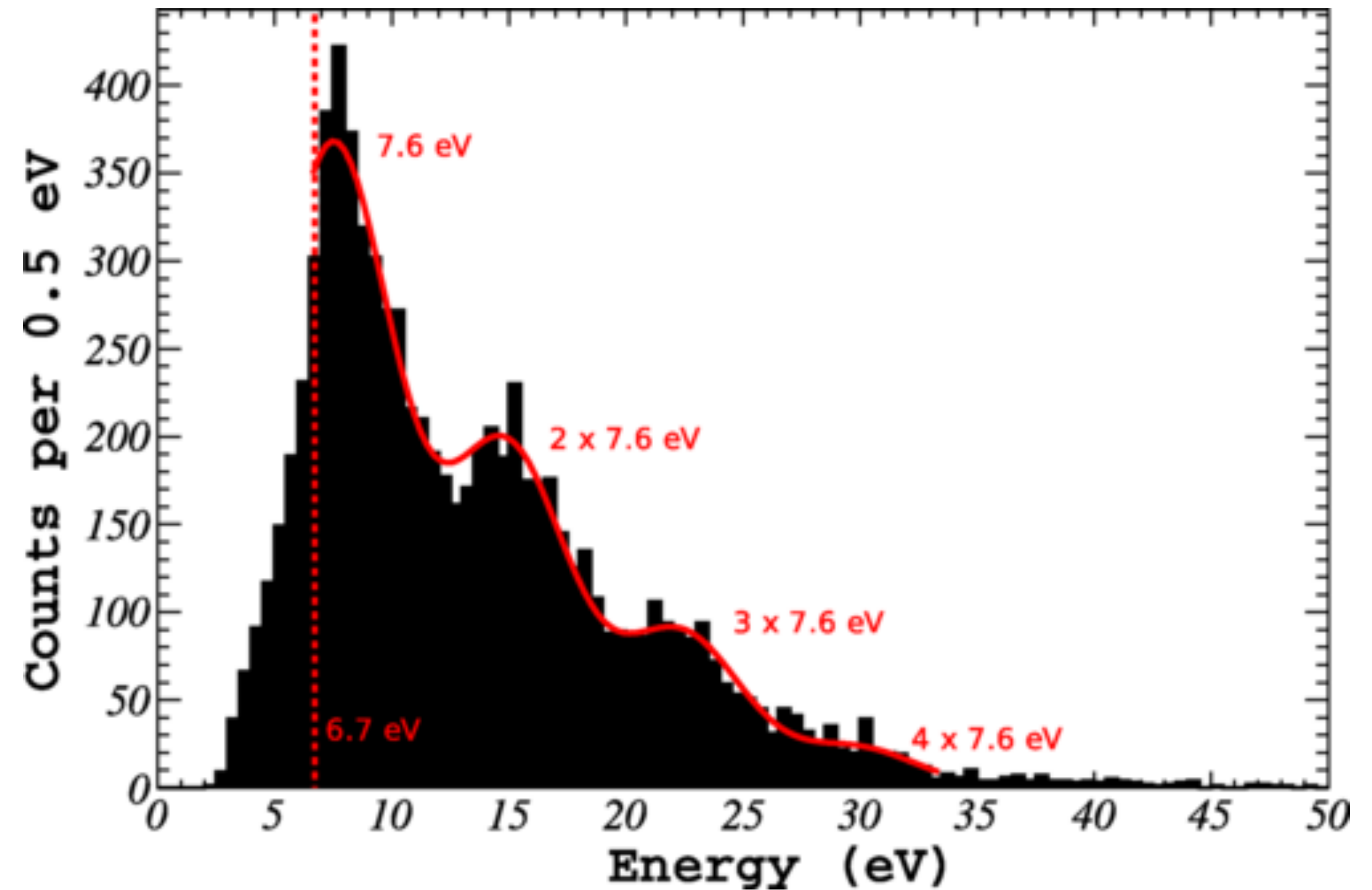
[arxiv.org:2506.09059](#)



De-excitation gamma of 7.7 MeV and Al nuclear recoil of 1144eV

First measurement of single photon in CRESST from sapphire de-excitation

Confirm validity of our calibration method done with ^{55}Fe source (5.9 and 6.5 keV) is valid down to threshold



Phys. Rev. D 110, 083038

Double TES approach

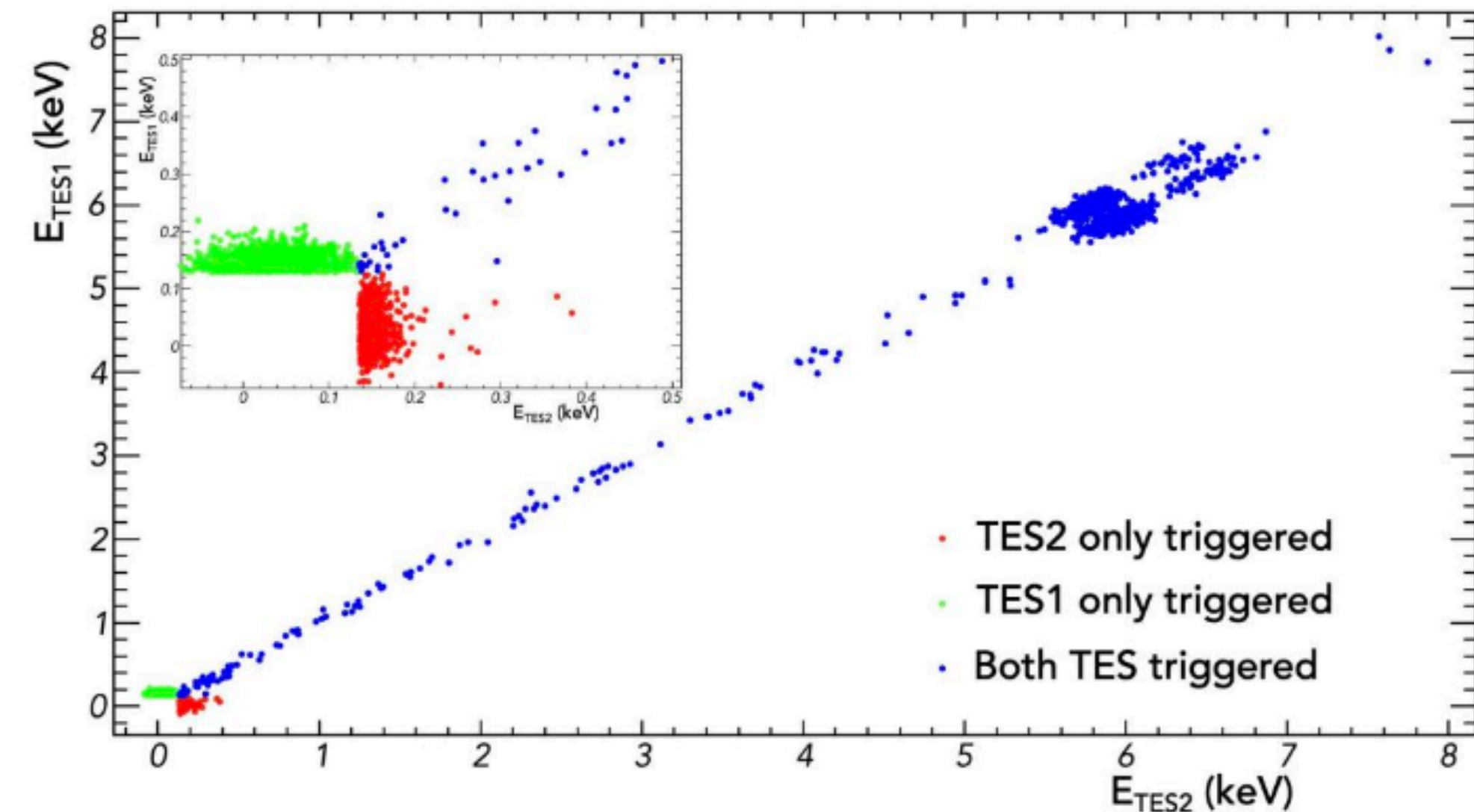
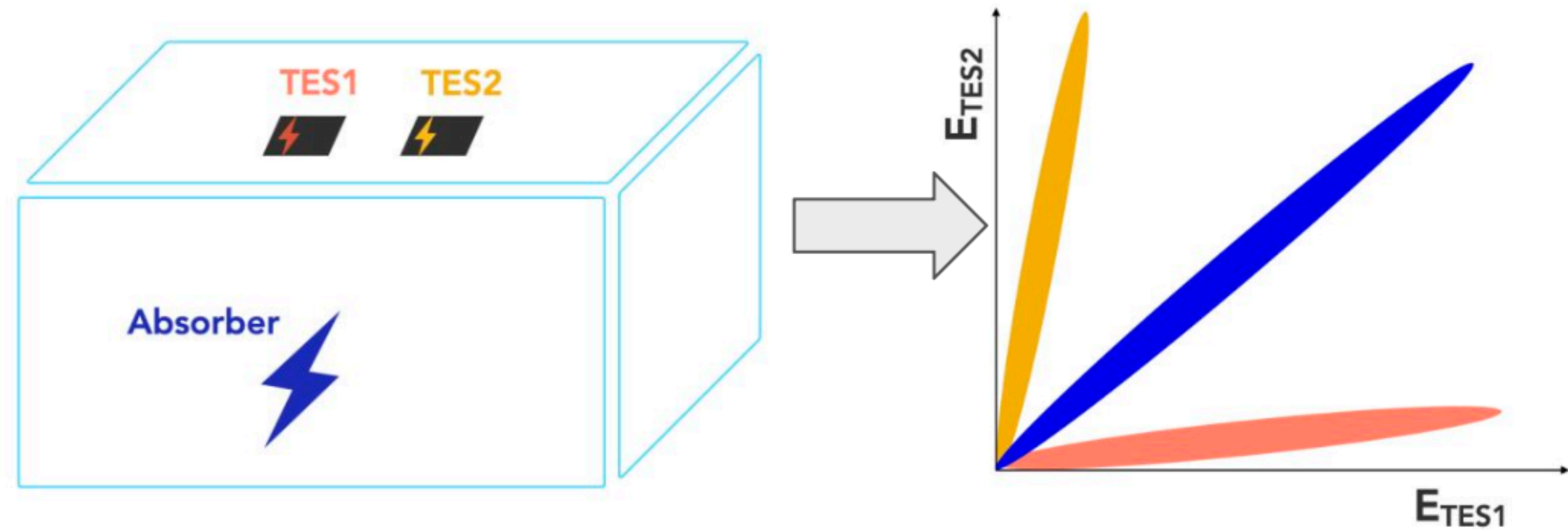
Probe the LEE origin

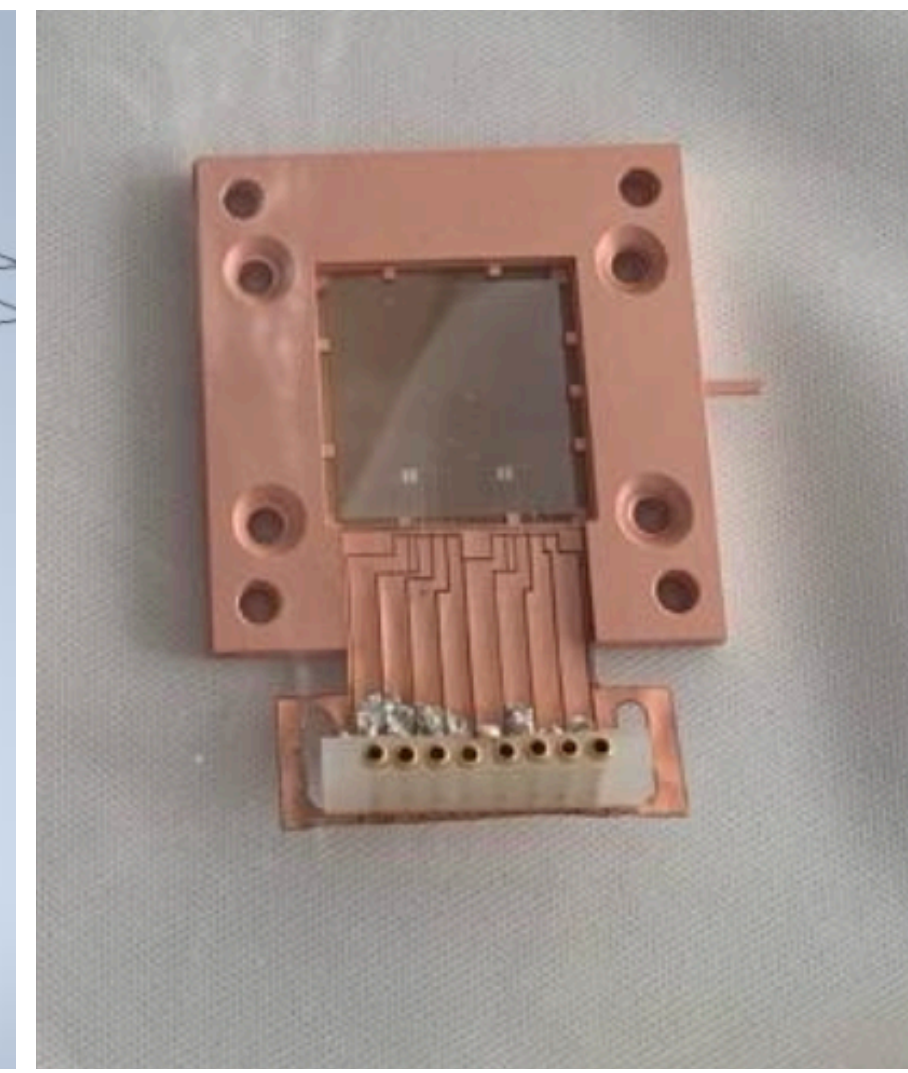
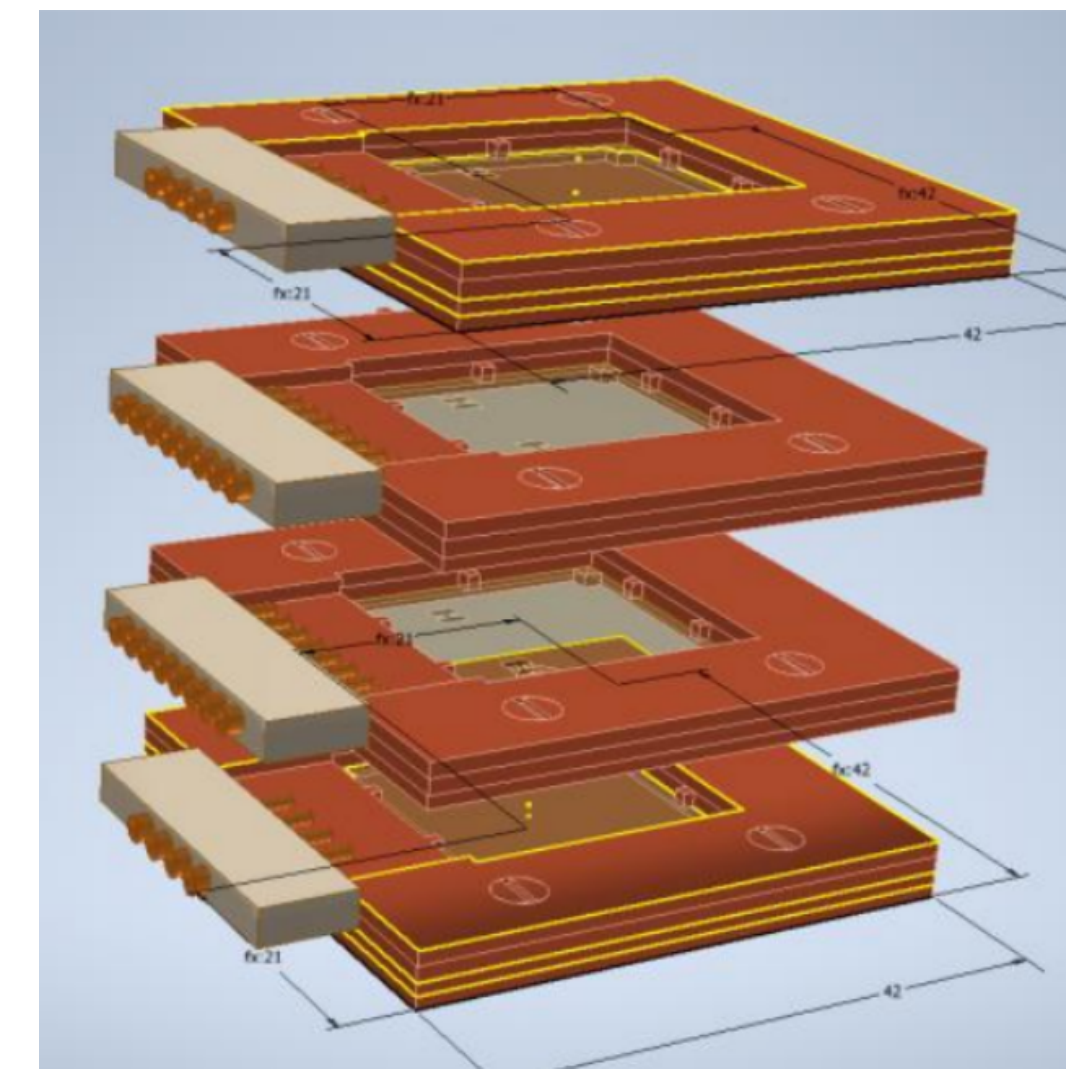
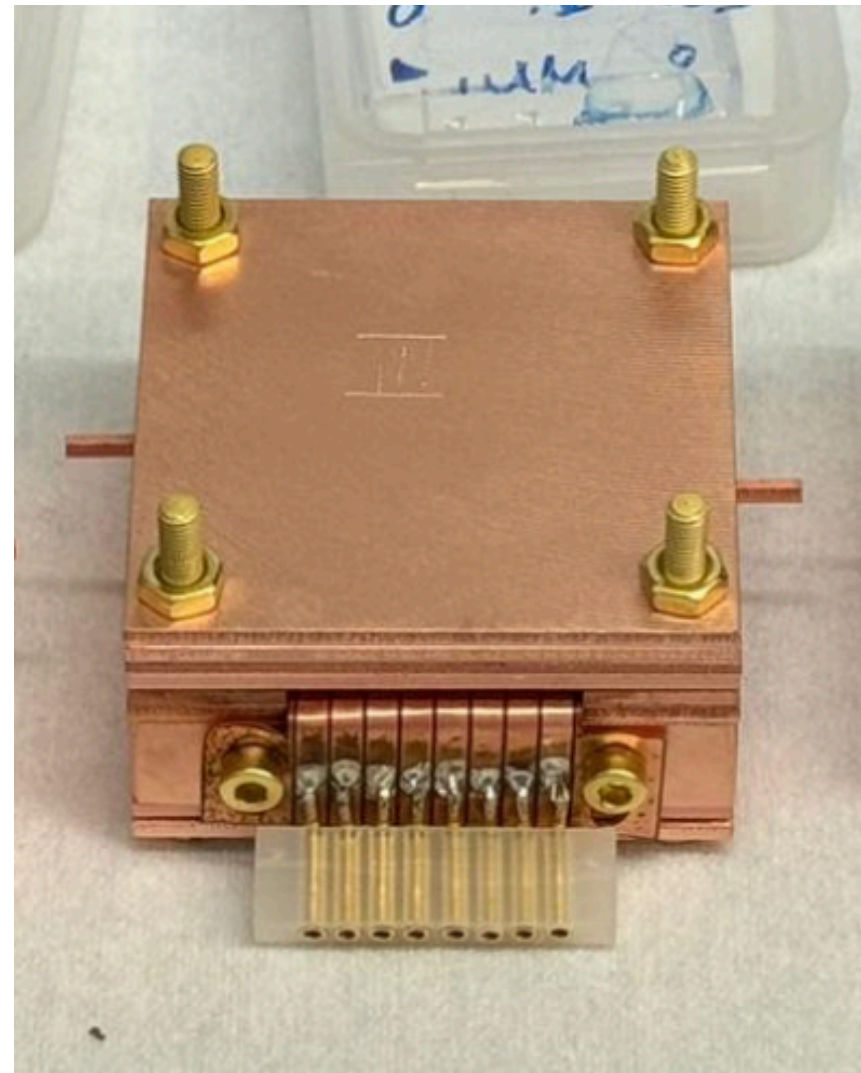
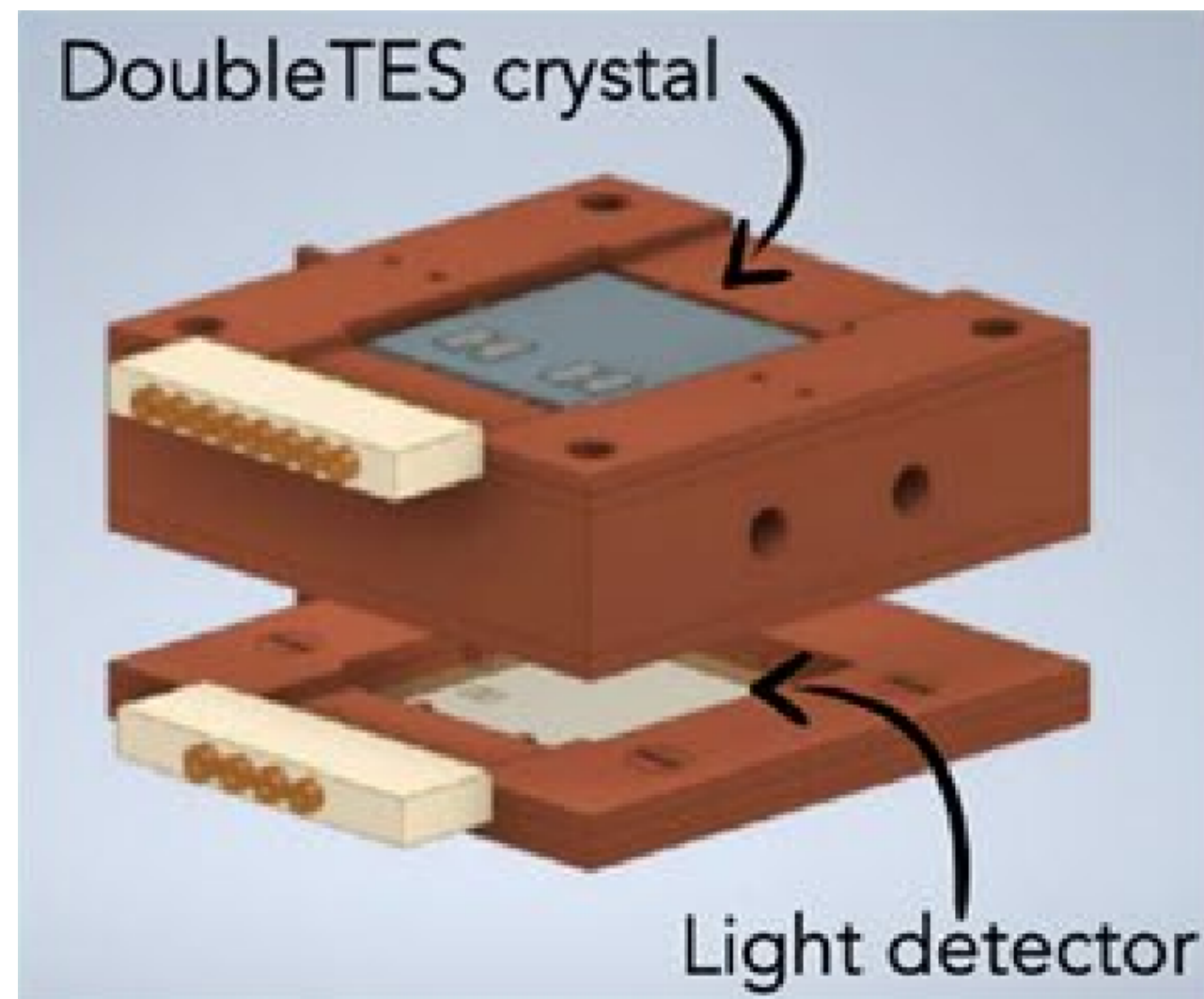
from inside TES or from interface
between TES and the crystal

Idea

- Events from the absorber are shared between the 2 TES
- Events close to single TES are seen only by one TES

First results from above ground
measurement





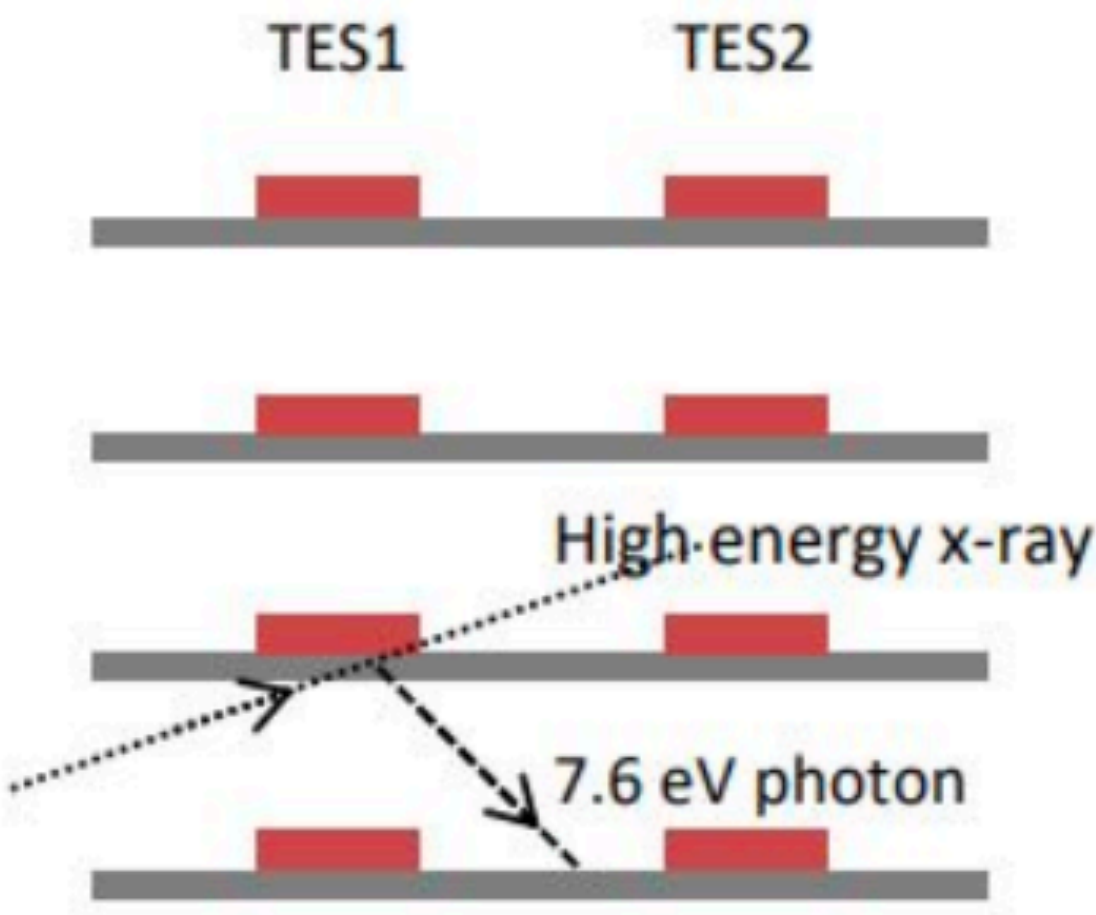
5x Double TES

- CaWO_4 crystal $20 \times 20 \times 10 \text{ mm}^3$
- Operated with two TESs
- Gravity-assisted holder
- Sapphire light detector

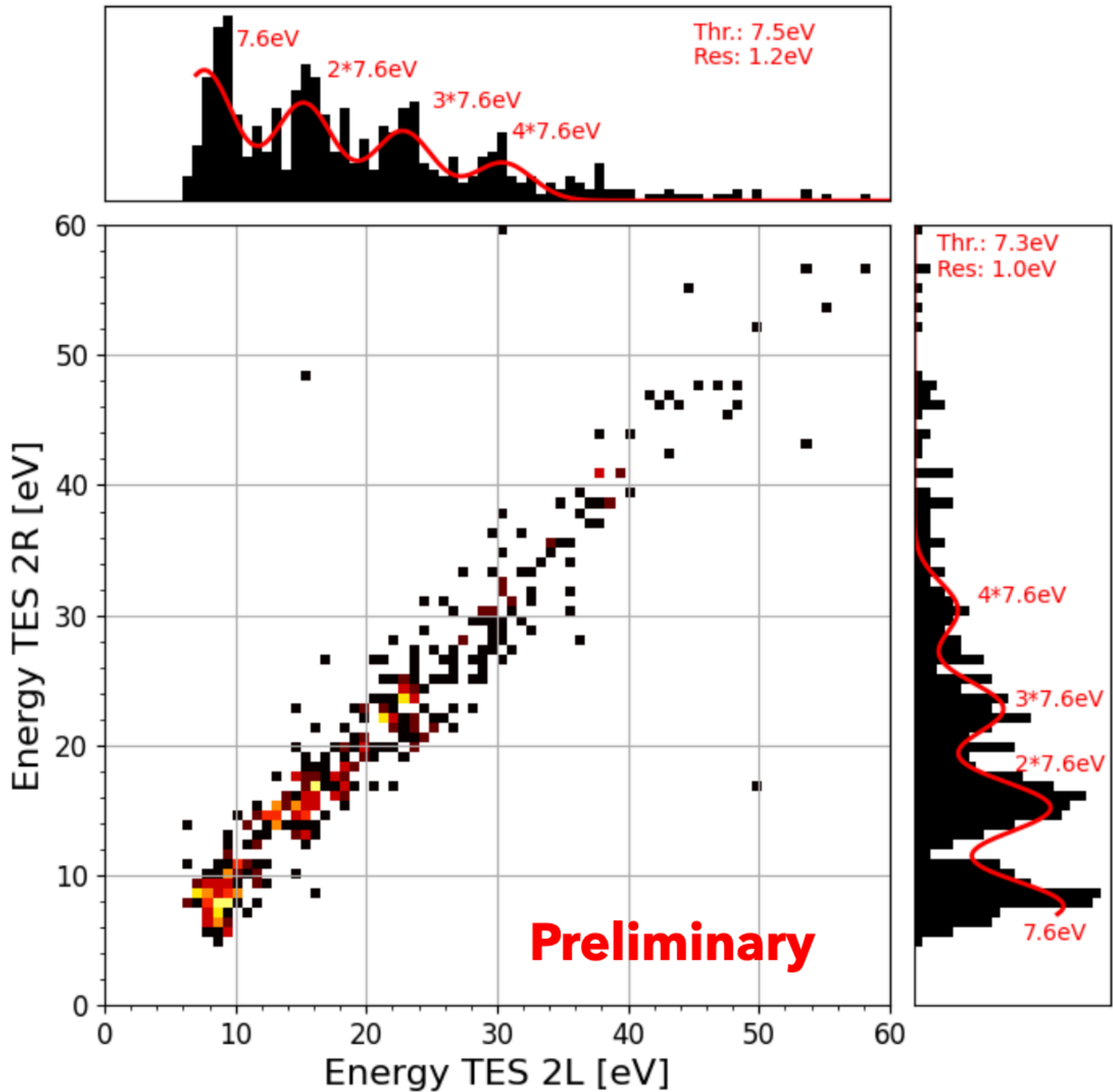
Stack

- 4x Silicon-On-Sapphire wafers $20 \times 20 \times 0.4 \text{ mm}^3$
- Operated with two TESs
- Gravity-assisted holder

- 8 TESs fully working
- Calibrated with an ^{55}Fe source
- Low-energy thresholds optimized to light dark matter
- Baseline resolution: ~ 1 eV
- Analysis threshold: ~ 7.5 eV
- Single photon detection

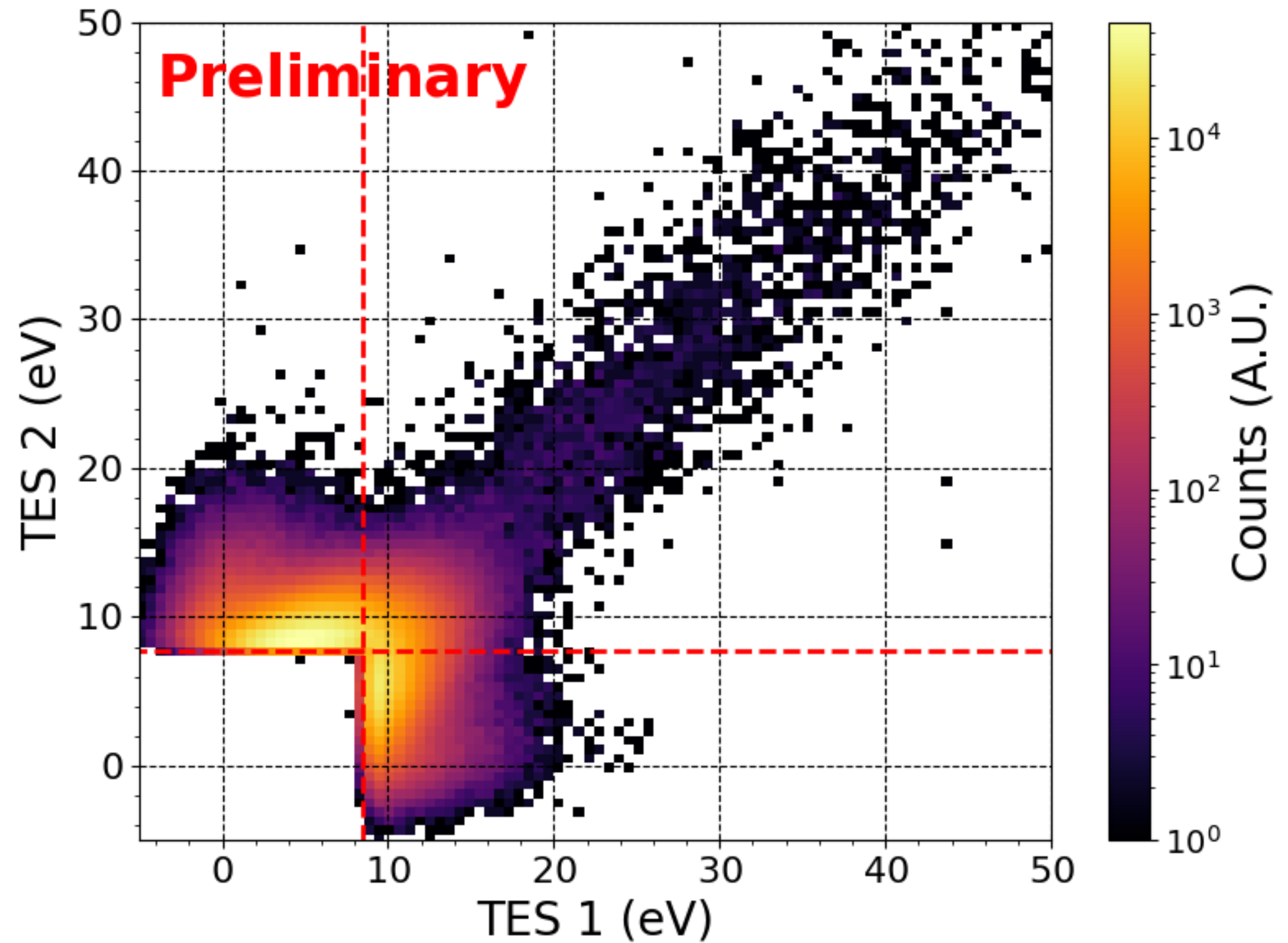
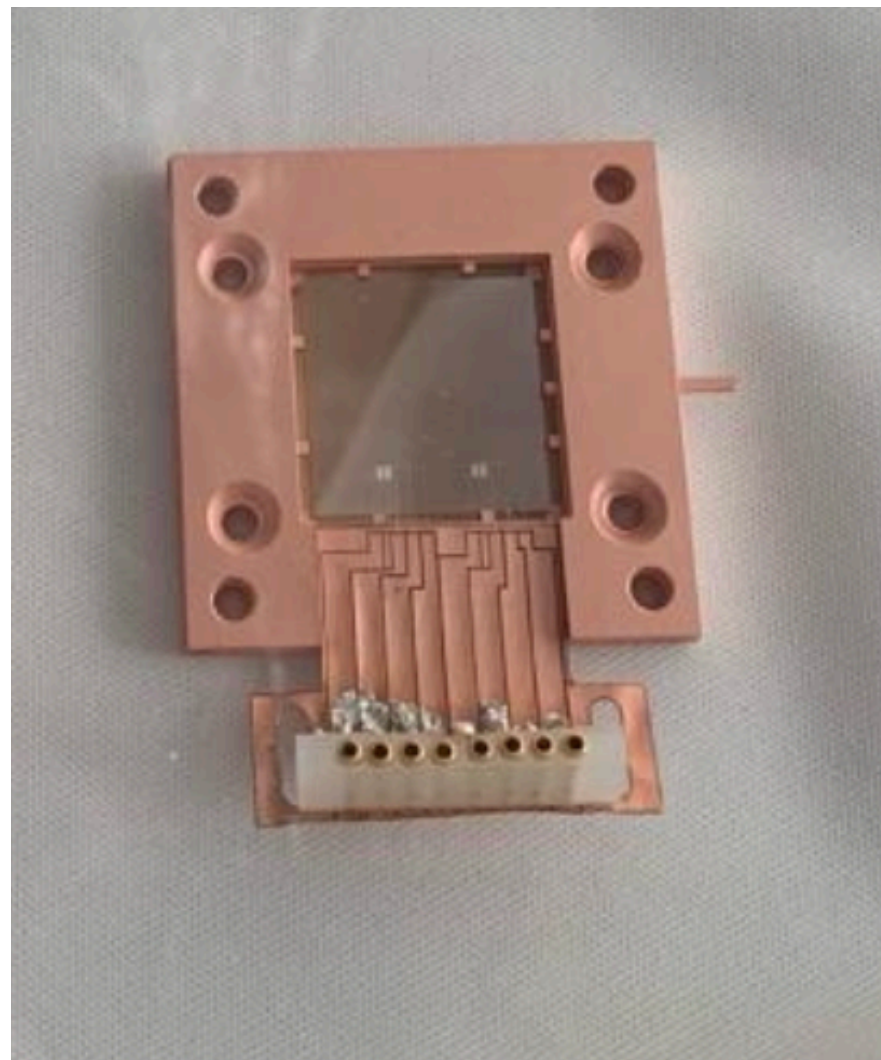


#	Baseline resolution TES1	Baseline resolution TES2
1	1.4 eV	1.1 eV
2	1.2 eV	1.0 eV
3	1.7 eV	1.5 eV
4	1.4 eV	3.9 eV



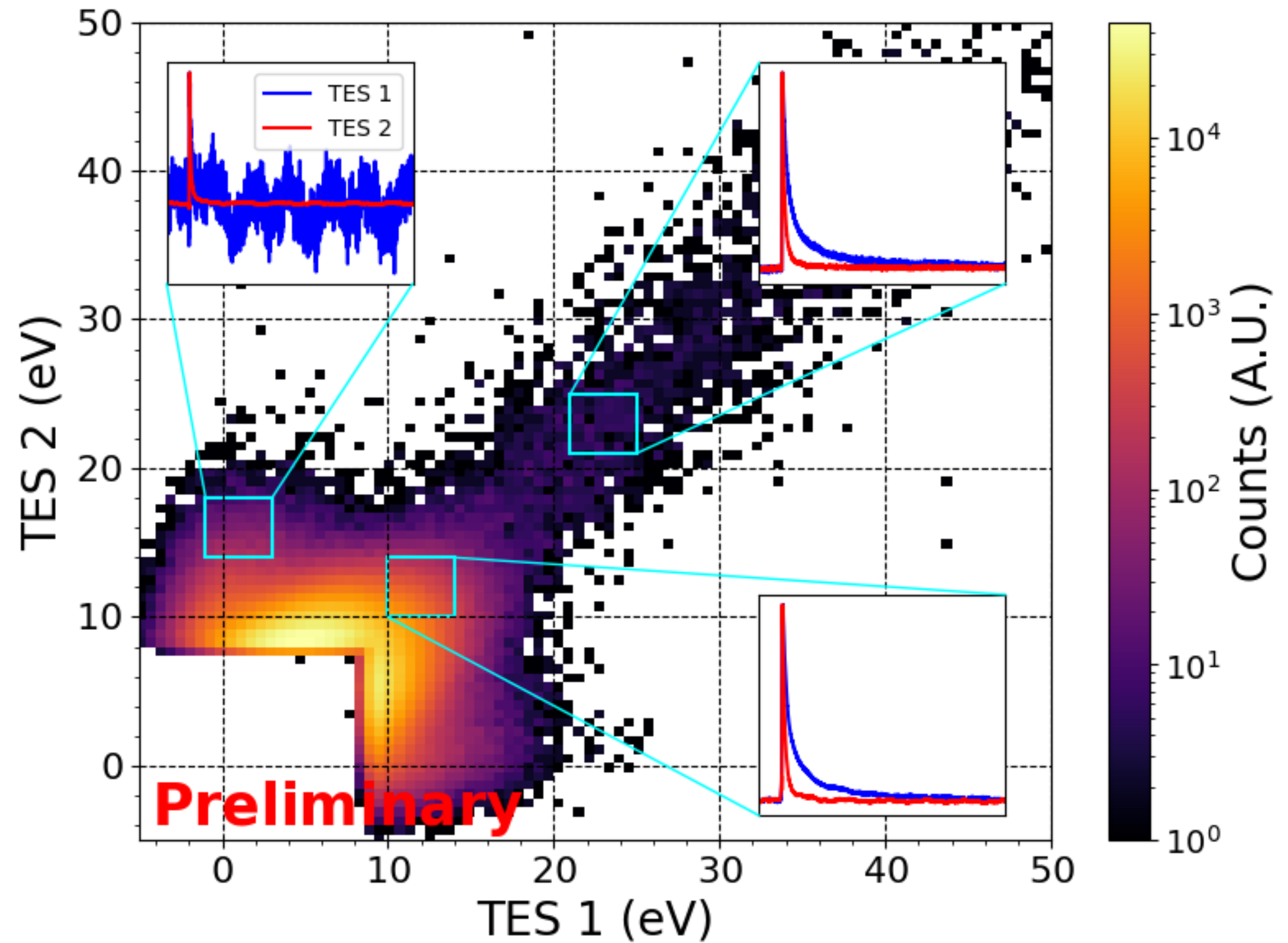
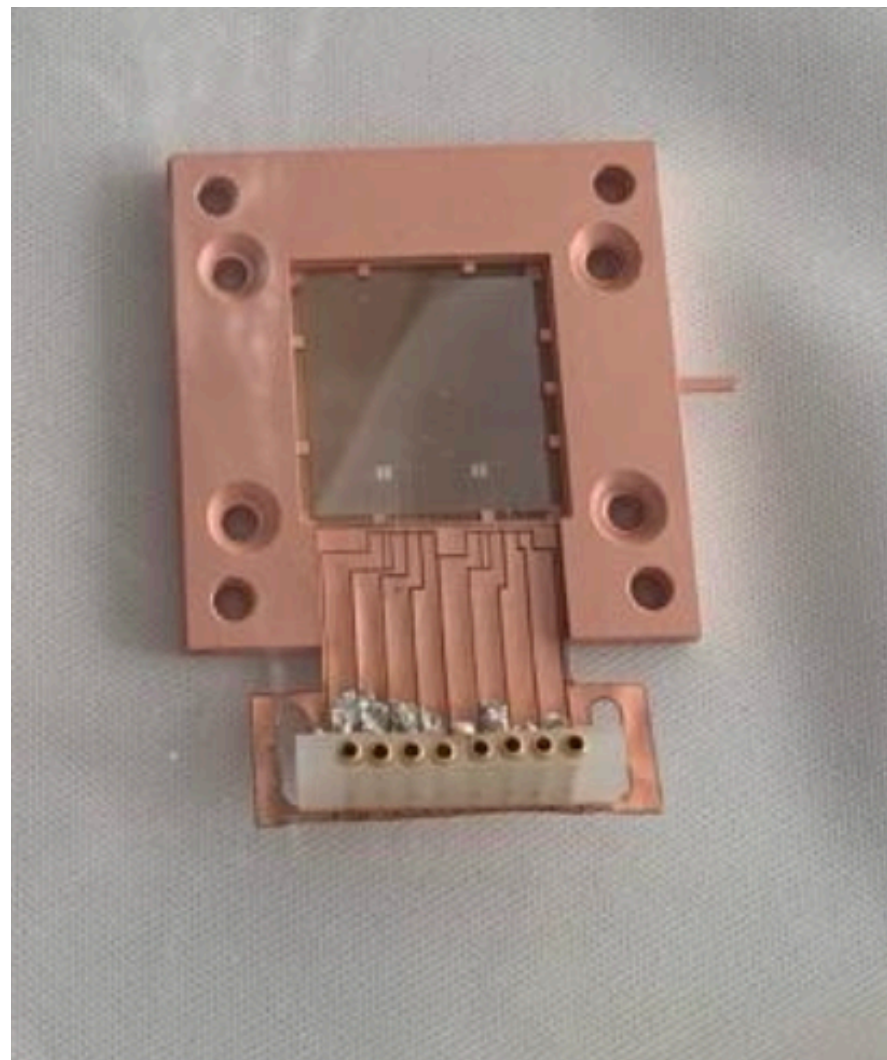
Stack 3

Analysis thresholds: 7.7 eV and 8.7 eV



Stack 3

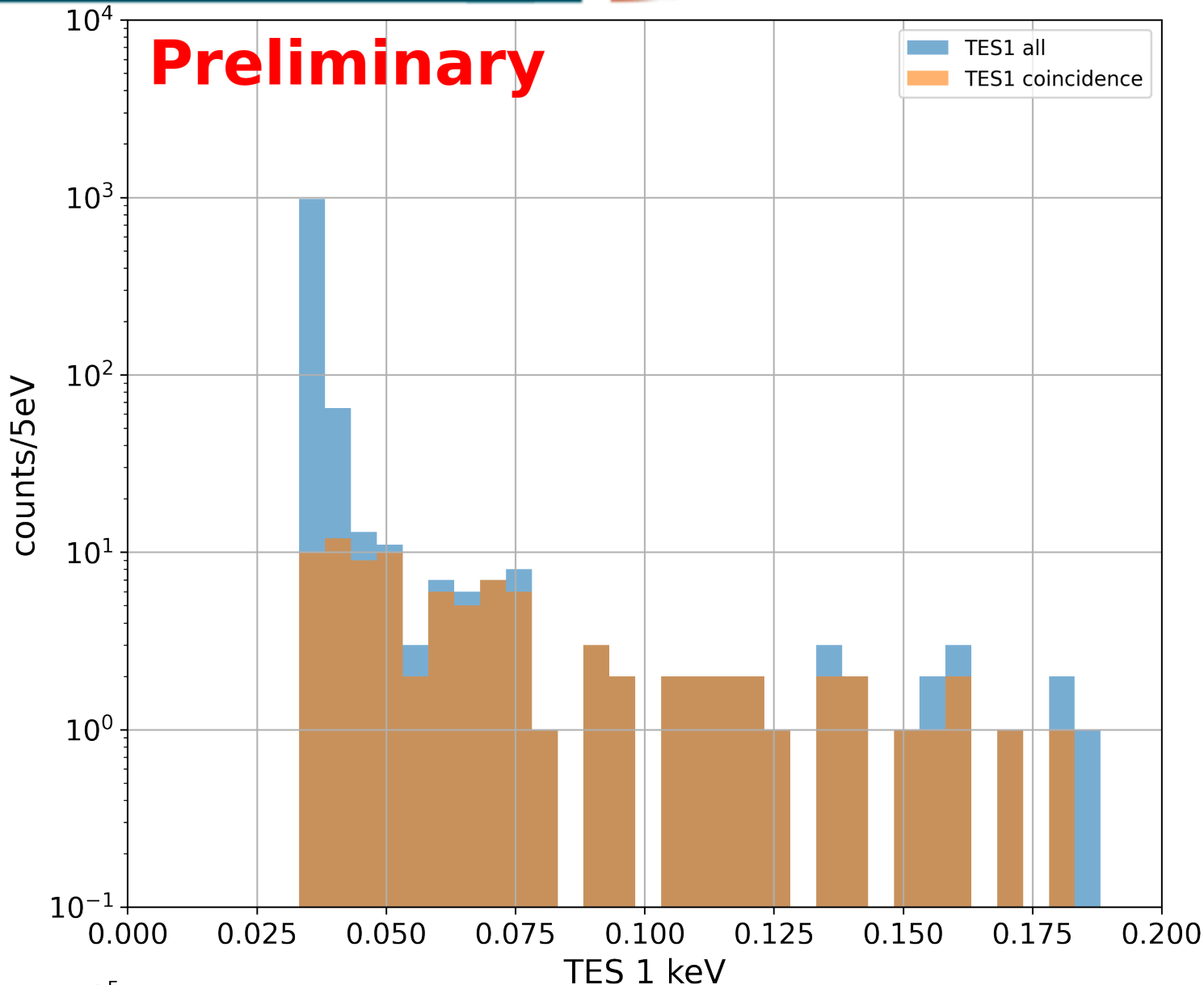
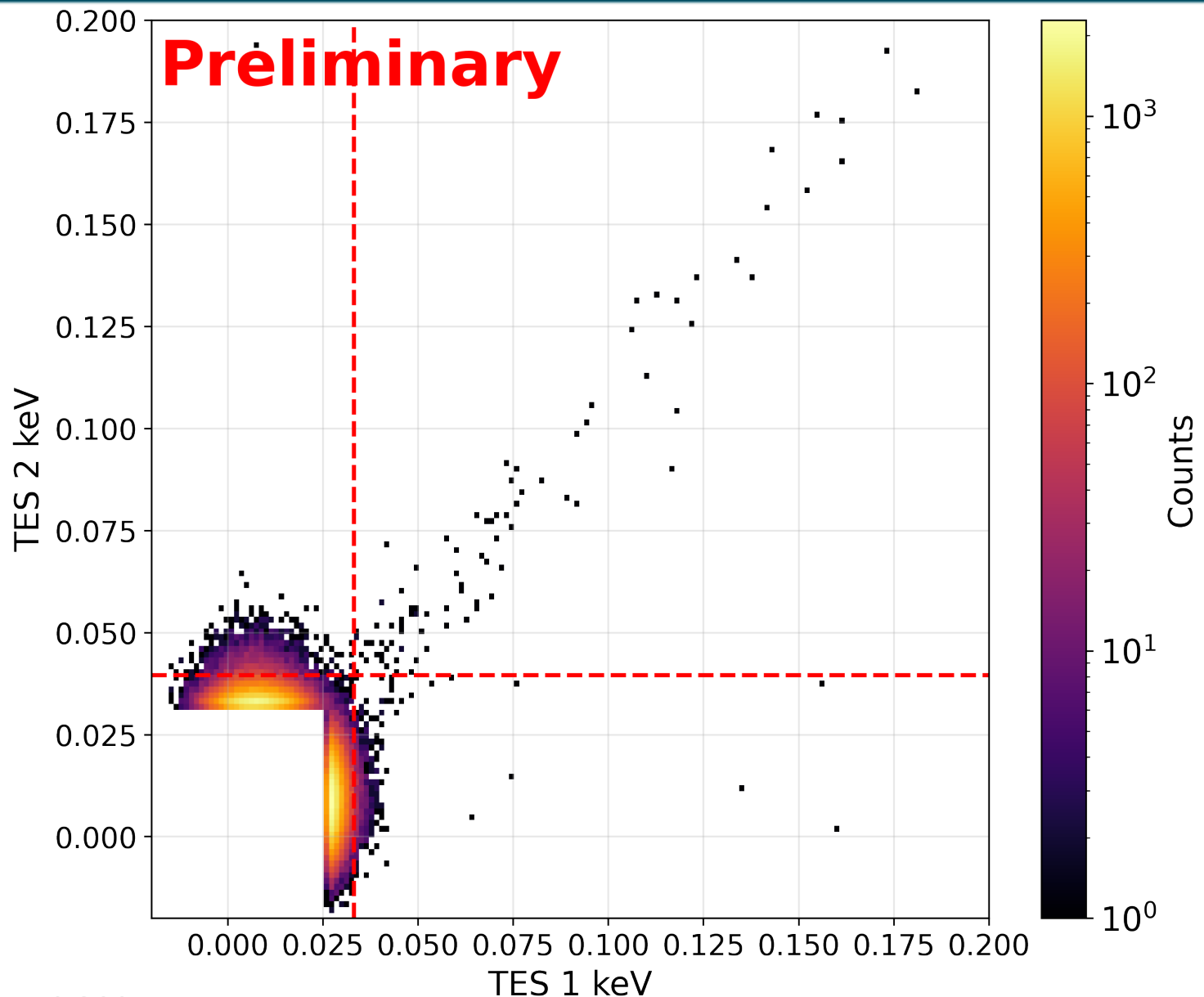
Analysis thresholds: 7.7 eV and 8.7 eV



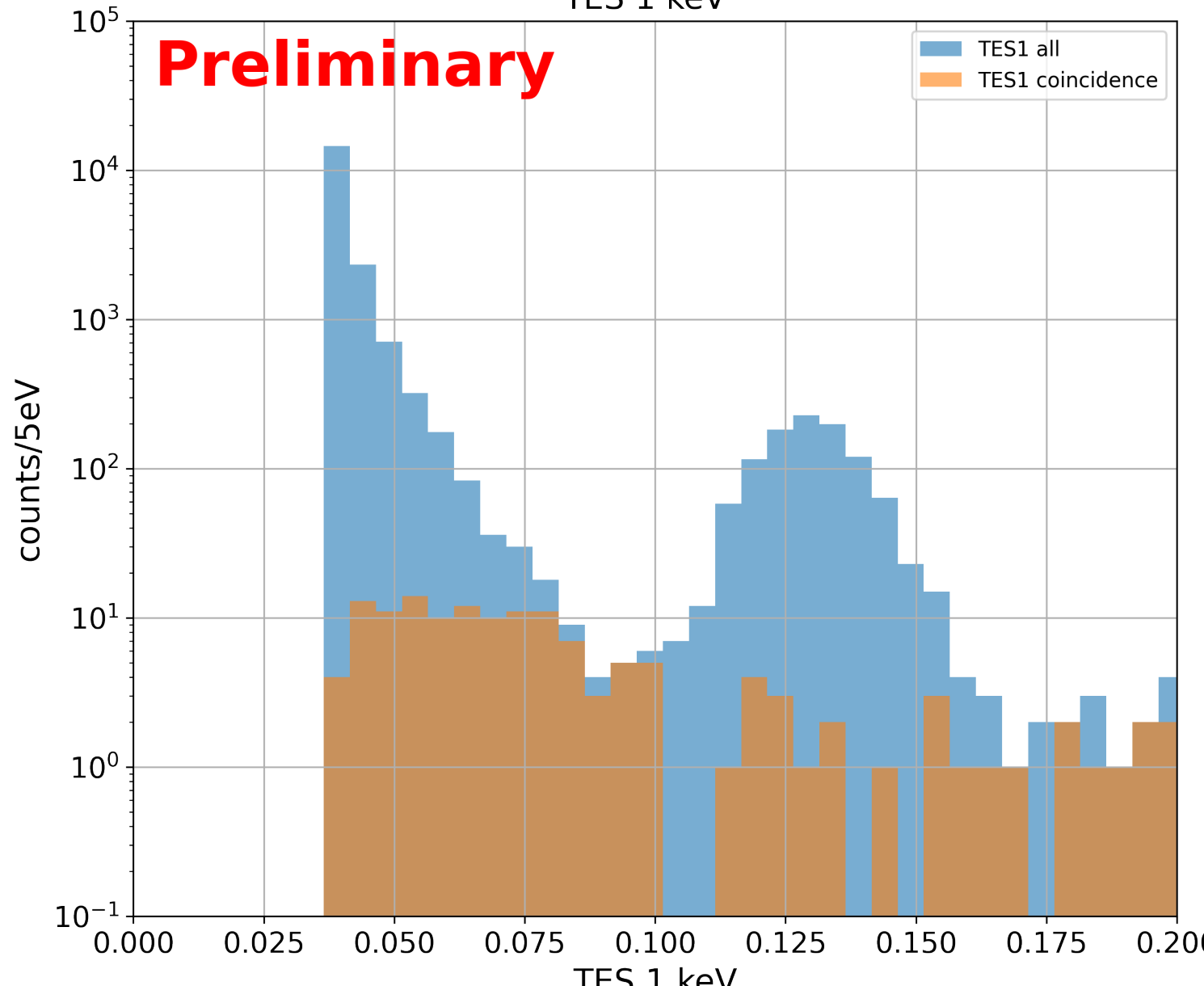
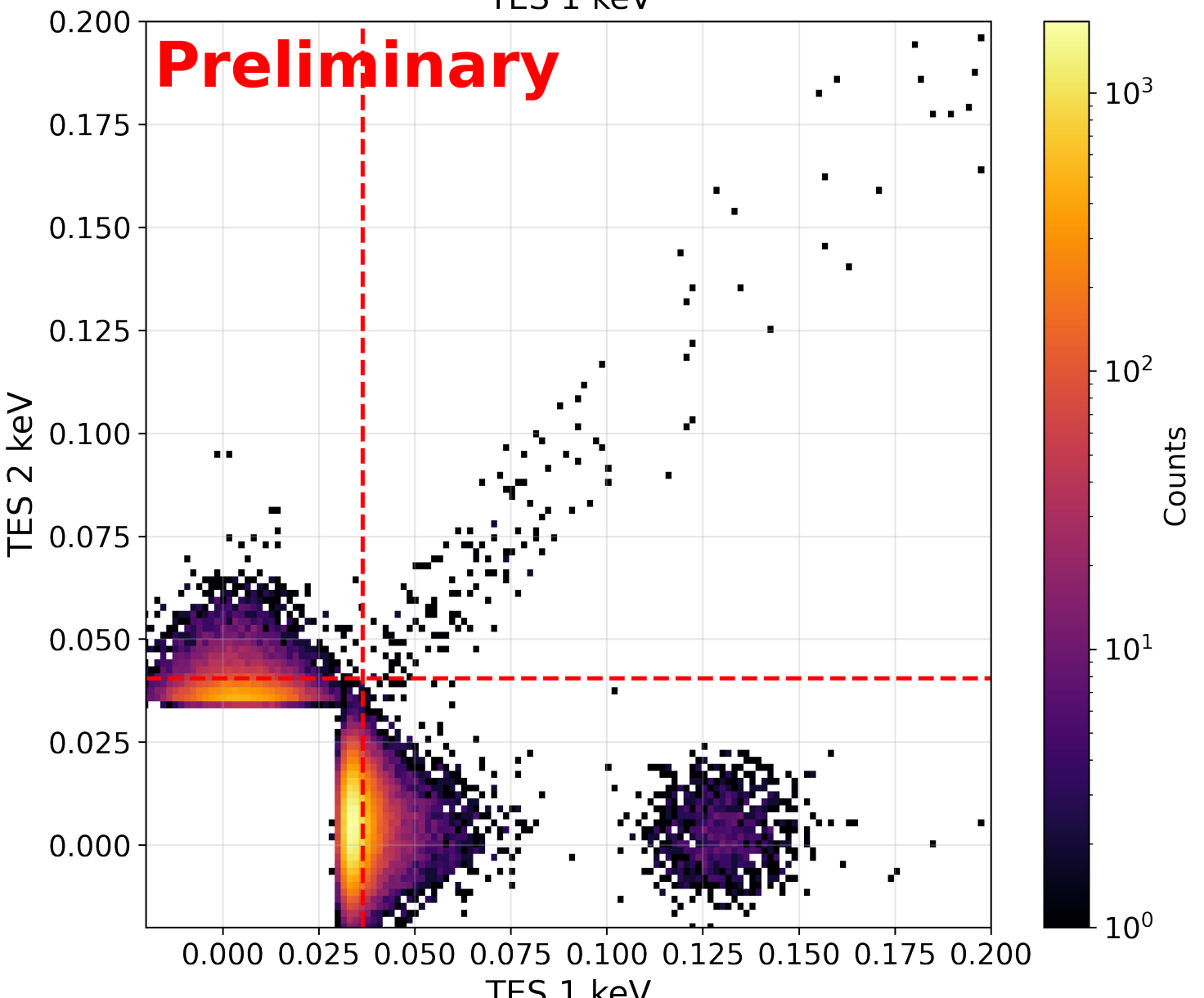
Double TES results

#	Baseline resolution TES1	Baseline resolution TES2	LD
2	6.6 eV	7.9 eV	✓
5	8.1 eV	7.3 eV	✓

DoubleTES2



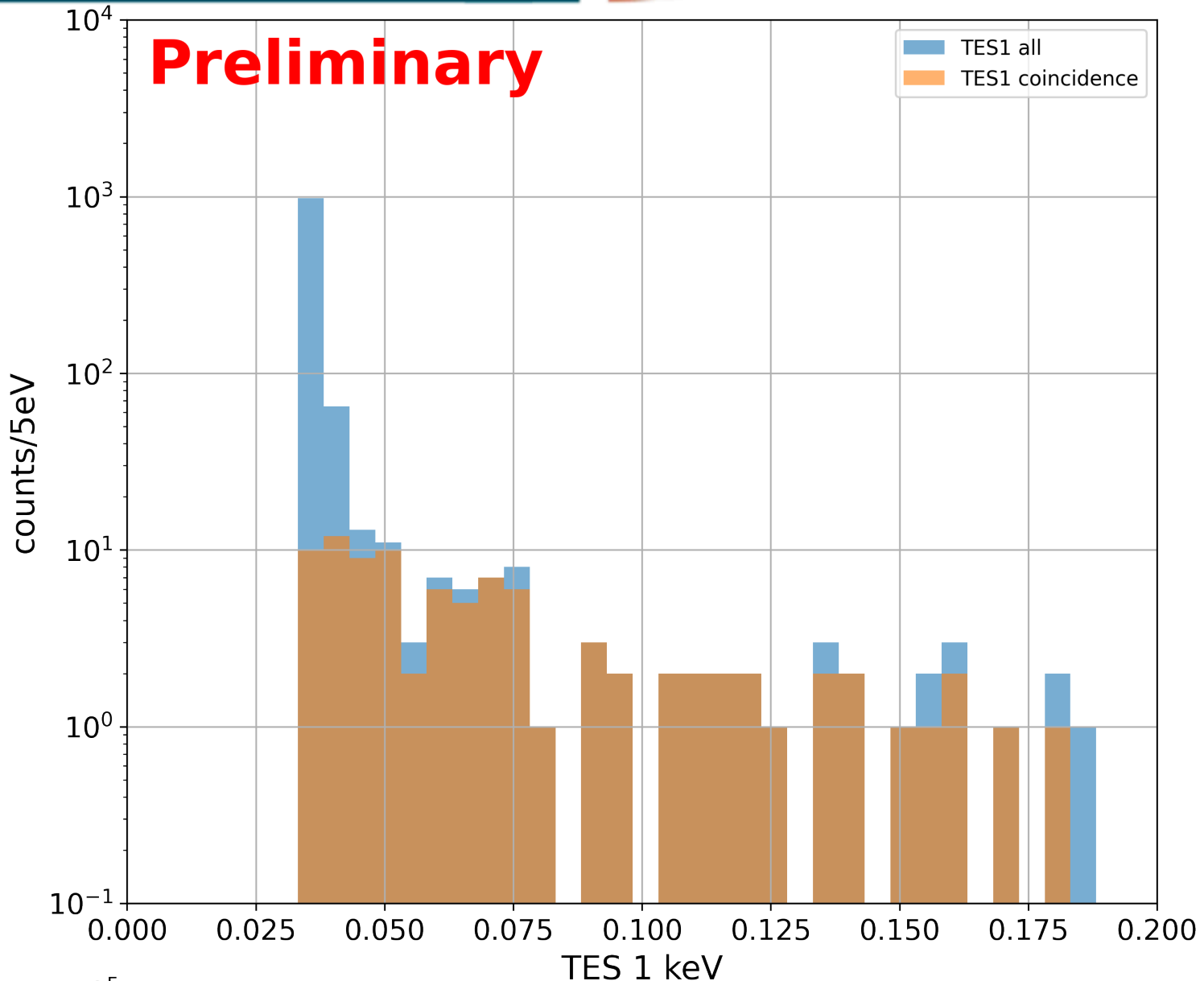
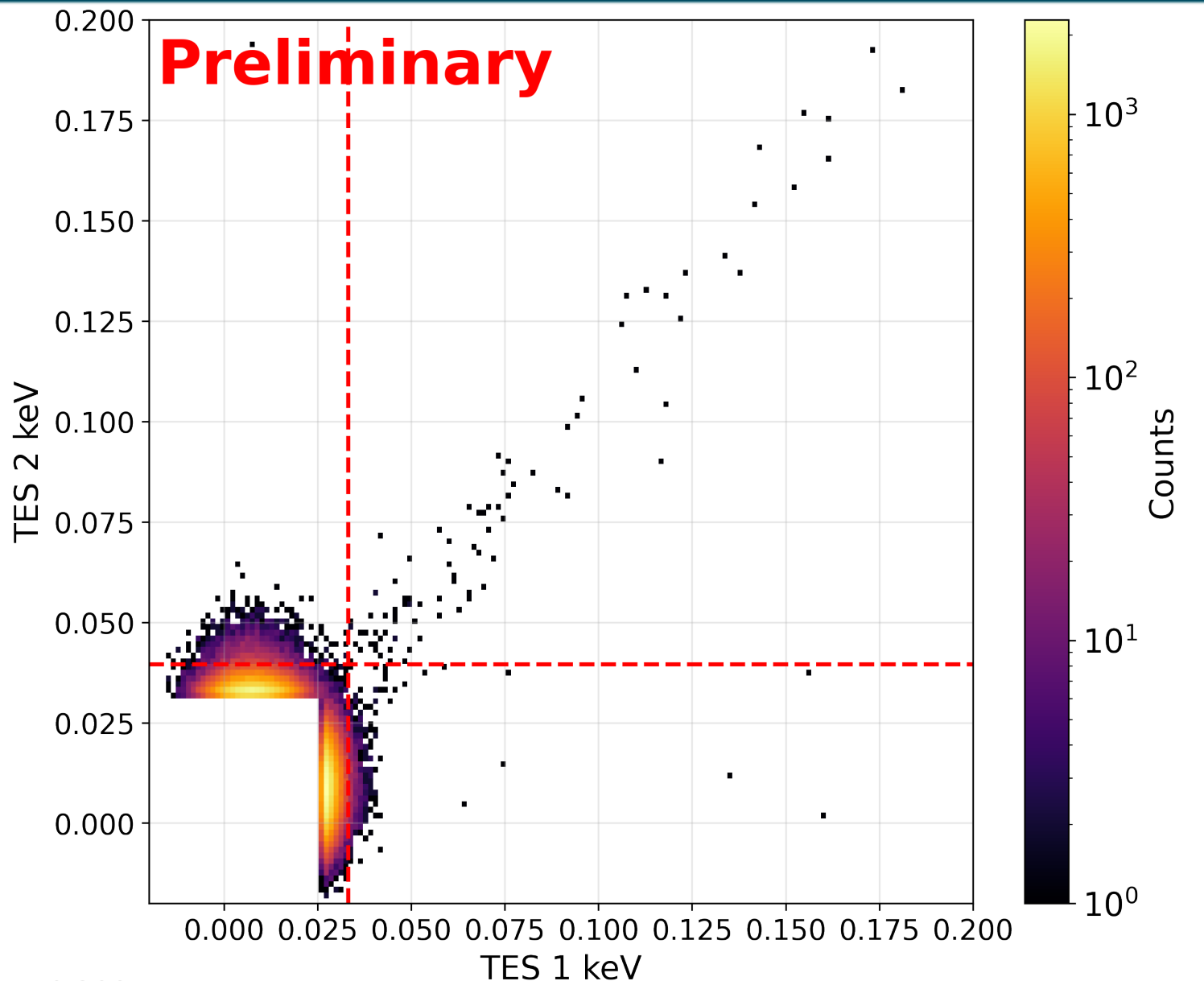
DoubleTES5



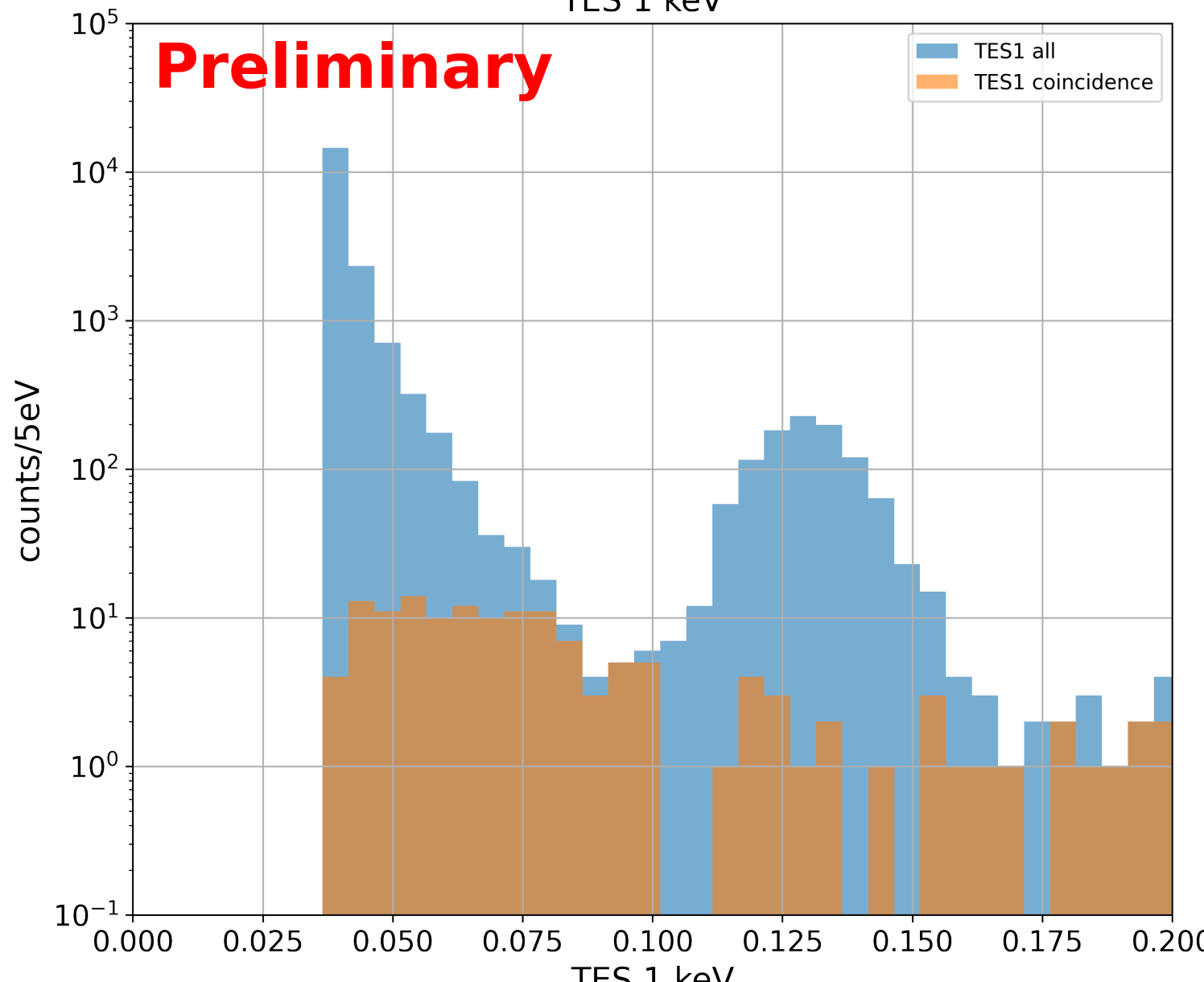
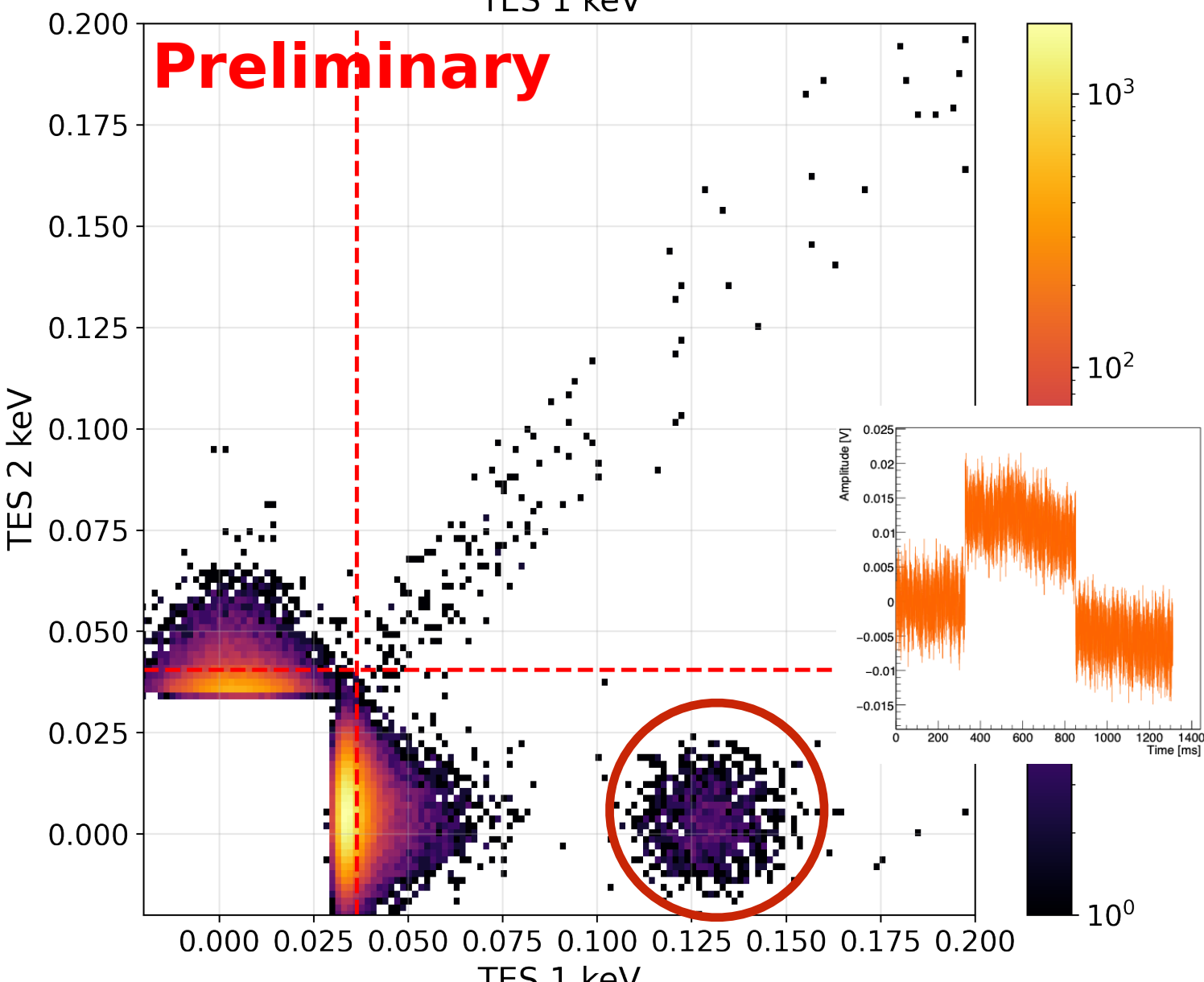
Double TES results

#	Baseline resolution TES1	Baseline resolution TES2	LD
2	6.6 eV	7.9 eV	✓
5	8.1 eV	7.3 eV	✓

DoubleTES2



DoubleTES5

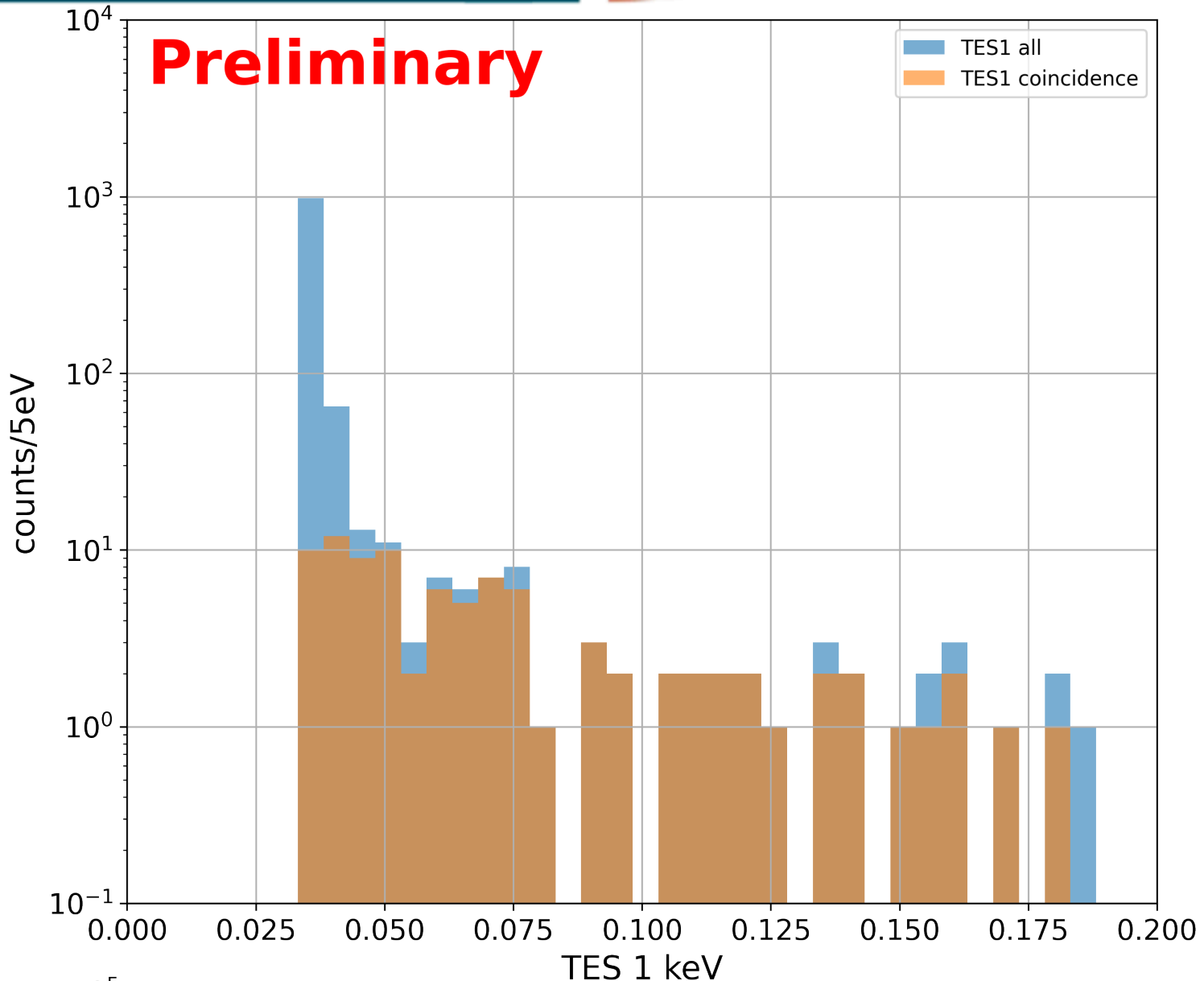
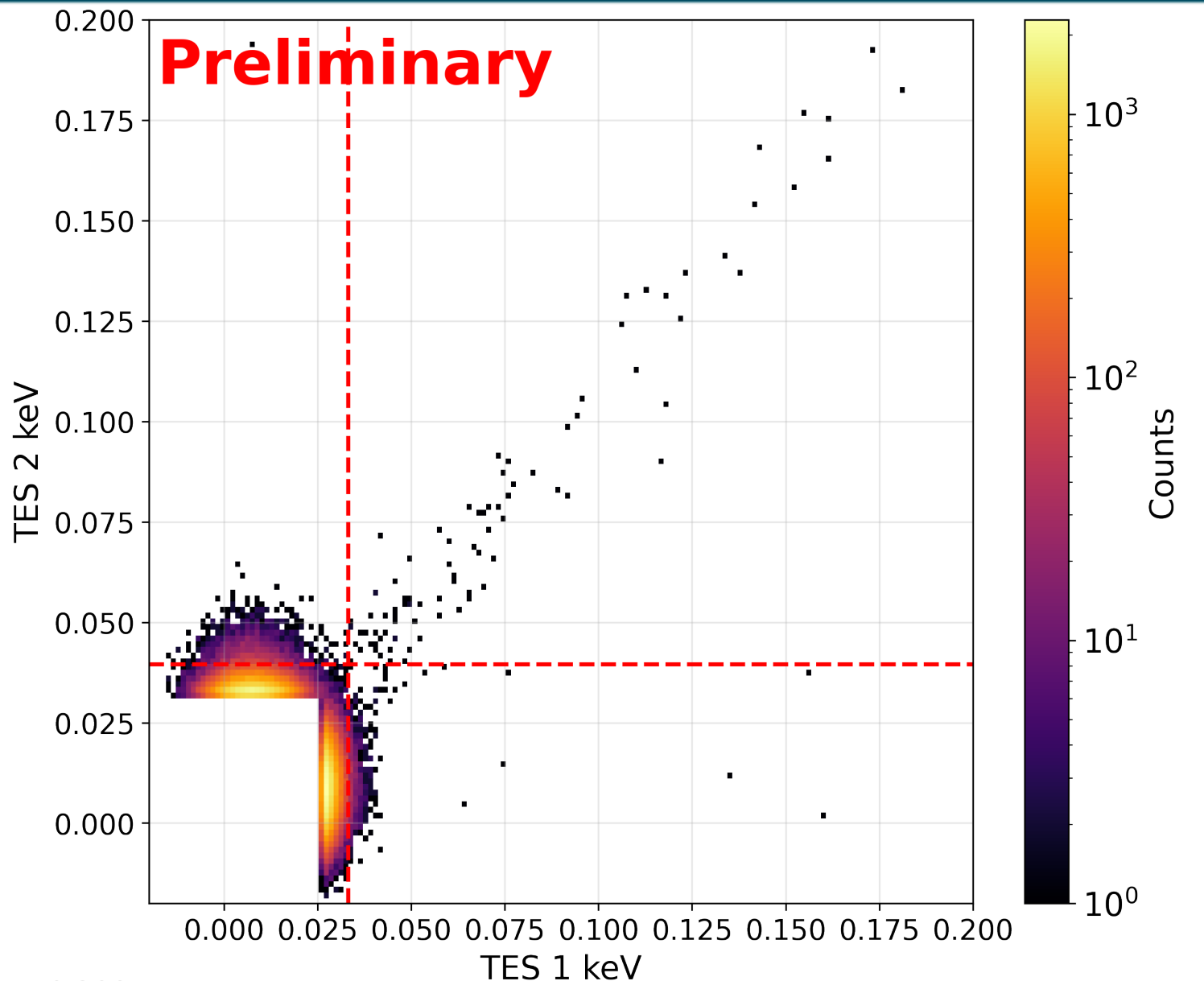


Double TES results

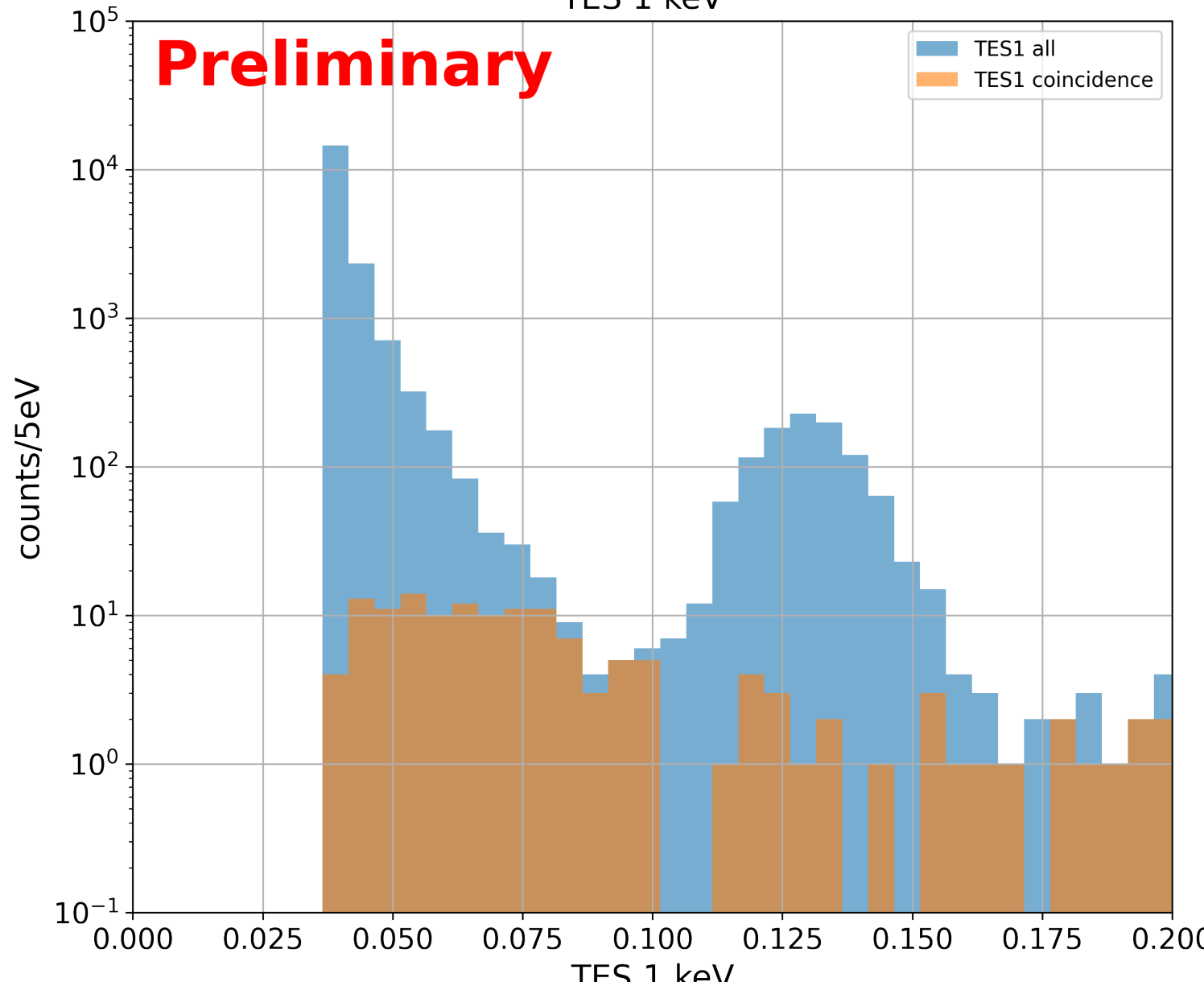
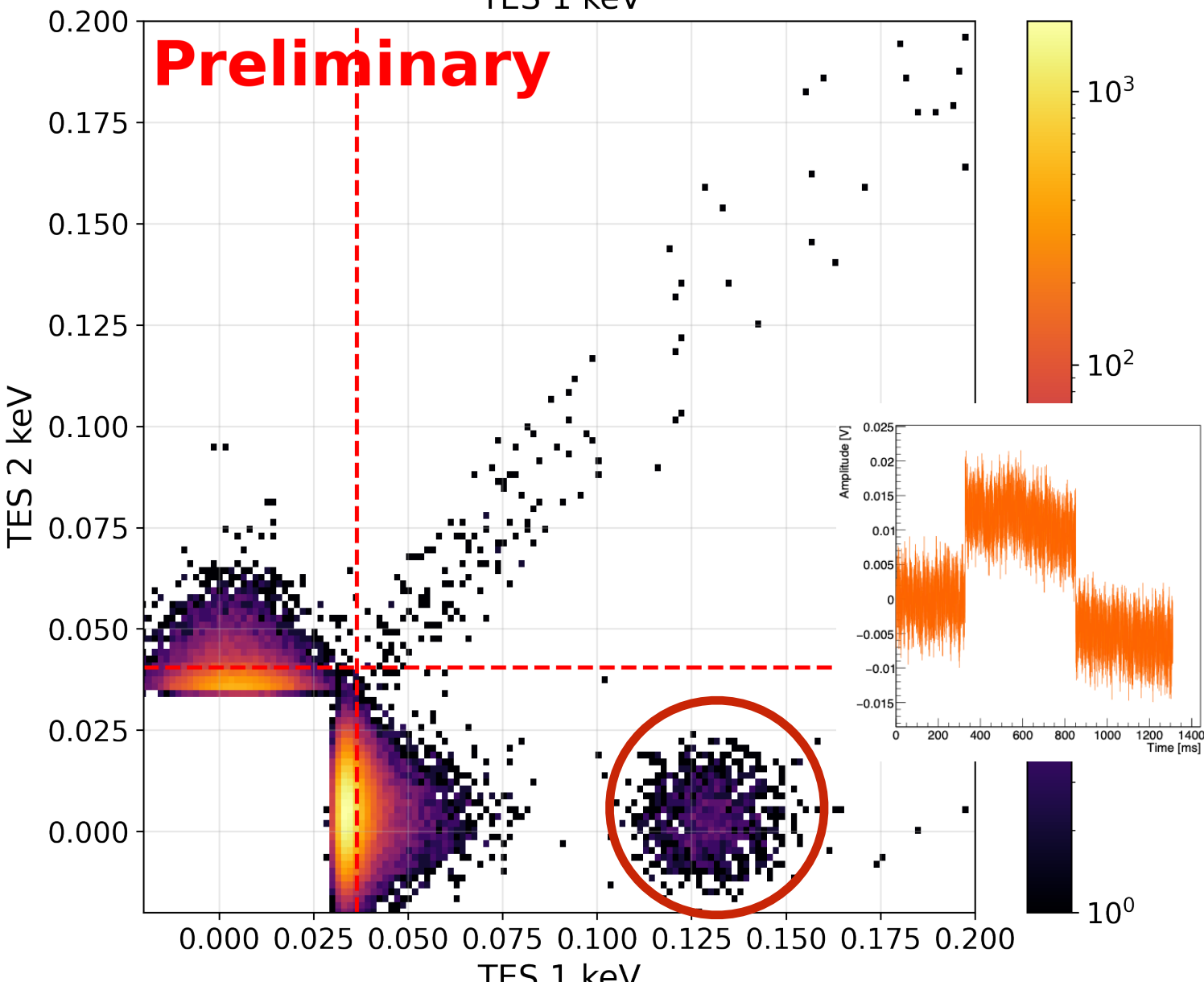
#	Baseline resolution TES1	Baseline resolution TES2	LD
2	6.6 eV	7.9 eV	✓
5	8.1 eV	7.3 eV	✓

Strong indication of LEE reduction

DoubleTES2

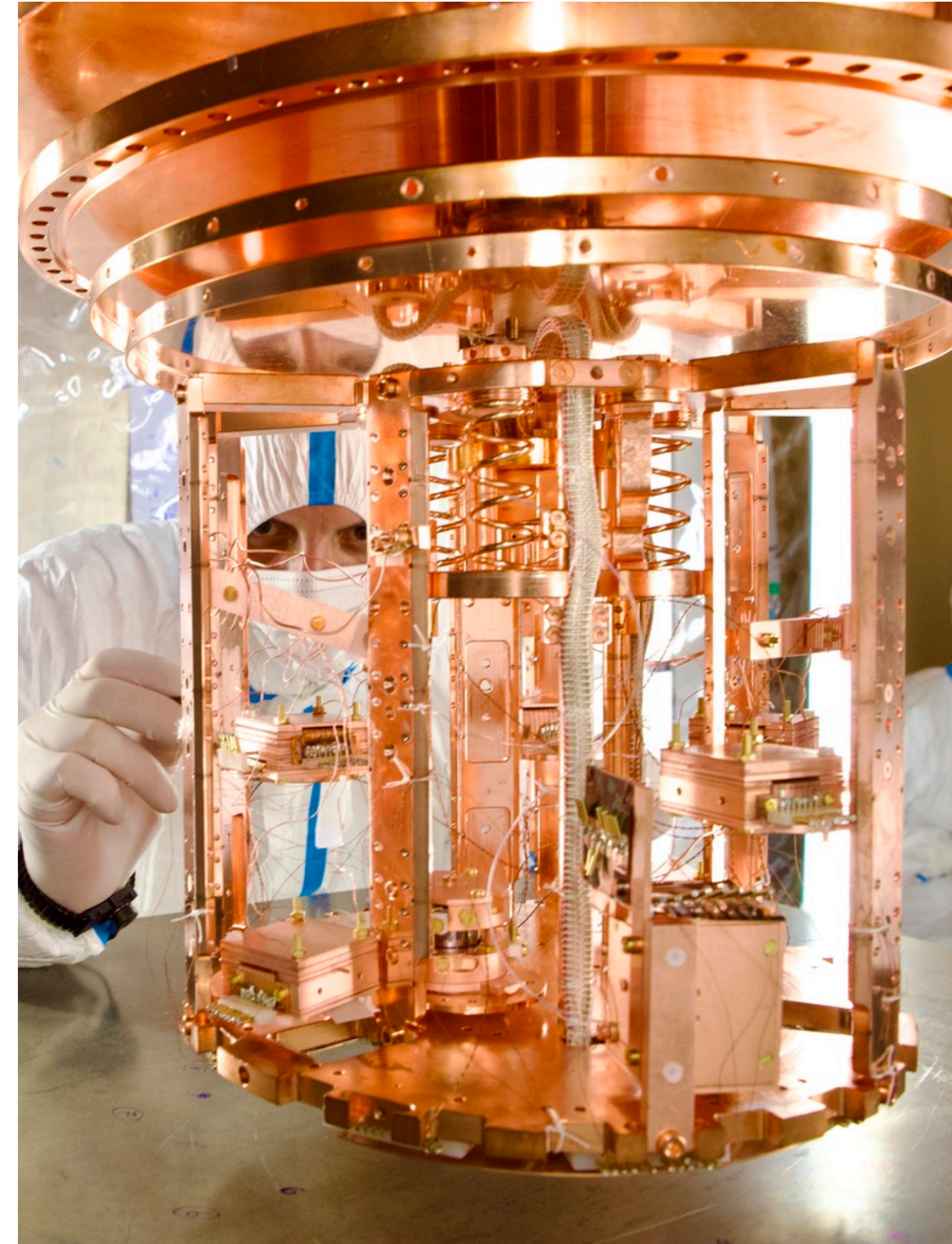


DoubleTES5

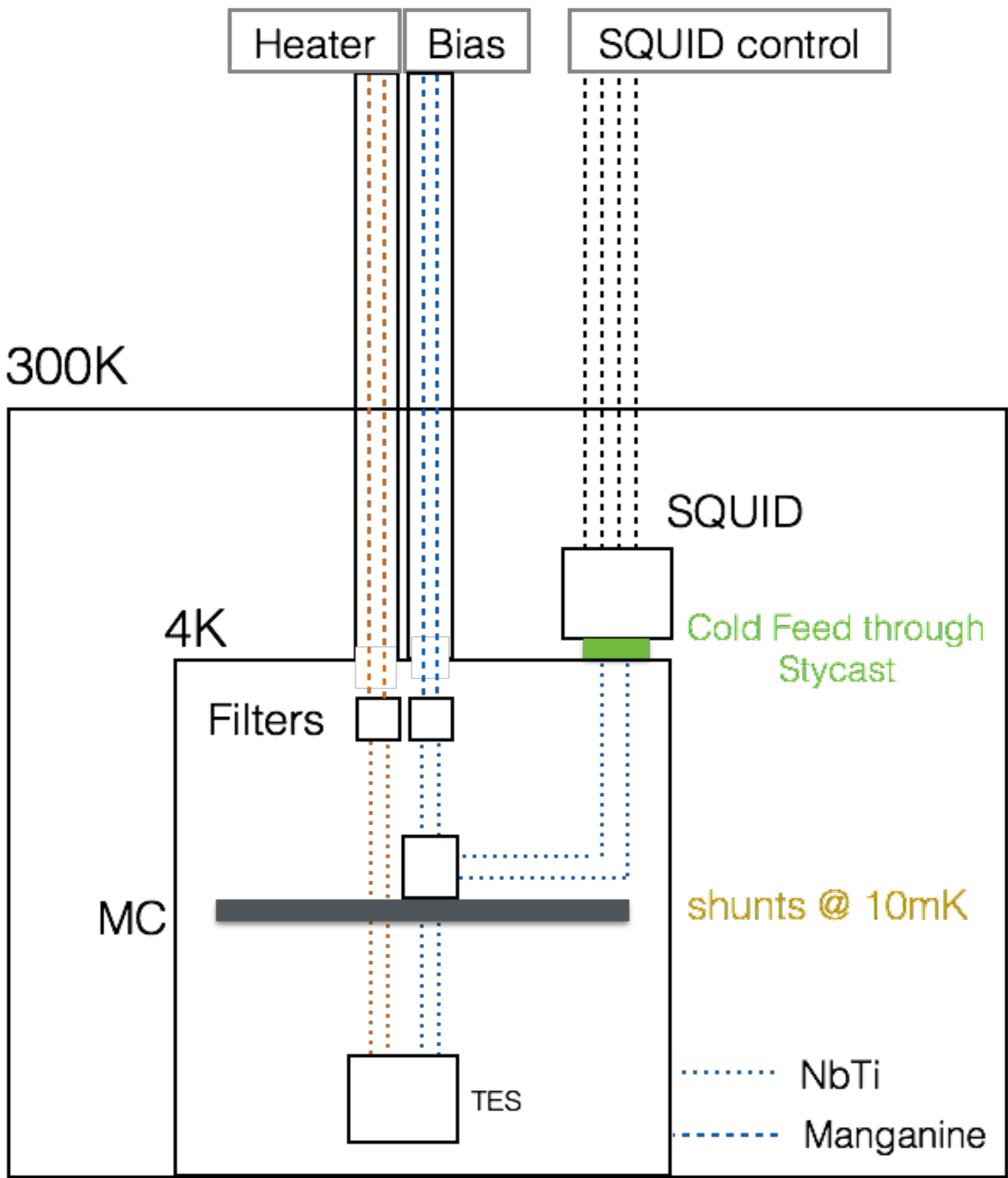
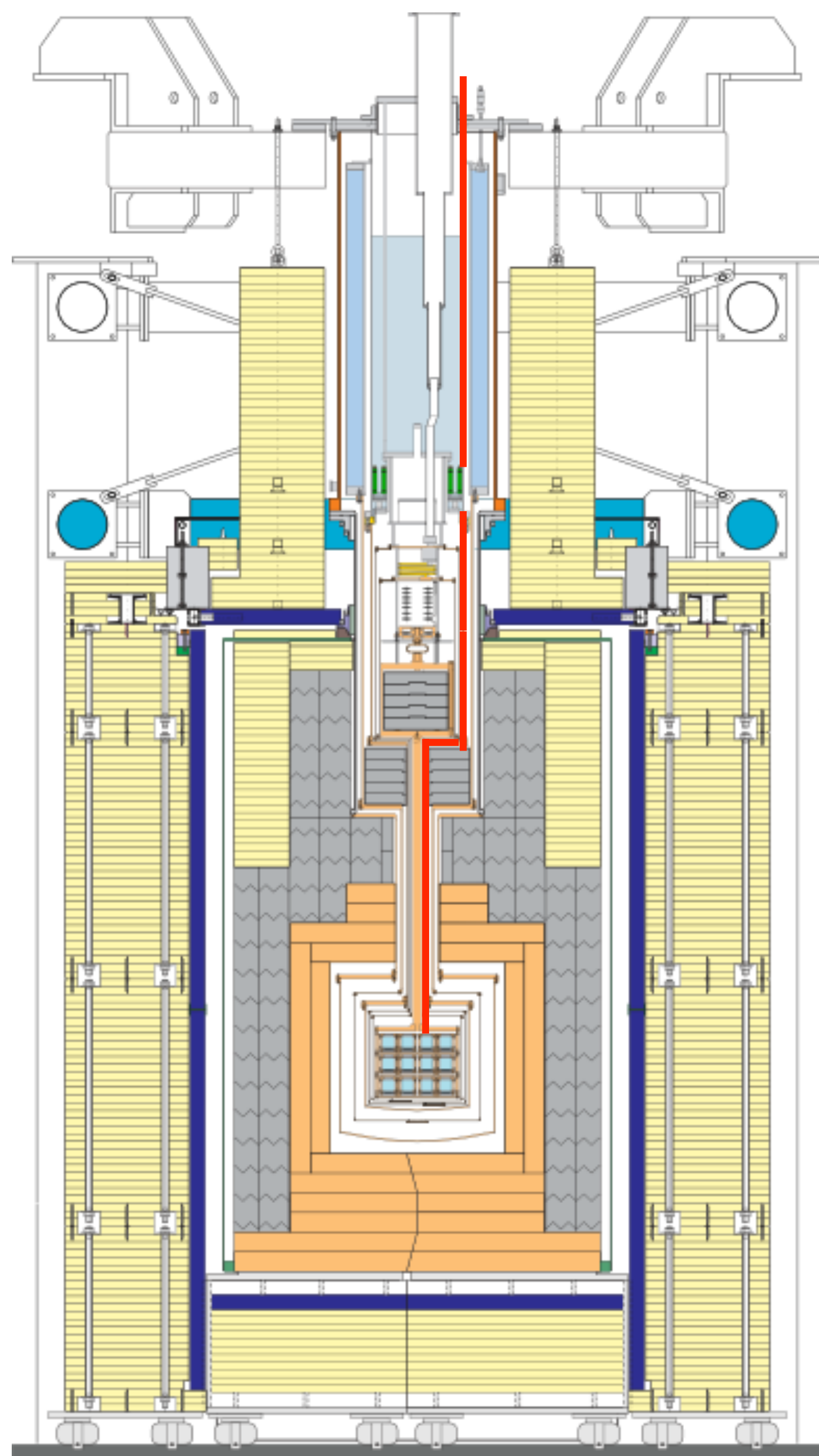


Conclusion

- CRESST energy thresholds $O(30-10\text{eV})$ allow for deep exploration of light DM mass region, extending sensitivity in the $\sim 100\text{ MeV}$ range for the DM mass
- CRESST's capability of using different target crystals allows investigation of both SI (Si , CaWO_4) and SD (LiAlO_2 , Al_2O_3) interactions
- Calibration from recoils induced by radiative capture of thermal neutrons allows calibration of CaWO_4 detectors without introducing contaminants in the CRESST setup
- Double TES preliminary results show, for the first time, a relevant reduction of LEE events. An analysis is ongoing to understand the behaviour of the different populations.
- Stay tuned for upcoming DM results!



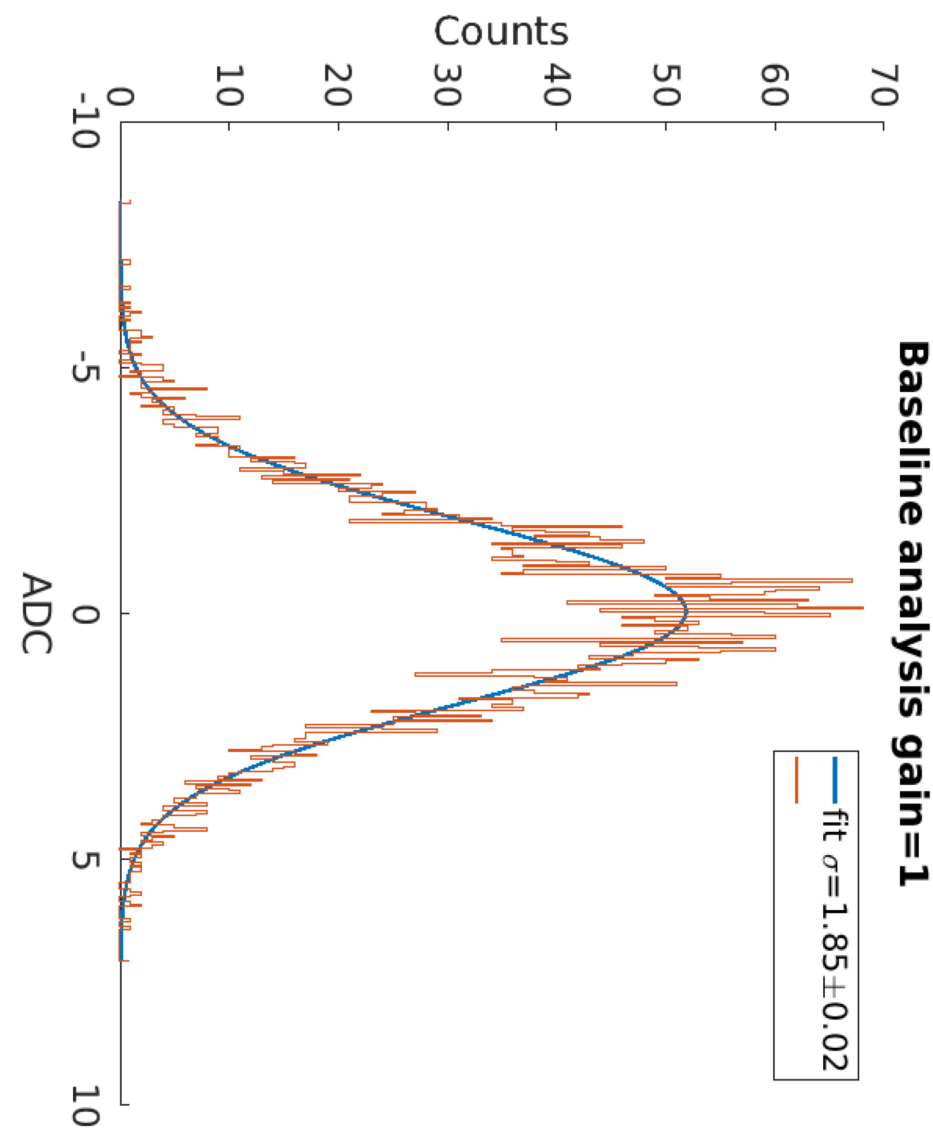
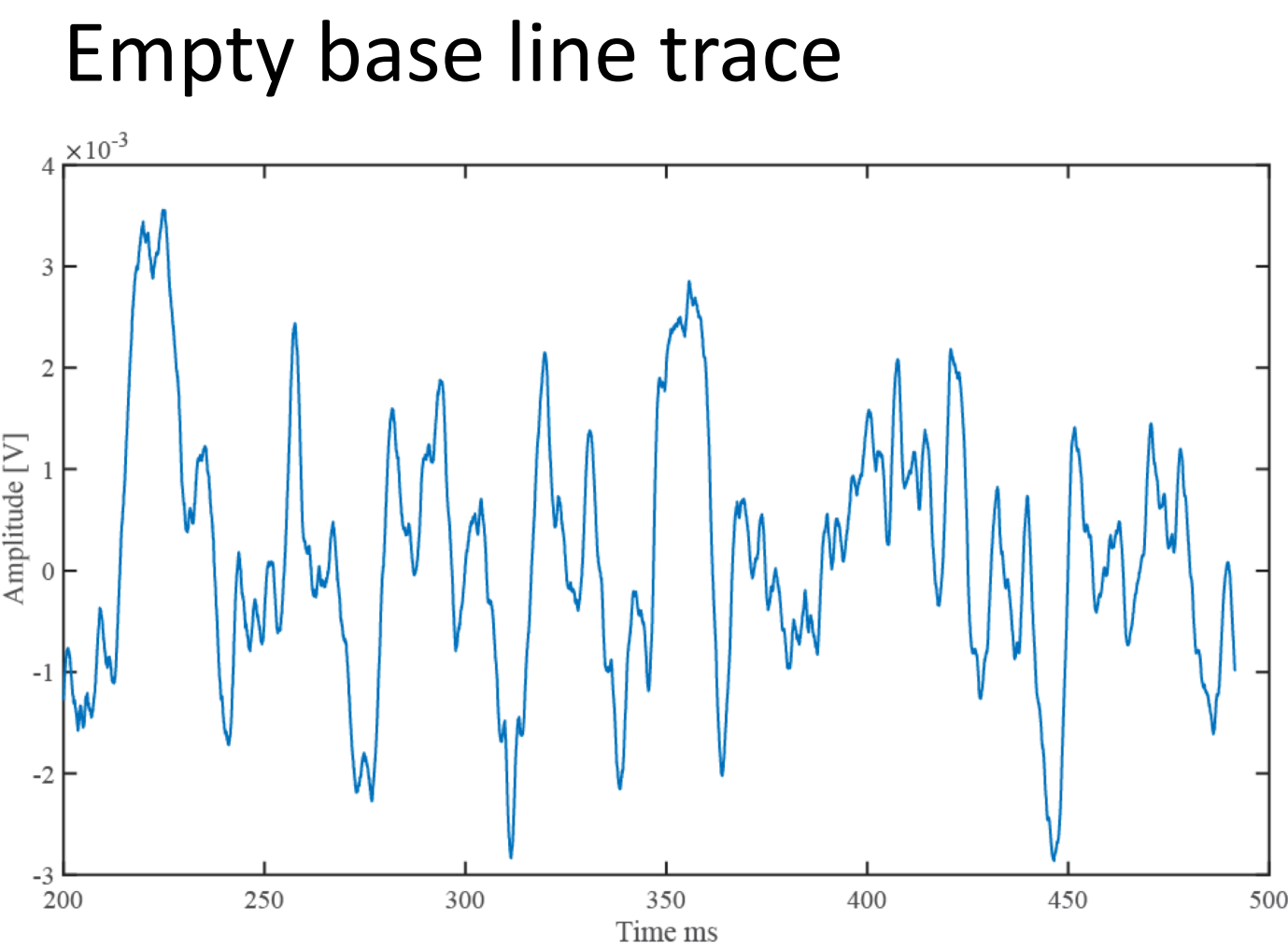
Backup slides



CRESST upgrade to 100 modules (288 readout channels) already in the procurement phase. Cryogenic infrastructure will remain mostly the same while readout chain is upgraded.

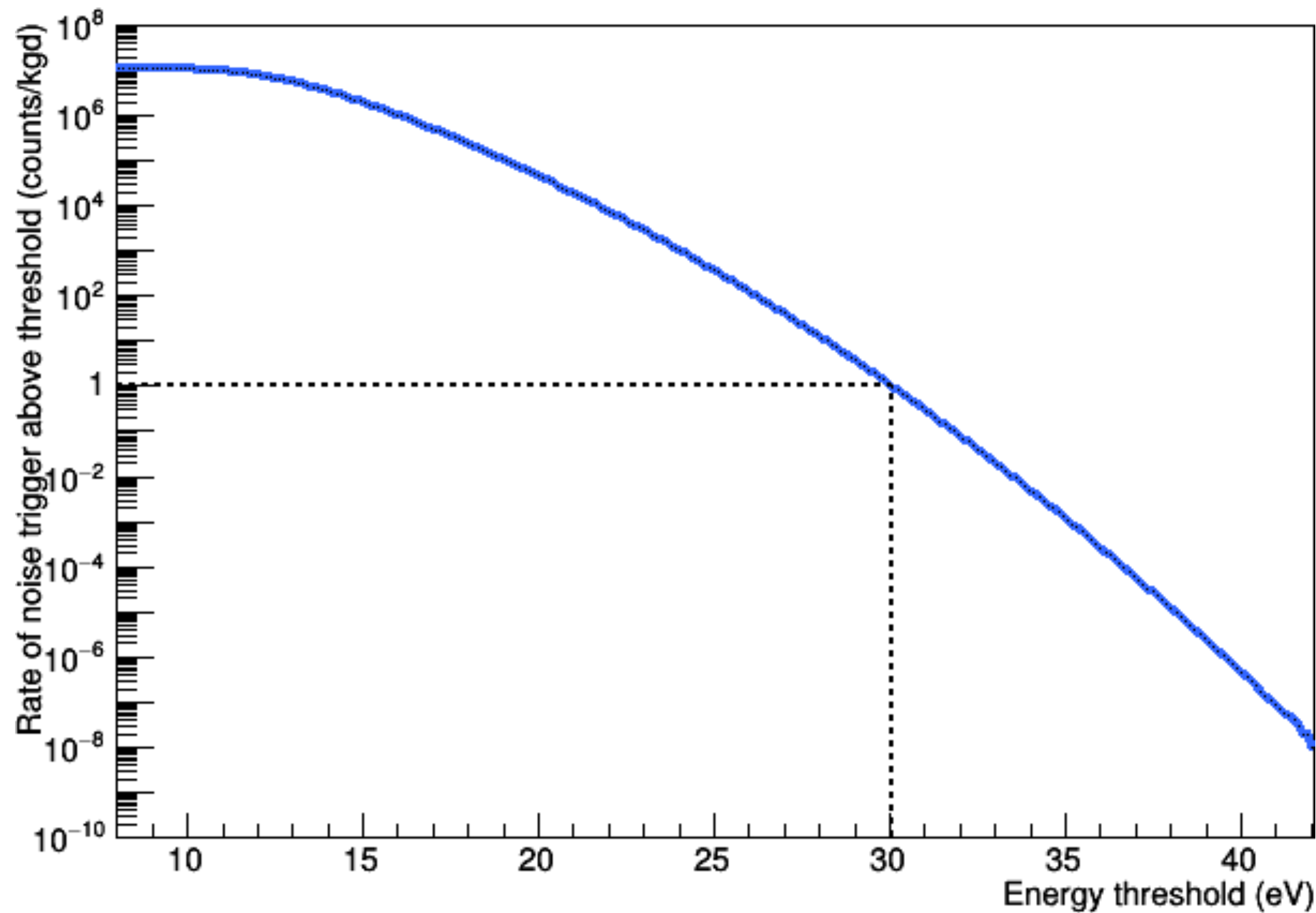
OPTIMUM TRIGGER THRESHOLD

Optimum filter for threshold analysis



Histogram of a typical
baseline trace

- Study the noise distribution after optimum filter in order to set the threshold
- Threshold optimised based on noise triggers in a given exposure

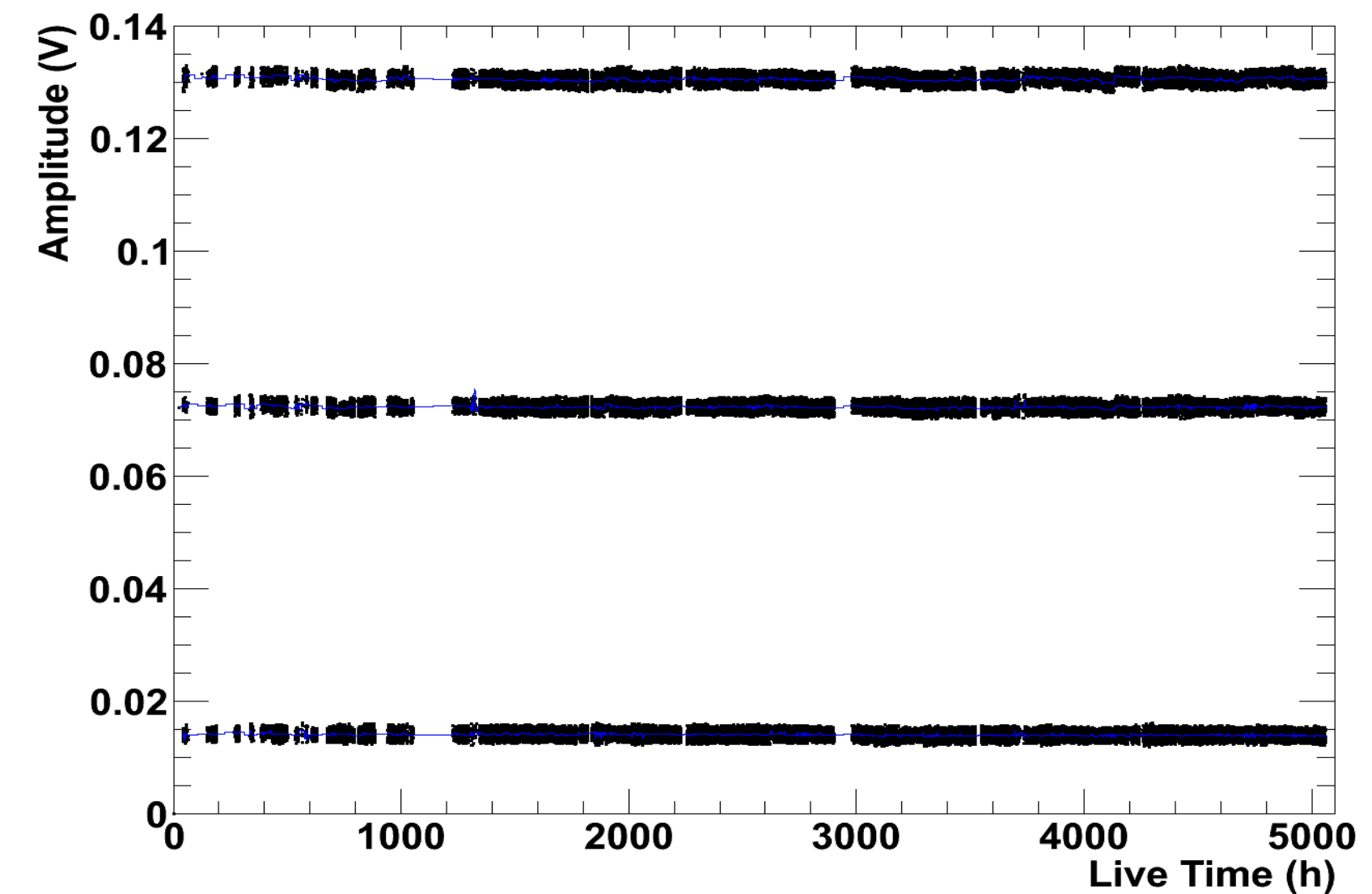
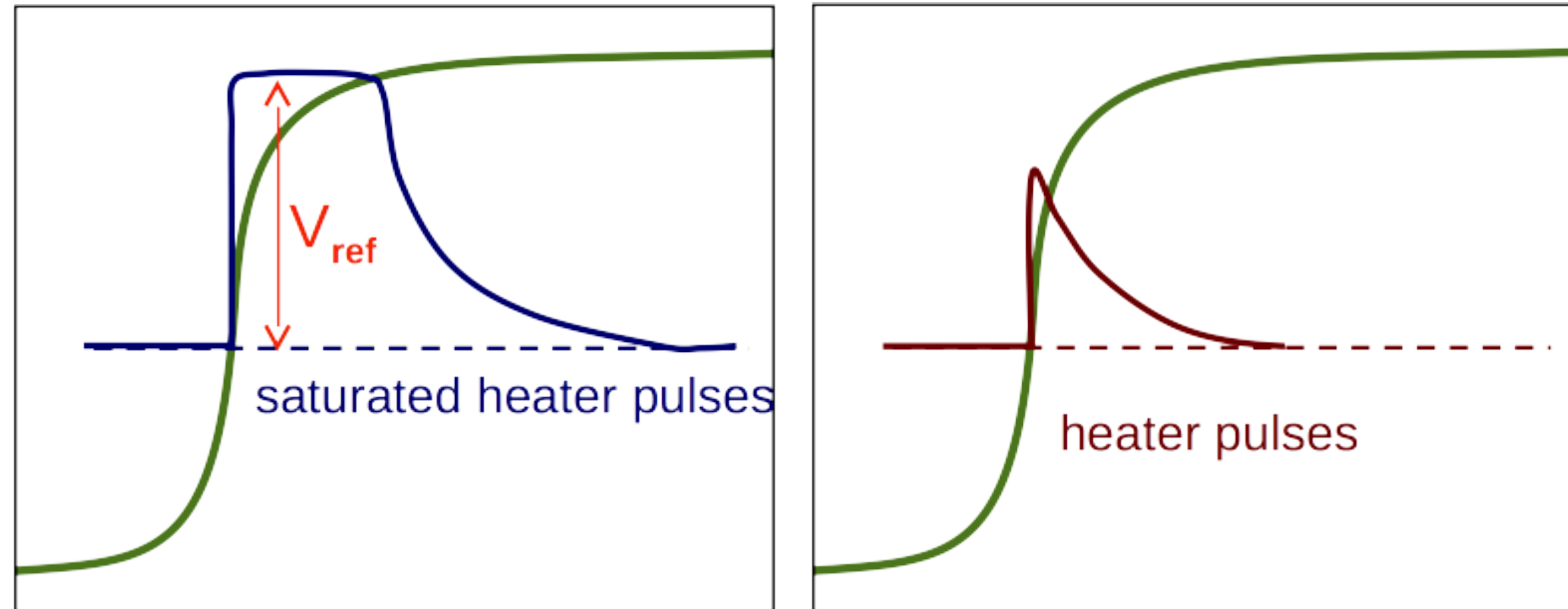
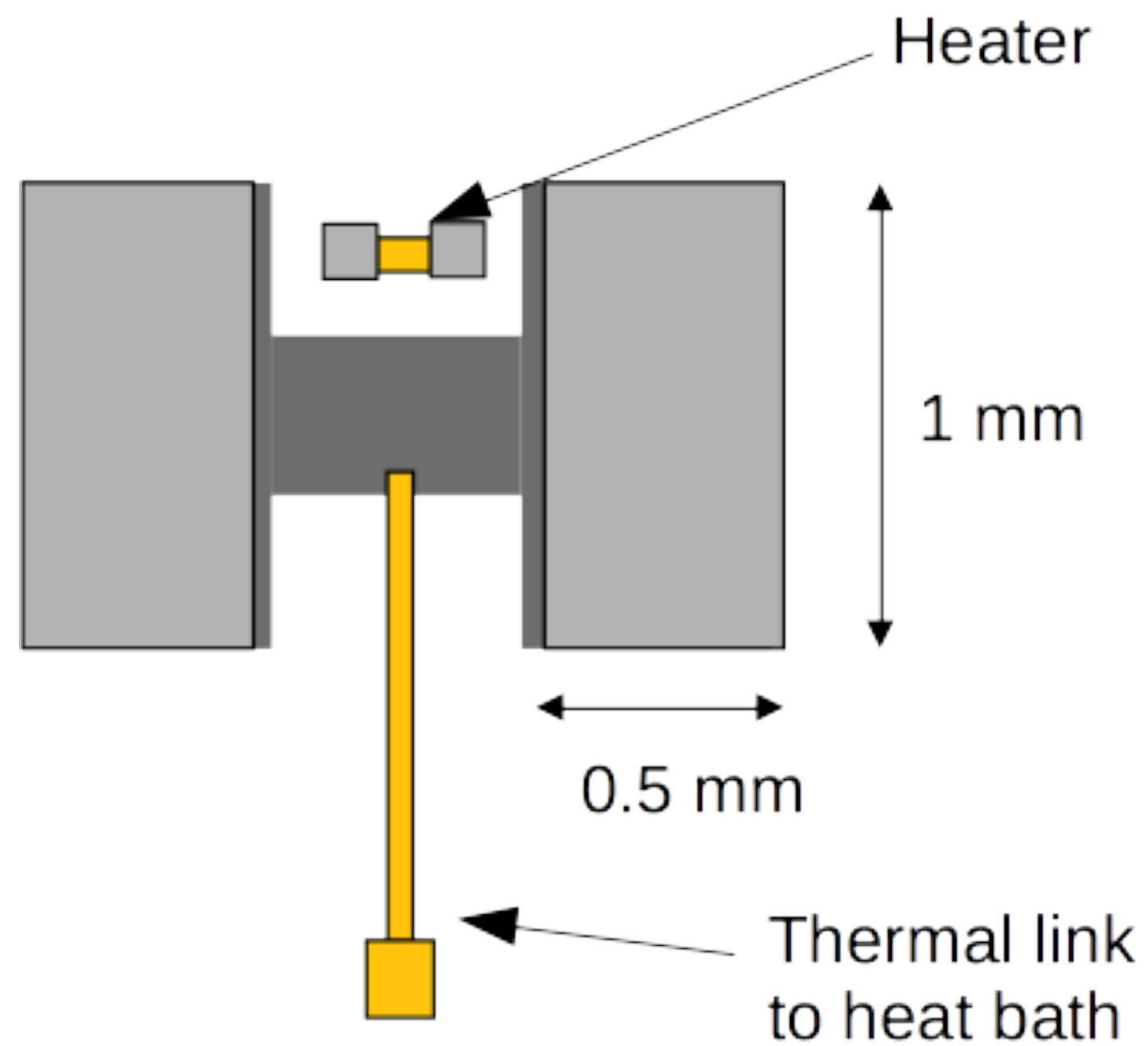


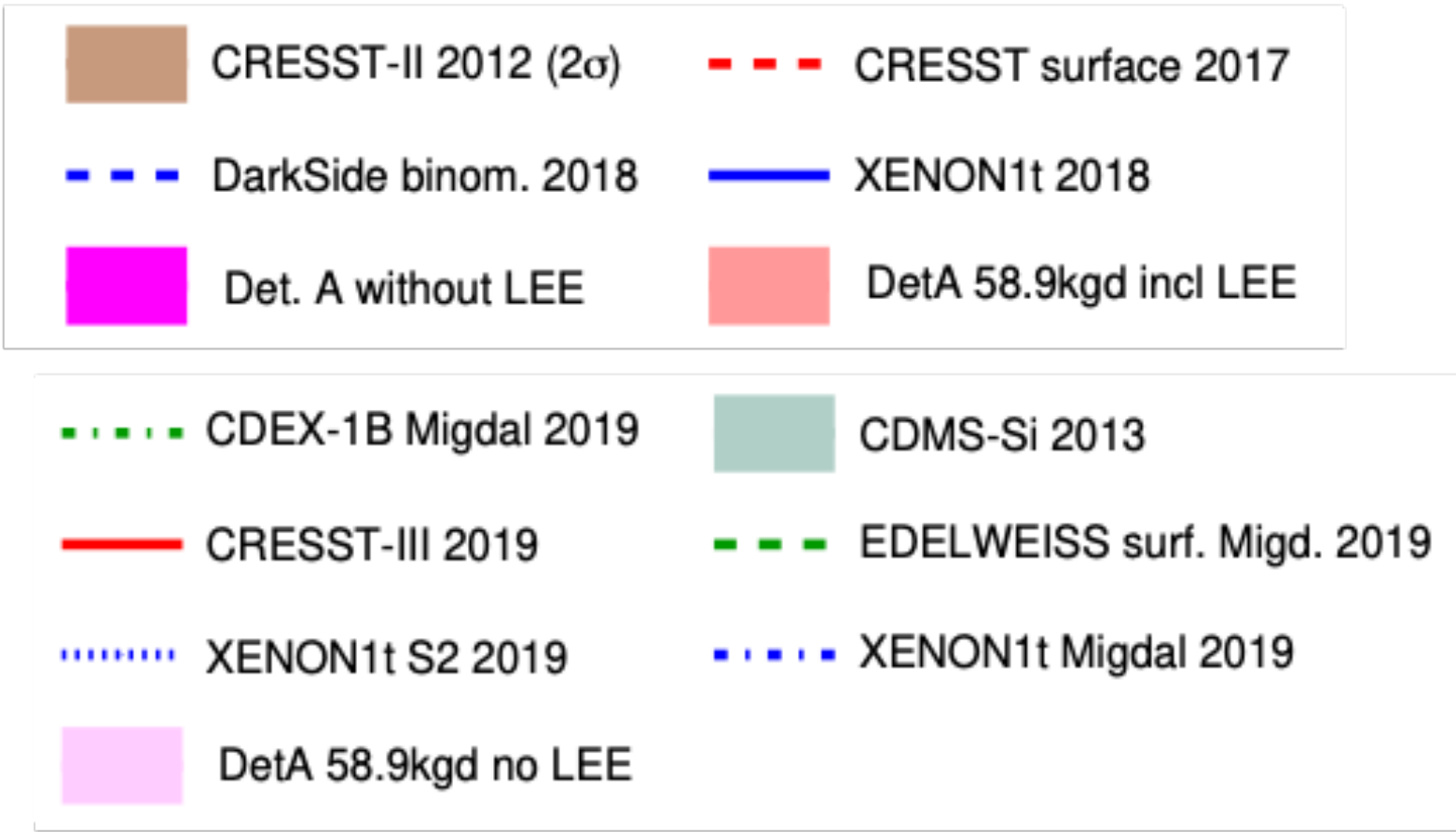
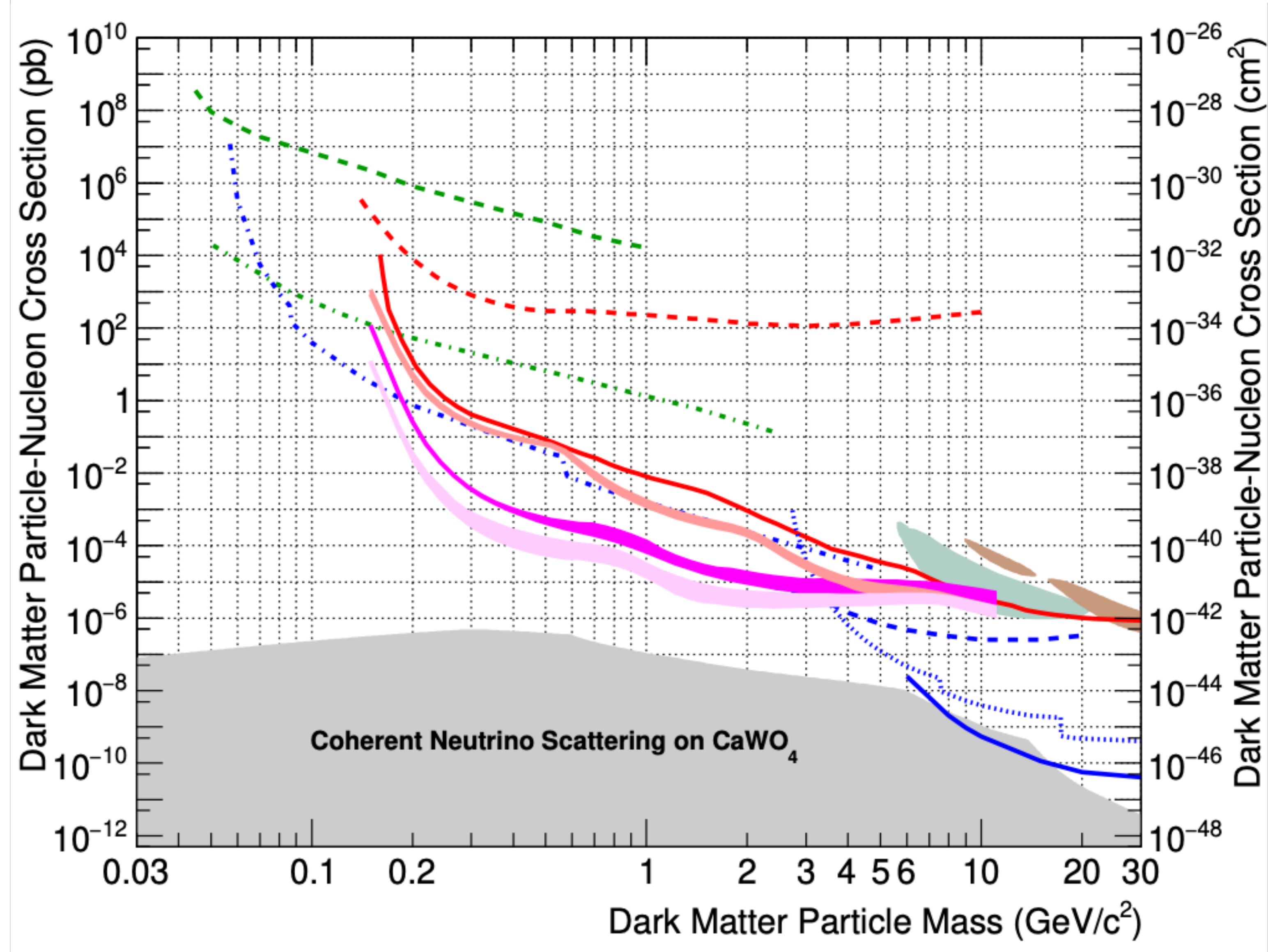
Analytical description of amplitude distribution in empty baselines

J Low Temp Phys (2019)
doi.org/10.1007/s10909-018-1948-6

W-TES equipped with heaters

- Stabilization of detectors in the operating point
- Injection of heat pulses for calibration and determination of trigger threshold





Sensitivity plots extending 2019 result to no-LEE scenario compared to 10-time more exposure.

LEE abatement is crucial to improve sensitivity.