

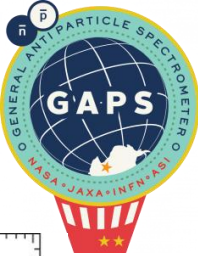
Search for Cosmic Ray Antinuclei from Dark Matter with the GAPS Antarctic Balloon Mission

Mengjiao Xiao

Shanghai Jiao Tong University

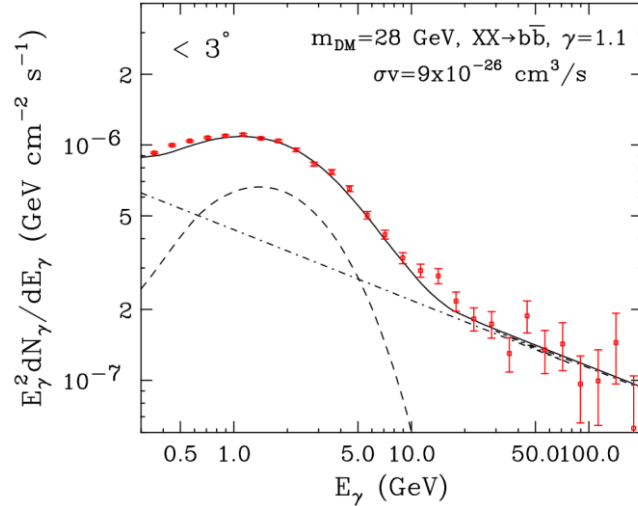
2025-08-27

(On behalf of the GAPS collaboration)

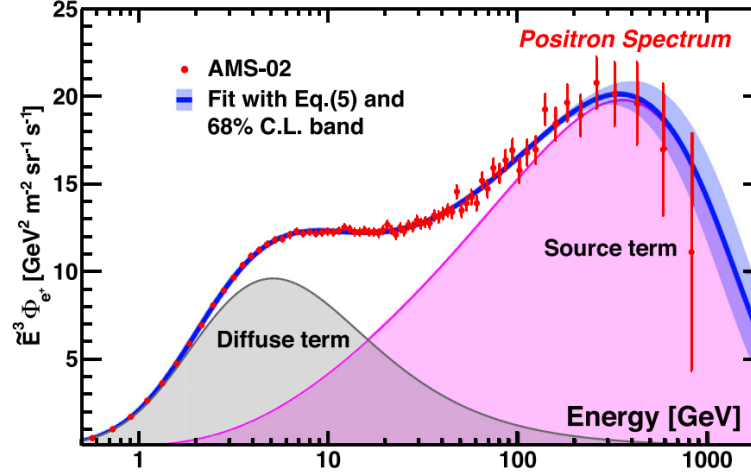


Dark Matter Hints in Cosmic Rays

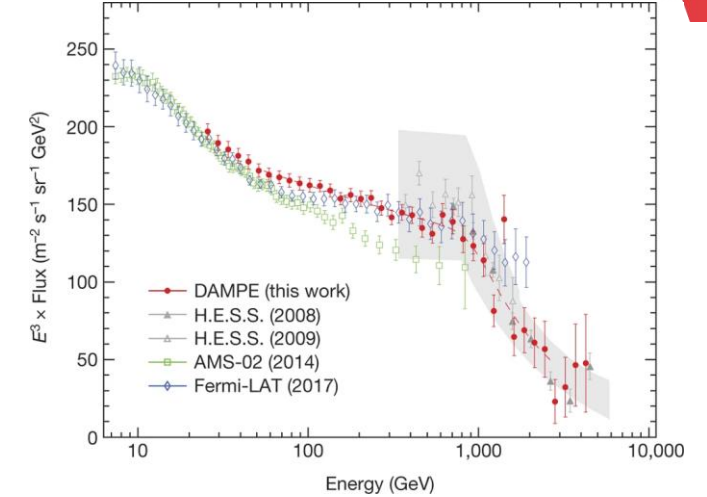
D. Hooper+ *PRD* (2011), I. Cholis+ *JCAP* (2009)



AMS-02: *Physics Reports*, Vol 894 (2021)

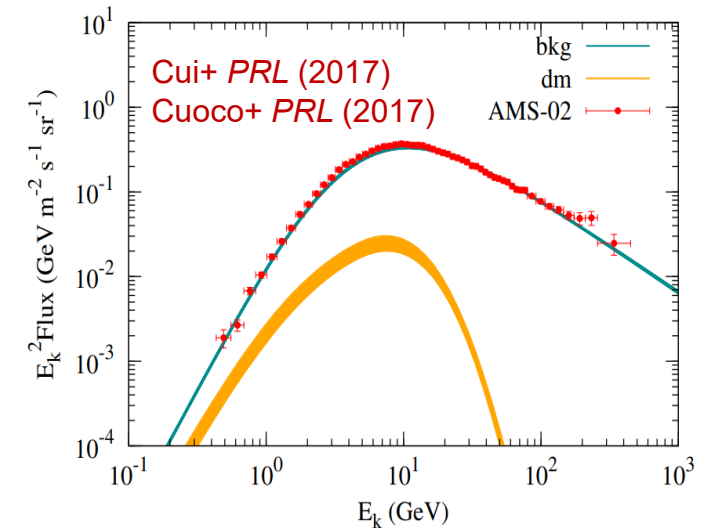


DAMPE, *Nature* (2017)



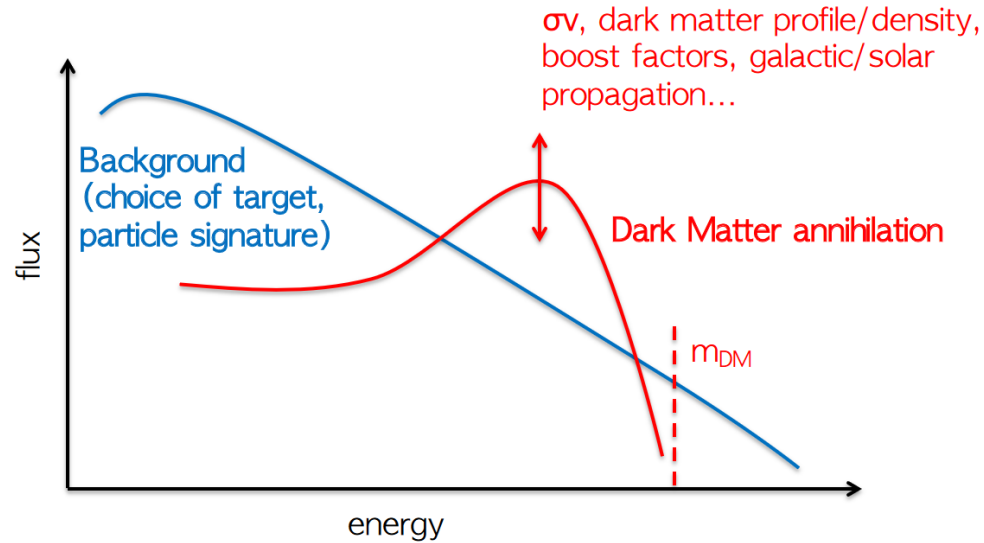
- GeV gamma excess at GC center by Fermi-LAT.
- Positron/Electron (~TeV) excess by PALEMA, AMS-02, Fermi-LAT, CALET, DAMPE.
- Antiproton excess in the 10-20 GV rigidity by AMS-02.
- Antihelium "candidates" by AMS-02.

Dark matter interpretation is complicated by astrophysical backgrounds and systematic uncertainties!



Dark Matter Detection with Cosmic Rays

❑ **Assumption:** dark matter annihilation/decay follows different kinematics (i.e. via physics BSM) than conventional productions.

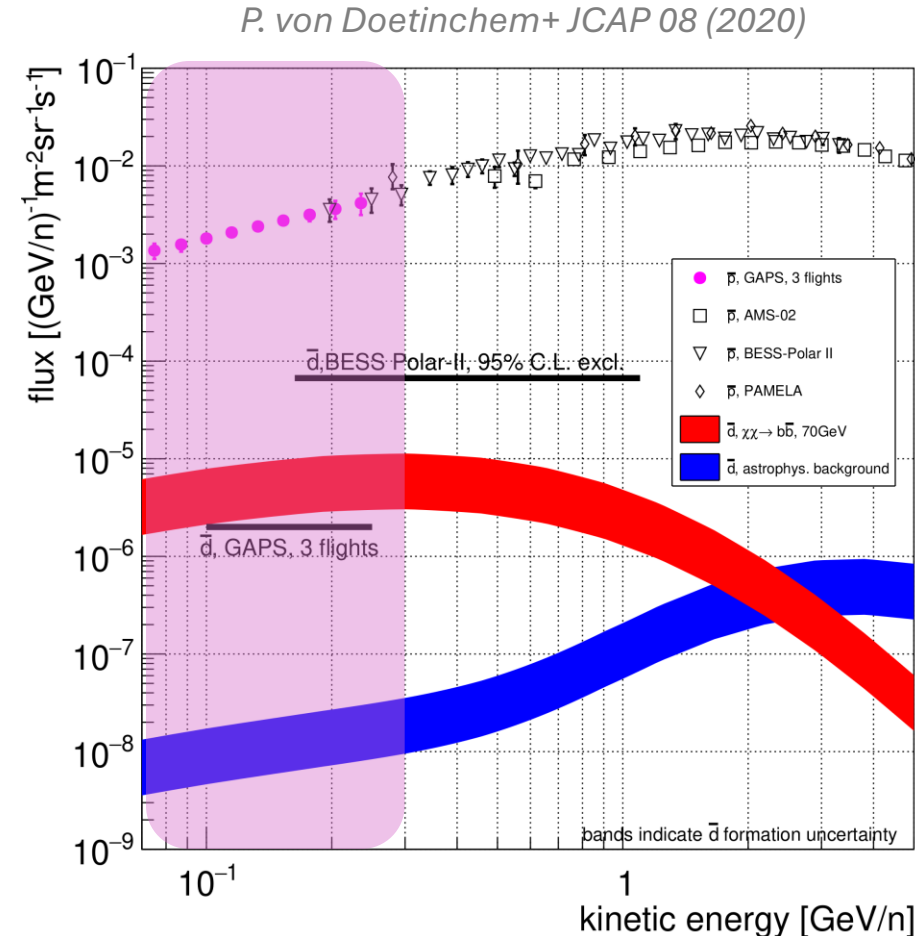
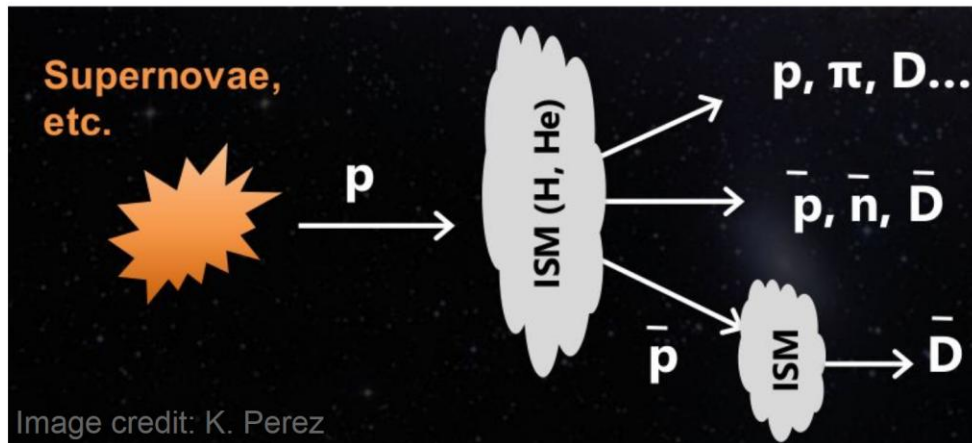
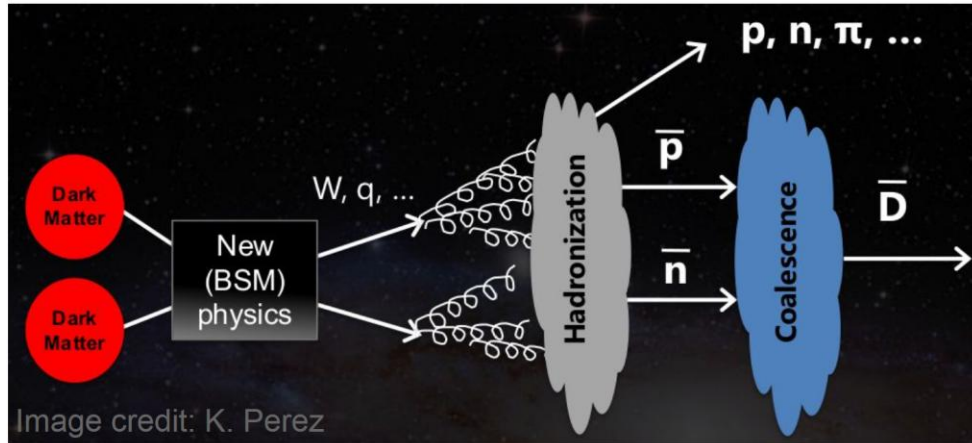


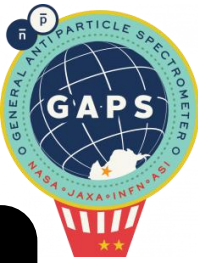
Common challenge (*FUN!*) =
minimize/constrain astrophysical bkg.
maximize predicted dark matter signal.

- ❑ Directly probes process that sets DM abundance!
- ❑ But, large systematic uncertainties
 - Cosmic ray propagation uncertainty
 - Hadronic interaction
 - Backgrounds from astrophysical sources
 - DM distribution profiles
 - DM annihilation final states
 -

Antideuterons as Dark Matter Signature

- Low-energy cosmic **antideuterons**: essentially background-free signature of dark matter, and **MEASURABLE!!**





The GAPS Balloon Mission

❑ **GAPS**=**G**eneral **A**nti**P**article **S**pectrometer

- Antarctic balloon experiment

❑ Unique sensitivity to **low-energy cosmic antinuclei** using novel exotic atom decay signatures: X-rays + charged particles

❑ Primary goal: low-energy ($KE \lesssim 0.25 \text{ GeV}/n$) **Antideuteron** as signature of new physics.

- Can probe many general dark matter models.

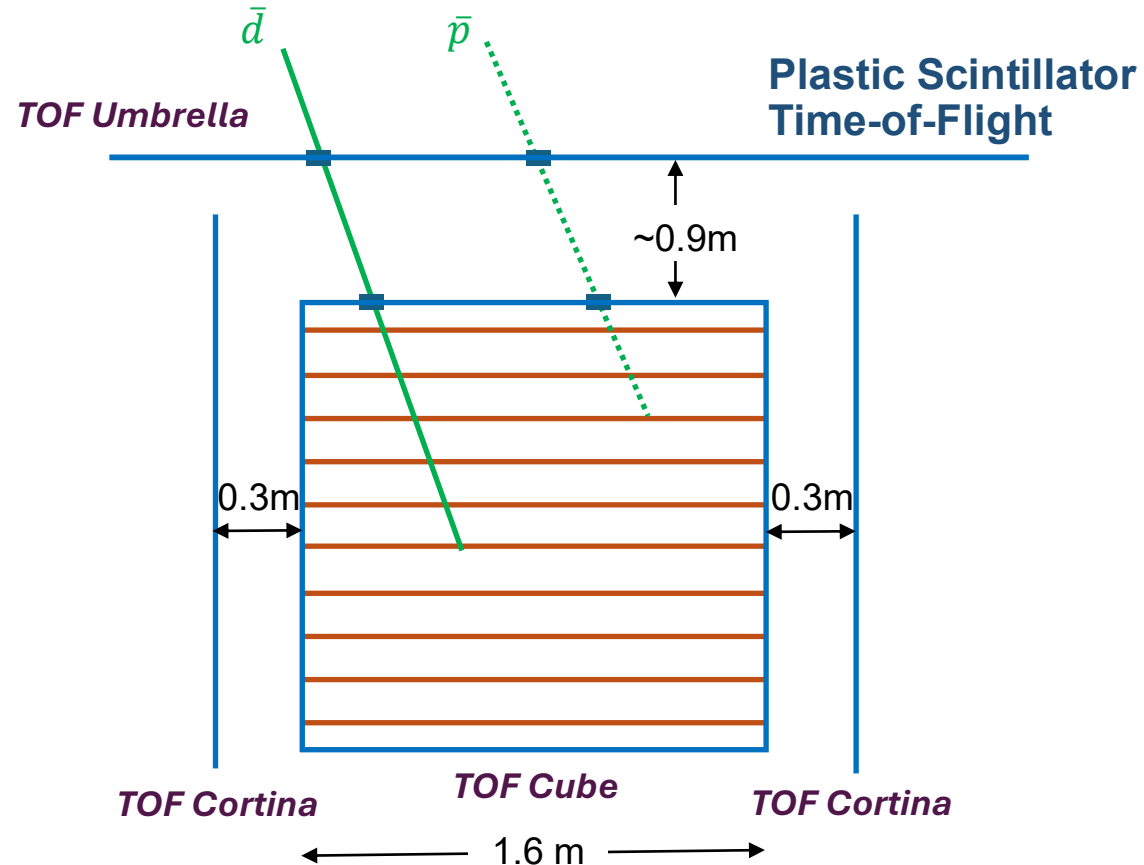
+ High statistics measurement of low-energy **Antiproton** and leading sensitivity to low-energy **Antihelium**.



➤ First Antarctic balloon flight late-2025, and two follow-up flights planned.



GAPS Novel Detection Technology



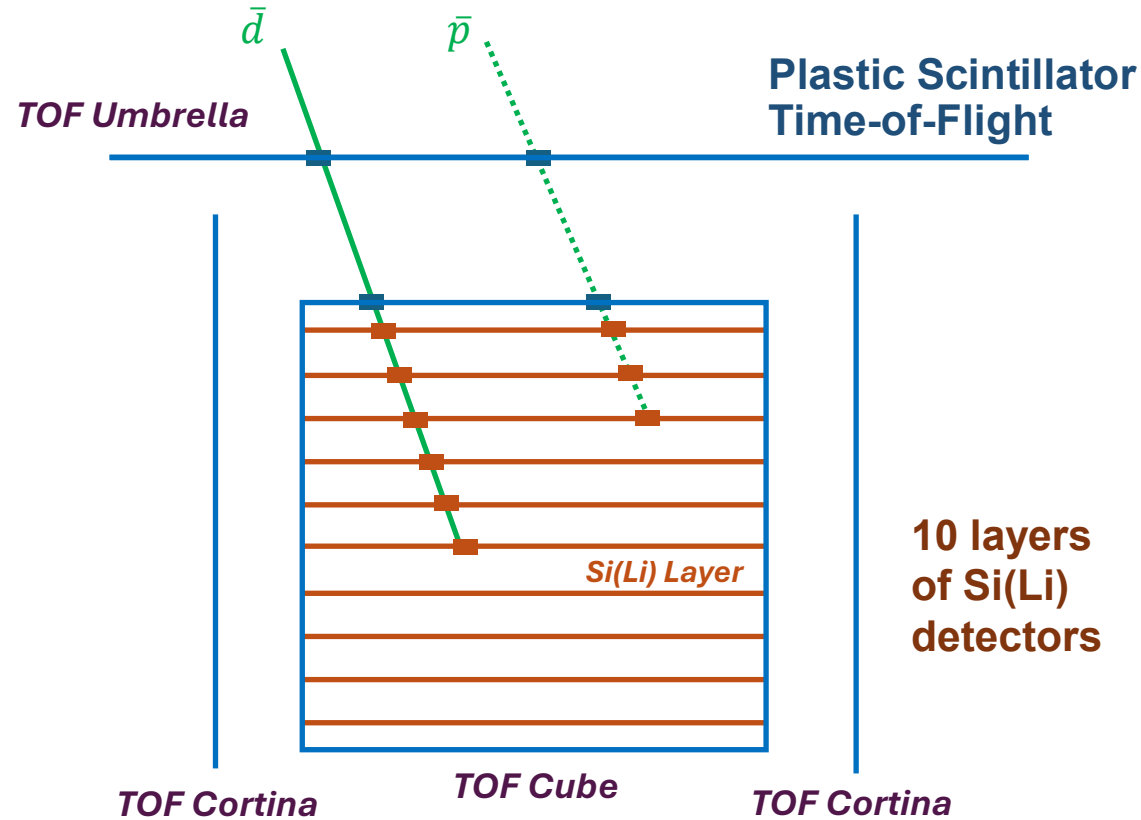
Time-of-flight system: measures velocity, incoming angle and dE/dx , fast trigger

Exotic atom technique verified at KEK: Aramaki+ Astropart.Phys. 49, 52-62 (2013)

GAPS sensitivity to antideuterons: Aramaki+ Astropart.Phys. 74, 6 (2016)



GAPS Novel Detection Technology



Time-of-flight system: measures velocity, incoming angle and dE/dx, fast trigger

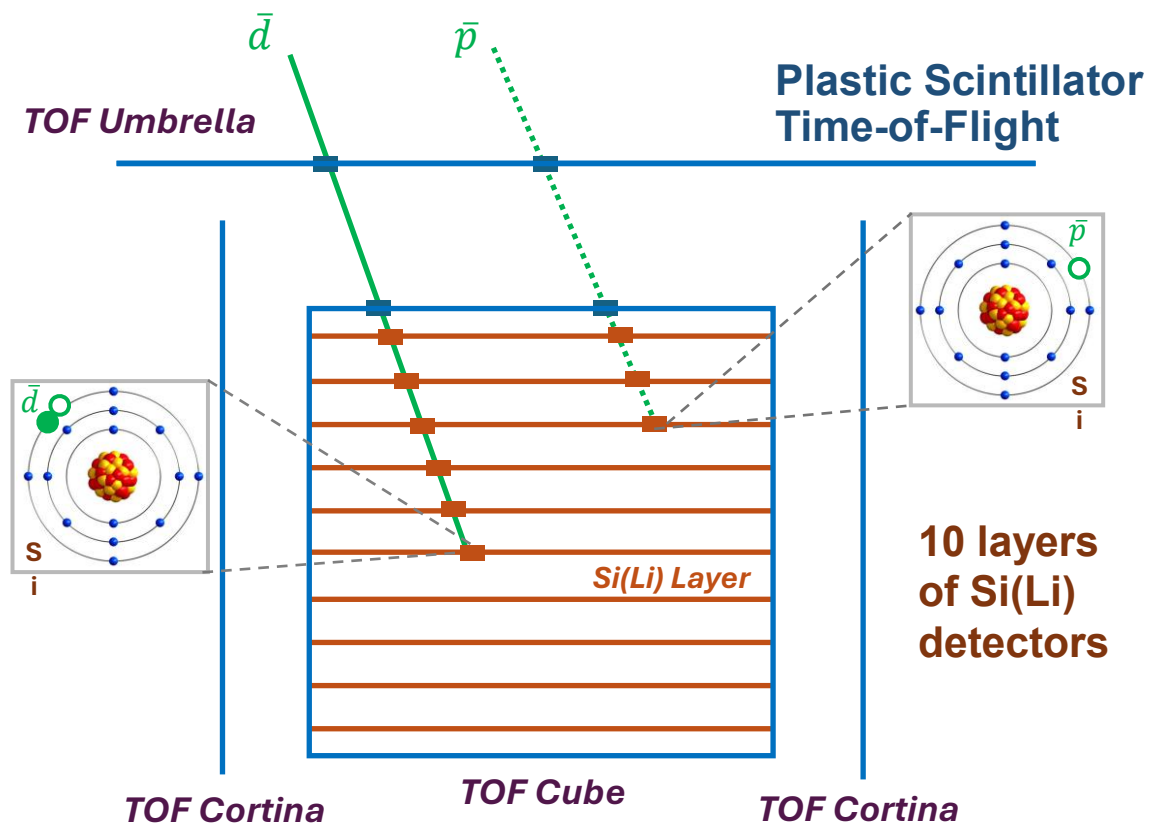
Si(Li) tracker:

- Slows/captures an incoming antiparticle

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GAPS Novel Detection Technology



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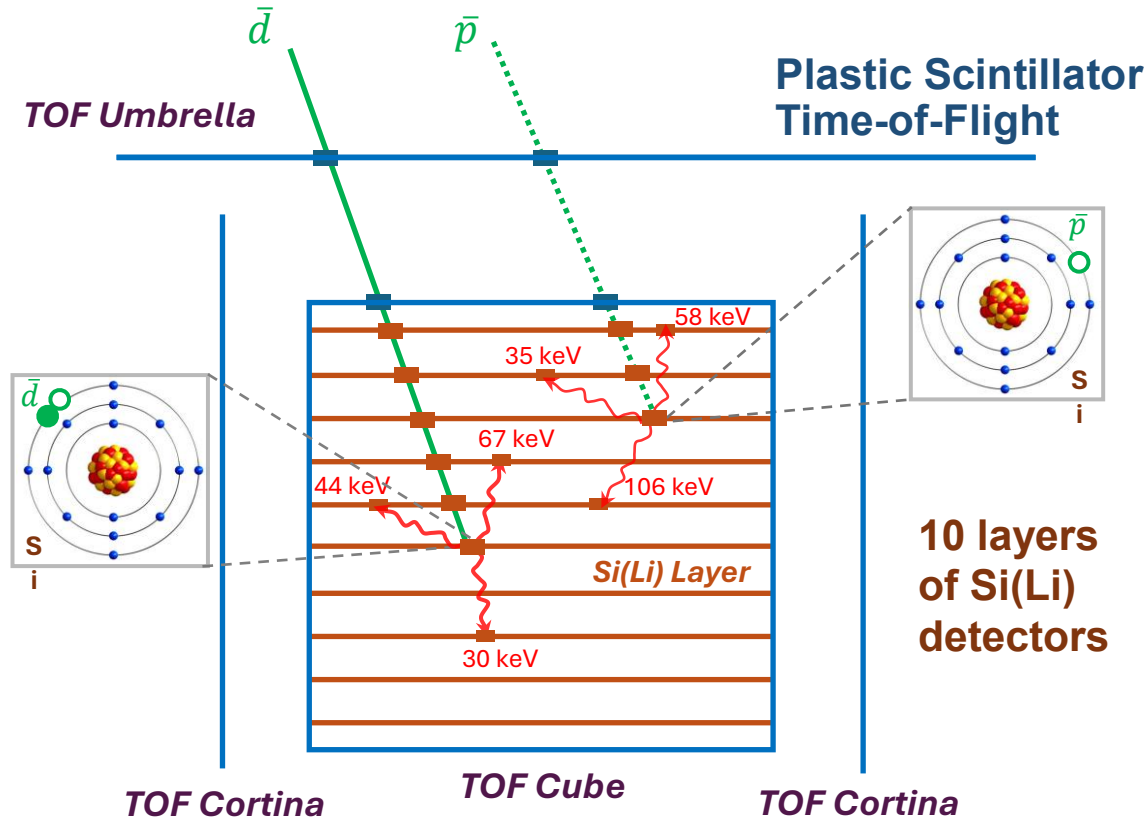
Si(Li) tracker:

- Slows/captures an incoming antiparticle into an **exotic atom**

Exotic atom technique verified at KEK: Aramaki+ *Astropart.Phys.* 49, 52-62 (2013)

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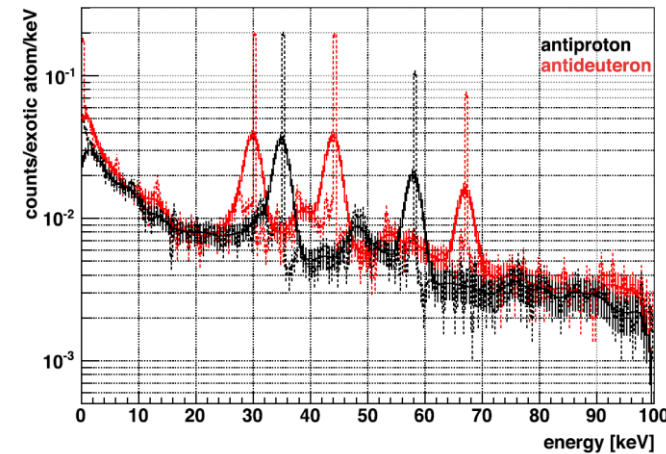
GAPS Novel Detection Technology



Time-of-flight system: measures velocity, incoming angle and dE/dx, fast trigger

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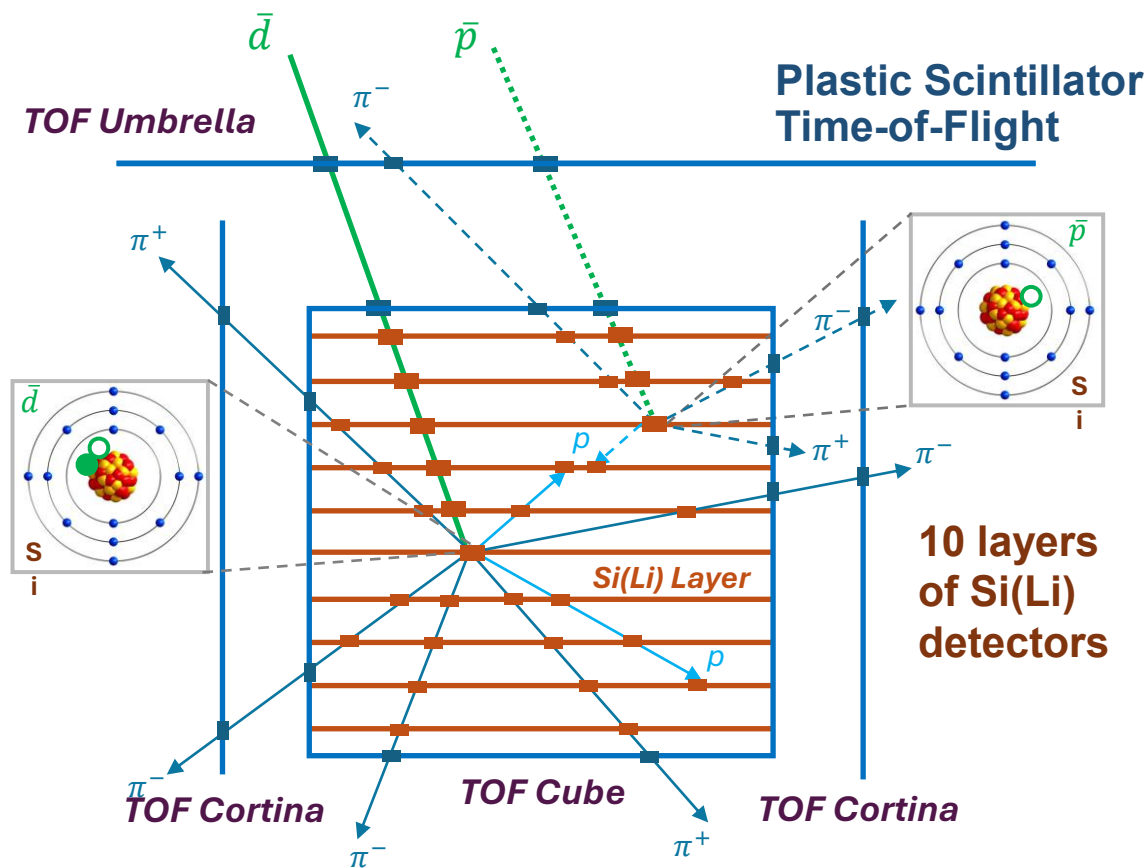
- Slows/captures an incoming antiparticle into an **exotic atom**
- Measures the decay **X-rays**



Exotic atom technique verified at KEK: Aramaki+ Astropart.Phys. 49, 52-62 (2013)

GAPS sensitivity to antideuterons: Aramaki+ Astropart.Phys. 74, 6 (2016)

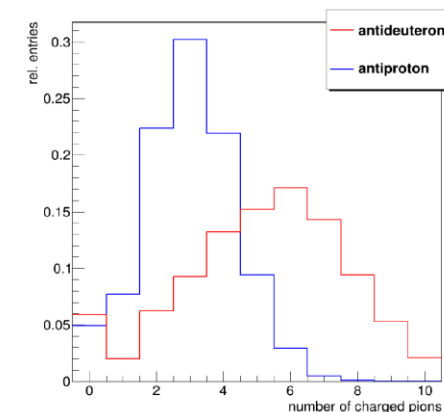
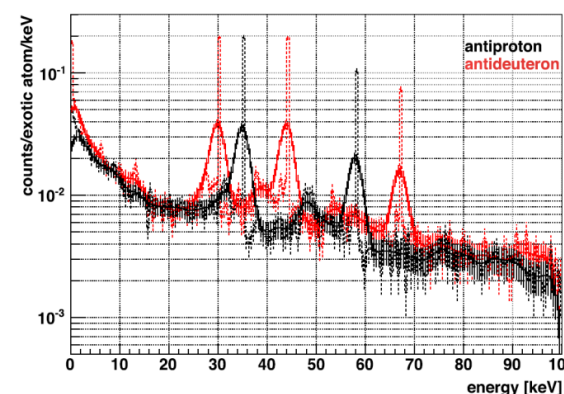
GAPS Novel Detection Technology



Time-of-flight system: measures velocity, incoming angle and dE/dx , fast trigger, tracks of outgoing particles

Si(Li) tracker:

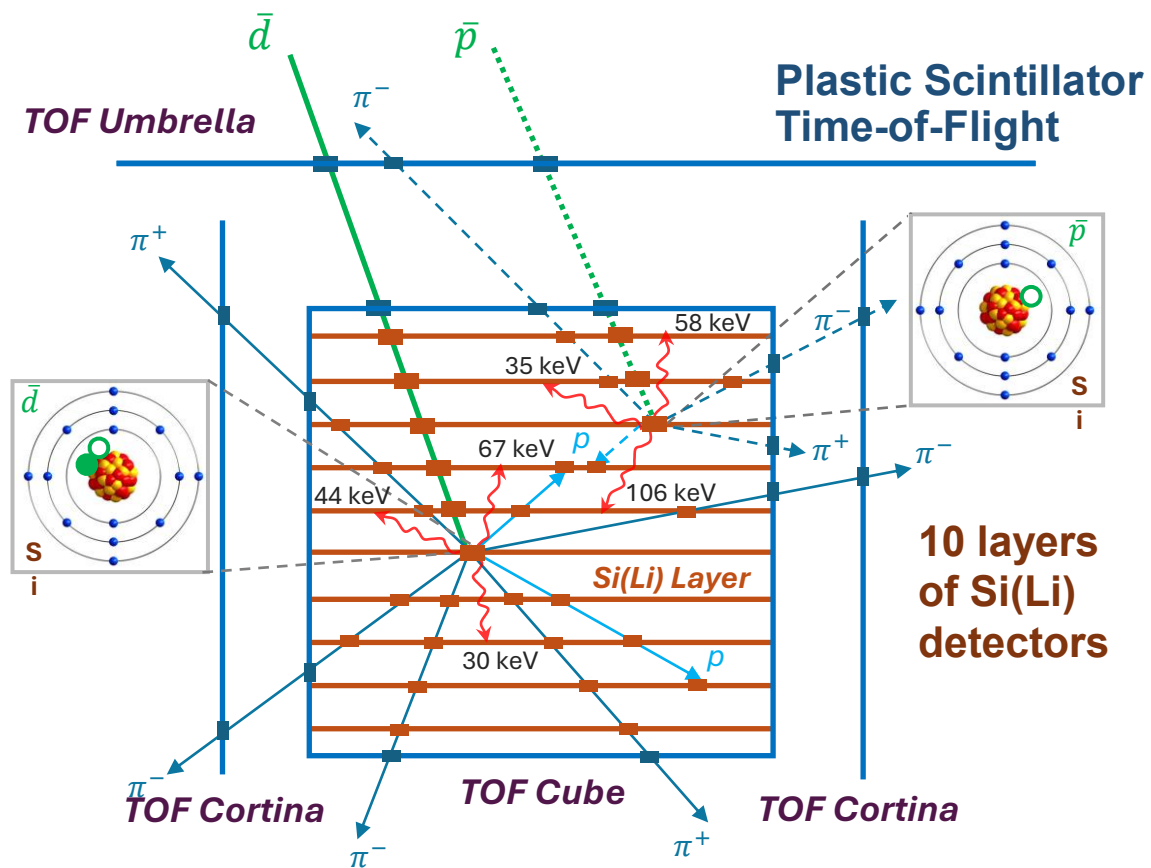
- Slows/captures an incoming antiparticle into an **exotic atom**
- Measures the decay **X-rays**
- Tracks the annihilated products (**charged π & p**)



Exotic atom technique verified at KEK: Aramaki+ Astropart.Phys. 49, 52-62 (2013)

GAPS sensitivity to antideuterons: Aramaki+ Astropart.Phys. 74, 6 (2016)

GAPS Novel Detection Technology



Time-of-flight system: measures velocity, incoming angle and dE/dx , fast trigger, tracks of outgoing particles

Si(Li) tracker acts as:

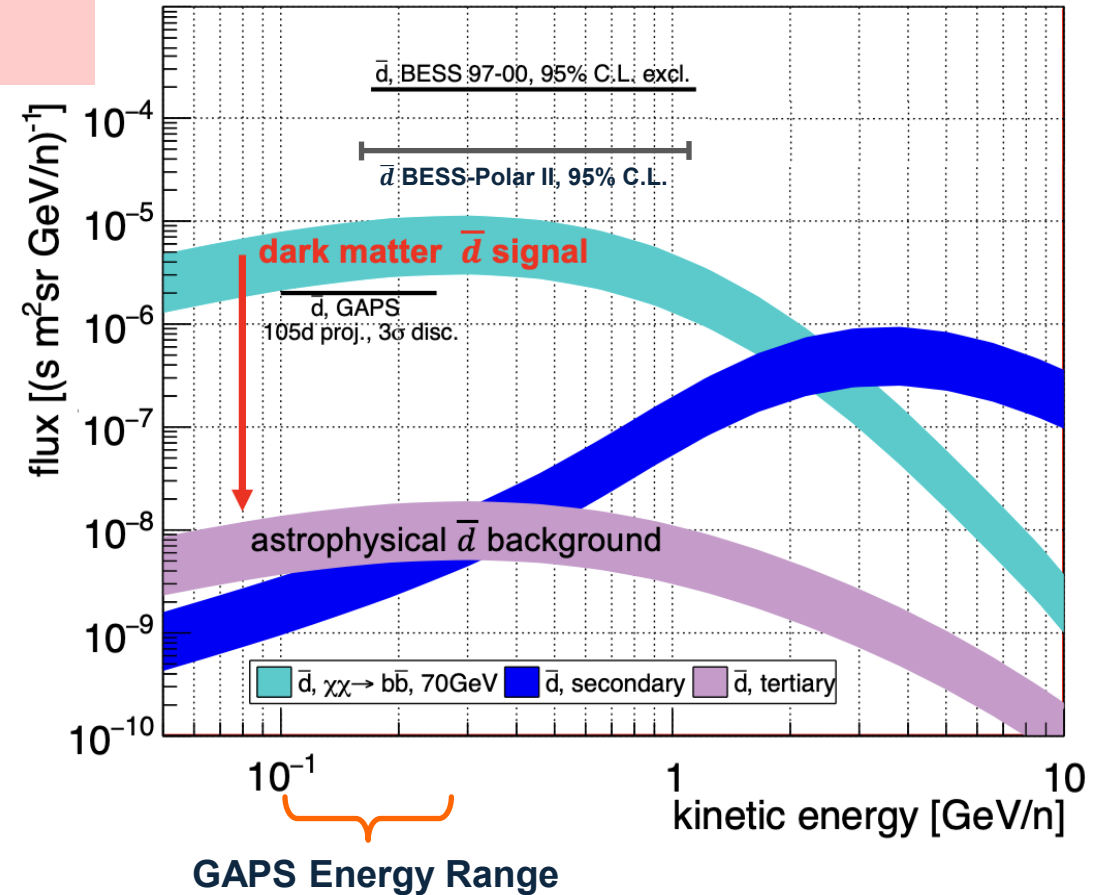
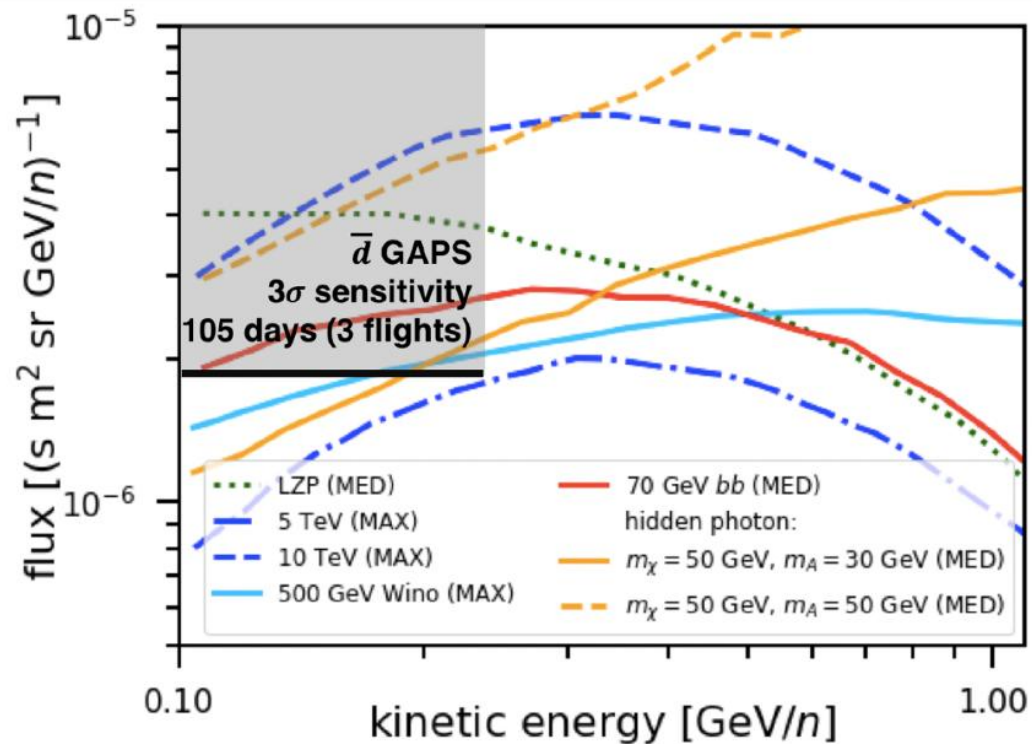
- **Target** to slow/capture an incoming antiparticle into an *exotic atom*
- **X-ray Spectrometer** to measure the decay X-rays
- **Particle Tracker** to measure the resulting dE/dX , stopping depth and annihilated *hadrons*

Exotic atom technique verified at KEK: Aramaki+ *Astropart.Phys.* 49, 52-62 (2013)

GAPS sensitivity to antideuteron: Aramaki+ *Astropart.Phys.* 74, 6 (2016)

GAPS Science: *Antideuteron*

- ❑ The GAPS antideuteron search is sensitive to a **wide range of generic DM models**:



Note: Any antideuteron signal needs to be compatible with antiproton constraints!



GAPS Science: *Antiproton*

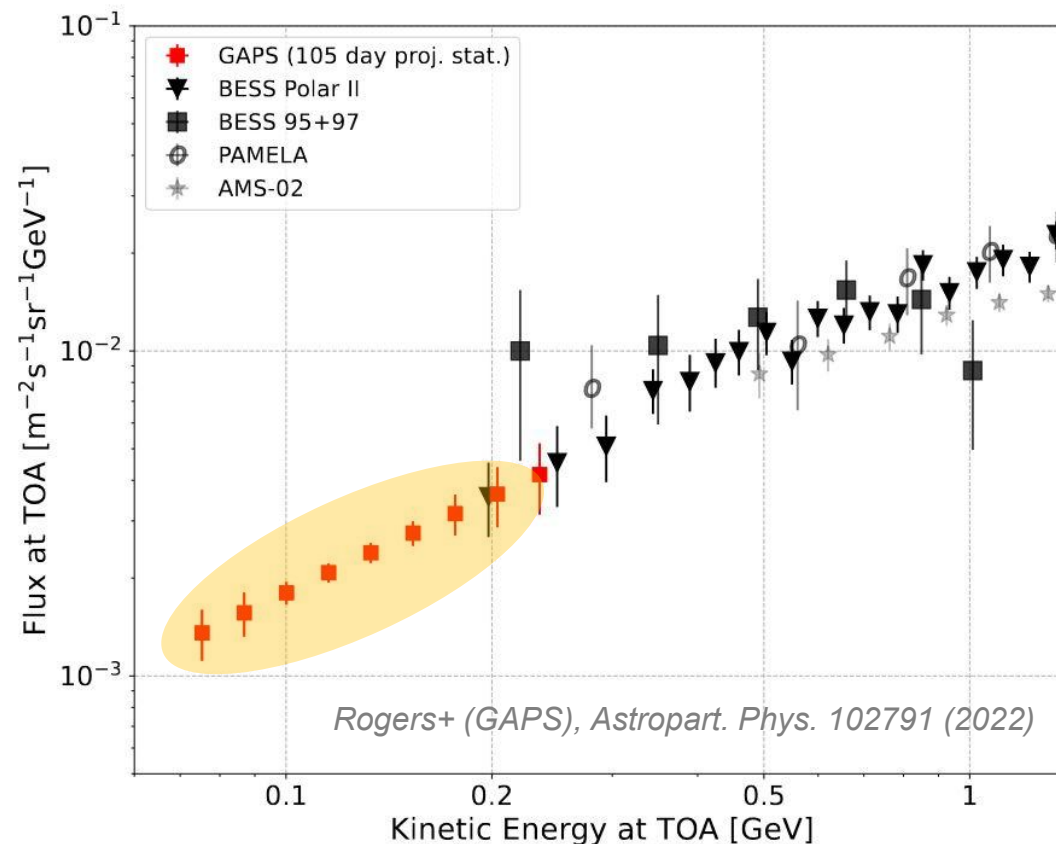
□ ~500 *antiprotons* ($\lesssim 0.25$ GeV/n)
for each balloon flight.

- BESS : 29 at ~0.2 GeV
- PAMELA: 7 at ~0.25 GeV
- AMS-02: $E > 0.25$ GeV



✓ Validate GAPS novel anti-nuclei
identification technologies.

➤ Reduce systematic uncertainties for
antideuteron search.



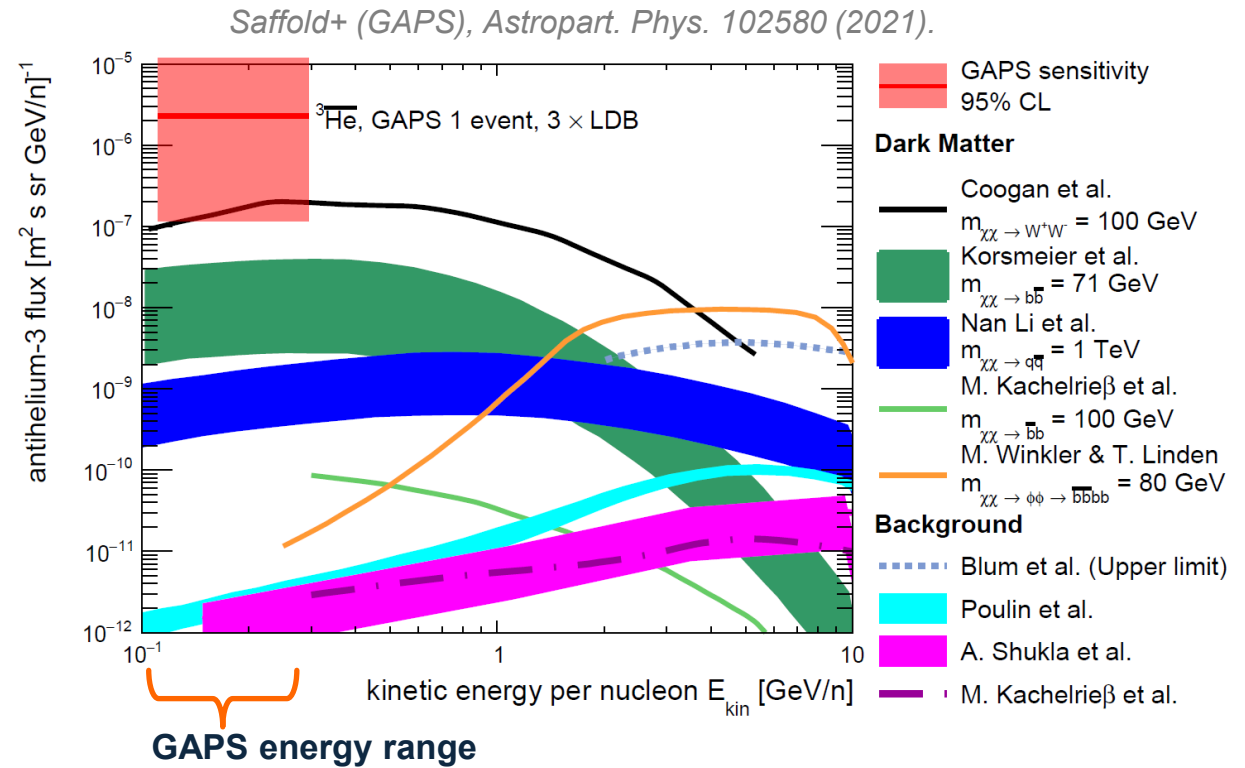
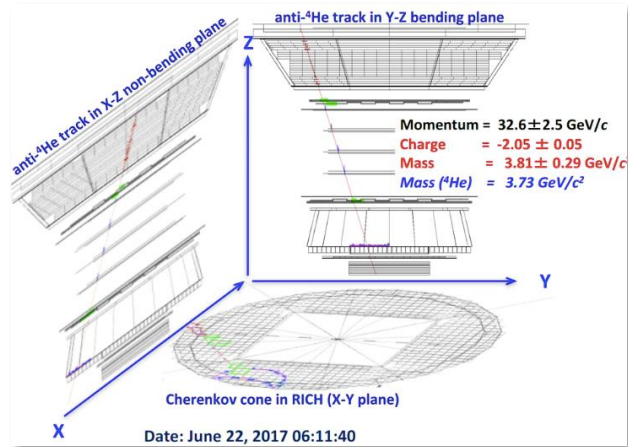
- Probe *light dark matter*, leading constraints on *primordial black hole* evaporation.
- Provide a novel insight on *cosmic ray propagation* models.

GAPS Science: *Antihelium*

□ GAPS flux sensitivity to antihelium-3 (three 35-day long duration flights).

- **2018:** “To date, we have observed eight events...with $Z = -2$. All eight events are in the helium mass region.”
– S. Ting (La Palma, AMS overview)

AMS Candidate Anti-He4 event ($p = 32.6$ GeV/c)



- Extends to lower energies (0.11-0.3 GeV/n), complementary to AMS-02.
 - Capable of confirming signal, orthogonal detection technique, uniquely low bkg.

❑ Time-of-Flight (TOF)

- Near-hermetic containment of tracker
- Velocity, trajectory and dE/dx measurement
- High-speed trigger and veto

❑ Si(Li) Tracker

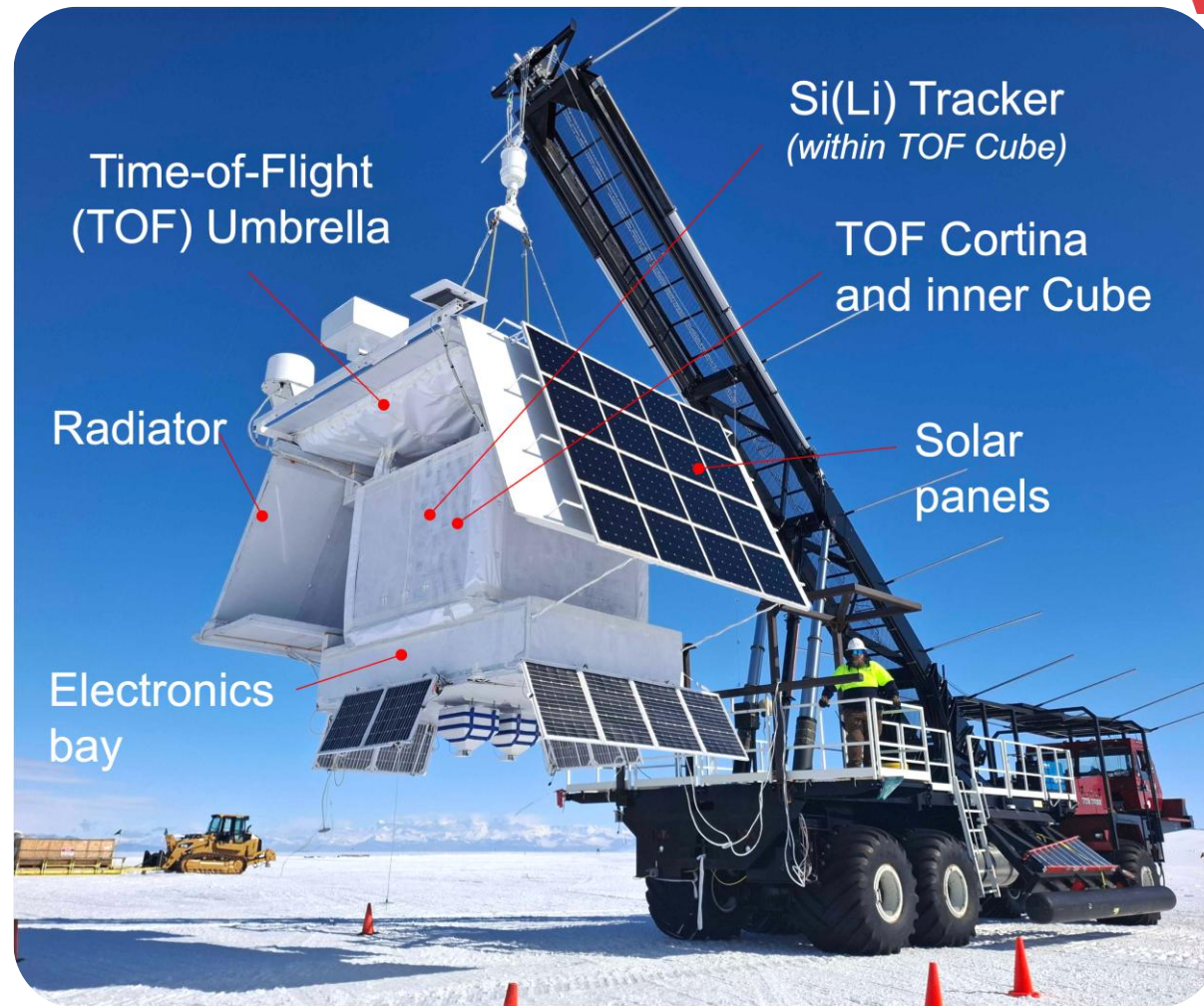
- Target to capture light nuclei $\lesssim 0.25 \text{ GeV}/n$
- Tracker for primary and secondary hadrons
- Spectrometer for de-excitation X-rays

❑ Thermal System

- Oscillating Heat Pipe for tracker cooling

❖ Support instrumentation

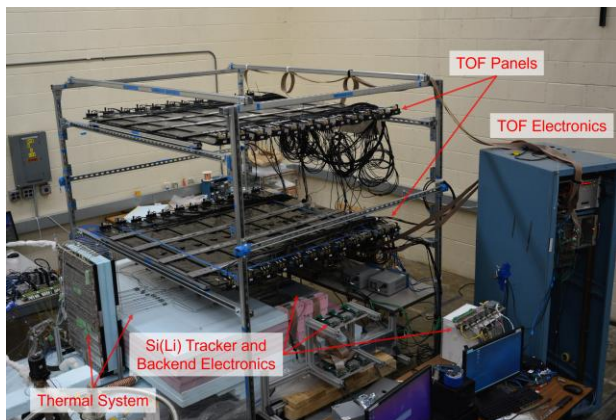
- Electronics, Solar panels, Gondola



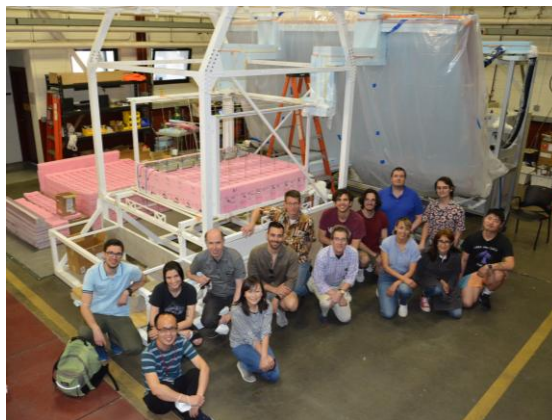


GAPS Instrument Integration

2020, GFP @MIT-Bates, MA



2021, initial integ. @MIT-Bates, MA



2022, full integ. @Berkeley-SSL, CA



6/2023, TVAC test
@NTS, CA



1/2024, upgrades
@CU-Nevis, NY



7/2024, compatibility testing
@CSBF, TX



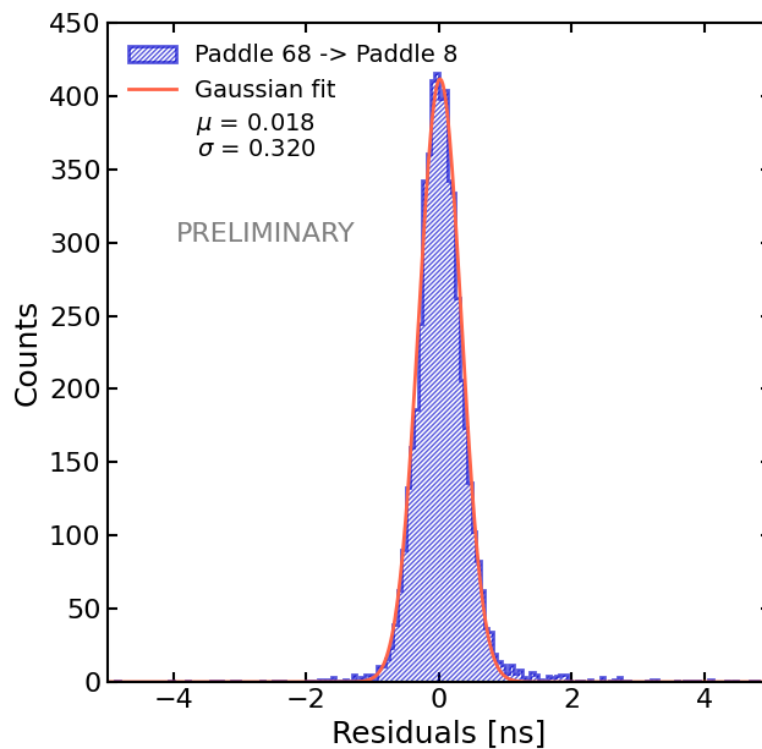
12/2024, commission & launch
attempts @McMurdo, Antarctic

TAUP @Xichang, Aug. 24-30, 2025

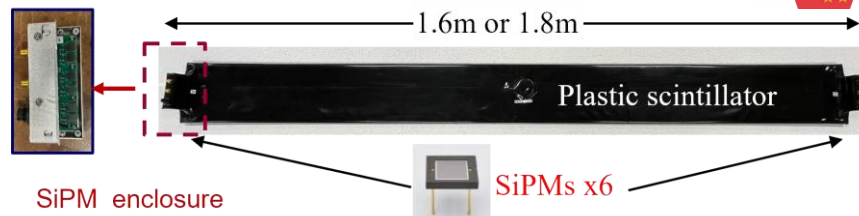
GAPS Instrument Performance: *ToF*

❑ ToF performance from the Antarctic commissioning.

- Design goal: <500 ps timing resolution for TOF system to be able to separate proton and deuteron.
- Cross-calibration with different paddle combinations to estimate timing resolution



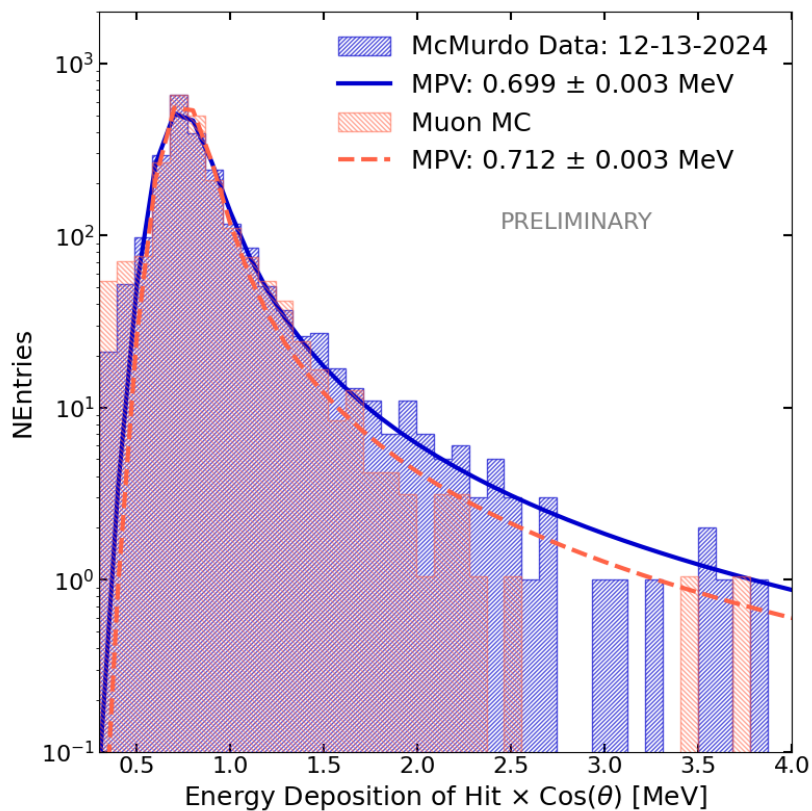
- Timing resolution



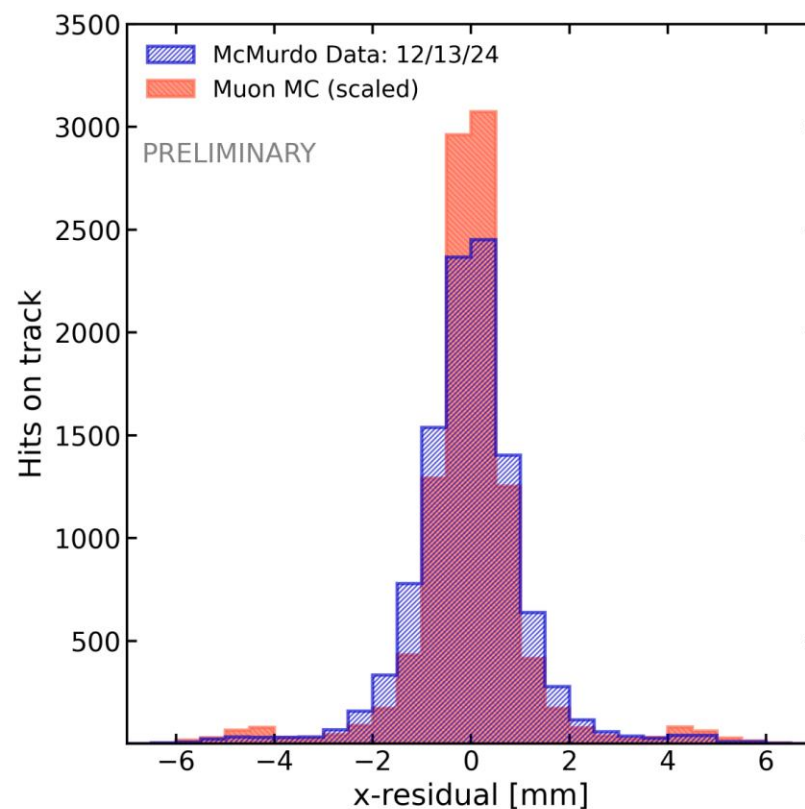
➤ Achieved: measured better than ~350 ps for all paddle combinations!

GAPS Instrument Performance: *Tracker*

❑ Tracker performance from the Antarctic commissioning.

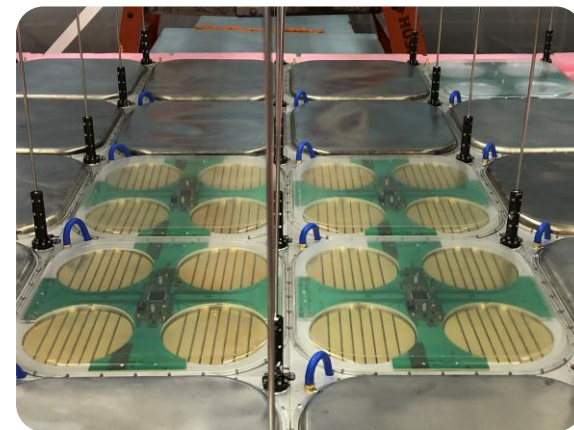
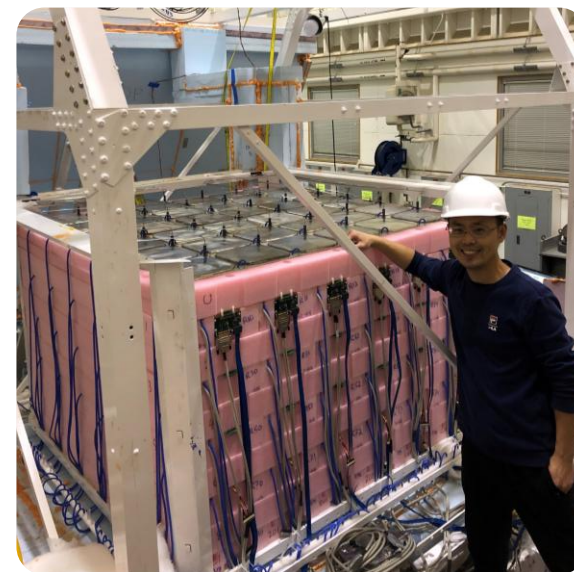


▪ Energy depositions on track



▪ Track position resolution

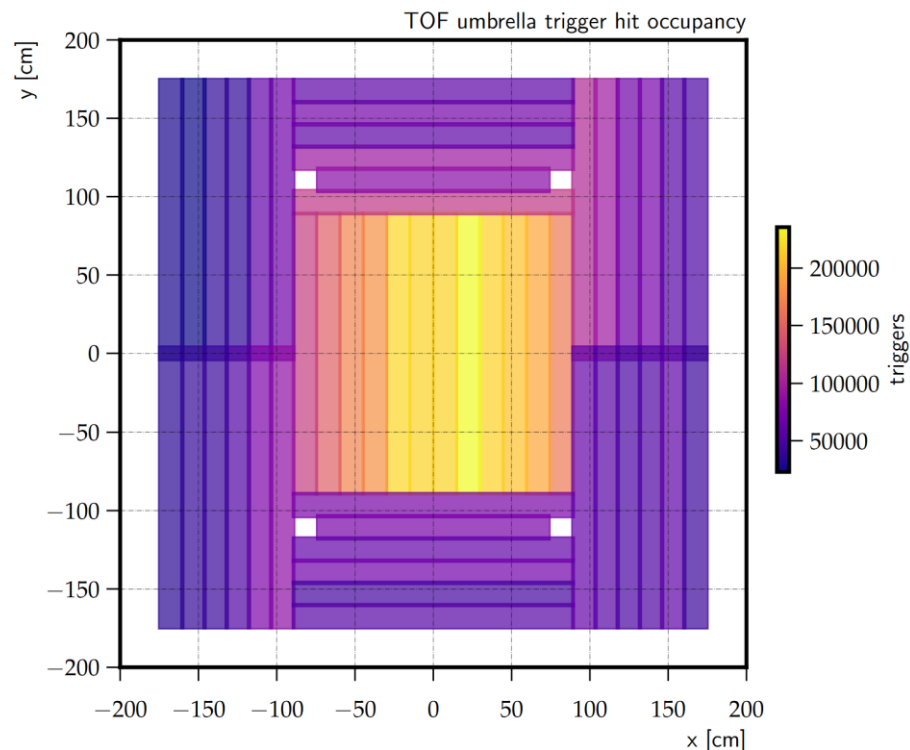
➤ Energy and track reconstruction validated!





GAPS Instrument Performance: *Trigger*

- ❑ Design goal: 450-500 Hz trigger rate (maxing out telemetry bandwidth), capability to trigger on multi-track events.

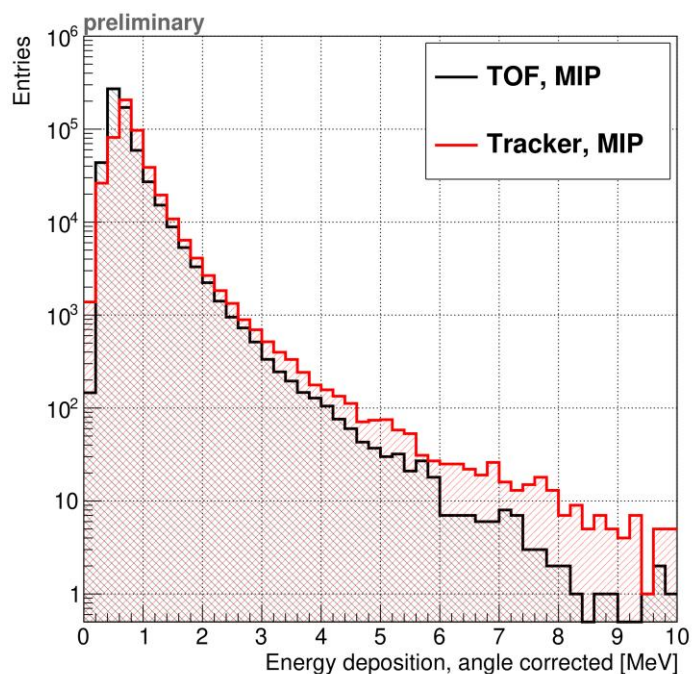


- GAPS trigger allows for 2 modes in parallel (multi + single track)
- Modes can be tuned with pre-scale factor
- On ground multi track ~few Hz, single track (w. prescale) ~450-500 Hz

- Achieved: stable operations at ground with tuned trigger at 450-500 Hz possible, can record multi track events, trigger acceptance verified with simulations!

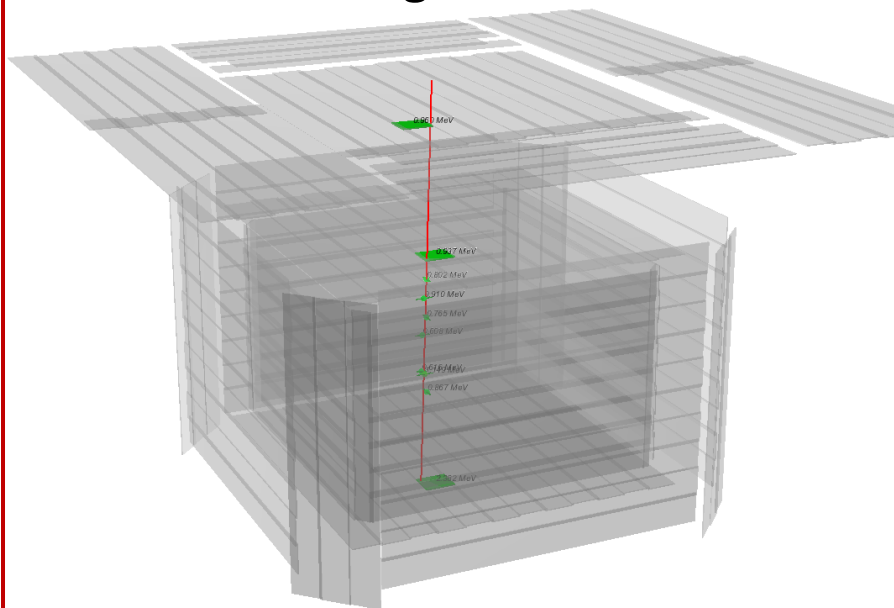
GAPS Instrument Performance: *System*

- ❑ ~10M muon events are collected from the on-ground testing in the Antarctica.
 - Event signatures are well understood and more detailed analysis is undergoing.



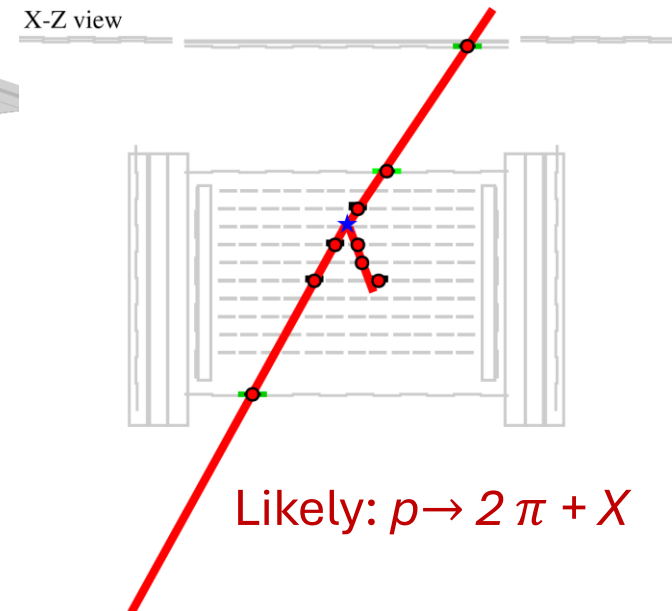
- Energy reconstruction

- Single track



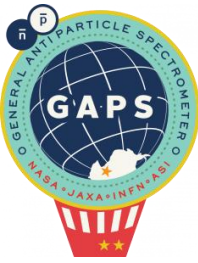
- Track Reconstruction

- Multi track





Summary & Conclusions

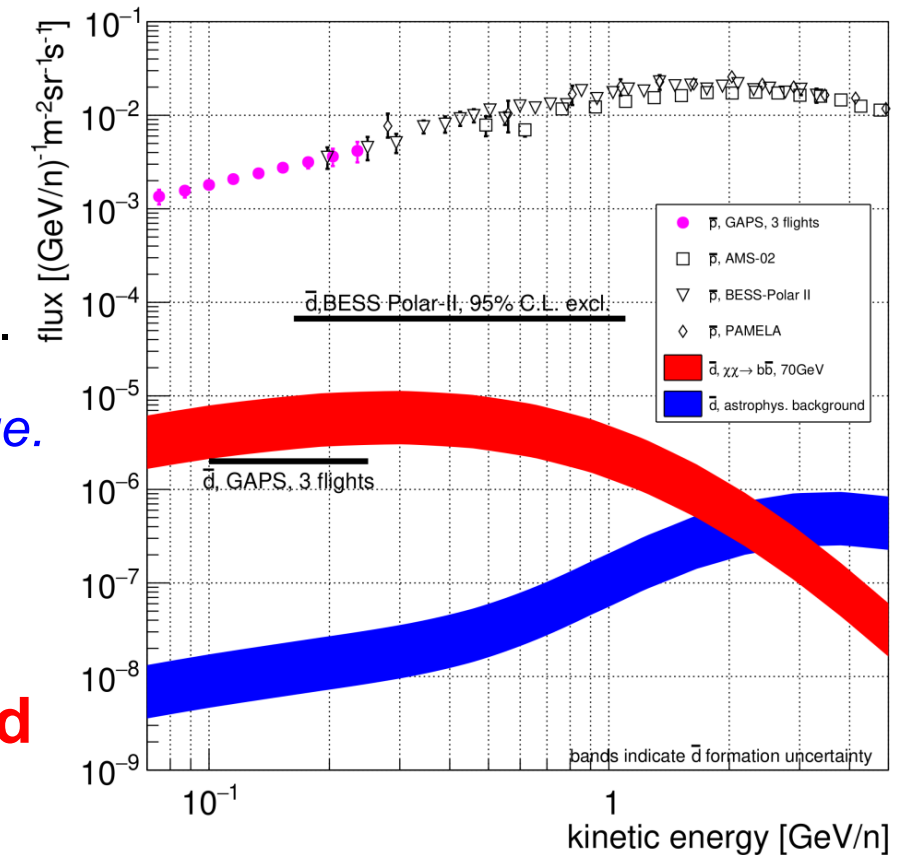


❑ Low-energy cosmic antinuclei are unexplored and unique for new physics (e.g. dark matter) searches.

❑ GAPS aims to deliver:

- *unprecedented \bar{d} sensitivity* by ~ 100 times below the current best limits, “smoking-gun” dark matter signature.
- *precision \bar{p} measurement in an unexplored energy range.*
- *leading sensitivity to low-energy cosmic anti-He.*

❑ **GAPS instrument is READY in Antarctica and planed for flight in late 2025, Stay tuned!!**



GAPS Collaboration



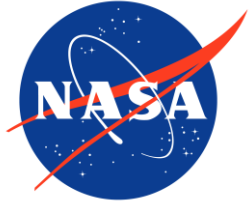
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THANK YOU!

