

Background simulation for the dark matter search experiment CRESST



Laboratori Nazionali del Gran Sasso



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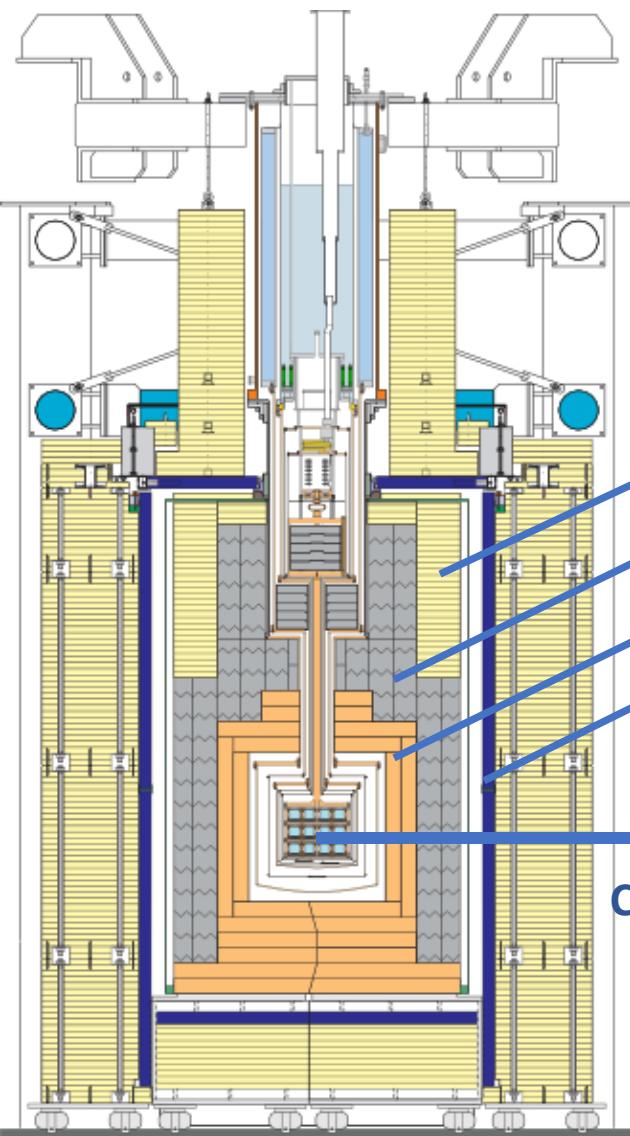


COMENIUS
UNIVERSITY
BRATISLAVA



TECHNISCHE
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WIEN

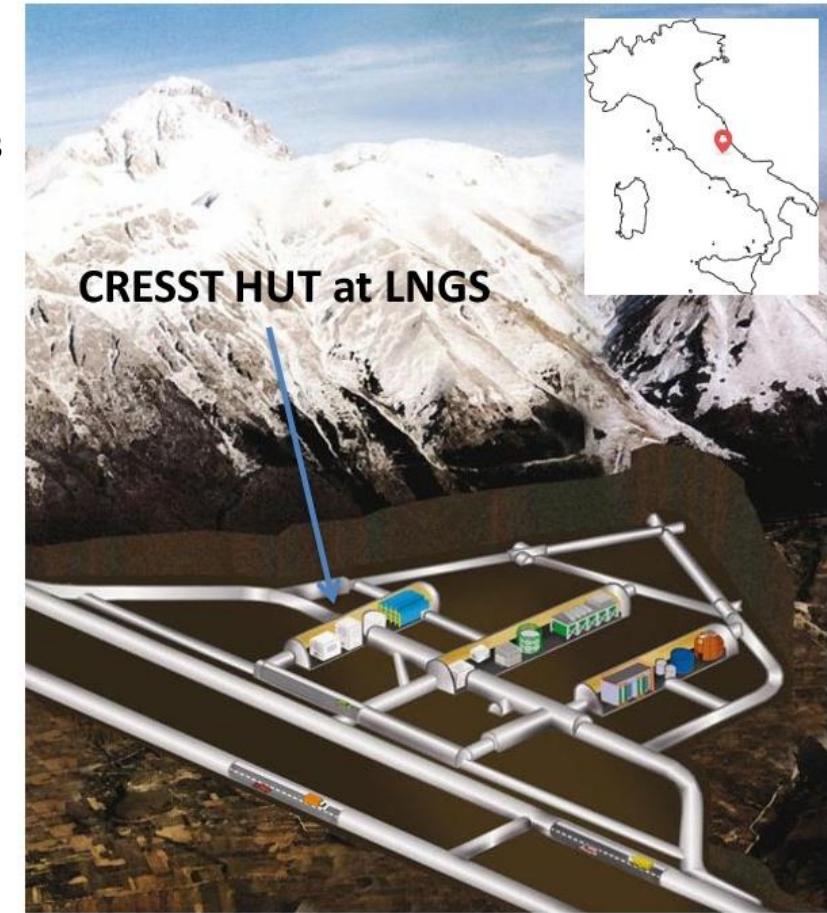
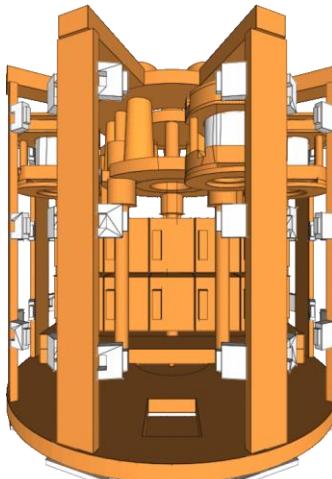
CRESST is located at LNGS (Laboratori Nazionali del Gran Sasso) in Italy



- Cryogenic scintillating calorimeter
- Target materials: CaWO_4 , Si, LiAlO_3 , Al_2O_3
- Read out channels: phonon and scintillation light

Shielding:

- polyethylene;
- lead;
- copper;
- muon veto system.



Experimental location:

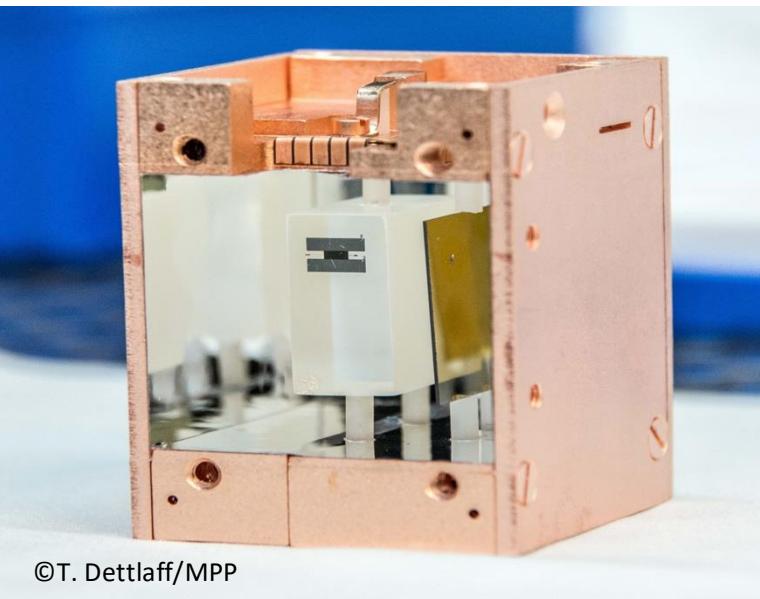
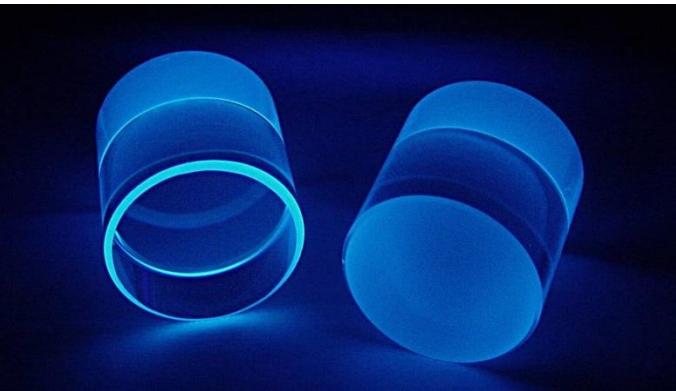
Average depth ~ 3600 m w.e.

Muon flux $\sim 2.6 \times 10^{-8} \mu/\text{s}/\text{cm}^2$

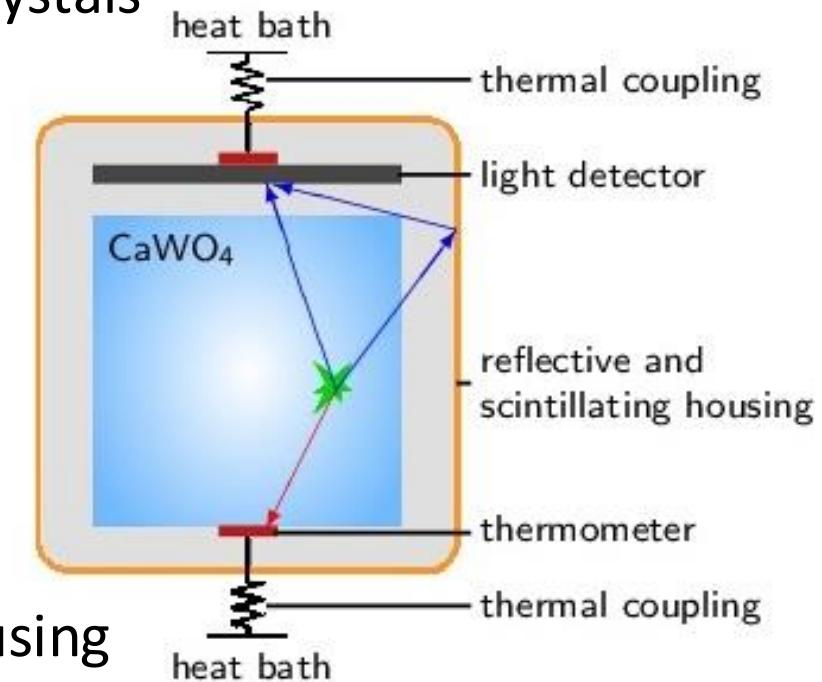
Neutrons < 10 MeV: $< 10^{-6} \text{n}/\text{s}/\text{cm}^2$

The CRESST experiment

Direct detection of dark matter particles via their scattering off target nuclei



- target material: CaWO₄ single crystals
- particle interaction
 - heat (phonon) signal
read-out with thermometer
 - light signal
read-out with light detector
- reflective and scintillating housing



Target crystals operated as cryogenic calorimeters (~15mK)

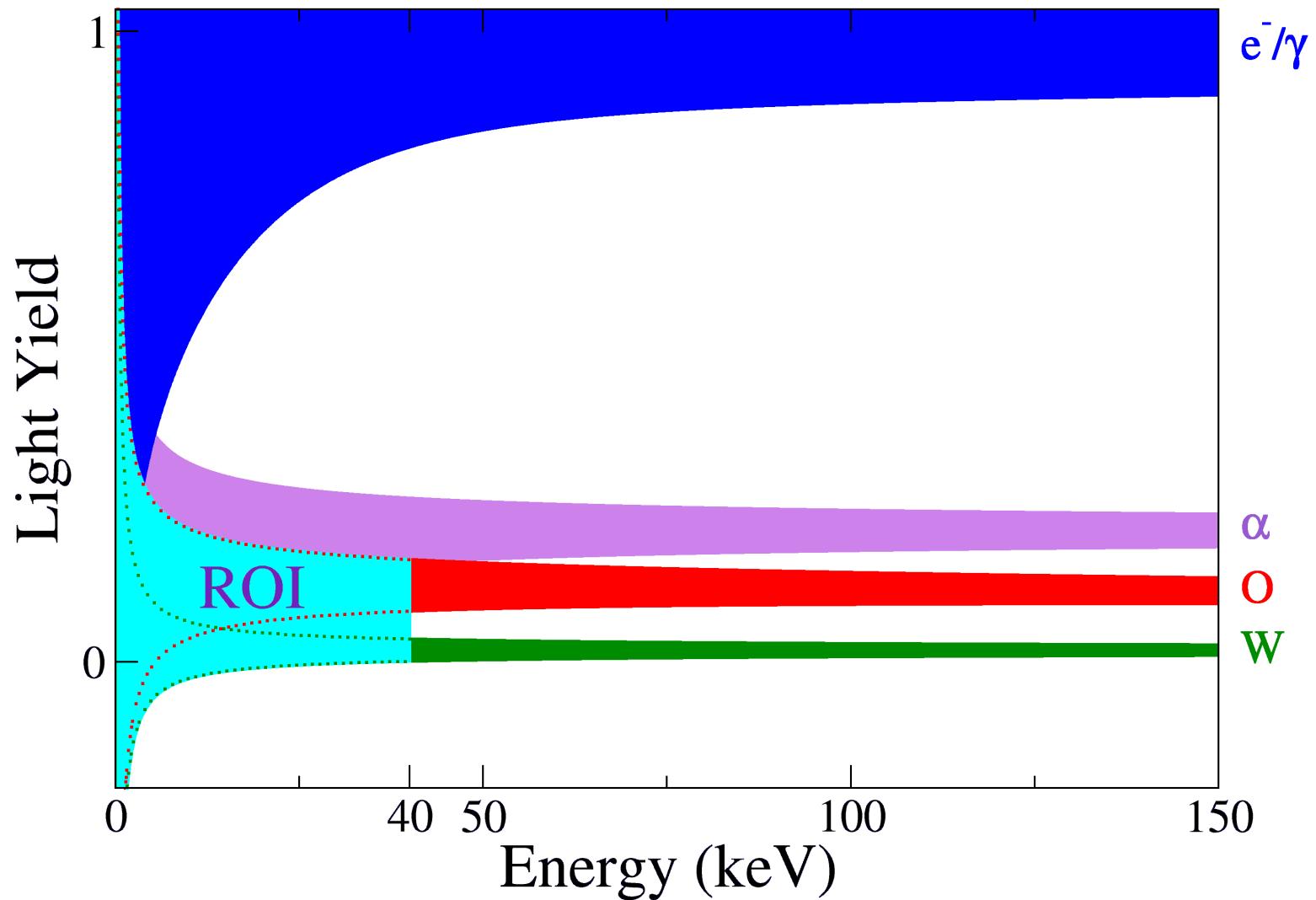
Event discrimination

$$\text{Light Yield} = \frac{\text{Light signal}}{\text{Phonon signal}}$$

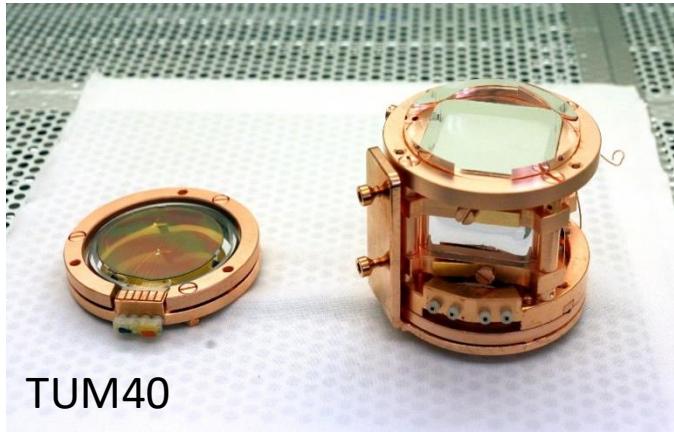
Characteristic of the event type

Excellent discrimination between potential signal events (**nuclear recoils**) and dominant radioactive background (**electron recoils**)

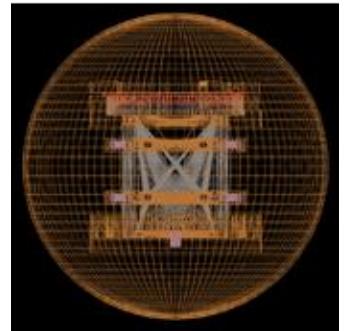
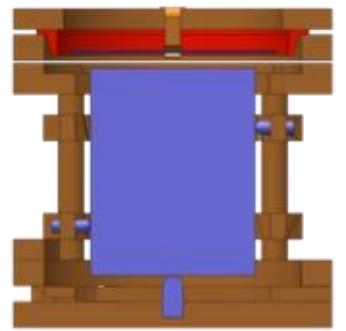
ROI: region of interest for dark matter search



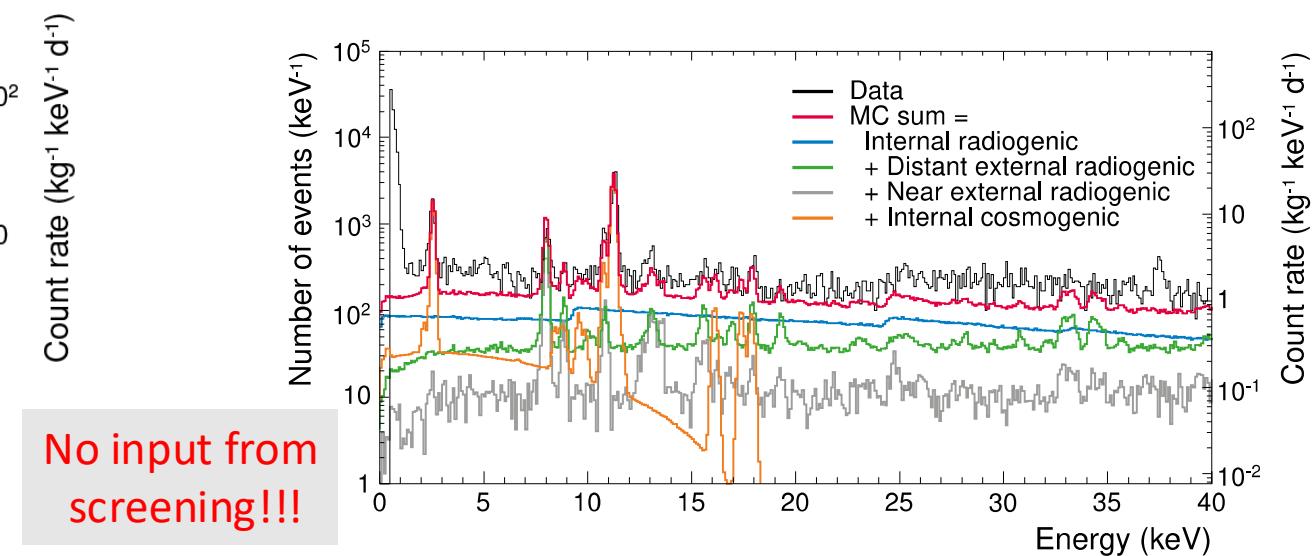
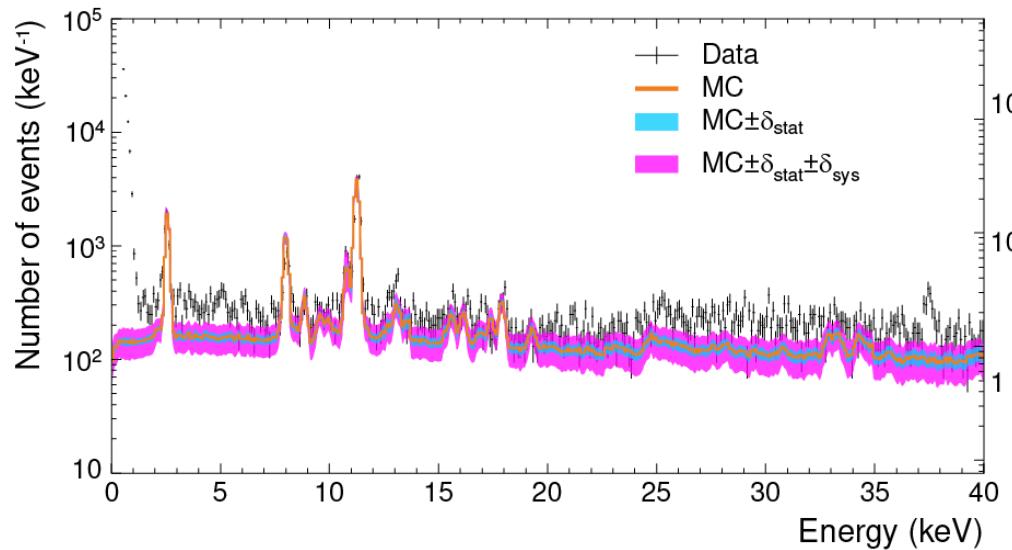
One module background simulation



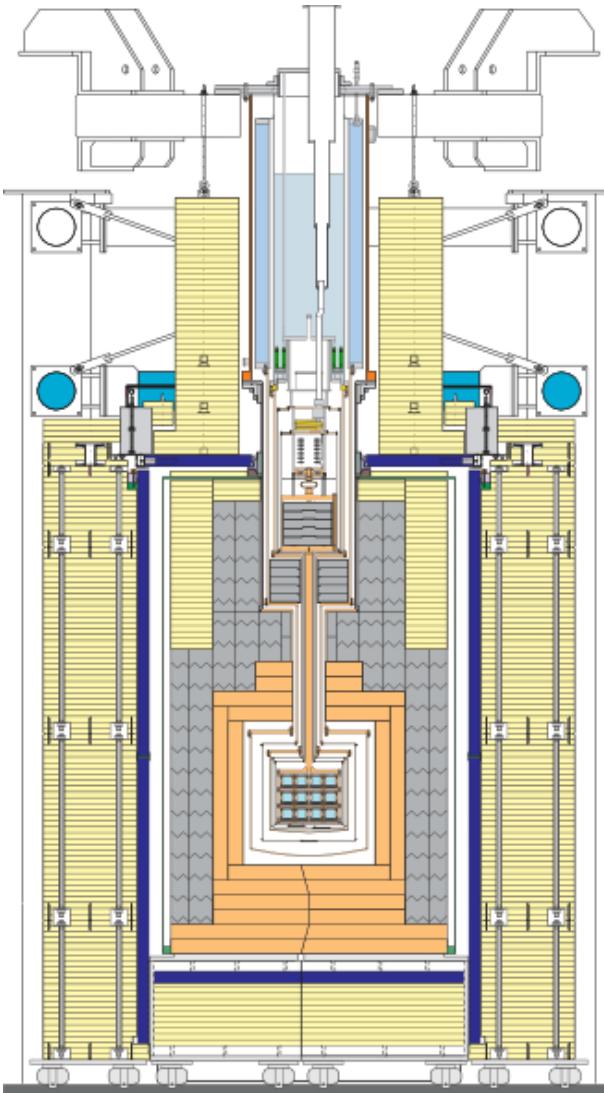
- Geant4 based electromagnetic background model for the CRESST experiment
- Study of cosmogenic activation of CaWO_4 crystal scintillator
- Simplified geometry reproduced already up to 68% background in ROI
- Foundation for more detailed models of the actual CRESST detector modules



Up to $68 \pm 16\%$ of background can be reproduced with simulations



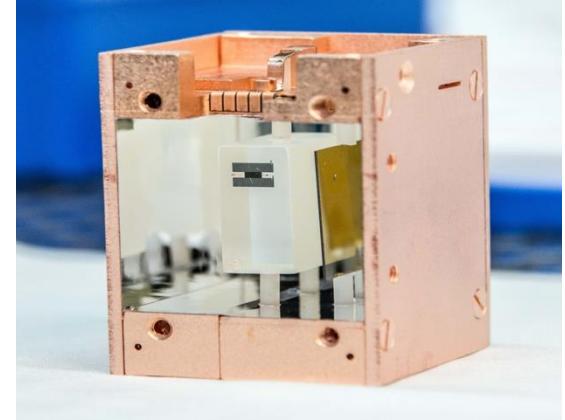
Screening campaign



CRESST set-up

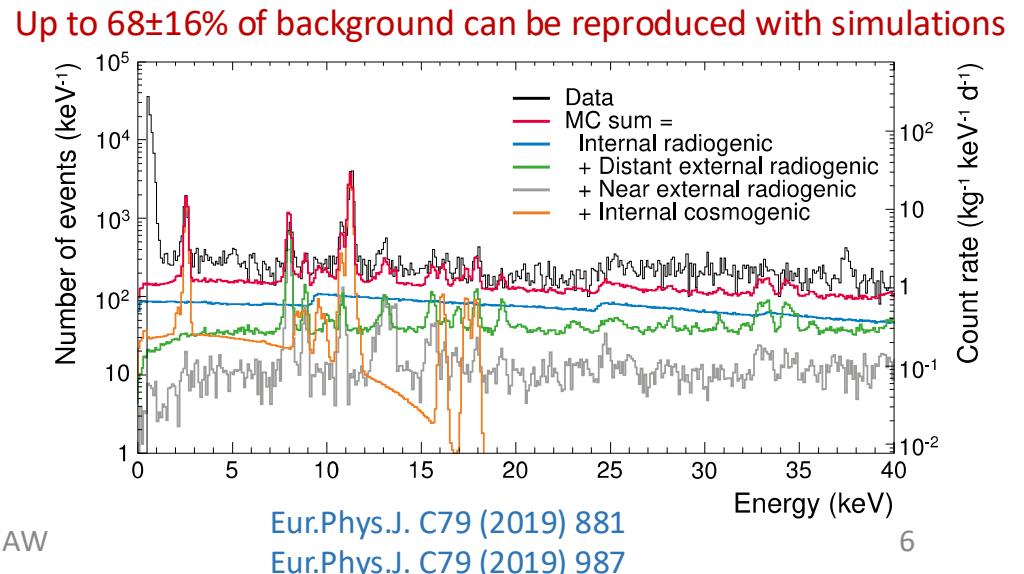
Materials:

- copper -> HPGe ✓ ICP-MS ✓ bulk ^{210}Po ✓ NAA ✓
- crystals (CaWO_4) -> bolometric meas. ✓
- reflective foil -> ICP-MS ✓
- bronze clams -> ICP-MS ✓
- polyethylene -> HPGe ✓ ICP-MS ✓
- lead -> ICP-MS ✓
- connectors -> HPGe ✓
- brass -> ICP-MS ✓
- pins -> ICP-MS ✓

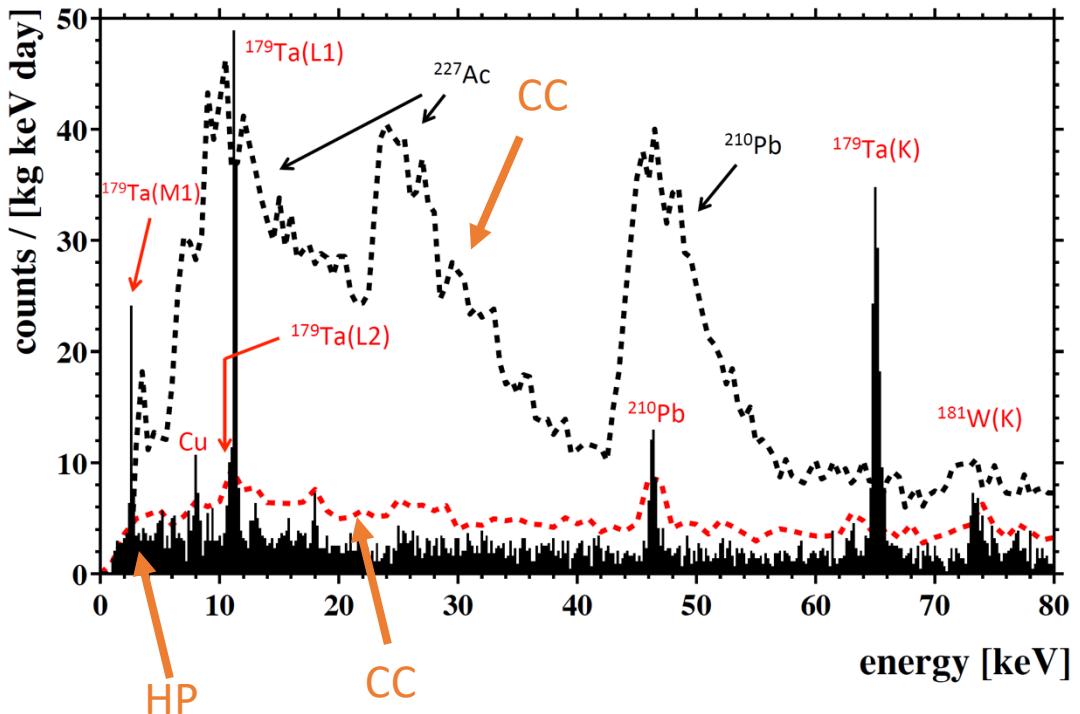


CRESST-III detector

Simulation



High-purity scintillating CaWO_4 crystals



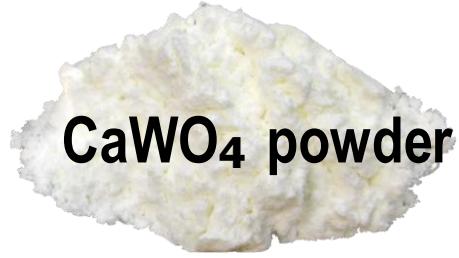
Commercial crystals (CC) → in-house production (HP)
A factor of 2-10 decrease in the background

Radiochemical Purification of CaCO_3

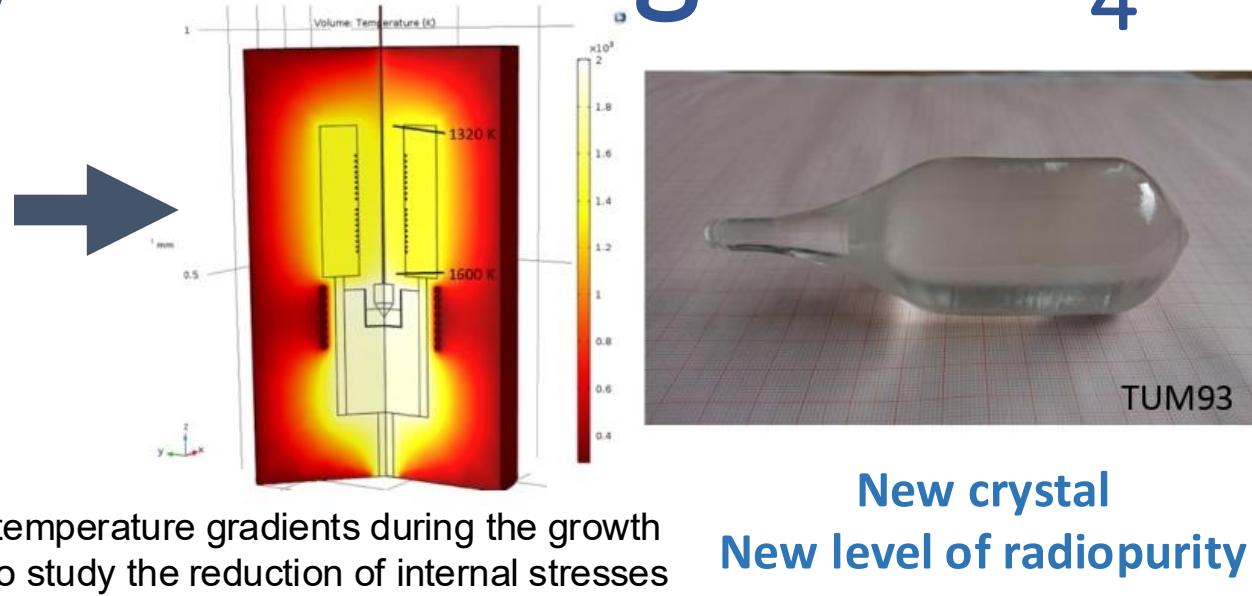
1. Transform CaCO_3 powder into aqueous solution of $\text{Ca}(\text{NO}_3)_2$
2. Mix solution with an extractor (TOPO) dissolved in n-Dodecan
→ Impurities move from the $\text{Ca}(\text{NO}_3)_2$ solution to the extractor solution
3. Extraction of $\text{Ca}(\text{NO}_3)_2$ solution
4. Remove precipitated CaWO_4
5. Washing with alkaline solution and water



High-purity scintillating CaWO_4 crystals



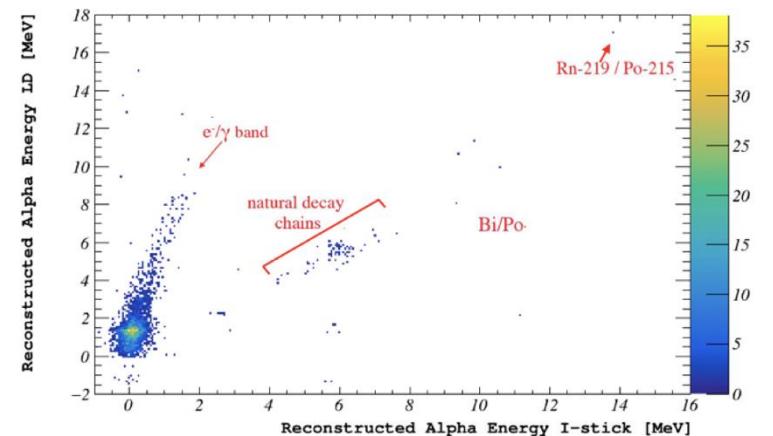
1.5 kg of purified CaWO_4 powder produced via co-precipitation using $(\text{NH}_4)_2\text{WO}_4$ solution (NEW!)



oxygen atmosphere at 1400°C for 20 h

New crystal
New level of radiopurity

Activity, (mBq/kg)	
Total alpha activity	0.535 ± 0.055
^{147}Sm and ^{180}W	0.046 ± 0.016
Single alpha lines	0.454 ± 0.051
Bi-Po cascades	0.029 ± 0.013
$^{219}\text{Rn} - ^{215}\text{Po}$ decay	0.006 (1 event)
Total alpha activity nat.	0.489 ± 0.053
decay chains	



Valentyna Mokina - HEPHY OEAW

Total α activity (3.08 ± 0.04) mBq/kg for TUM40 [1] and for TUM93 [2] crystal is (0.489 ± 0.05) mBq/kg.

Increase by a factor of 6.3 ± 0.7

[1] R. Strauss et al., JCAP 2015 06, 030 (2015)

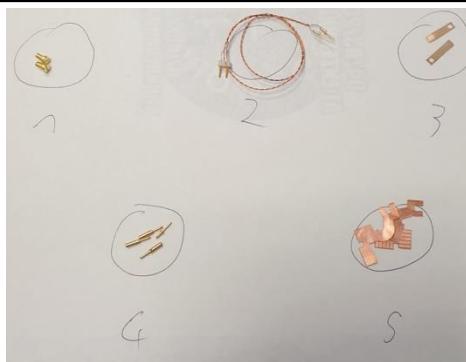
[2] A. Kinast et al., J LTPhys. (2022) 10.1007/s10909-022-02743-7

Screening campaign

Sample	Th [pg/g]	U [pg/g]
1. brass screw	70 ± 21	14 ± 4
2. superconducting cable with copper matrix	77 ± 23	84 ± 25
3. bronze clamp	162 ± 49	690 ± 207
4. pins	20 000 ± 6 000	176 000 ± 53 000
5. copper circuit	734 ± 220	283 ± 85

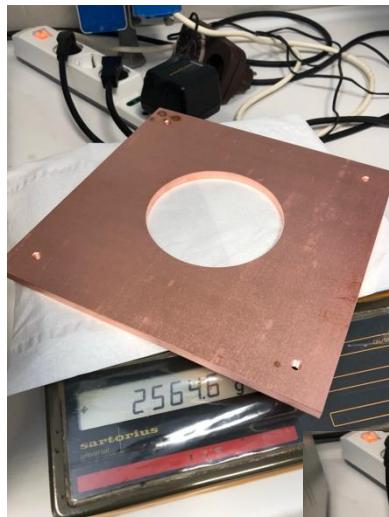


ICP-MS at LNGS

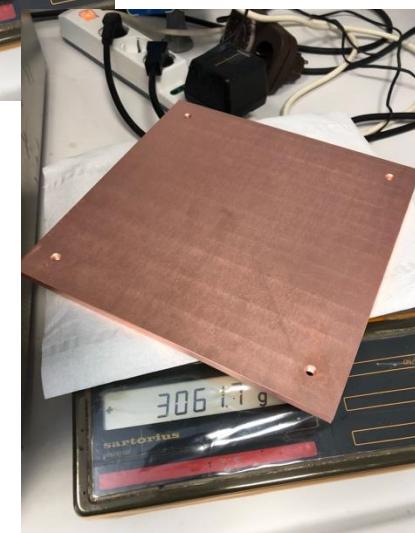


Sample	^{232}Th [$\mu\text{Bq/kg}$]	^{235}U [$\mu\text{Bq/kg}$]	^{238}U [$\mu\text{Bq/kg}$]
Copper*	< 2	< 0.43	< 6.2
Copper CUORE	< 2	< 0.46	< 65

* more sensitive result



16 plates



18 plates

Chain	Nuclide	Activity, (mBq/kg)
^{232}Th	^{228}Ra	<0.024
	^{228}Th	<0.021
^{235}U		<0.05
^{238}U	^{234}Th	<3.5
	$^{234\text{m}}\text{Pa}$	<0.76
	^{226}Ra	<0.02
	^{40}K	<0.19
	^{137}Cs	<0.0056
	$^{46}\text{Sc}(83.8\text{d})$	0.029 ± 0.006
	$^{48}\text{V}(15.97\text{d})$	<0.04
	$^{54}\text{Mn}(312\text{d})$	0.051 ± 0.009
	$^{59}\text{Fe}(44.5\text{d})$	0.042 ± 0.011
	$^{56}\text{Co}(77.2\text{d})$	0.054 ± 0.008
	$^{57}\text{Co}(272\text{d})$	< 0.14
	$^{58}\text{Co}(70.9\text{d})$	0.5 ± 0.05
	$^{60}\text{Co}(5.28\text{y})$	0.046 ± 0.006

HPGe at LNGS

ICP-MS and HPGe measurements of shielding materials

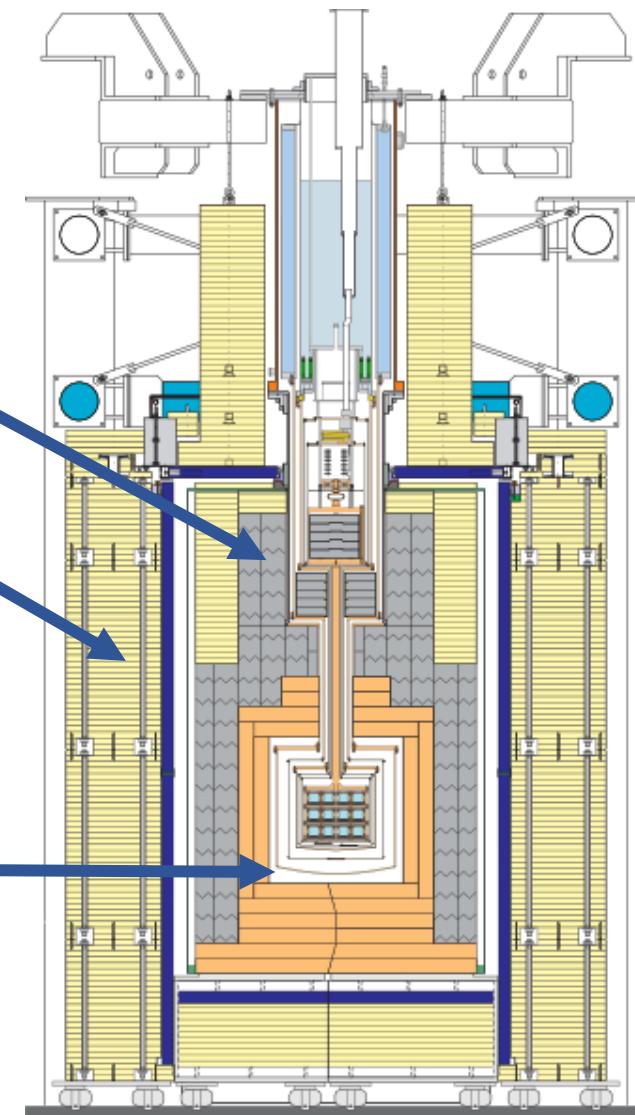
ICP-MS at LNGS

Needs an update!!

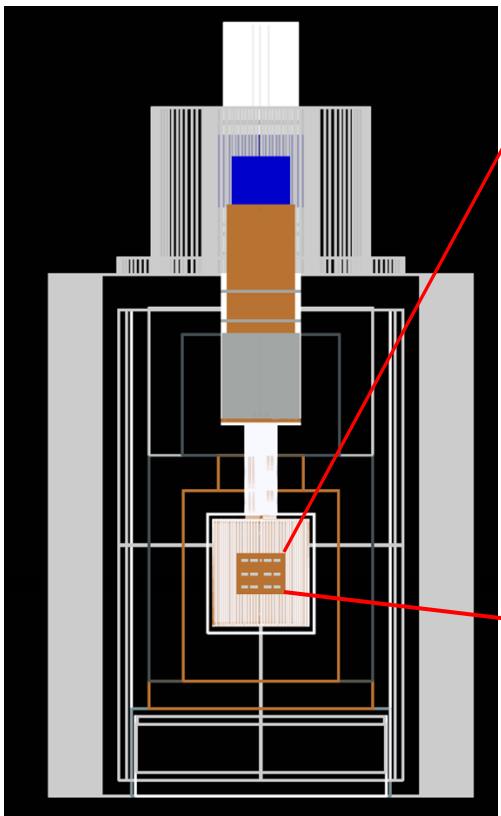
Sample	Th [pg/g]	U [pg/g]
lead	< 200	< 200
polyethylene outer	< 1300	< 500

HPGe at LNGS

Sample	Th [pg/g]	U [pg/g]
polyethylene inner	100 ± 30	82 ± 9



From one module to full setup simulation



DAWN visualization of the Carousel with detectors as implemented in Geant4

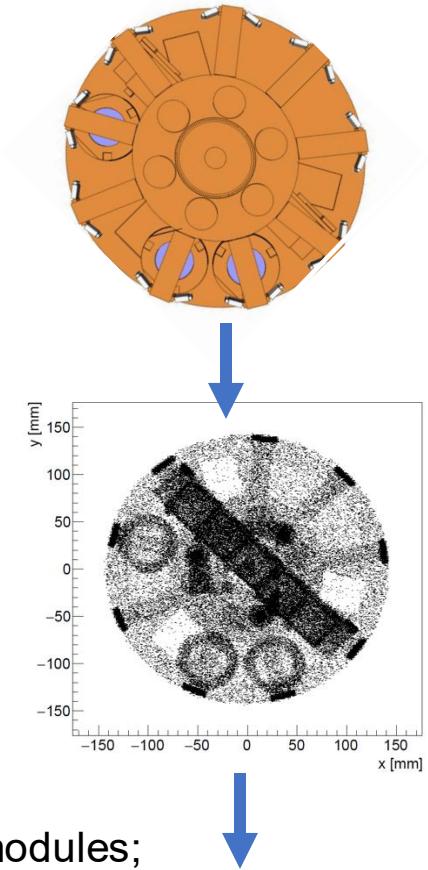


CRESST-II
phase 1
730kg result
Eur.Phys.J.
C72(2012)1971

CRESST-II
phase 2
TUM40, Lise
Eur.Phys.J.
C74(2014)3184,
C76(2016)25

CRESST-III
phase 1
Detector A
Phys.Rev.
D100(2019)102002

CRESST-III
phase 1 Run 2
and so on...



- Implementation of full geometry (each Run, all detectors);
- Adaption of the e.m. background model to the actual CRESST detector modules;

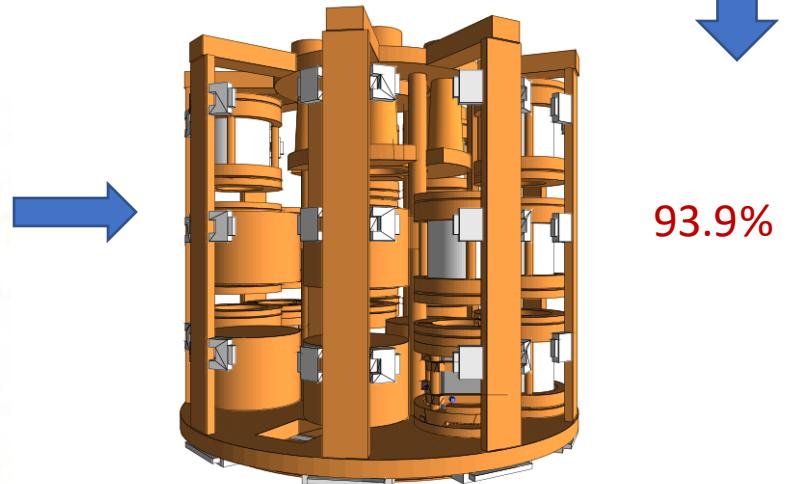
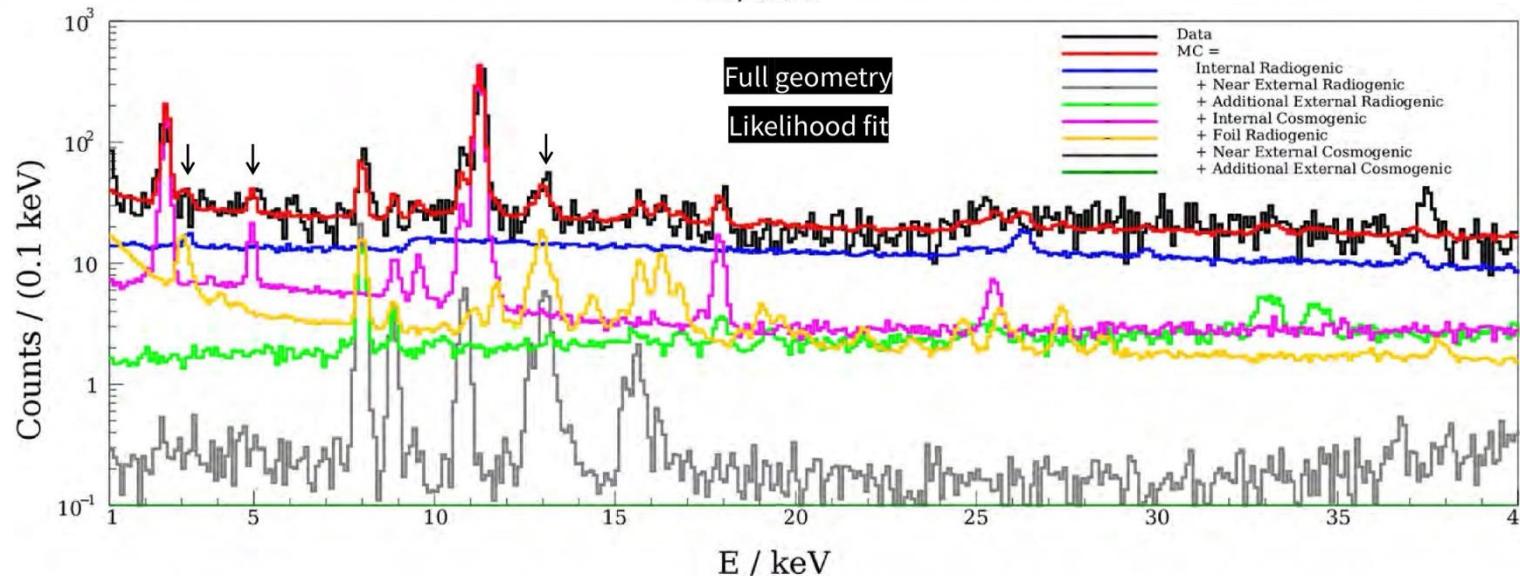
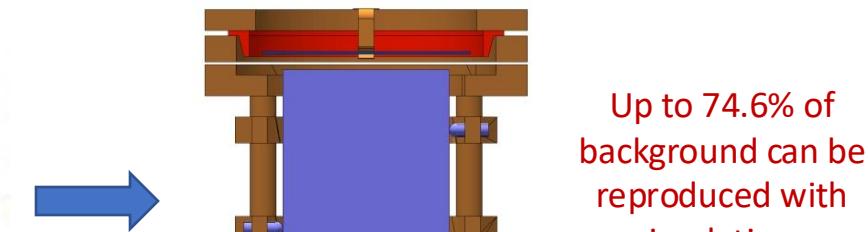
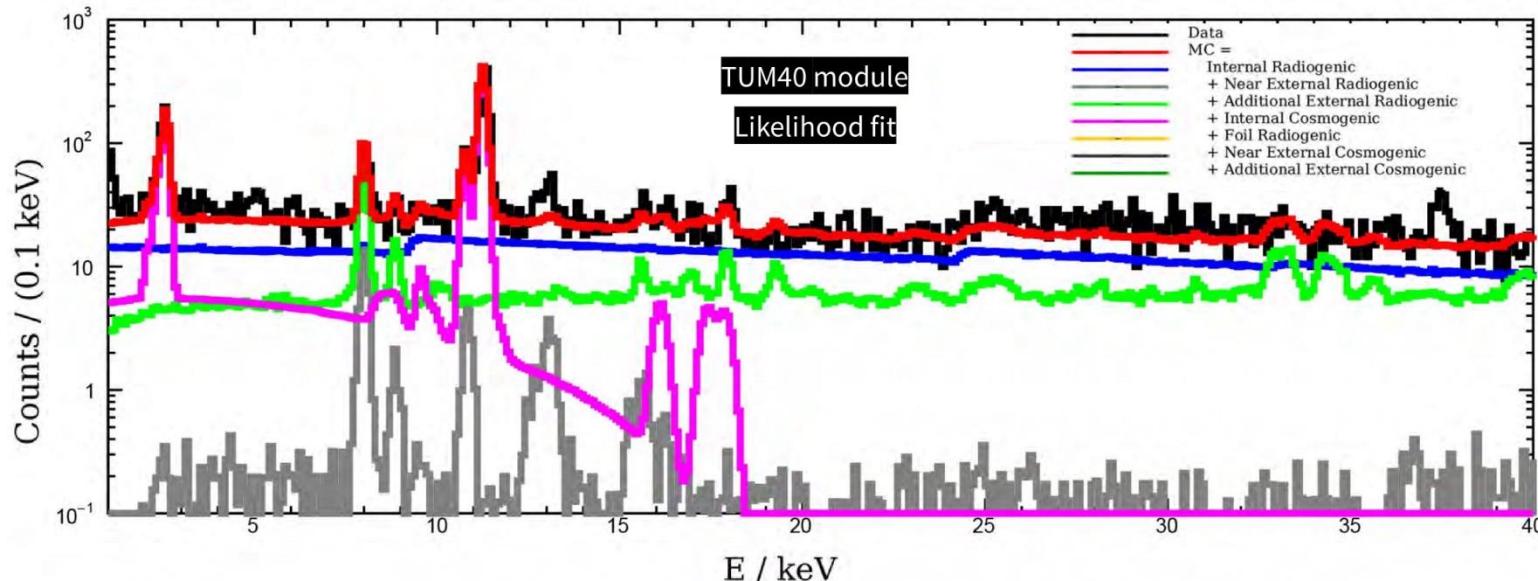
CRESST set-up



- Contamination levels from material assays conducted within;
- Simulation of neutron background;
- Study of cosmogenic activation of CaWO₄ crystal scintillator;
- Surface contaminations studies.

Simulating a homogeneous contamination in all parts made of Cu inside the Carousel.

Likelihood model of background



Conclusions

- CRESST operates a new generation of TUM-grown crystals with improved radiopurity due to chemical purification of their raw materials
- The screening campaign is ongoing to understand the activity concentration of different isotopes in materials used in the experiment
- The results of these studies are used as an input for simulation of the background of the CRESST experiment (development of sub-keV Monte Carlo model)
- Likelihood model of the background is developed and allows to reproduce up to 94%

Waiting for dark matter



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