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## A Unified Framework for Dwarf Galaxy Cores and Cluster Substructure Cusps via Self-Interacting Dark Matter with Mass Segregation

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Self-interacting dark matter (SIDM) provides a promising solution to small-scale structure anomalies, yet a unified explanation across mass scales remains challenging. Recent dwarf galaxy surveys favor a relatively small cross section of  $\sim 0.3 {\rm cm}^2/{\rm g}$ —insufficient to induce core collapse—while strong lensing signals in galaxy clusters require dense inner subhalos typically associated with much larger cross sections. We demonstrate that this tension can be resolved in two-component SIDM models, where even modest inter-component interactions lead to efficient mass segregation in massive halos. This enhancement occurs without requiring large self-interaction rates, as the strength of SIDM effects scales with the scale density  $\rho_s$  and radius  $r_s$  as  $\sigma_{\rm eff}/m~r_s\rho_s^{3/2}$ . Consequently, the impact is amplified in cluster subhalos, where both  $\rho_s$  and  $r_s$  are large, while remaining insignificant in dwarf galaxies. Using a set of cosmological zoom-in simulations of clusters and high-resolution controlled simulations of dwarf halos, we show that this framework can simultaneously reproduce cored dwarfs and cuspy cluster substructures. Our results highlight two-component SIDM with mass segregation as a unified and testable framework for explaining structural features of dark matter halos that are otherwise puzzling.

## Collaboration you are representing

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