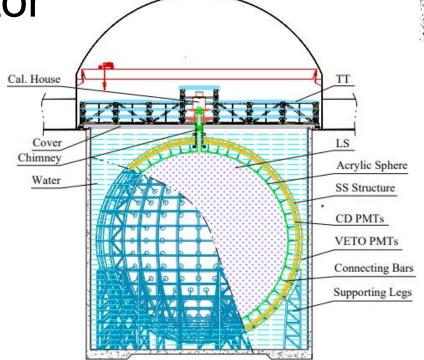
# R&D program for neutrinoless double beta decay search at JUNO

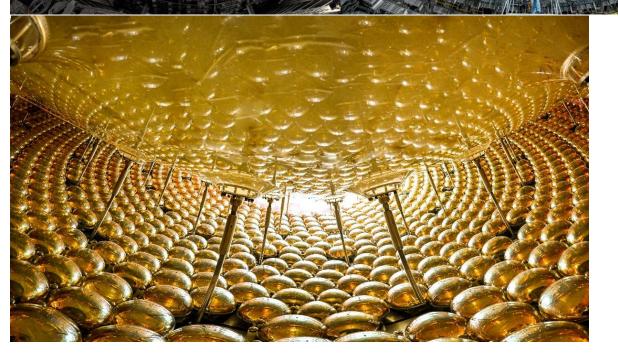




#### **JUNO Detector**







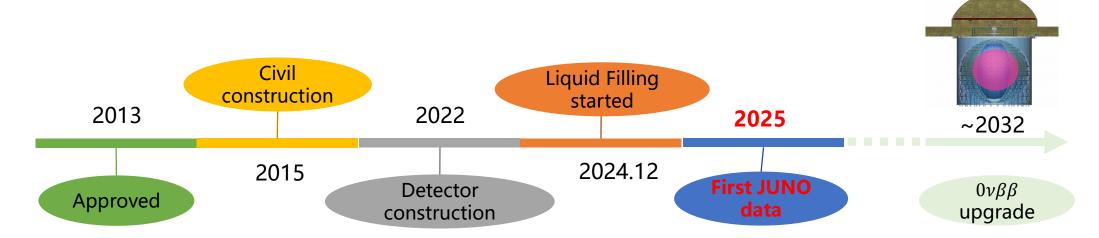
- Largest LS detector ever (20 kton), designed for NMO determination with reactor  $\overline{\nu_e}$
- 650m rock overburden underground
- 35.4m acrylic vessel covered by 20"+3" PMTs (75%+3% photocoverage)



#### JUNO Status



- 10 years endeavor of construction of the experiment
- Ready to start physics data taking in August
- Determine NMO with 6.5 years of data
- Upgrade for 0vbb decay search after NMO determination

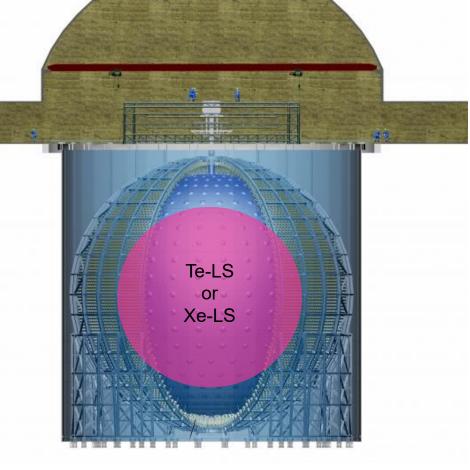




JUNO- $0\nu\beta\beta$  upgrade

 JUNO offers an unique opportunity to search for 0vββ

- 20 kton LS → 100-ton scale isotope loading
  - e.g., <u>Tellurium OR Xenon</u>
- Low background
- Energy resolution < 3% @ 1 MeV →</li>
   2.4x better than KamLAND-Zen
- Critical R&D
  - Te route: Isotope loading technology
  - Xe route: cost effective enrichment & clean balloon
  - Background control



Concept of the experiment



## TeLS R&D program



2020: Azeotropic distillation approach developed → >5% Te mass loading

2021: Absorbance at 430 nm unchanged for 0.6% Te loading

2022~2023: R.T. synthesis approach established to decrease safety risks

2024: Synthesis and purification protocol established for Te-LS production

Exploratory stage

Technical Breakthrough

Application-oriented

2017 - 2019

2020 - 2024

2025 -

2017: inorganic Te + surfactant

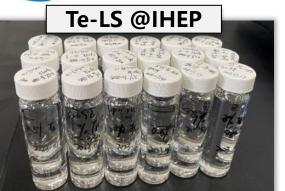
2018: liquid-liquid extraction

2019: organic Te-diol compounds

2025: 100 kg production demonstration

<u>Lab studies</u>: L.Y. optimization, stability test (>3
years), Te-LS transparency characterized by A.L.,
develop methods for radioactive impurities removal

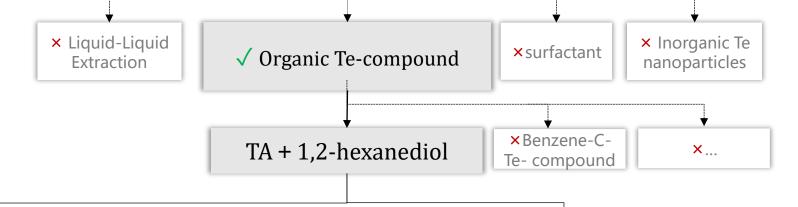






Excep. high Te-doping conc.(3~5%)
No Impact on LS Optical Perf.
Ultra-low rad. Background
Stability >10 years





Technical Route 1: Azeotropic distillation approach

**Innovation Highlights:** water-free reaction → enhanced solubility and stability

**Advantages**: Te>>5%, Uniform

Transparent, and Stable

Patent: ZL 202011370855.5

Publication: NIM A 1049 (2023) 168111

Technical Route 2: Room temp. Syn. approach

#### **Innovation Highlights:**

No heating required, R.T. Synthesis

**Advantages**: Performance unchanged,

greener & more scalable

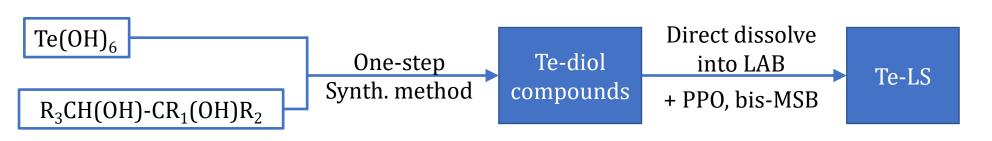
Patent: applied

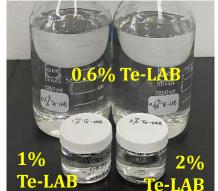
Publication: *in preparation* 



## Route 1: Azeotropic distillation approach







- Water-free environment for the synthesis
- No extra water introduced into the reaction system
  - Solid state telluric acid used directly instead of being dissolved in the water
- Water generated removed continuously by azeotropic distillation
  - Azeotropic solvent → Acetonitrile, for its relatively lower boiling point (81.6 °C) and higher water content (16%) in azeotrope

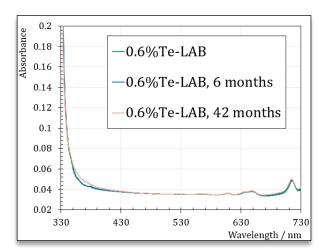


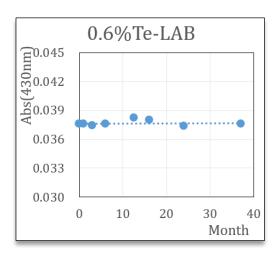
#### Route 1: performance



- Ultrahigh Te Solubility
  - miscible with LAB, capability of Te loading in JUNO LS: > 5%
- Exceptionally High Transparency & Long-Term Stability
  - limited influence on the absorbance of scintillation solvent and stable for 3.5 years
- A one-step synthesis
  - simple, the product can be used for preparing Te-LS directly without post-processing, easily Scalable for mass production
- Broad applicability to diverse diols
  - a serials of diols with more sophisticated structures can be used since the reaction is greatly facilitated by the removal of water





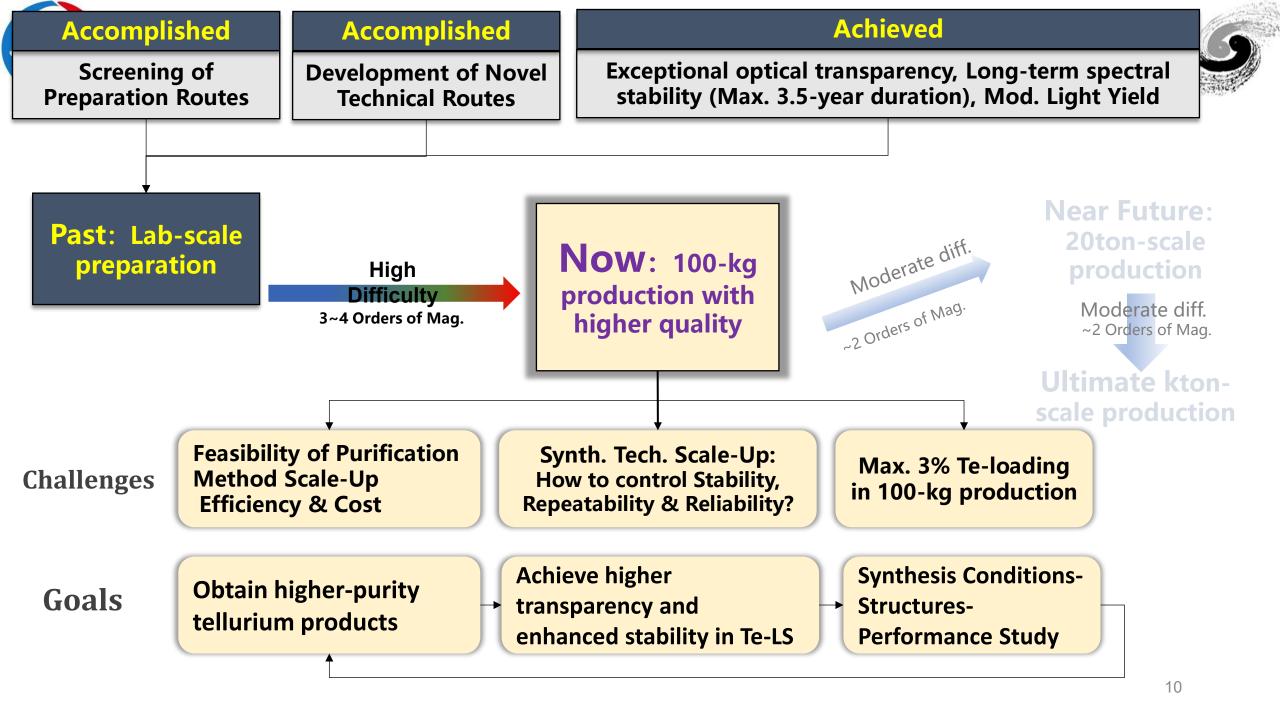




## Route 2: Room temperature Synthesis approach



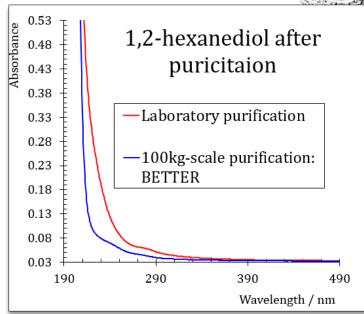
- **Disadvantage** of Azeo. **Dist. Approach**: safety risks associated with large-scale heating and distillation of low-boiling-point solvents (acetonitrile, 81.6 ∘C)
- Developed a Room-Temperature Synthesis approach
  - one-step synthesis at room temperature
- Excellent properties of Te-samples:
  - Exceptional optical transparency (ΔAbs(430nm)/%Te ≤ 0.0003)
  - Long-term spectral stability exceeding or approaching 1 year till now for both 1% and 3%
     Te formulations
  - Outperform even those produced by the Azeo. Dist. Method
- A green, efficient alternative for large-scale Te-LS production for nextgeneration neutrinoless double-beta decay experiments

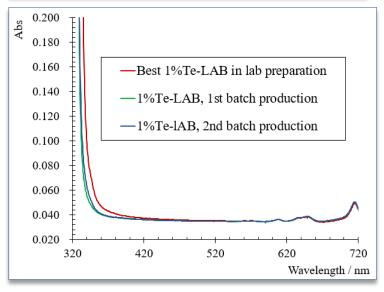




#### Progress of 100-kg production

- Bulk purification methods for raw materials with corresponding QC standards and protocols developed
  - Extensive research on optical purification method
  - A combined purification approach for both purification efficiency and cost
  - Dramatic cost reduction with superior purity compared to expensive lab purification
- Dual-Temperature (Room Temp./High Temp.) Reactor was designed and customized
- 100kg-scale production in 10 batches in progress
  - Higher-transparency Te-LAB samples were obtained, >50 kg
  - Te-LAB samples from 1<sup>st</sup> and 2<sup>nd</sup> production exhibit absorption spectra far superior to the best lab samples.
  - Attenuation length of Te-LAB (0.5% doping) reaches 20m @ 430nm
  - L.Y. measurement is ongoing



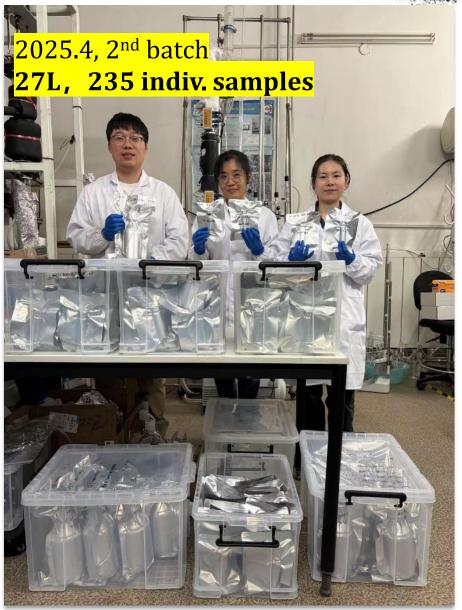




#### Te samples w/ diff. formulations were prep. for mult. research purposes









#### Backgrounds



- Full background evaluation
  - <sup>136</sup>Xe loading as an example in 2016
  - 130Te loading estimate ongoing
- Advantage of large LS based detector 

   negligible external material background
- Background contributions:
  - Cosmogenic isotope
    - Updates on spallation background on Te/Xe
  - B solar ν-e scattering
  - 2νββ
  - Internal LS background

Table 5. Summary of the projected backgrounds in the  $0\nu\beta\beta$  ROI. For light cosmogenic isotopes, the values are from GEANT4 MC, while for FLUKA MC the total residual background would increase  $0.07/\text{ROI/(ton}^{136}\text{Xe)/yr}$ .

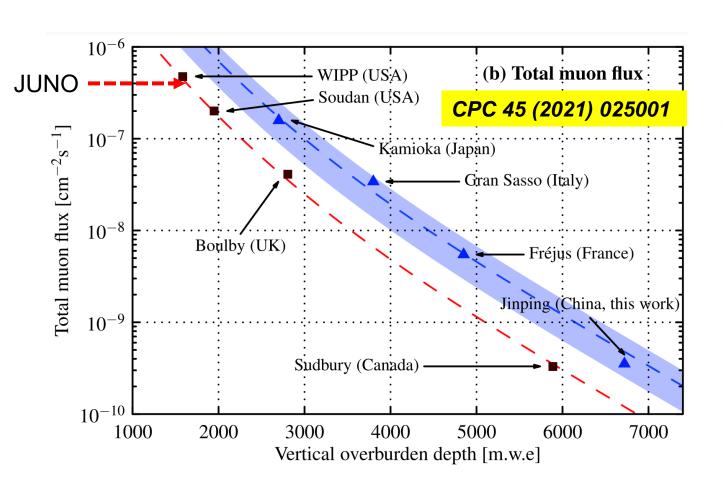
summary of backgrounds in 0νββ ROI				
$[ROI\cdot(ton\ ^{136}Xe)\cdot yr]^{-1}$				
2νββ	0.2			
$^8{ m B~solar}~\nu$	0.7			
cosmogenic background				
$^{10}\mathrm{C}$	0.053			
$^6\mathrm{He}$	0.063			
$^8{ m Li}$	0.016			
$^{12}\mathrm{B}$	$3.8 \times 10^{-4}$			
others $(Z \leq 6)$	0.01			
$^{137}\mathrm{Xe}$	0.07			
internal LS radio-purity $(10^{-17} \text{ g/g})$				
<sup>214</sup> Bi ( <sup>238</sup> U chain)	0.003			
$^{208}$ Tl ( $^{232}$ Th chain)	_			
<sup>212</sup> Bi ( <sup>232</sup> Th chain)	0.03			
external contar	external contamination			
<sup>214</sup> Bi (Rn daughter)	0.2			
total	1.35			



## Cosmogenic background on <sup>12</sup>C



- JUNO: 650 m rock overburden
- Long-lived  $\mu$ -spallation isotope could become background



**Table A9.** The estimated rates for cosmogenic isotopes in JUNO LS by FLUK! simulation, in which the oxygen isotopes are neglected. The decay modes and Q value are from TUNL Nuclear Data Group [475].

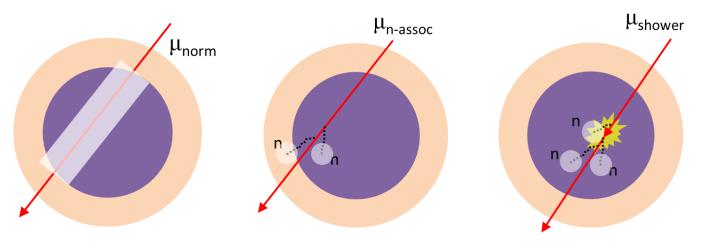
Isotopes	Q (MeV)	$T_{1/2}$	Rate (per day)
<sup>3</sup> H	0.0186 (β <sup>−</sup> )	12.31 year	$1.14 \times 10^{4}$
<sup>6</sup> He	3.508 ( <i>β</i> <sup>−</sup> )	0.807 s	544
<sup>7</sup> Be	$Q_{EC} = 0.862 (10.4\% \ \gamma, E_{\gamma} = 0.478)$	53.22 d	5438
<sup>8</sup> He	10.66 ( $\beta^-\gamma$ : 84%), 8.63 ( $\beta^-n$ : 16%)	0.119 s	11
<sup>8</sup> Li	16.0 (β <sup>-</sup> )	0.839 s	938
в	16.6 (β <sup>+</sup> )	0.770 s	225
<sup>9</sup> Li	13.6 ( $\beta^-$ : 49%), 11.94 ( $\beta^- n$ : 51%)	0.178 s	94
°C	15.47 ( $\beta^+p$ : 61.6%, $\beta^+\alpha$ : 38.4%)	0.126 s	31
<sup>10</sup> Be	0.556 (β <sup>−</sup> )	1.51e6 year	1419
<sup>10</sup> C	2.626 ( $\beta^{+}\gamma$ )	19.29 s	482
11Li	20.55 ( $\beta^- n$ : 83%, $\beta^- 2n$ : 4.1%)	0.00875 s	0.06
<sup>11</sup> Be	11.51 ( $\beta^-\gamma$ : 96.9%), 2.85 ( $\beta^-\alpha$ : 3.1%)	13.76 s	24
<sup>11</sup> C	0.960 (\beta^+)	20.36 min	$1.62 \times 10^4$
<sup>12</sup> Be	11.708 ( $\beta^-\gamma$ , $\beta^-n$ : 0.5%)	0.0215 s	0.45
$^{12}B$	13.37 $(\beta^-\gamma)$	0.0202 s	966
<sup>12</sup> N	16.316 ( $\beta^{+}\gamma$ )	0.0110 s	17
$^{13}B$	13.437 $(\beta^-\gamma)$	0.0174 s	12
<sup>13</sup> N	1.198 ( <i>β</i> <sup>+</sup> )	9.965 min	19
$^{14}B$	20.644 ( $\beta^-\gamma$ , $\beta^-n$ : 6.1%)	0.0126 s	0.021
<sup>14</sup> C	0.156 ( <i>β</i> <sup>−</sup> )	5730 year	132
<sup>15</sup> C	9.772 (β <sup>−</sup> )	2.449 s	0.6
<sup>16</sup> C	8.010 (β <sup>-</sup> n: 99%)	0.747 s	0.012
<sup>16</sup> N	$10.42~(\beta^-\gamma)$	7.130 s	13
<sup>17</sup> N	8.680 ( $\beta^-\gamma$ : 5%), 4.536 ( $\beta^-n$ : 95%)	4.173 s	0.42
<sup>18</sup> N	13.896 ( $\beta^-\gamma$ : 93%), 5.851 ( $\beta^-n$ : 7%)	0.620 s	0.009
Neutron			155 000



### Background on <sup>12</sup>C after Veto



- Excellent  $\mu$  tagging and tracking capability
- Dedicated veto strategies for different types of muons
- Major isotopes can be efficiently rejected



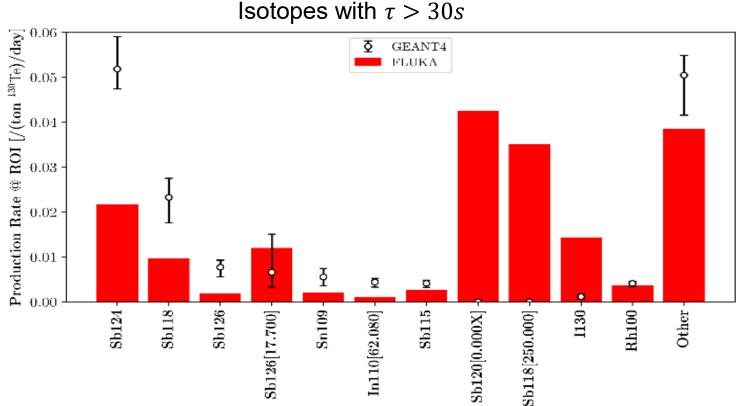
Refs: arXiv:2006.11760, Chin. Phys. C 45 (2021) 023004 arXiv:1610.07143, Chin. Phys. C 41 (2017) 053001

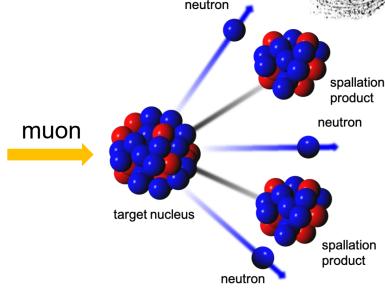
Cosmogenic Isotopes	Background Index unit: ROI-1 (ton 136Xe)-1 yr-1		
	No veto	w/ veto	
<sup>10</sup> <b>C</b>	16.4	0.053	
<sup>6</sup> He	4.9	0.063	
<sup>8</sup> Li	1.5	0.016	
<sup>12</sup> B	1.9	3.8e-4	
<sup>137</sup> Xe	2.3	0.07	
Others (Z≤6)	0.51	0.01	
Total		0.2	



## Te spallation isotopes yield calculation

- Muon induced spallation isotopes from Te
- Isotope yield calculation with FLUKA and GEANT
  - Variations among models
  - Identification of major long lived isotopes



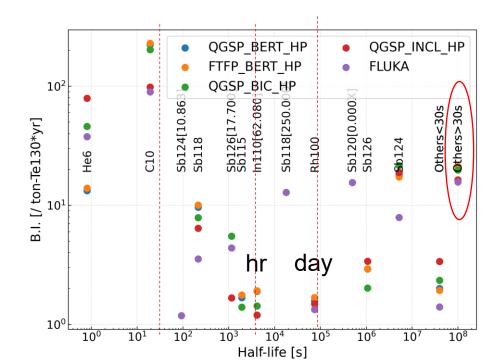


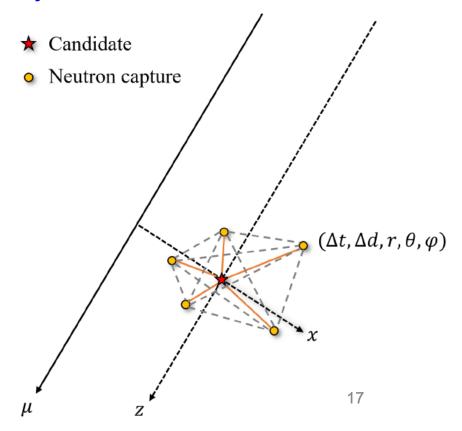


## Development of Rejection method



- Understanding of shower and attempts to reject background
  - Using accompanying neutrons information
  - Powerful rejection achieved for short lived isotopes
    - 98% bkg. rej. @ ~80% sig eff. with BDT
    - Advanced ML like GNN and transformer expected to improve further
  - Long lived isotopes are more difficult for event-by-event rejection
    - · Large number of preceding muons, convection et al
    - Multi-site discrimination will help

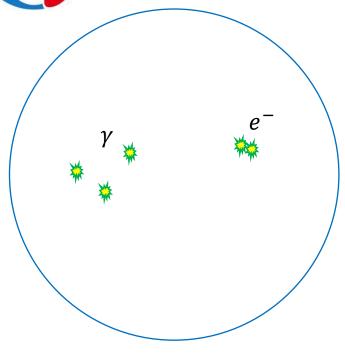




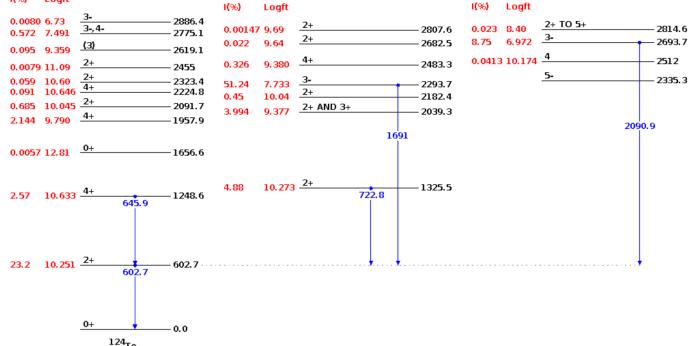


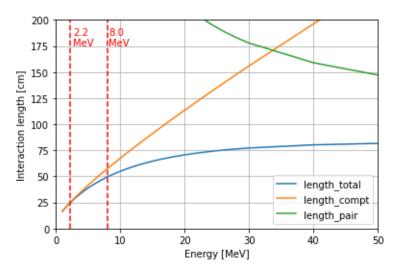
#### Multi-site discrimination











- Interaction length of ~1-2 MeV gammas in LAB 10-30 cm
   → 0.5-1.5 ns time of flight
- Timing resolution: ~1ns (Dynode PMTs)
- Further suppression of ultra-long-lived isotope is promising
- Analysis ongoing



#### Summary



- After years of efforts, promising TeLS technical routes have been established
  - High loading concentration (>5% Te), excellent optical performance (20m A.L. 0.5% Te) and stability (>3 years 0.6% Te) achieved
  - Core technologies on enhancing transparency and stability mastered
  - Light yield optimization is our next goal
- We made comprehensive FLUKA/GEANT study on cosmogenic spallation isotope background and developed efficient rejection strategy
- We are updating the sensitivity evaluation with more realistic and up-to-date JUNO detector information
- JUNO has great potential to explore the  $|m_{\beta\beta}|$ -meV region w/ >100 tons of  $0\nu\beta\beta$  isotope with clear route for technologies R&D
- Stay tuned!