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# Search for Double Electron Capture in Sn-112 using gamma-ray TES

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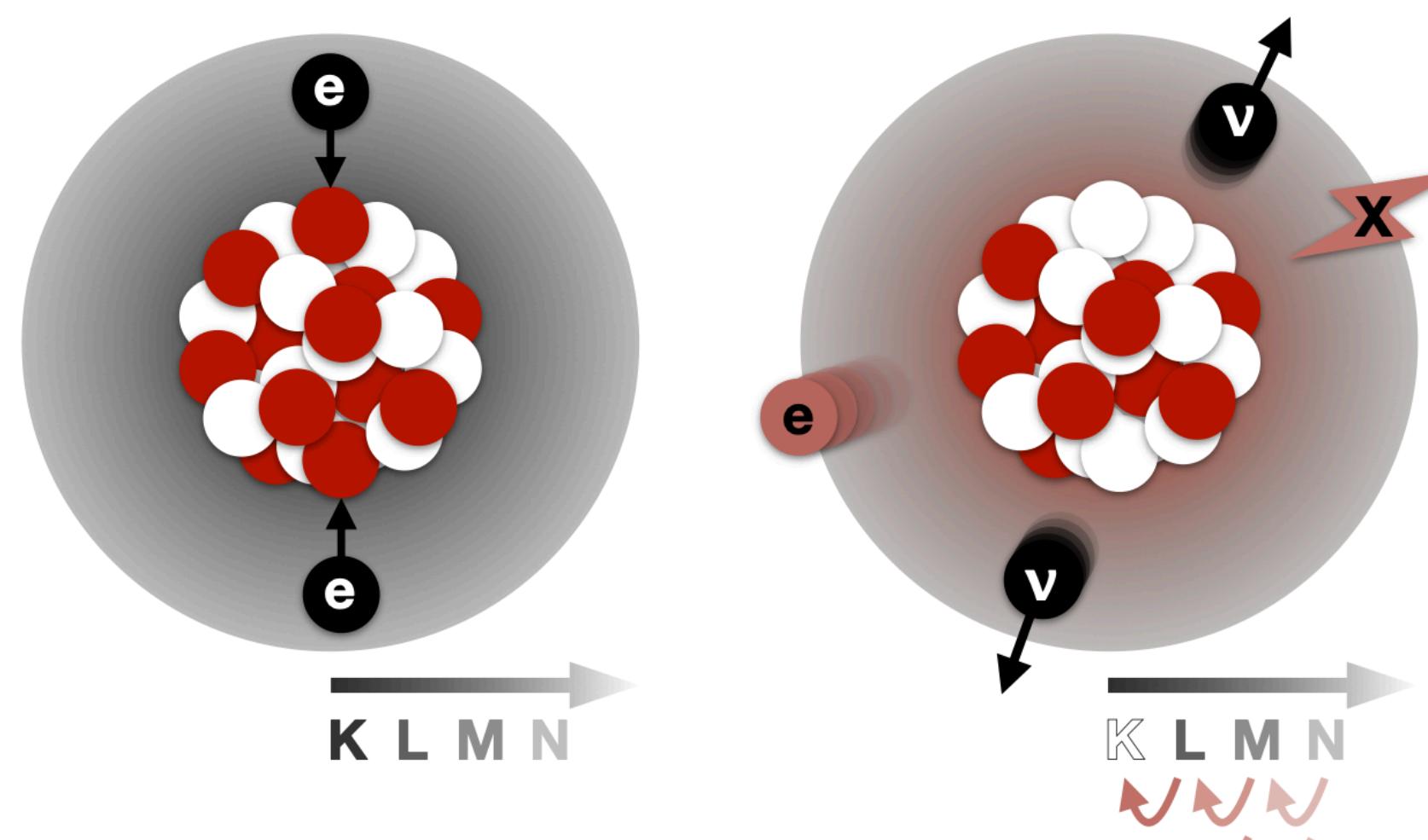
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# Motivation

# Double Electron Capture

## Double Electron Capture (ECEC)

- $(A, Z) + 2e^- \rightarrow (A, Z-2) + 2\nu_e$  (+X-rays, Auger electrons, O(10 keV))
    - 34 nuclei are known to be candidates for ECEC.
    - Direct observation : Only in  $^{124}\text{Xe}$  ( $T_{1/2} = 1.8 \times 10^{22}$  year)[1]
- Observation of  $2\nu$ ECEC would provide valuable information



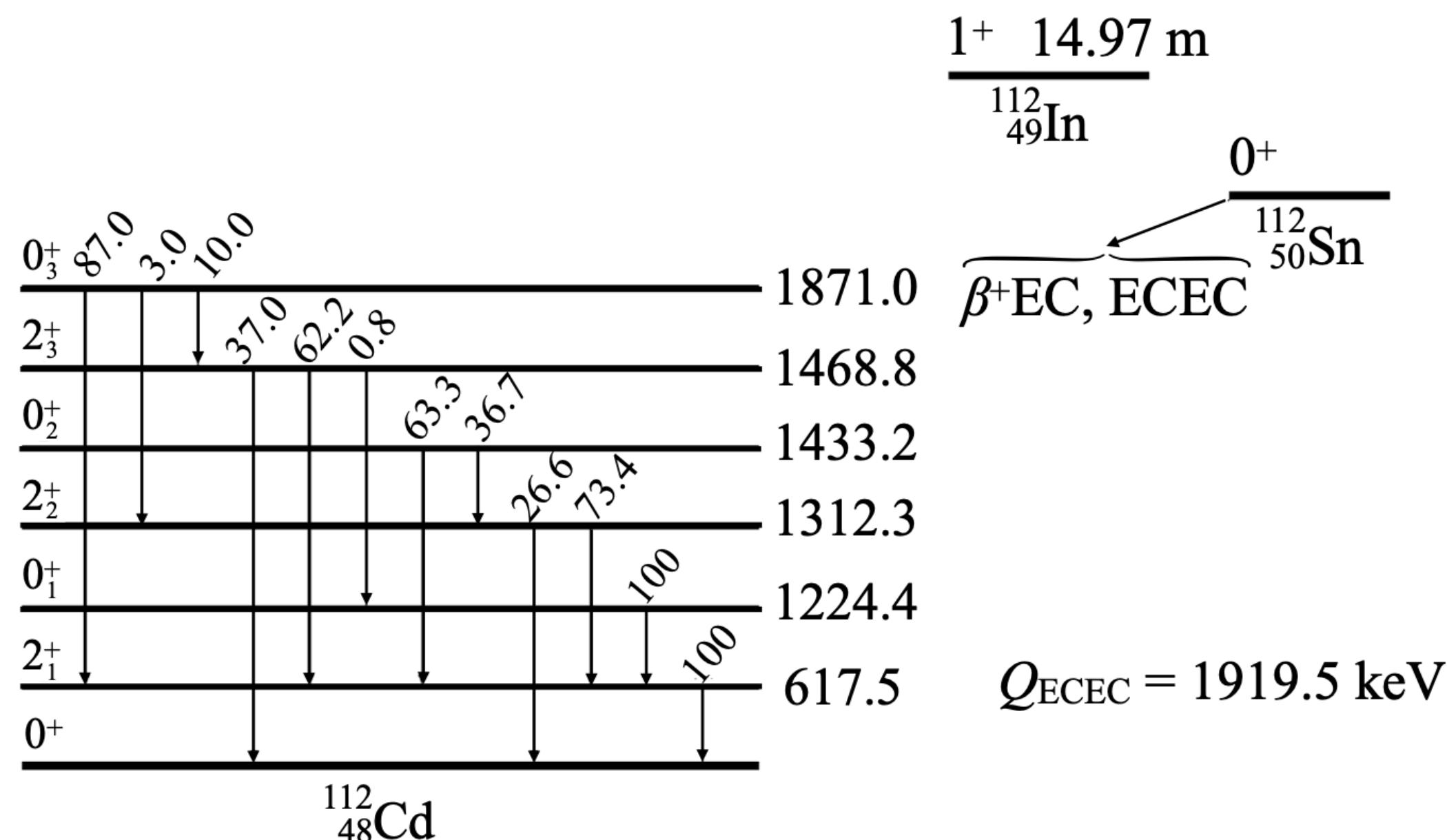
- If neutrino is Majorana particle,  $0\nu$ ECEC is also allowed :  $(A, Z) + 2e^- \rightarrow (A, Z-2)$
- Can be an access to physics BSM.

# Double Electron Capture in $^{112}\text{Sn}$

$^{112}\text{Sn}$  : Stable isotope of tin with the natural abundance of 0.0097

- Decay to  $^{112}\text{Cd}$  with ECEC or  $\beta^+ + \text{EC}$  ( $Q=1.92 \text{ MeV}$ )
- Previous study : Searched for  $\gamma$ -rays from de-excitation of  $^{112}\text{Cd}$  w/ HPGe detector[2]
  - Longer half-lives than decay to the ground state
  - Low signal efficiency : typically a few percents, at highest 13.6%

We use gamma-ray TES to overcome these difficulties.

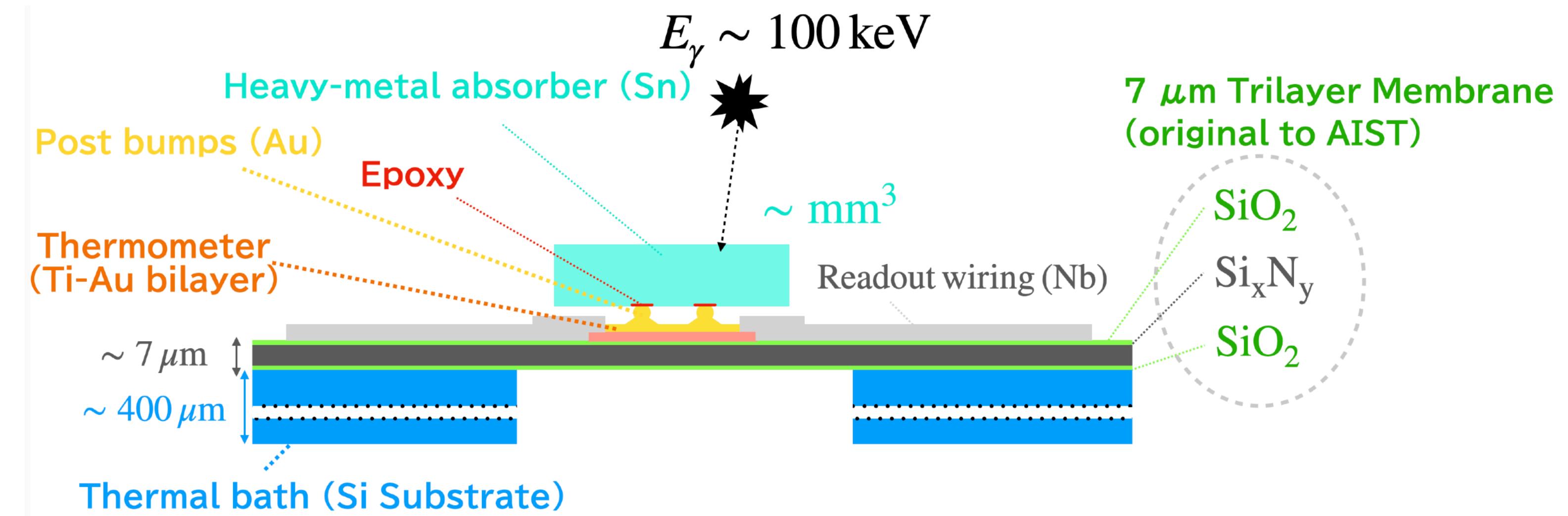
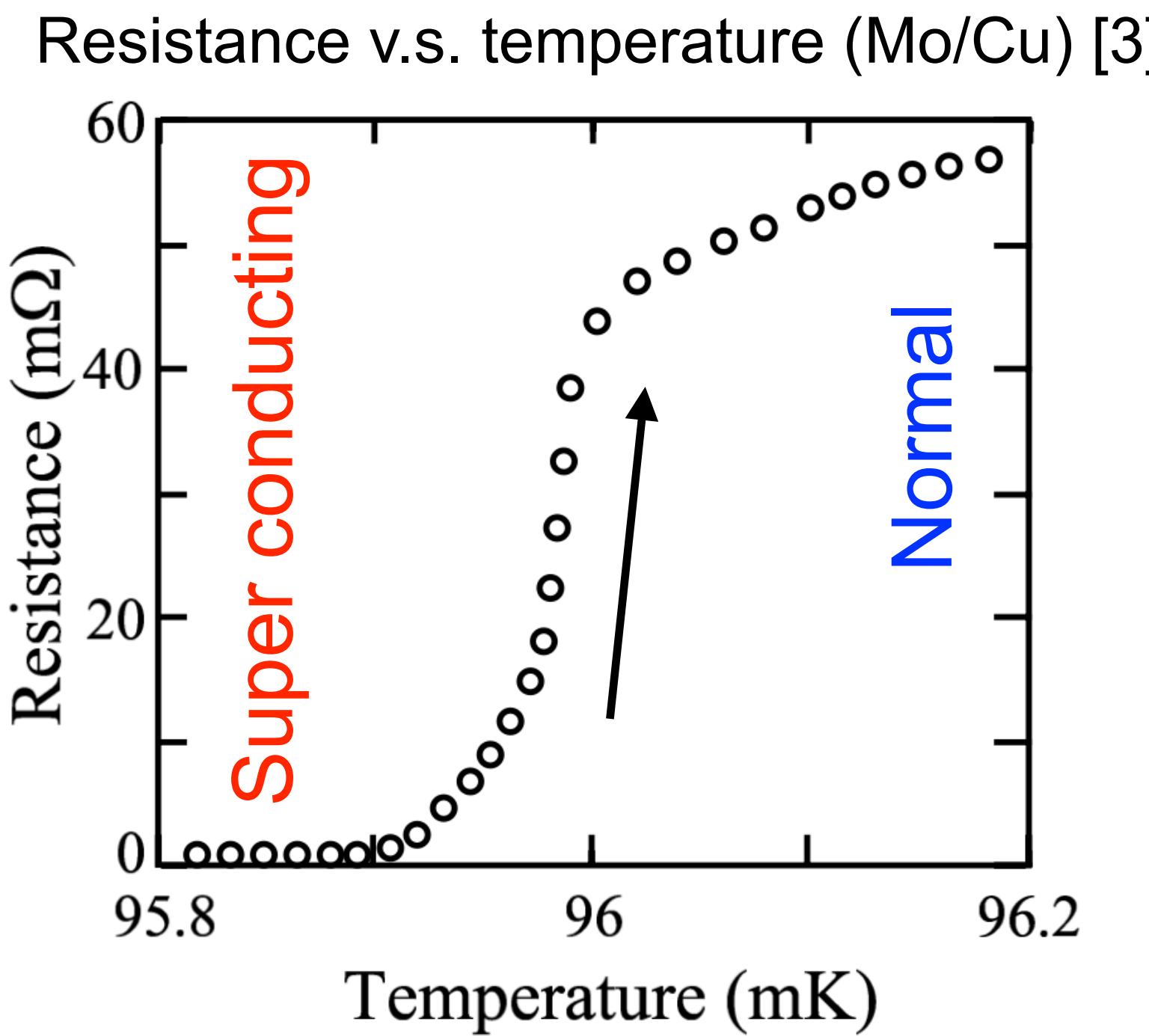


States of $^{112}\text{Cd}$	Half-lives (Theory) [yr]
g.s.	$1.7 \times 10^{22}$
$0_1^+$	$7.4 \times 10^{24}$
$2_1^+$	$4.9 \times 10^{28}$
$2_2^+$	$1.9 \times 10^{32}$

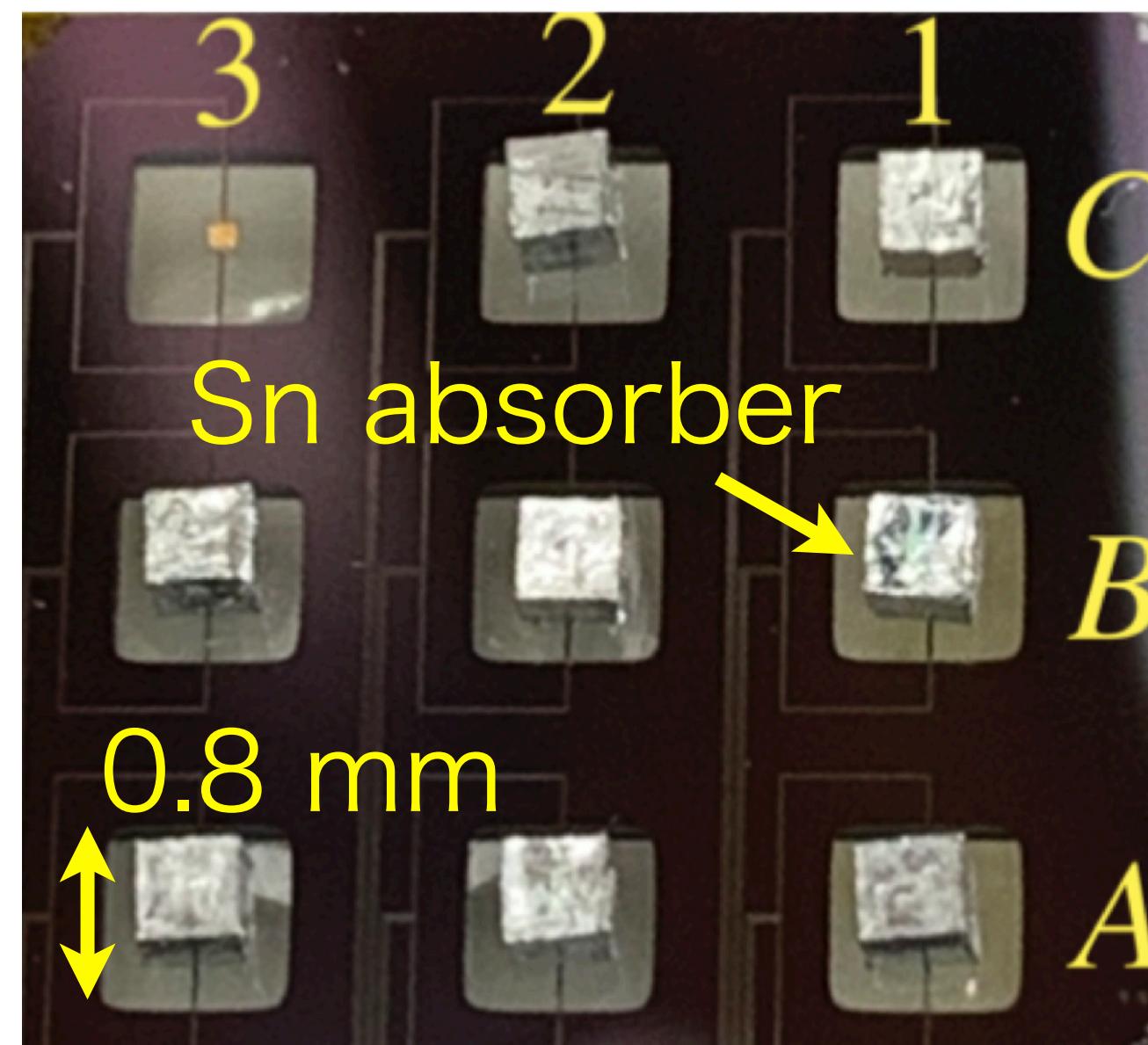
# What is gamma-ray TES?

## Gamma-ray TES:

- TES specialized in gamma-ray spectroscopy
  - TES + heavy-metal made absorber
- gamma-ray TES developed at AIST\* : 8-pixels TES array with *nat.*Sn energy absorber  
(Absorber size : 0.8 mm,  $3.7 \times 10^{-3}$  g/pixel *nat.*Sn)

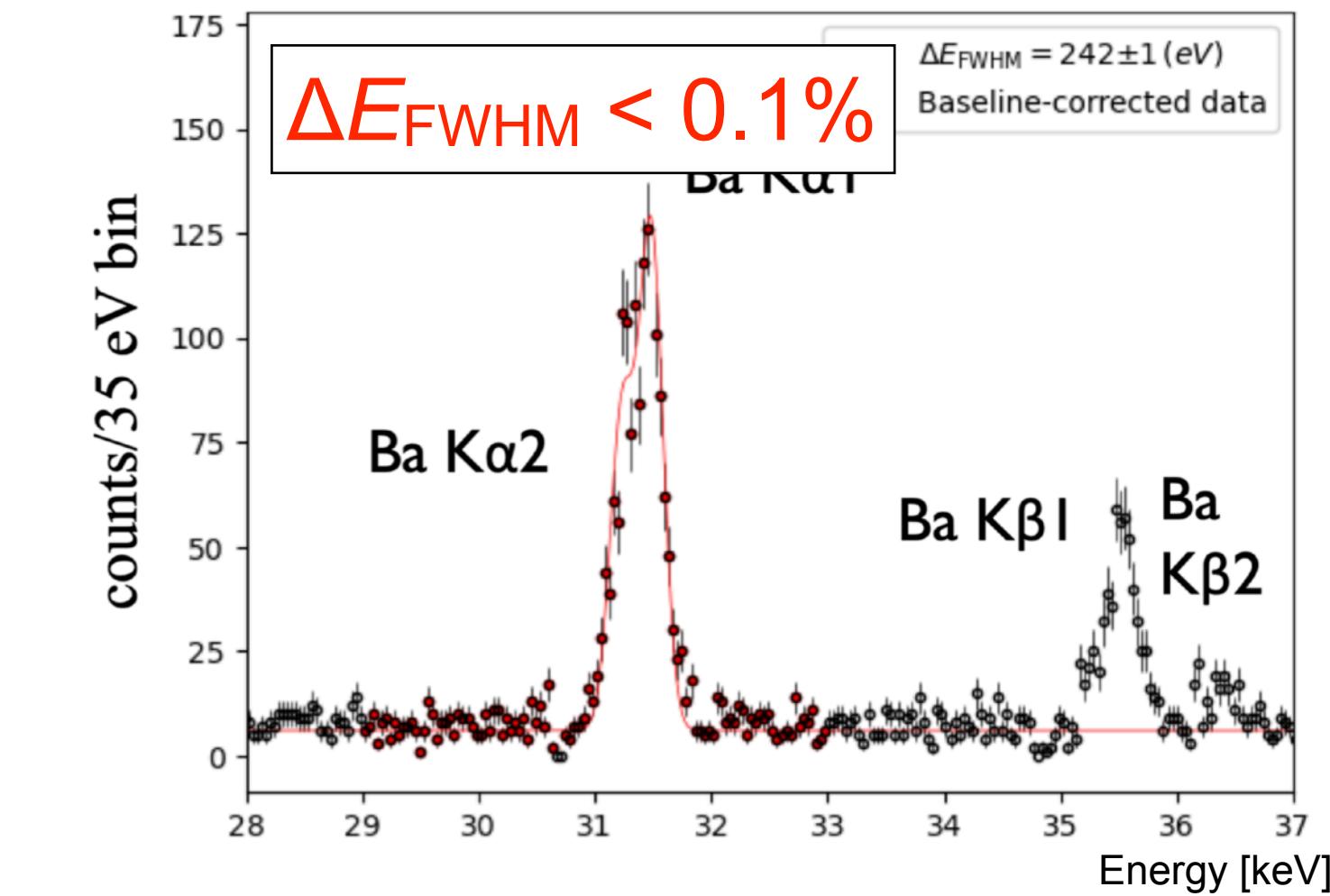
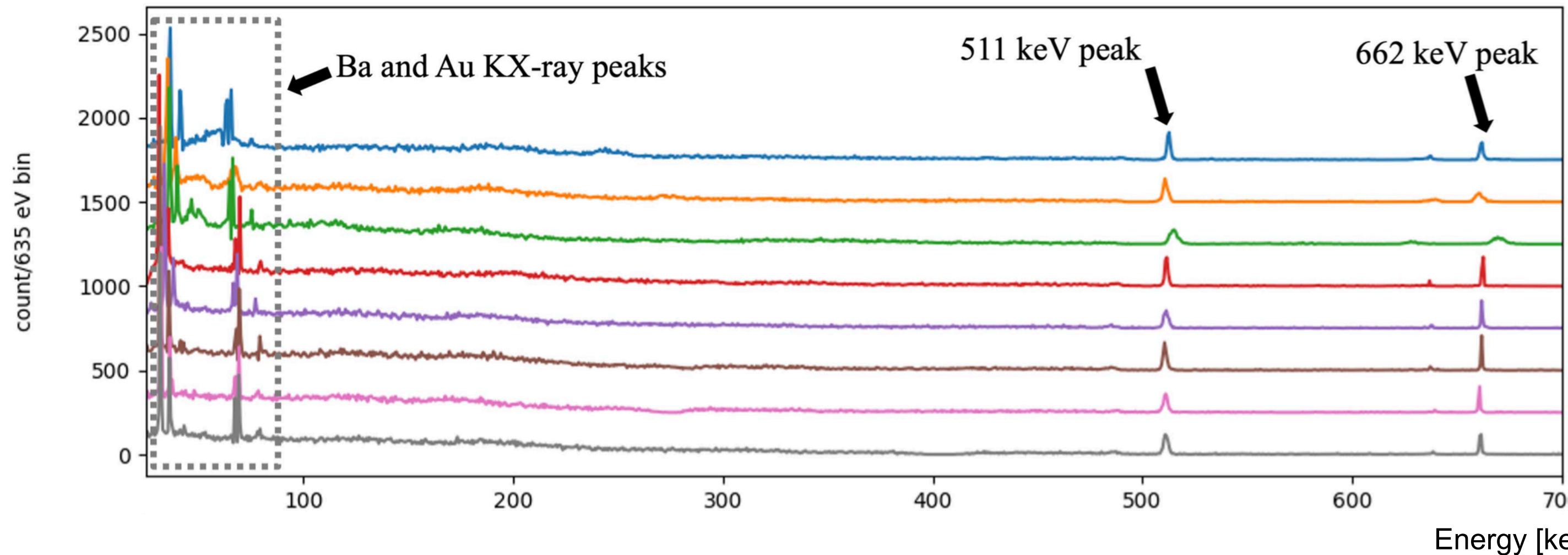


# Gamma-ray TES Array @ AIST



## Advantages in $^{112}\text{Sn}$ ECEC search

- Absorber = reaction source  
→ high detection efficiency
- Available in the energy range of a few ten keV to several hundred keV with high energy resolution  
→ search for ECEC to  $^{112}\text{Cd}$  ground state by looking for X-ray and Auger electron



# Sensitivity Study

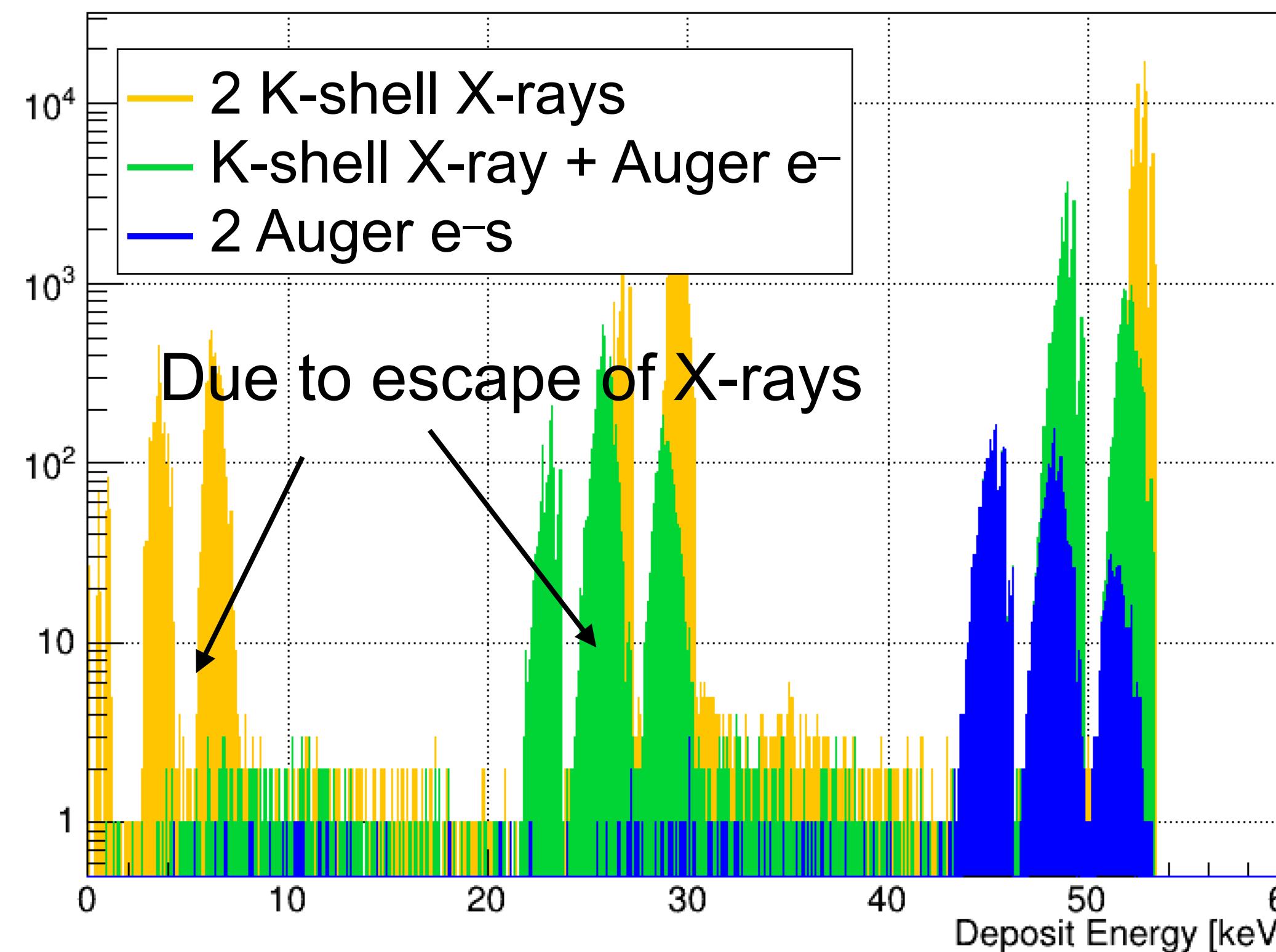
# Energy Spectrum

[2]Phys.Rev.C C 83, 045503 (2011)

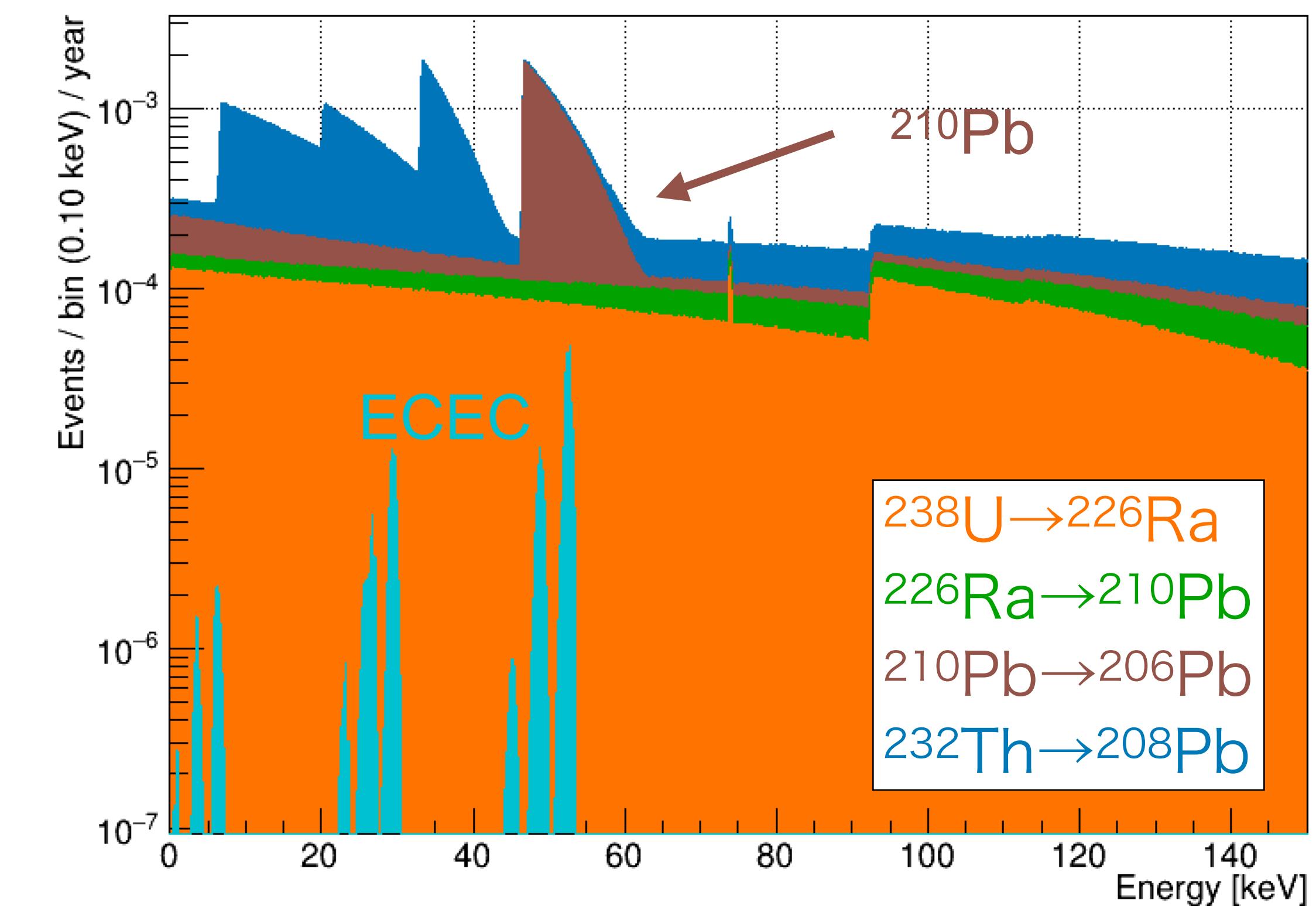
Energy deposit of ECEC and BG in Sn absorbers are simulated using Geant4

- TES with 0.8 mm x 0.8 mm x 0.8 mm cubic absorber, energy resolution of 0.1 keV
- ECEC : Two Cd atoms with a single K-shell vacancy
- BG :  $^{238}\text{U}$  : 1.8 mBq/kg,  $^{232}\text{Th}$  : 1 mBq/kg (U.L. from [2]) and their daughters

Energy deposit of X-rays and Auger electrons of  $^{112}\text{Cd}$



BG spectrum (stacked)

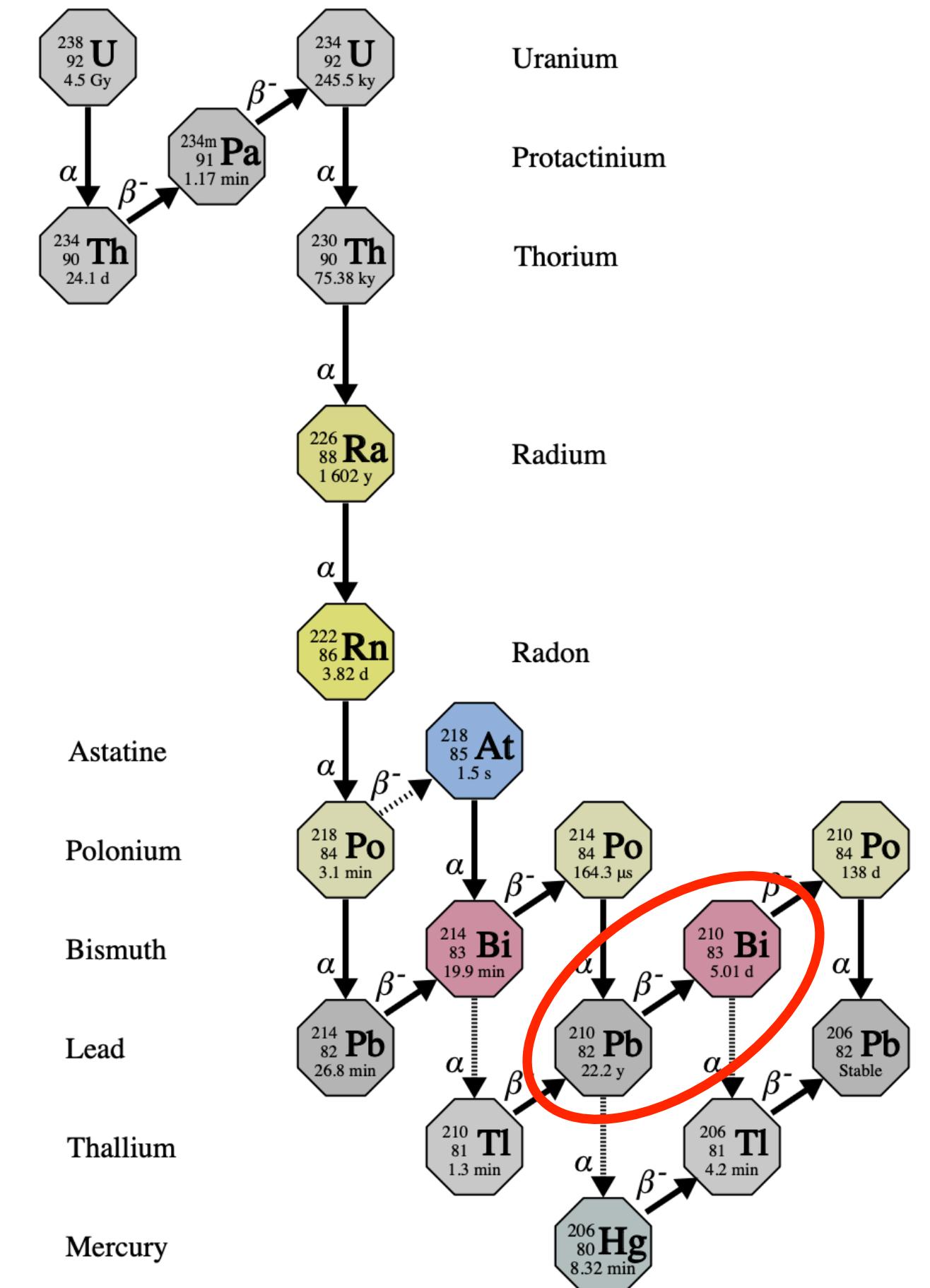


# Sensitivity Study : Assumptions

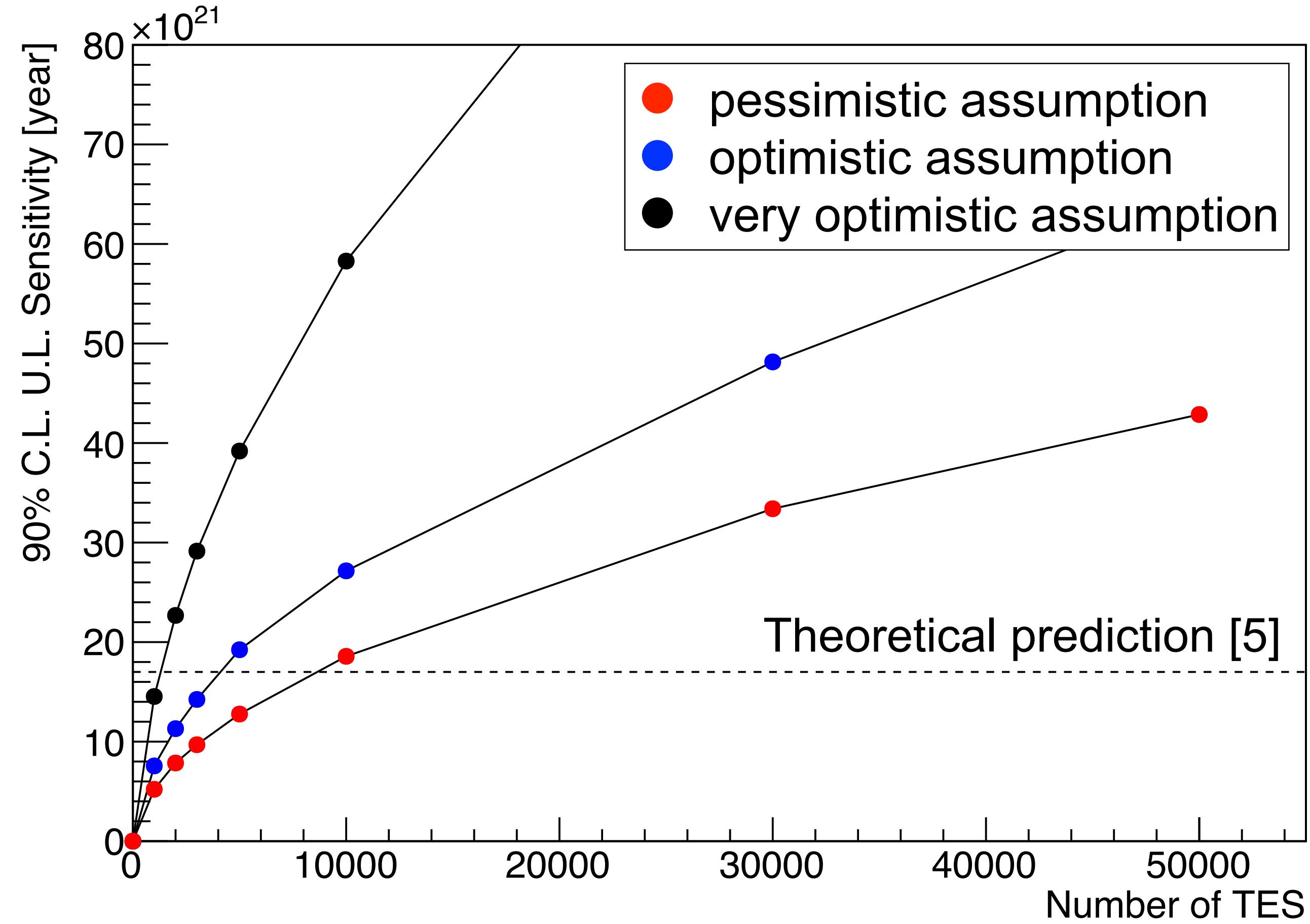
- 90% C.L. rejection sensitivity is estimated using 10000 ToyMC experiments.

## Assumptions

- 5 years observation using TES with  $(0.8 \text{ mm})^3$  enriched  $^{112}\text{Sn}$  absorber
- $^{238}\text{U}$ ,  $^{232}\text{Th}$  contamination ... 3 assumptions
  - Pessimistic ...  $^{238}\text{U} : 1.8 \text{ mBq/kg}$ ,  $^{232}\text{Th} : 1 \text{ mBq/kg}$  (U.L. from [2])
  - Optimistic ... pessimistic case  $\times 0.5$
  - Very optimistic ... pessimistic case  $\times 0.1$
- Background rejection by delayed coincidence
  - to remove  $^{210}\text{Pb}$
  - consecutive signals in the same TES pixel within one month will be vetoed.  
→ veto : 98.4% rejection efficiency @ 12.1% signal sacrifice



# Sensitivity Study : Result



<https://simonsobservatory.org/gallery/>



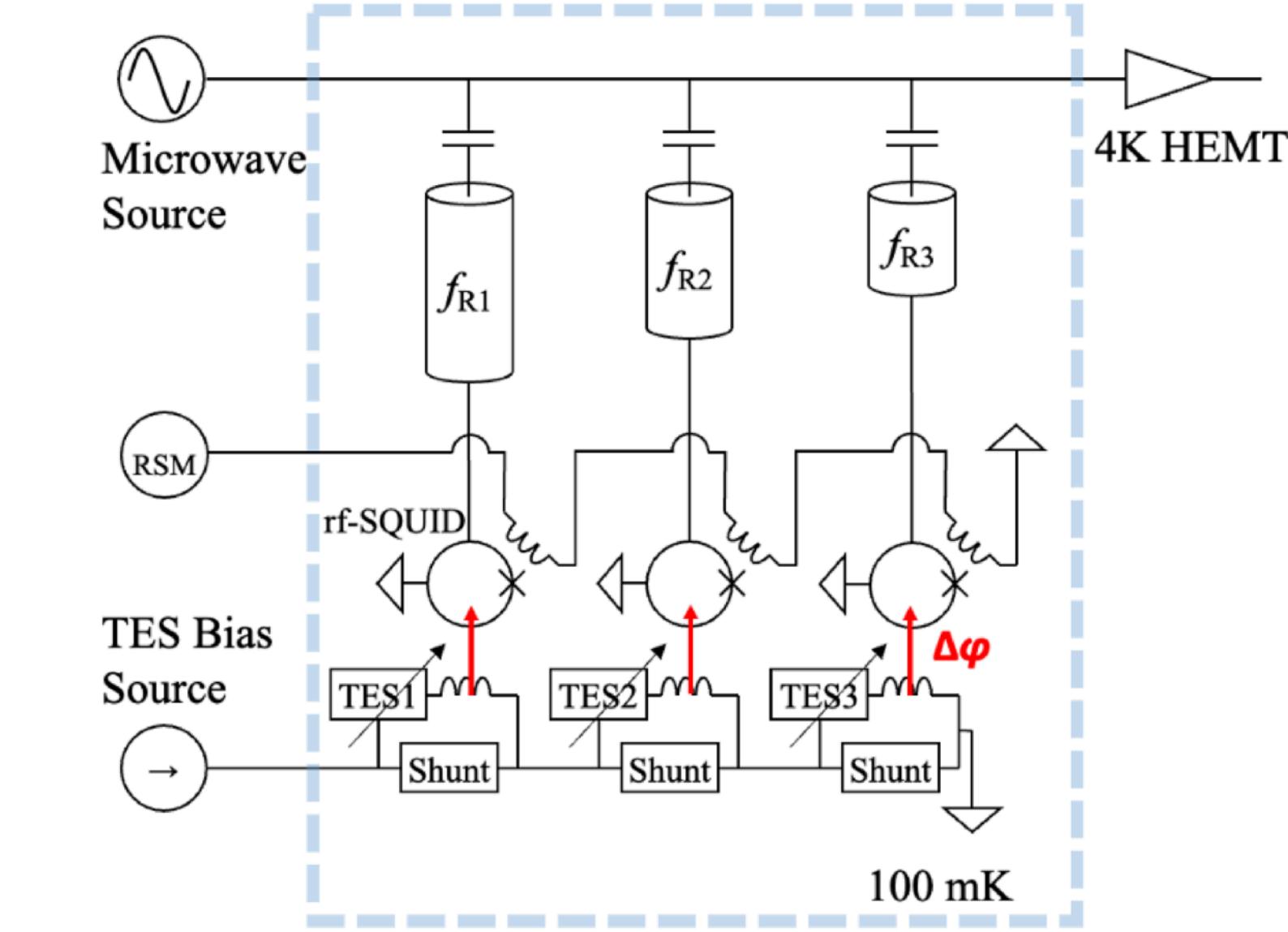
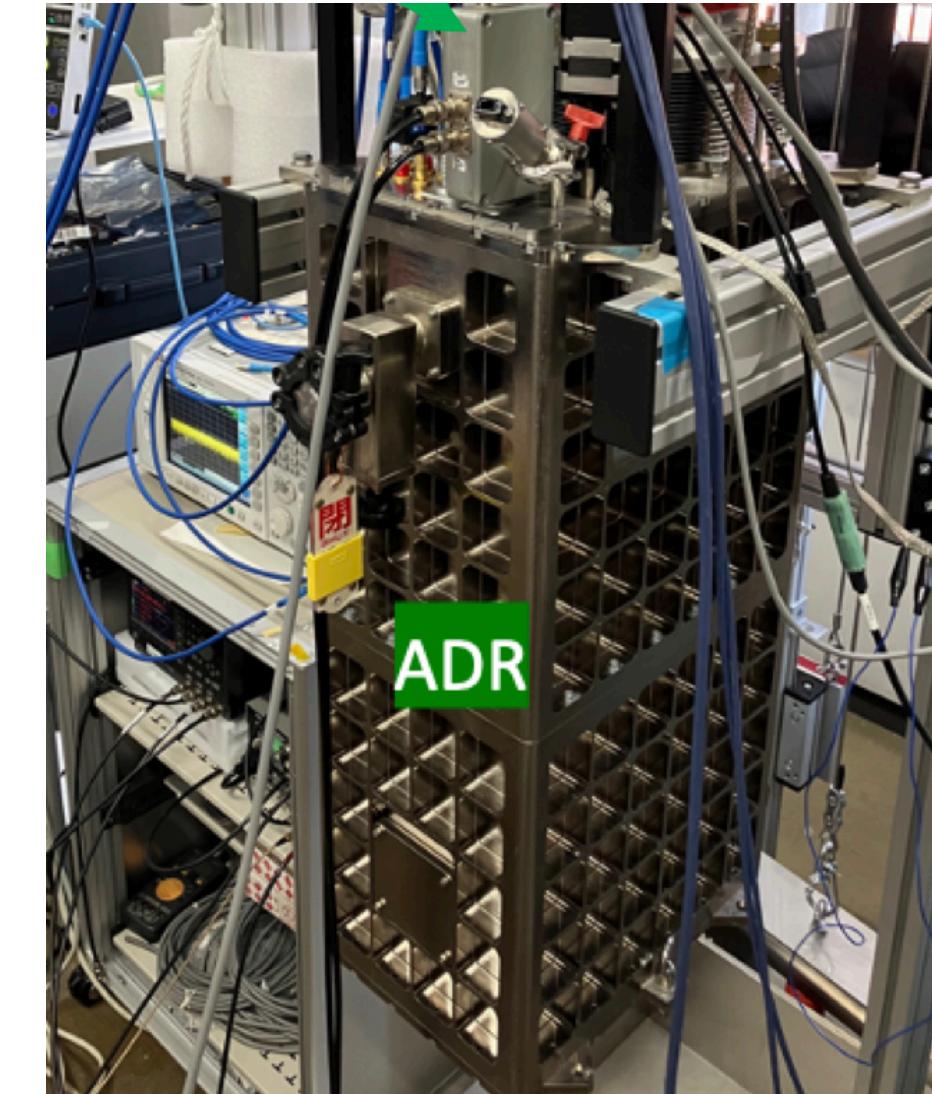
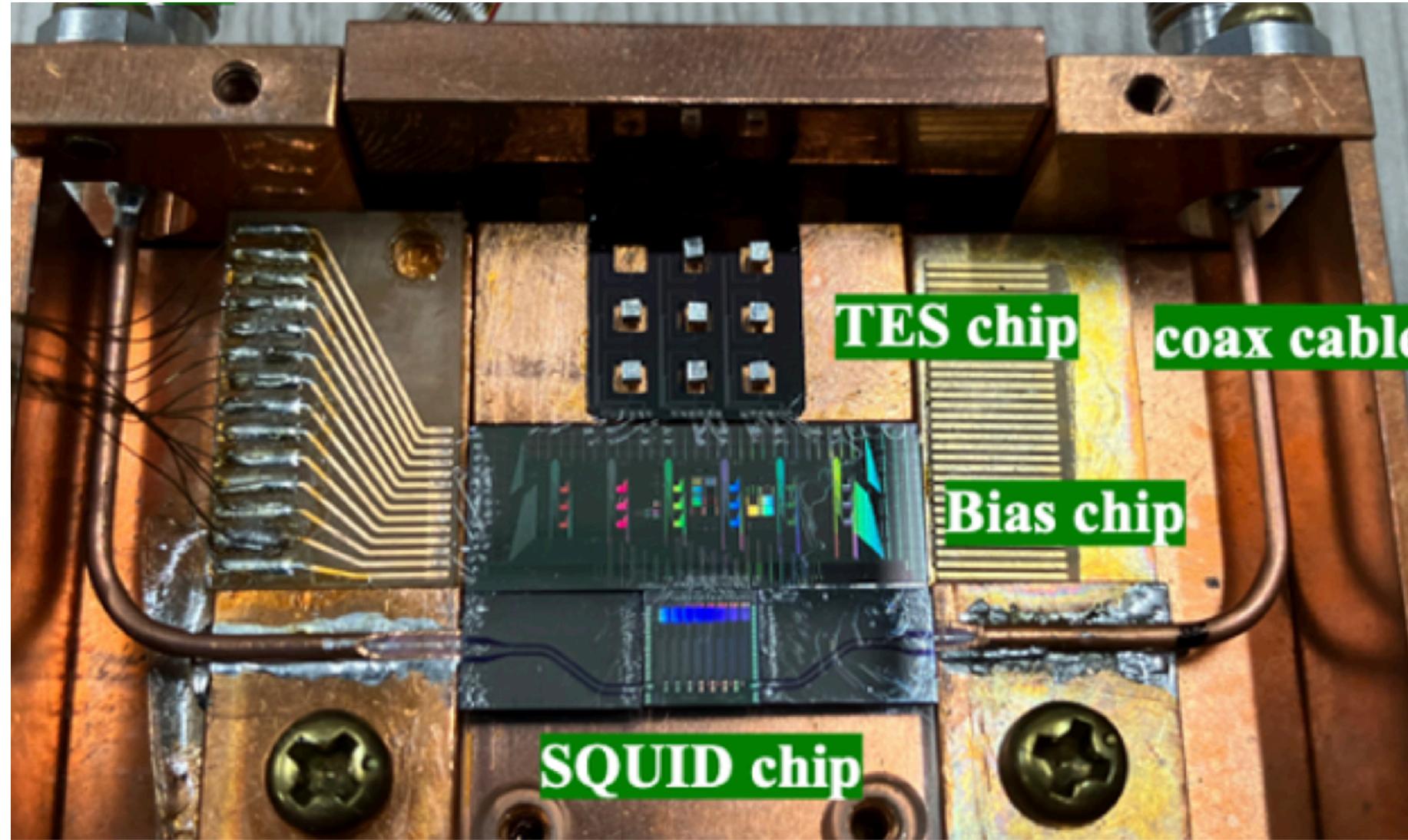
Simons Observatory (CMB experiment)  
~12000 TESs per telescope

We can test a theoretical prediction using realistic numbers of TES and 5 years observation.

- ~10000 TESs even under pessimistic assumption
- < 5000 TESs under optimistic assumptions

# **Demonstration of ECEC Search Using gamma-ray TES Array**

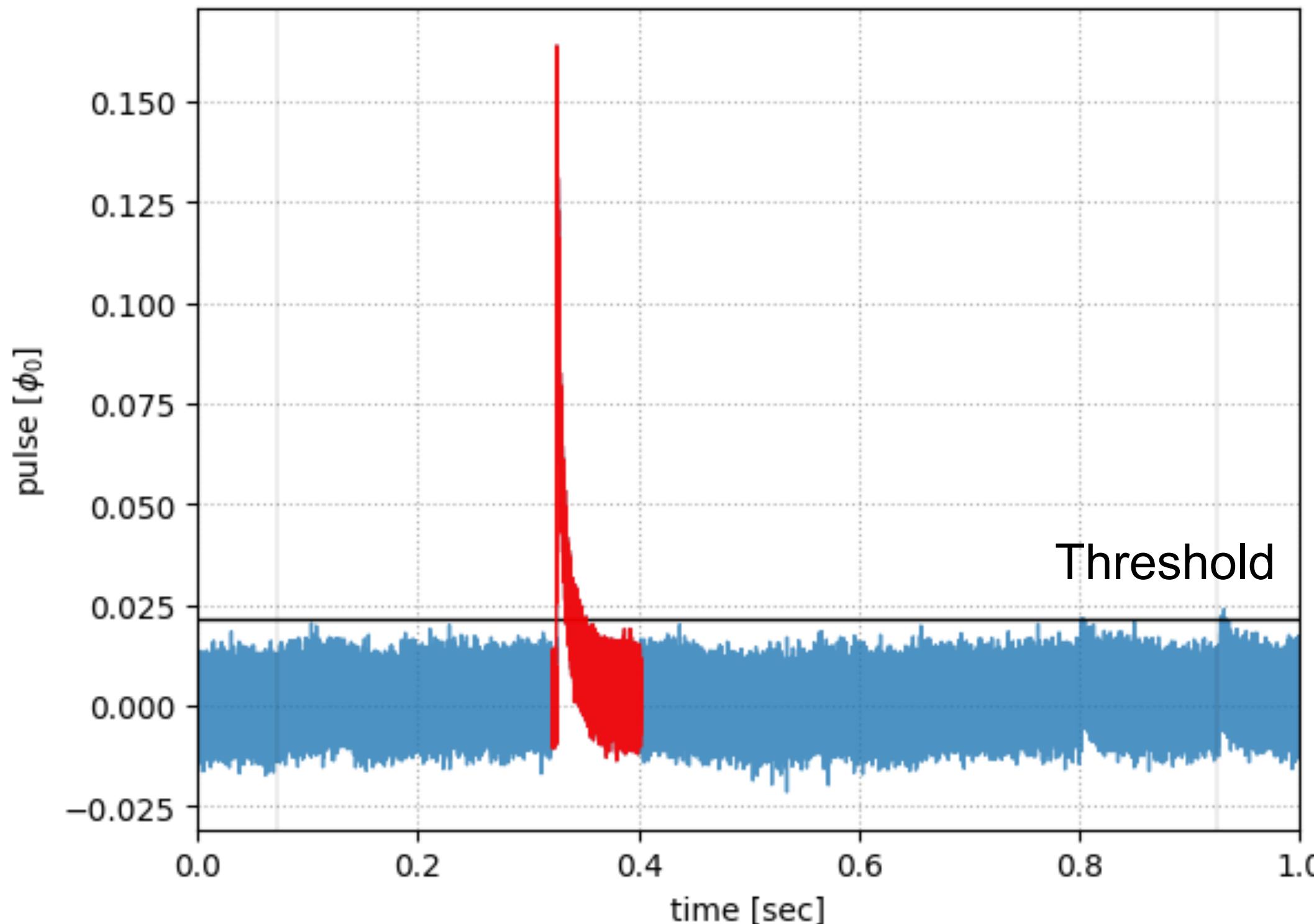
# Measurement



- TES module was cooled using adiabatic demagnetization refrigerator to lower than 100 mK
- 1-sec-duration waveforms were recorded repeatedly using microwave multiplexing.
- $^{130}\text{Ba}$  source calibration data and ECEC search data are taken.
- Measurement time for ECEC search : 22200 sec in total
  - Data from six TES pixels are used for analysis. Exposure :  $1.3 \times 10^{-10} \text{ kg}^*\text{yr}$  ( $^{112}\text{Sn}$ )

# Analysis Method

- Signal waveform is extracted from 1-sec-duration data
  - Analytical threshold to pulse height :  $\sim 10\text{keV}$
- Pulse Height Amplitude (PHA), that is proportion to the event energy, is calculated.
  - optimal filter method is used
- Absolute energy calibration with  $^{130}\text{Ba}$  81 keV peak



## Calculation of PHA using optimal filtering

$$\text{PHA} = C \int_{-\infty}^{\infty} D(t) \mathcal{F}^{-1} \left( \frac{M(f)}{|N(f)|^2} \right) dt$$

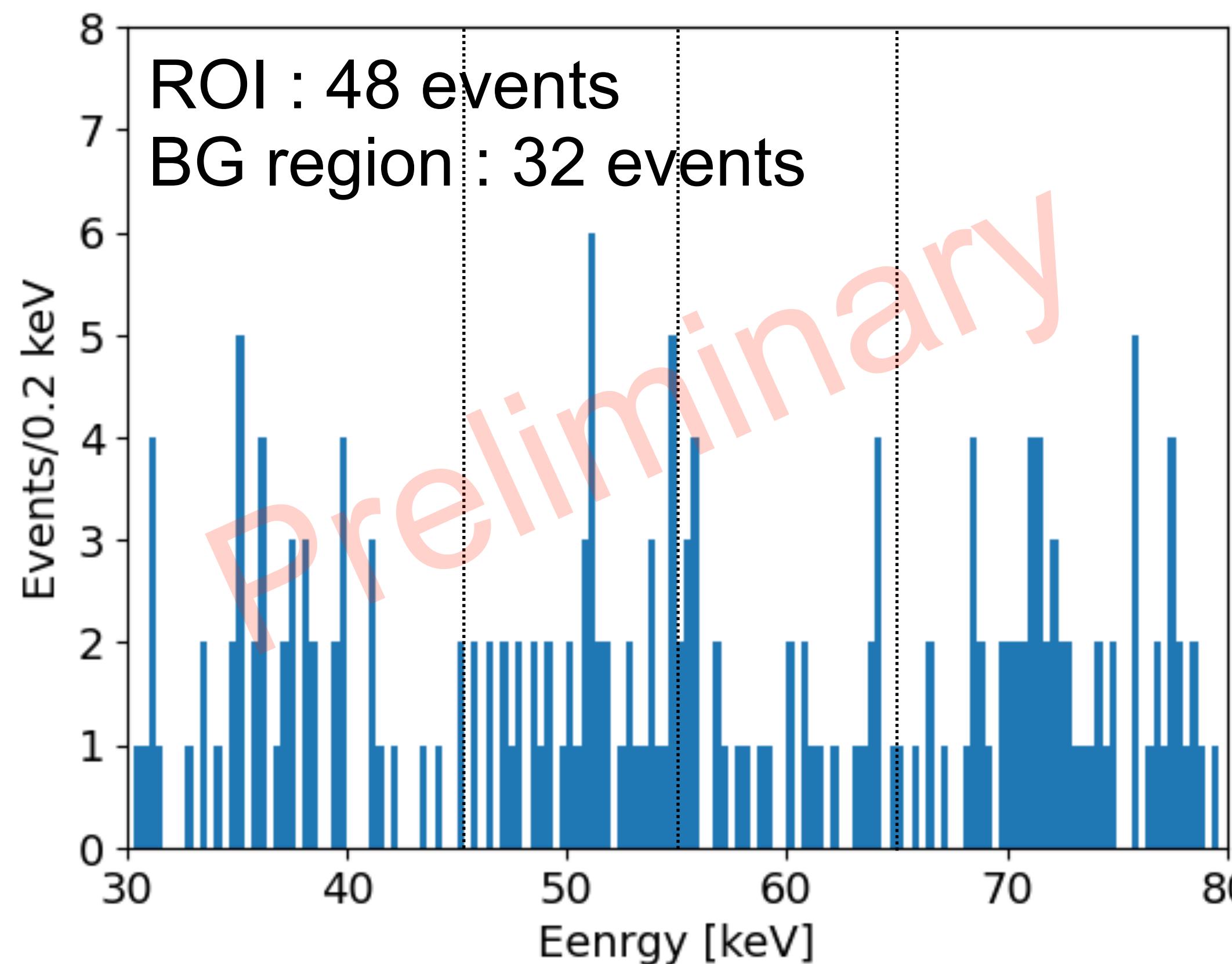
$C$  : Normalization,  $D(t)$  : signal waveform,  $\mathcal{F}^{-1}$  : IFT  
 $M(f)$  : frequency spectrum of signal template,  
 $N(f)$  : noise spectrum

pulse area( $\propto$ event energy) calculation suppressing the influence of noise.

# Limit on ECEC

Set the limit on the number of ECEC events by defining  $45 \leq \text{Energy} < 55 \text{ keV}$  and  $55 \leq \text{Energy} < 65 \text{ keV}$  region as signal region and BG region, respectively.

- Energy selection efficiency : 73%
- Uniform distribution of BG is assumed.
- Feldman Cousins' method was used to calculate 90% C.L. U.L.



90% C.L. U.L. : 28.5 events  
→  $T_{1/2}(\text{ECEC}) > 1.5 \times 10^{12} \text{ yr}$  (Theory :  $1.7 \times 10^{22} \text{ yr}$ )

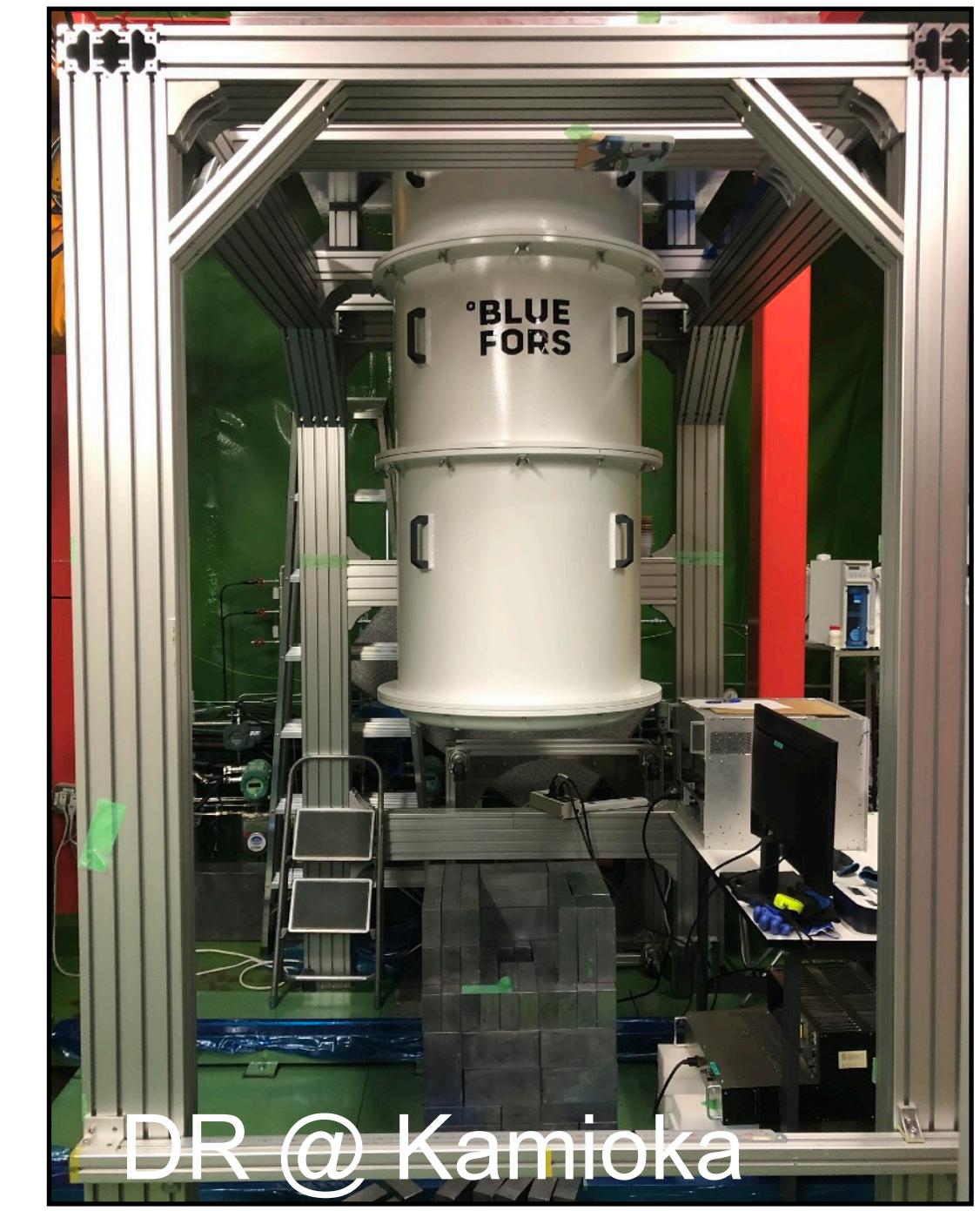
First demonstration of ECEC search  
using gamma-ray TES

# Prospects

神岡極稀事象研究拠点

## Developments in Kamioka Extremely Rare-phenomena and Neutrino-research Laboratory (KERNEL)

- KERNEL : launched in the Kamioka mine as a inter-institute collaboration hub.
- $O(10^5)$  muon flux reduction
- super clean environment
- BG measurement using HPGe detector[6] and measurement using DR in KERNEL



# Summary

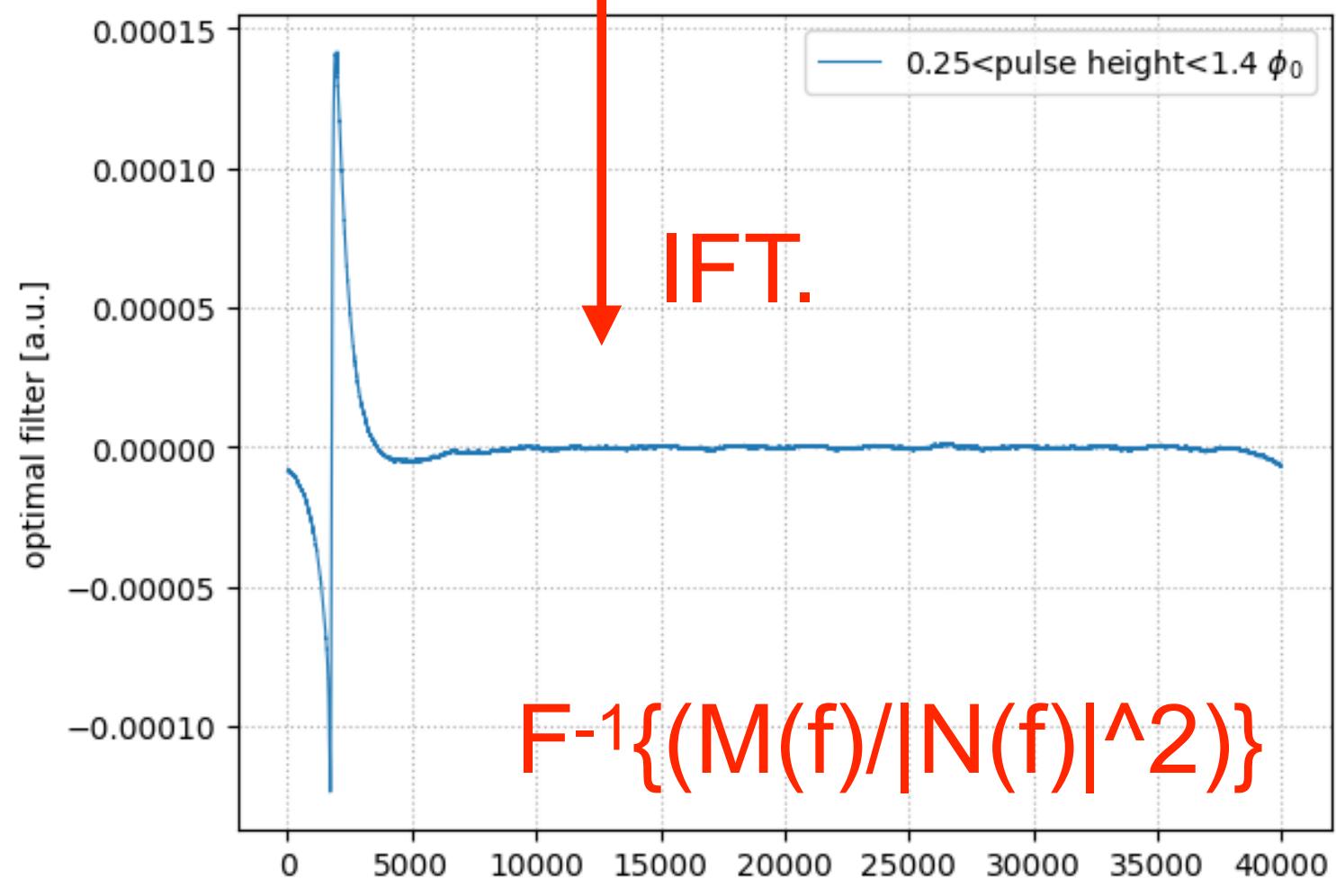
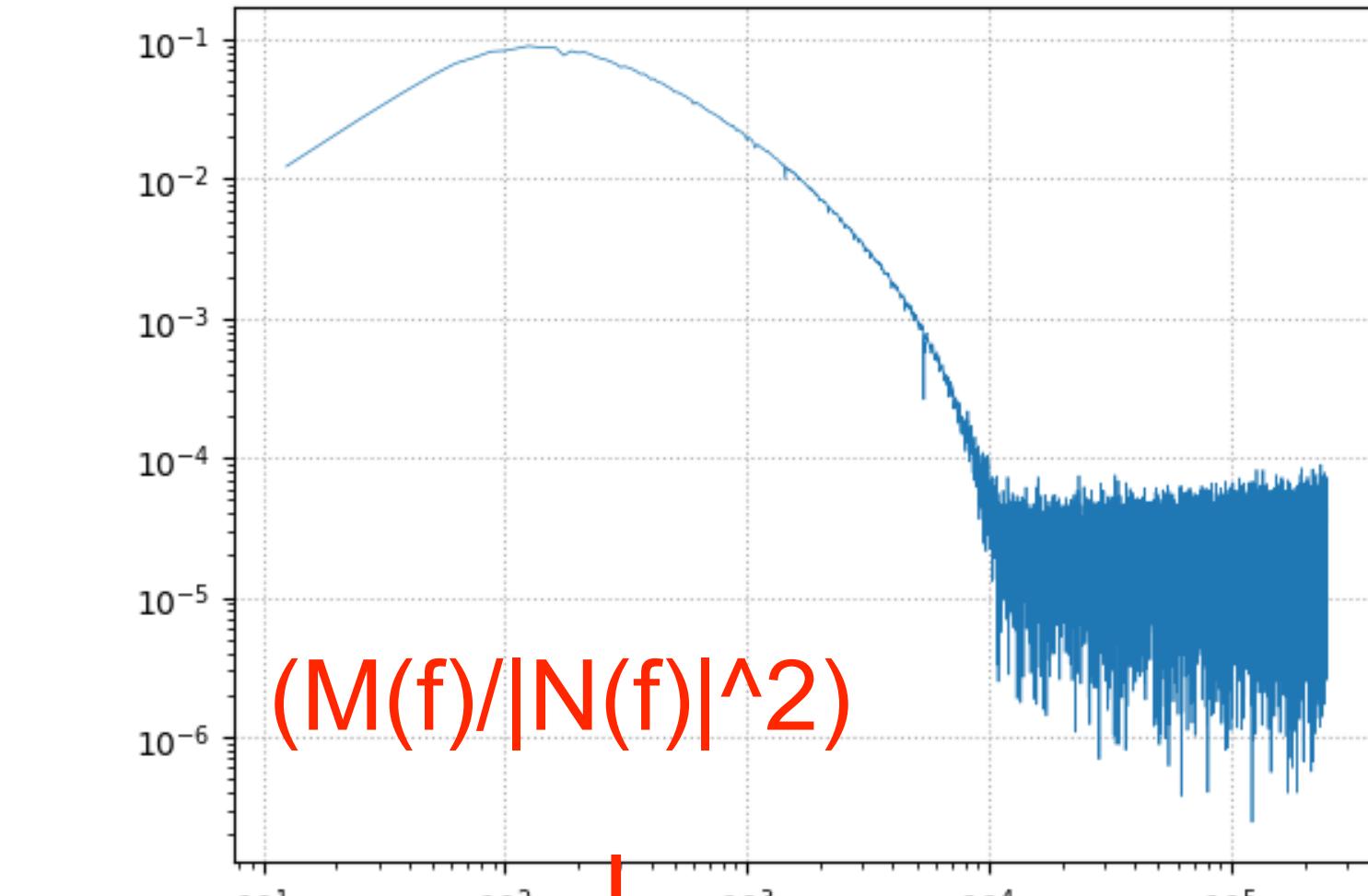
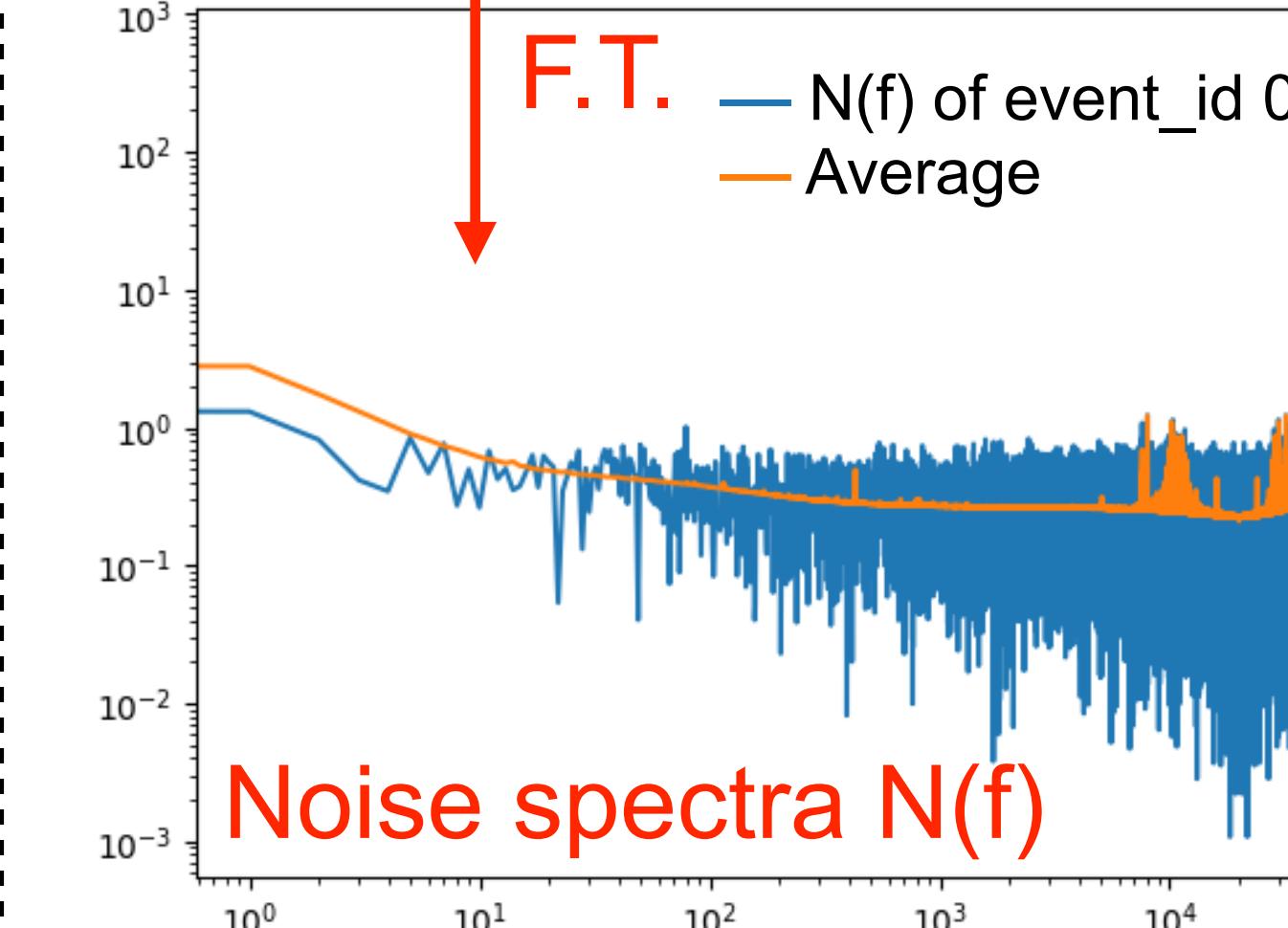
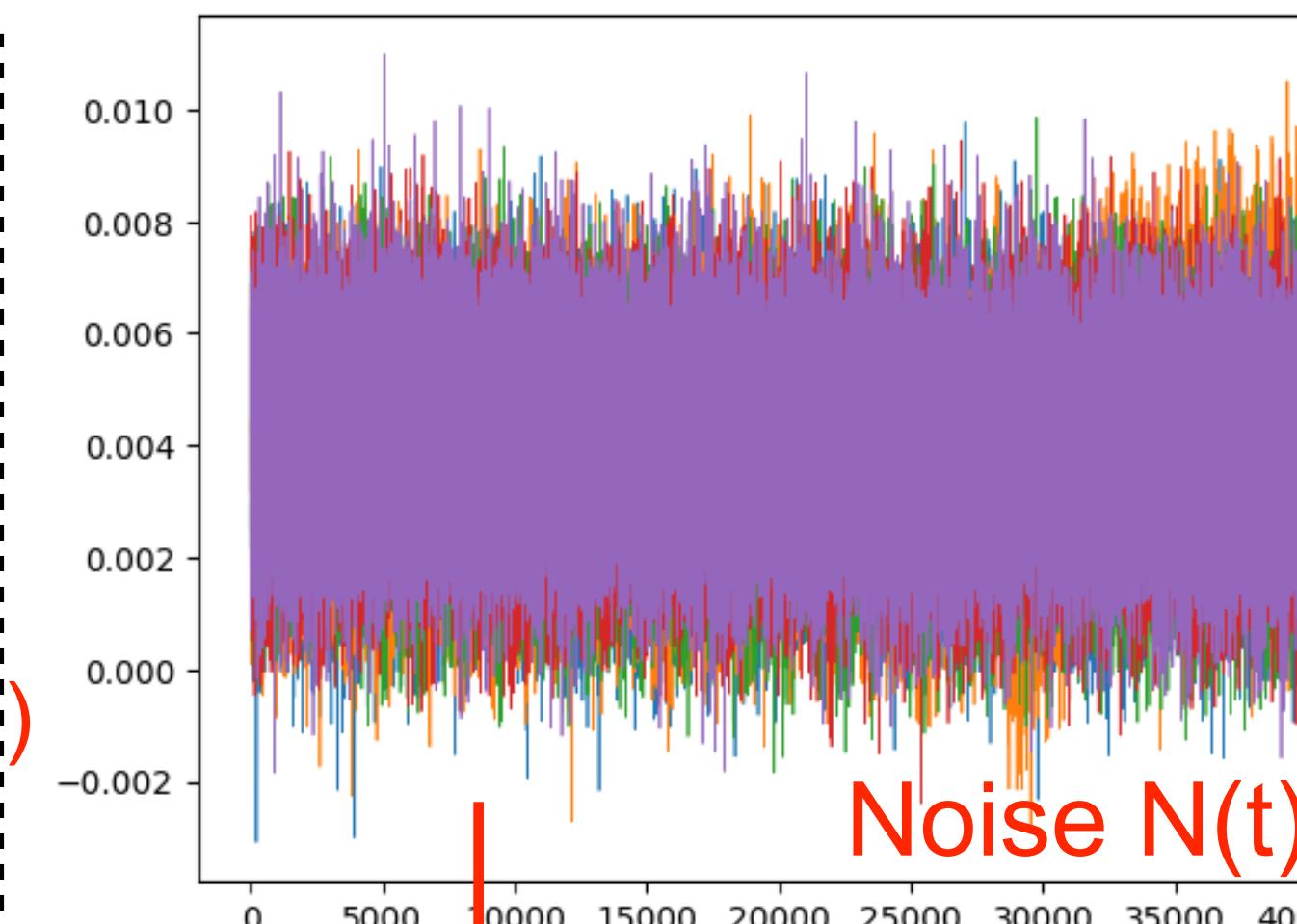
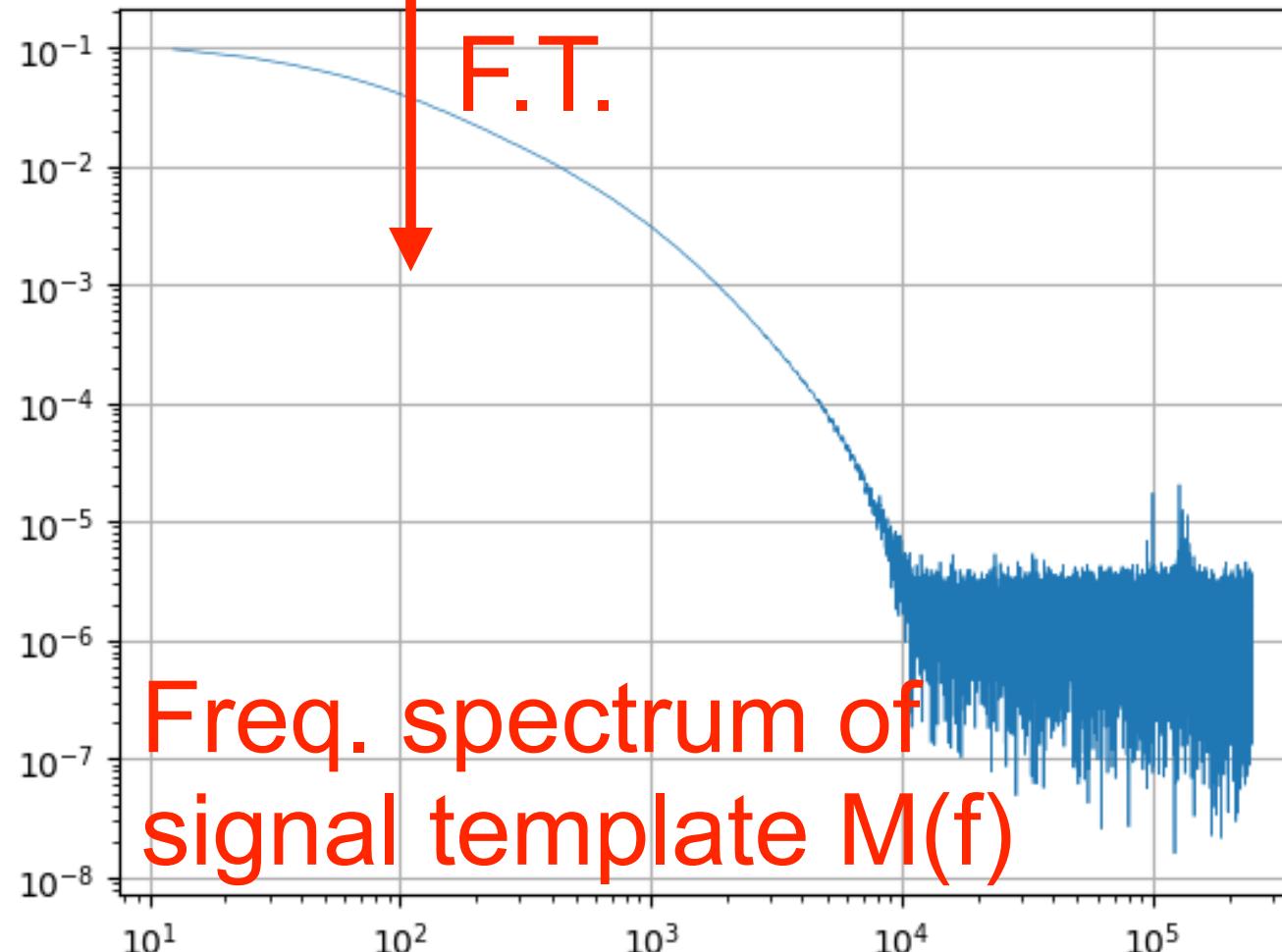
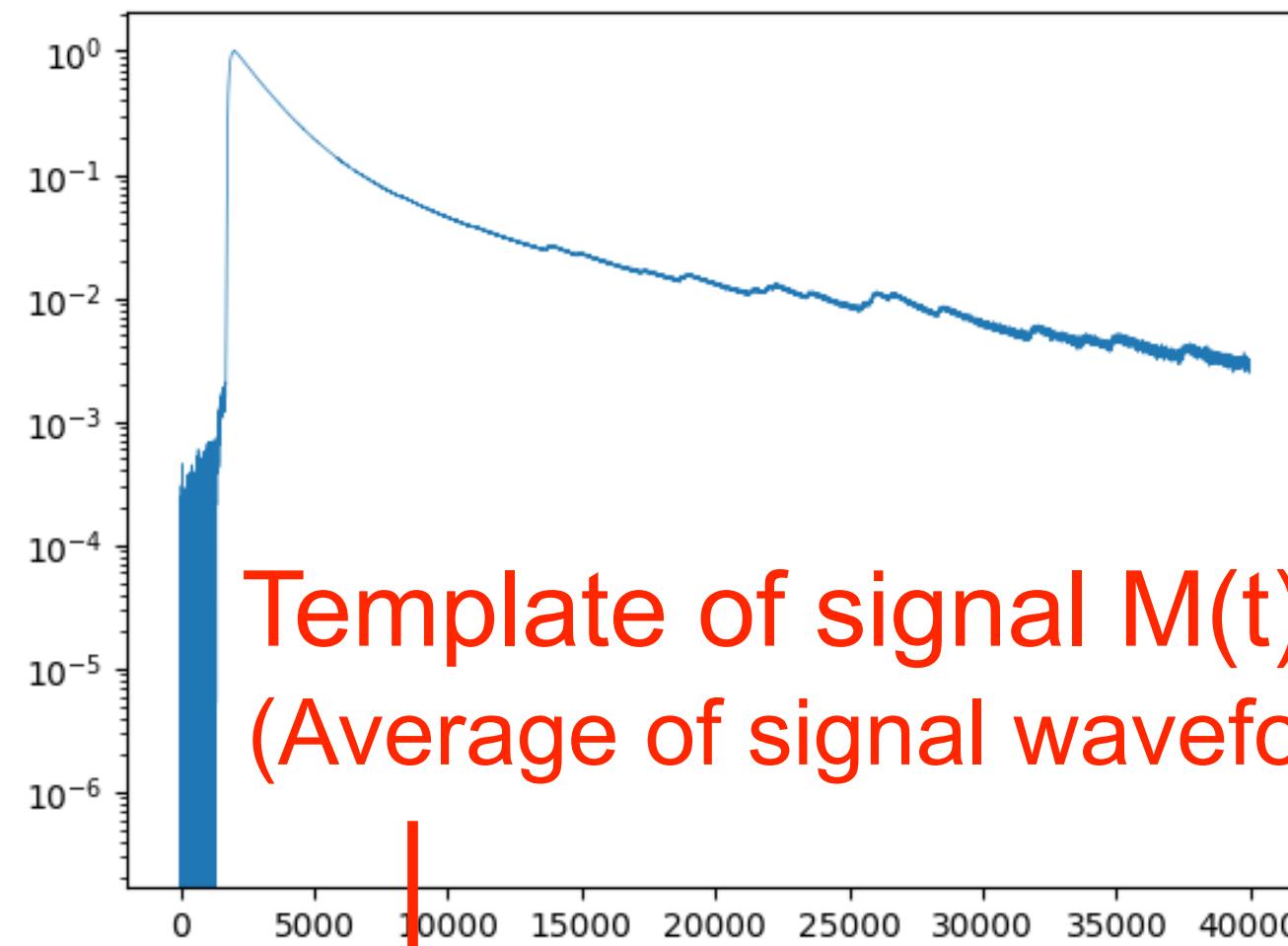
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- Double electron capture can be an input to theoretical study of nuclear structure and a key to understand Majorana nature of neutrinos.
- Simulation study shows that we can test a theoretical prediction of ECEC half-life using realistic number of TES pixels and measurement time.
- We demonstrated ECEC search using gamma-ray TES array.
  - limit on half-life :  $\underline{T_{1/2}(\text{ECEC}) > 1.5 \times 10^{12} \text{ yr}}$
- We are planning developments in KERNEL, underground laboratory in Kamioka.

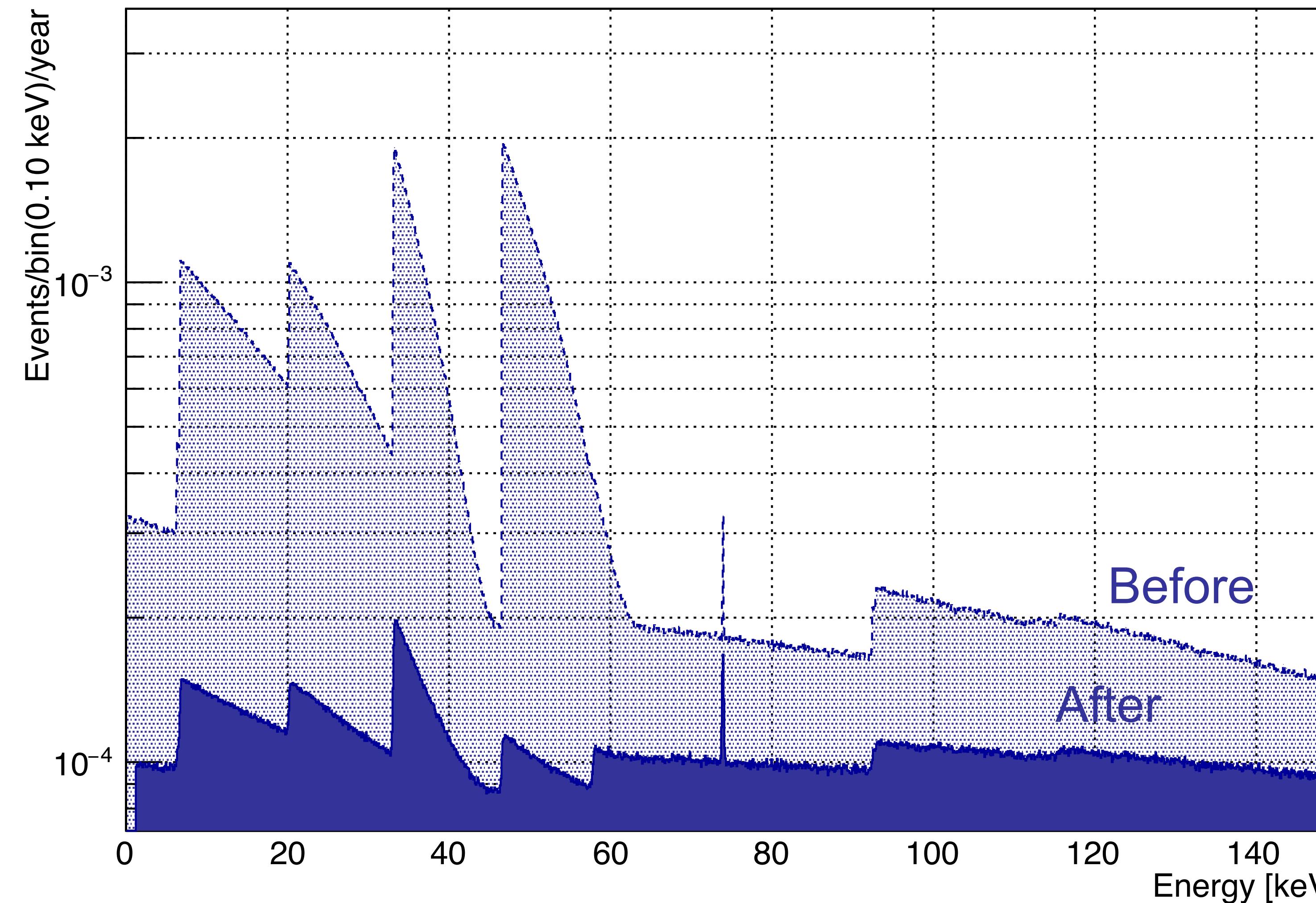
**Thank you for your attention**

# Energy Reconstruction

$$PHA = C \int_{-\infty}^{\infty} D(t) \mathcal{F}^{-1} \left( \frac{M(f)}{|N(f)|^2} \right) dt$$



# BG reduction by Delayed Coincidence



# Details of Simulation

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Energy deposit of ECEC and BG decays in Sn absorbers are simulated using Geant.4

## Detector assumption

- Size of absorber : 0.8 mm x 0.8 mm x 0.8 mm
- Energy resolution : 0.1 keV

## Signal

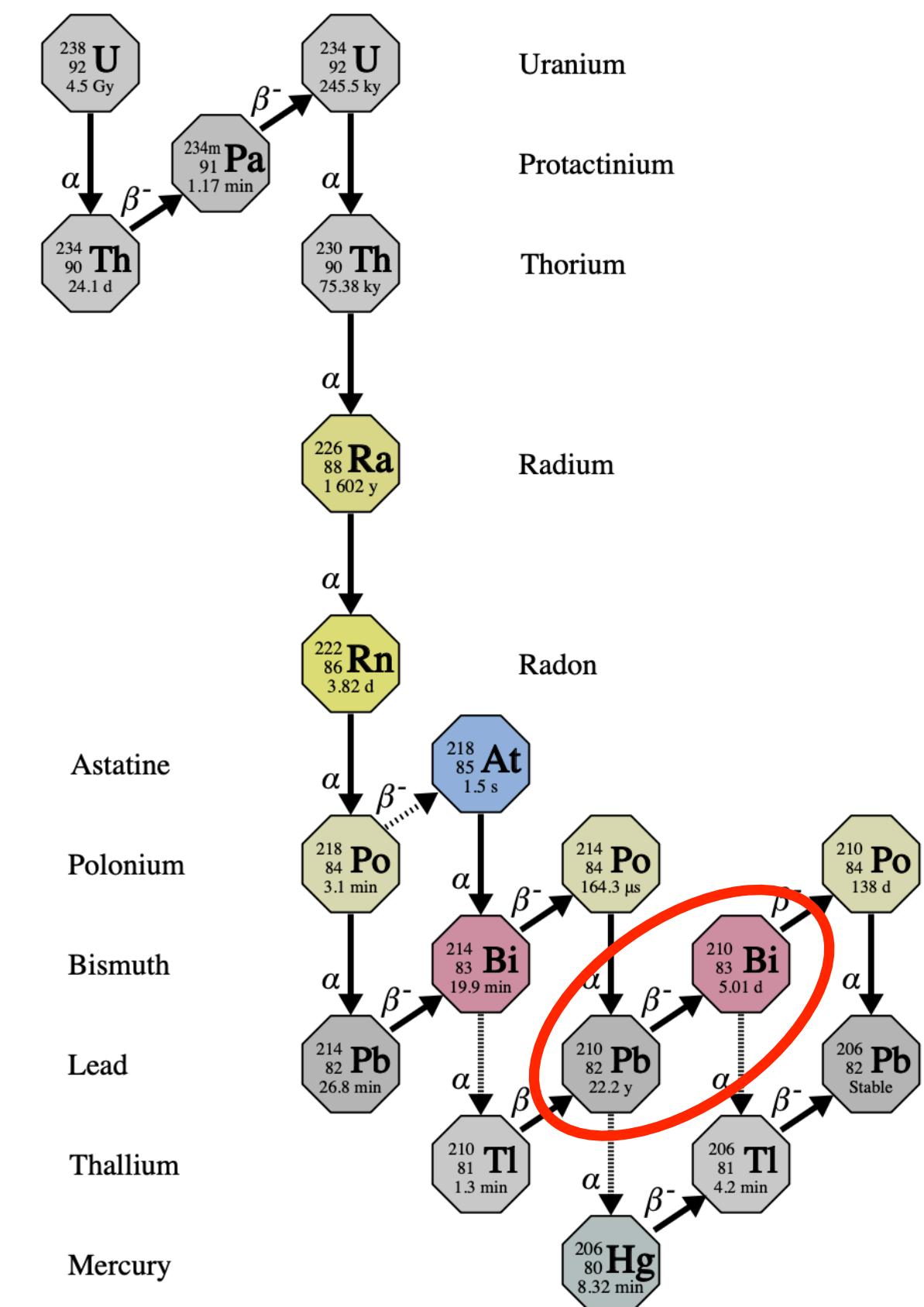
- ECEC : Two Cd atoms with a single K-shell vacancy
  - Atomic relaxation package in Geant.4 (G4AtomicDeexcitation)
  - Note : relative capture ratio of two K-shell vacancy : 73.4% (up to N5 shell)[4]
- BG : Daughter nuclei in  $^{238}\text{U}$  and  $^{232}\text{Th}$  series
  - Radioactive decay in Geant.4
  - Density : Upper limit from [2] ... U : 1.8 mBq/kg, Th : 1 mBq/kg
  - Radiation equilibrium

# Sensitivity Study

- 90% C.L. rejection sensitivity is estimated.
  - 10000 Toy MC events are generated based on the simulation spectrum.
  - 90% C.L. U.L. of signal rate is calculated for each Toy MC events.
    - Median of 10000 trials is adopted as sensitivity.

## Assumptions

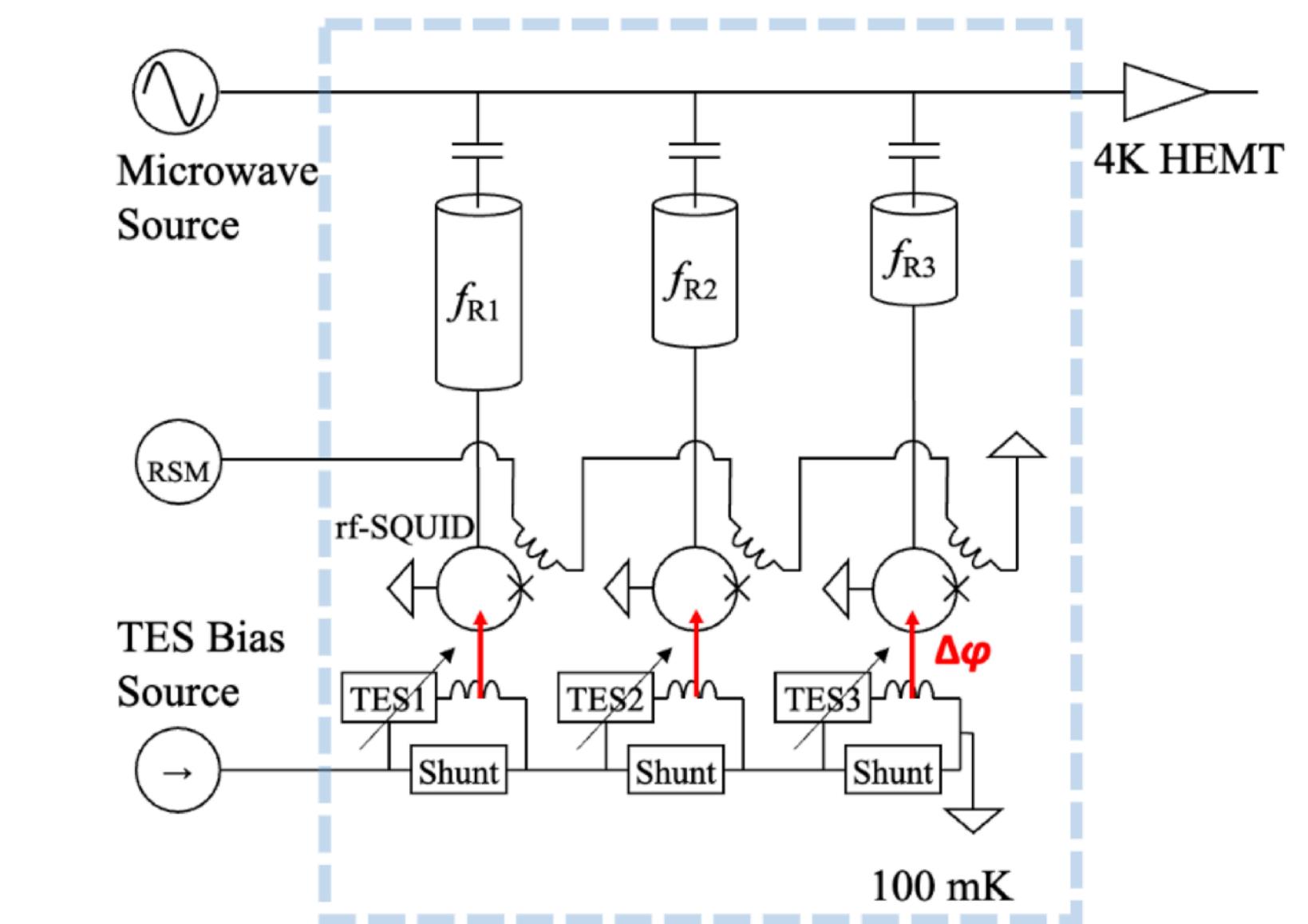
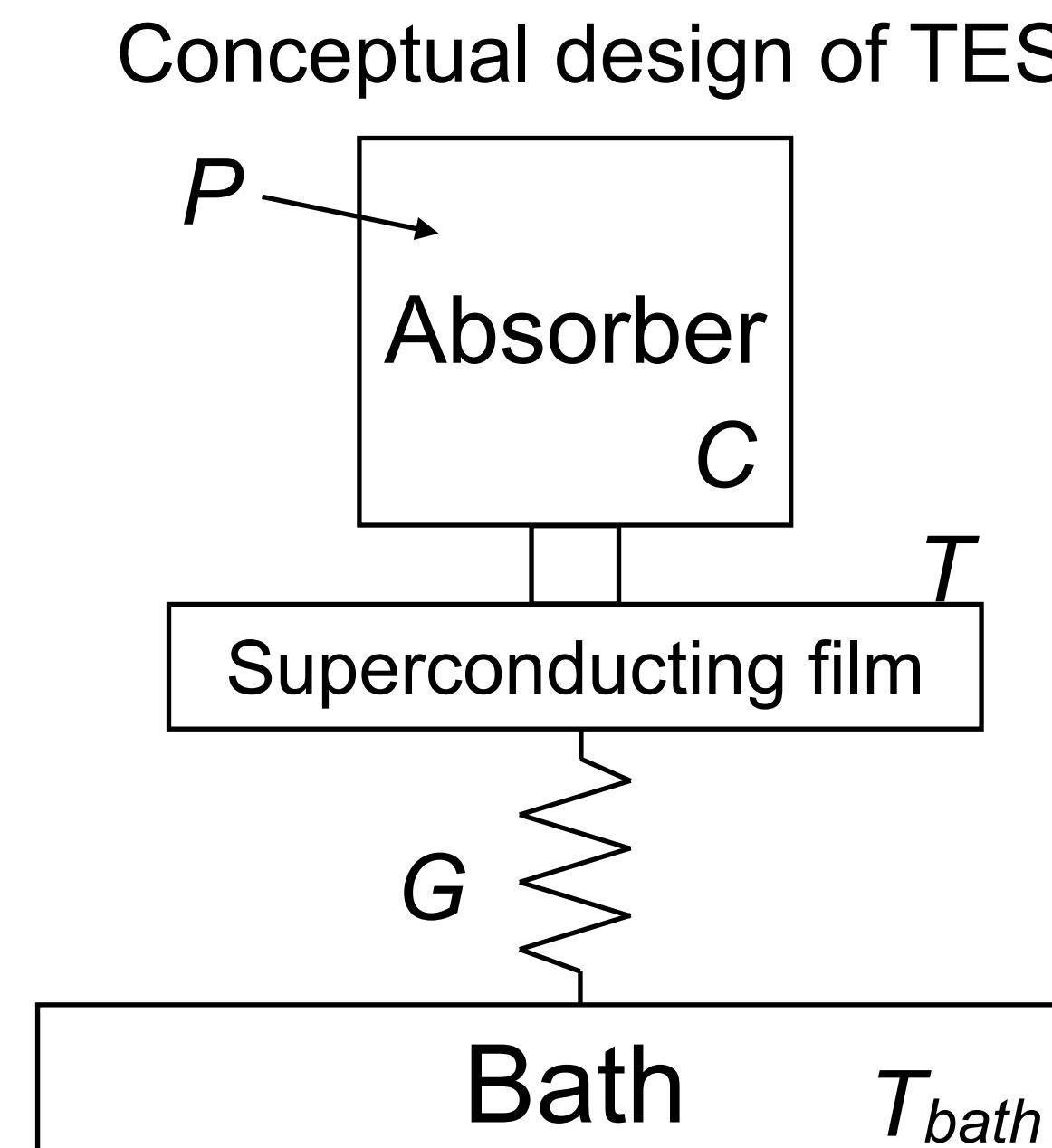
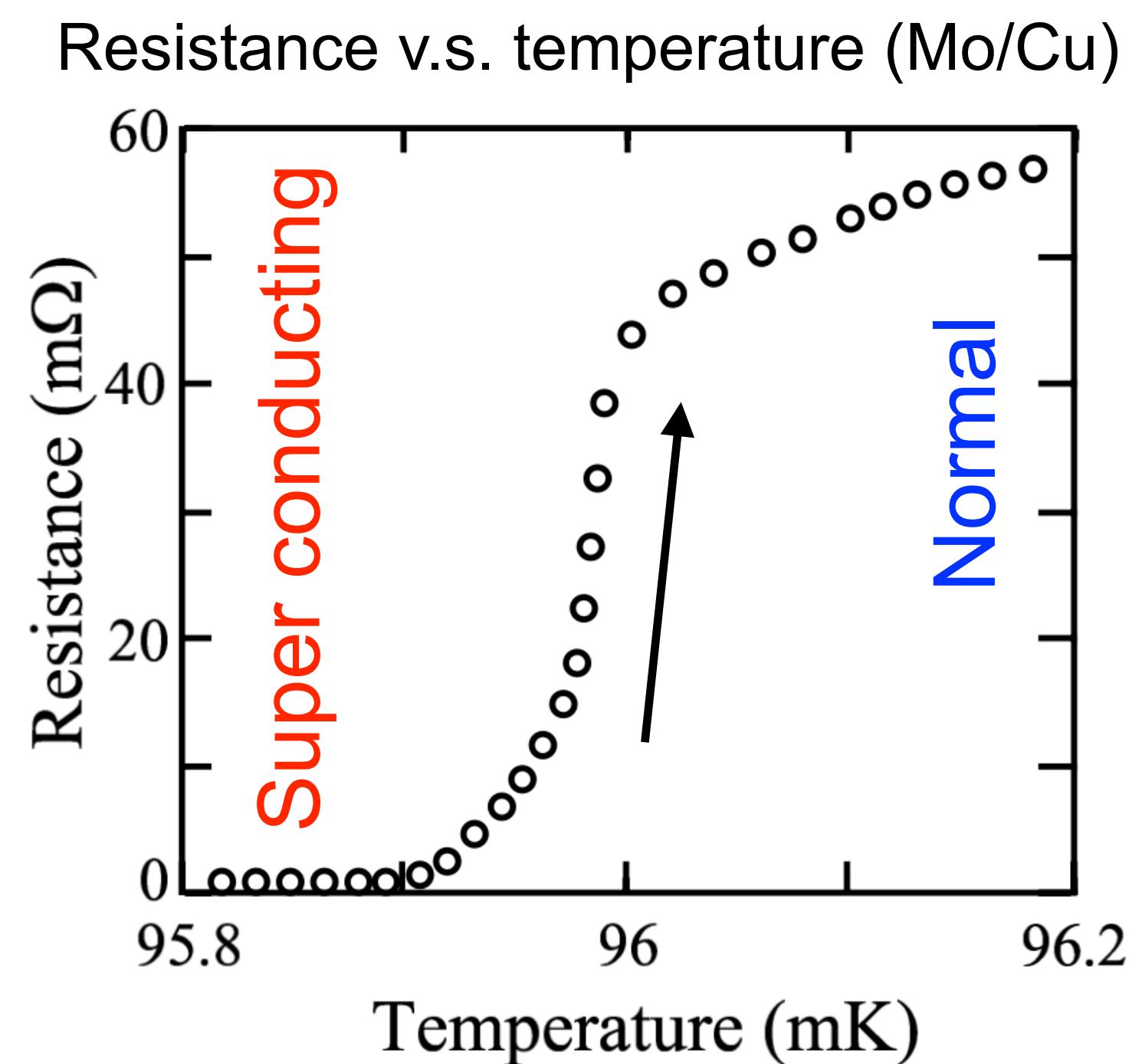
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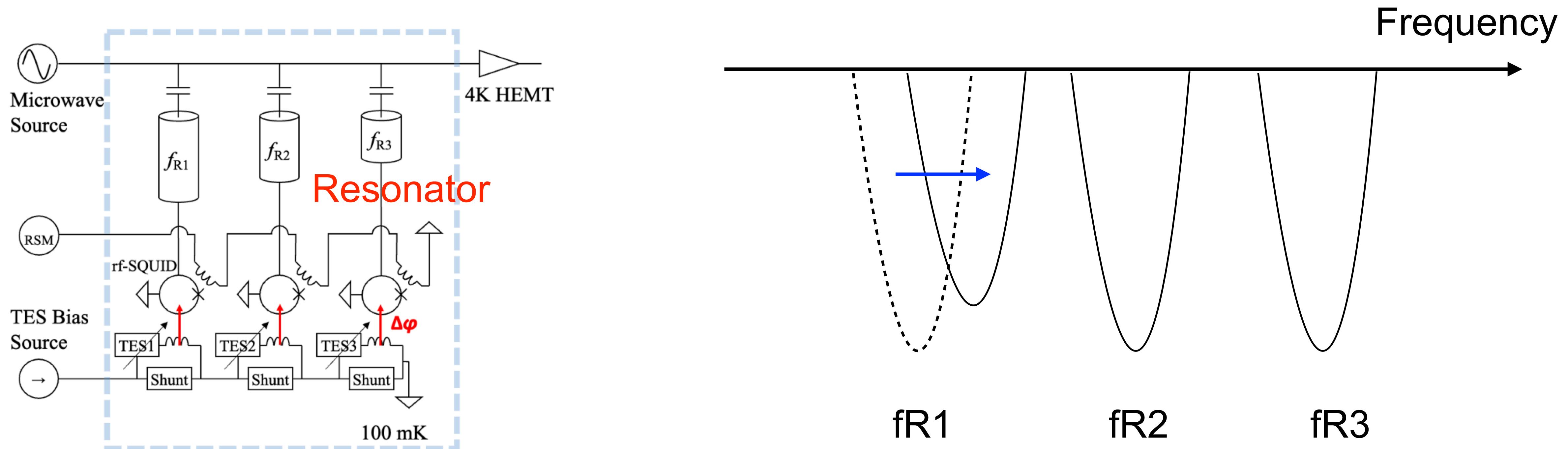
# What is TES?

## Transition Edge Sensor (TES)

- A thermometer that operates on the sharp change in resistance of superconducting materials.
  - Energy deposit at absorber is detected as the change of resistance.
  - Technically, the change in resistance is detected as a change in current using a SQUID.
- Each absorber is small ( $\sim 1\text{mm}$ ) → multi pixelization



- Multiplexed readout system
- Input coils are connected to TESs, and the change of current due to change of resistance is detected by SQUID as the magnetic signal  
→ The signal is detected as the change go resonant frequency



# Energy Calibration

Energy spectra is obtained by multiplying calibration factor to PHA spectra.

- Calibration factor is determined by looking at the  $^{130}\text{Ba}$  81 keV peak in calibration data.
- Energy linearity is assumed.

