

Vertex Reconstruction of Shower Muons in Large Liquid Scintillator Detectors using PMT Waveforms

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Introduction

- The Jiangmen Underground Neutrino Observatory (JUNO), located in southern China, is a next-generation neutrino experiment[1]. Construction of JUNO is nearing completion, with liquid scintillator filling finished and has started collecting data.

JUNO detector[2]:

- 20k ton LS
- 35k ton water
- 3% energy resolution @1MeV
- 17612 20-inch CD Large PMTs(CD LPMTs)
- 25600 3-in CD Small PMTs
- 2400 20-inch WP Large PMT

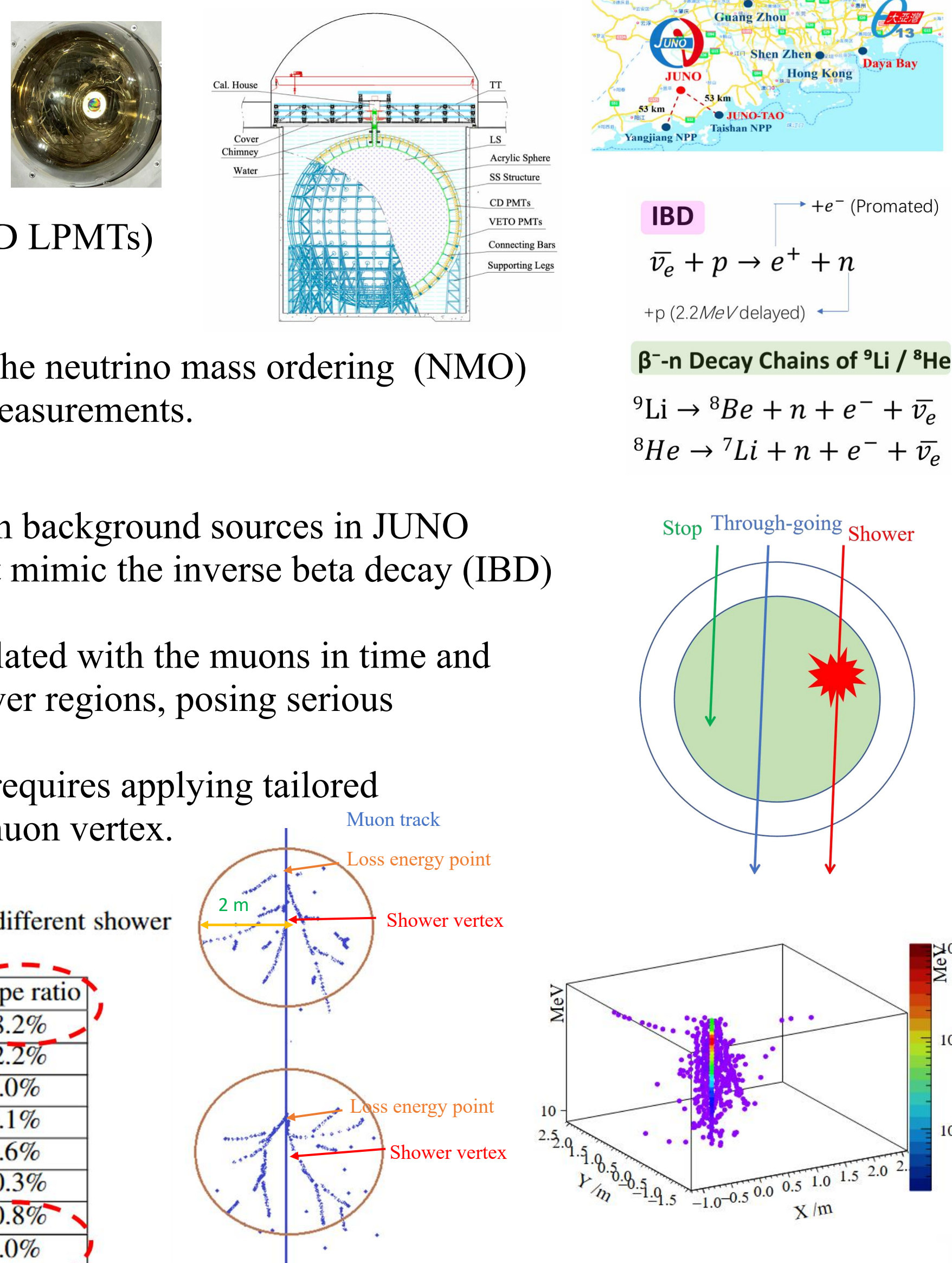
JUNO's primary goal is to determine the neutrino mass ordering (NMO) through reactor neutrino oscillation measurements.

- Cosmic muons are one of the main background sources in JUNO because they produce isotopes that mimic the inverse beta decay (IBD) signal.
- Cosmogenic isotopes tightly correlated with the muons in time and space, especially abundant in shower regions, posing serious background to signal events.

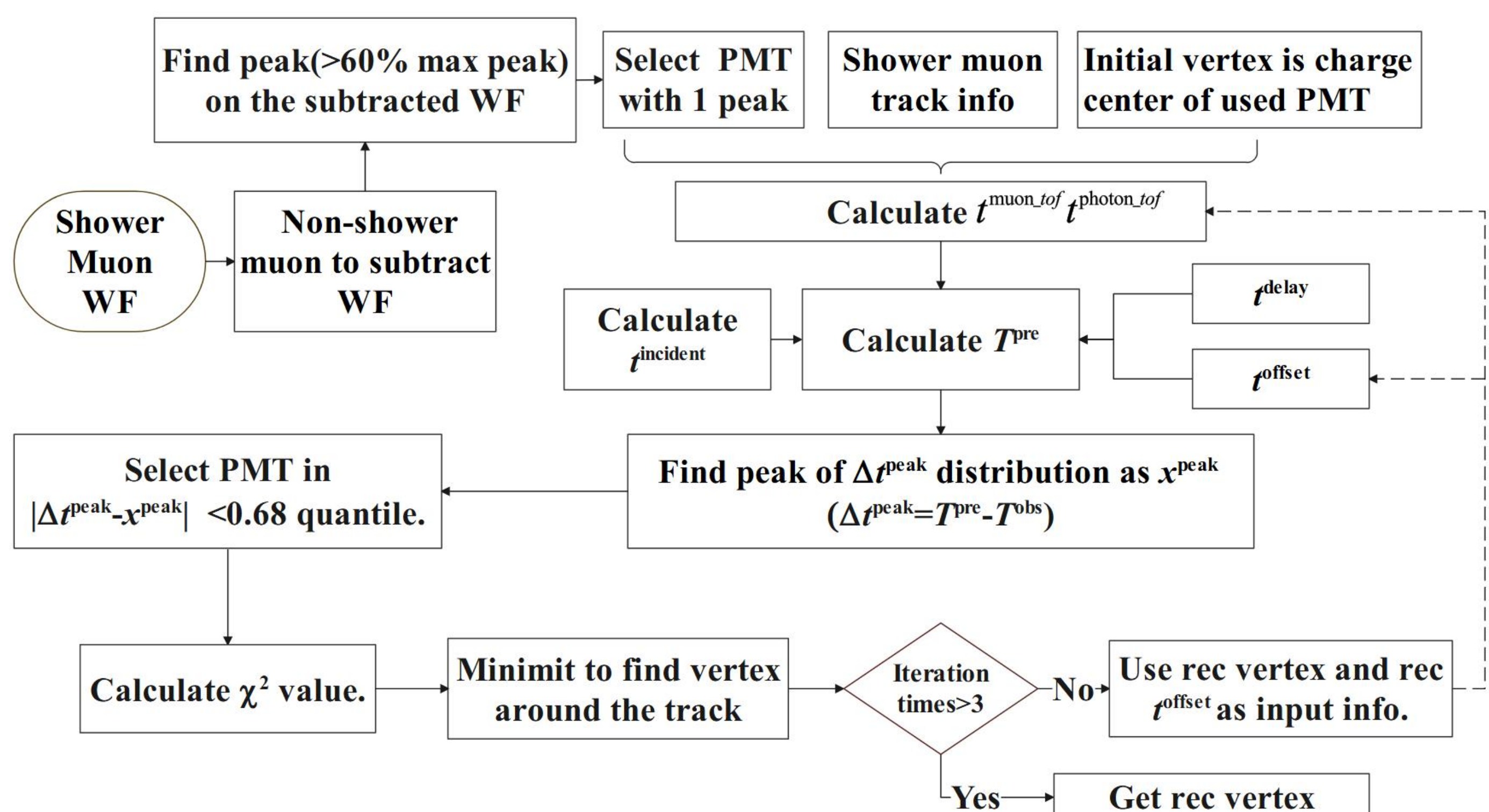
Efficient background rejection requires applying tailored reconstruction strategies for shower muon vertex.

Table 1. The proportion of isotopes produced by different shower energies of single muon and by non-shower muons.

Muon type	Muon ratio	Isotope ratio
All	19.8%	88.2%
Shower muon		
1 - 3 GeV	11.0%	12.2%
3 - 5 GeV	4.5%	7.0%
5 - 7 GeV	1.5%	5.1%
7 - 9 GeV	0.8%	3.6%
> 9 GeV	2.0%	60.3%
Through-going	76.3%	10.8%
Stop	3.9%	1.0%



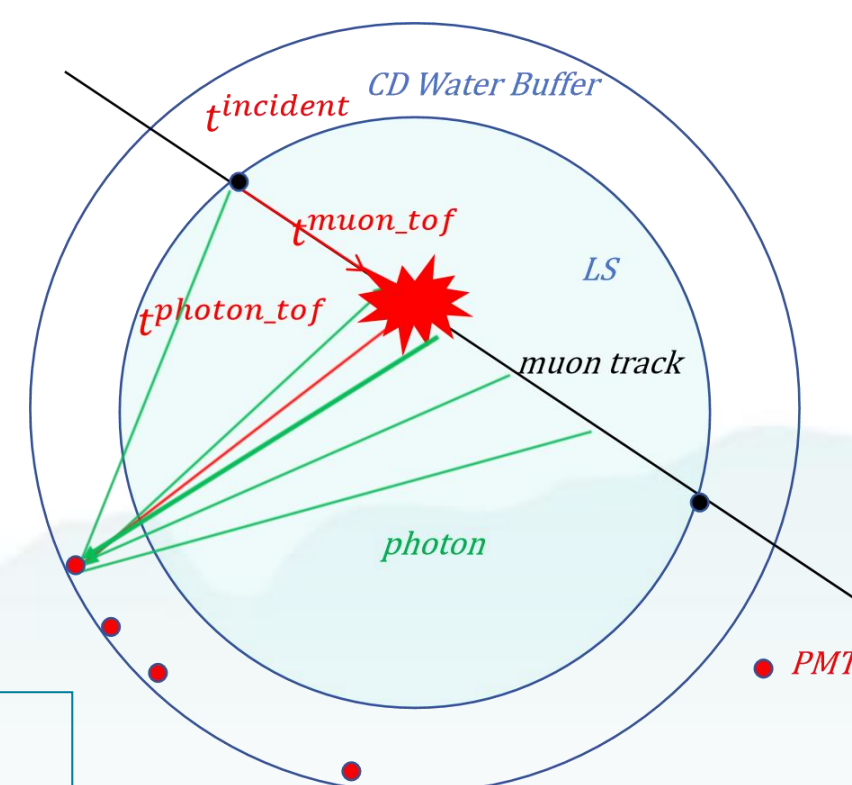
Methods



$$T_{pre} = t_{incident} + t_{muon_tof} + t_{photon_tof} + t_{delay} + t_{offset}$$

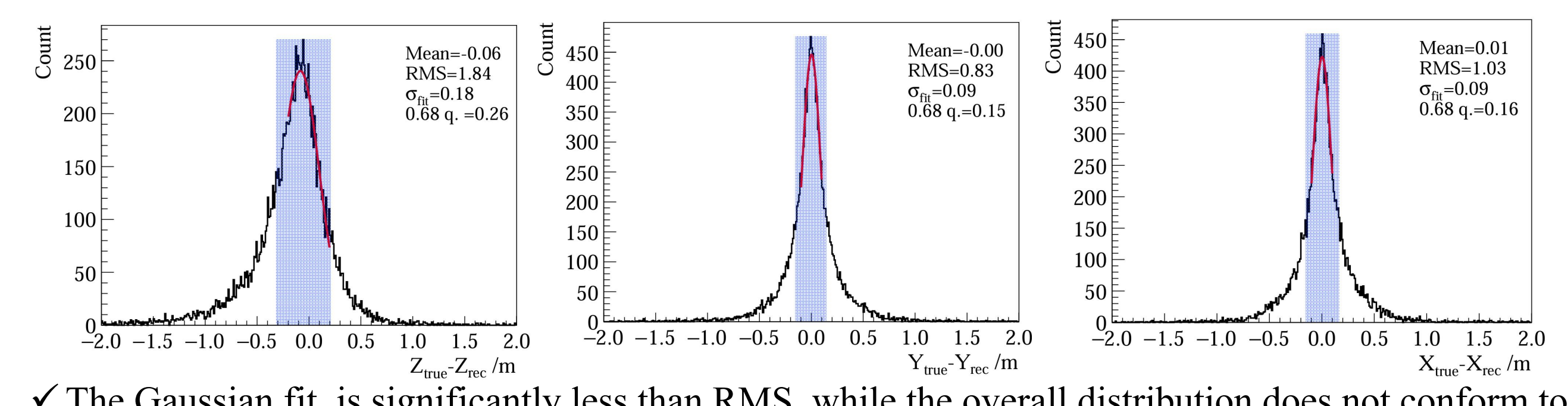
$$\chi^2 = \sum_i \frac{(T_{pre} - T_{obs})^2}{\sigma_{tts}^2}$$

- $t_{incident}$: The time of muon incident into CD from the earlist PMT
- t_{muon_tof} : Time of flight of muon
- t_{photon_tof} : Time of flight of photons considering effective refractive index
- t_{delay} : The delay time caused by the PMT readout electronics and TT (transit time) of PMT, can be calibrated using laser source
- t_{offset} : Describing the rise time from FHT to peak of waveform
- σ_{tts} : TTS of PMT

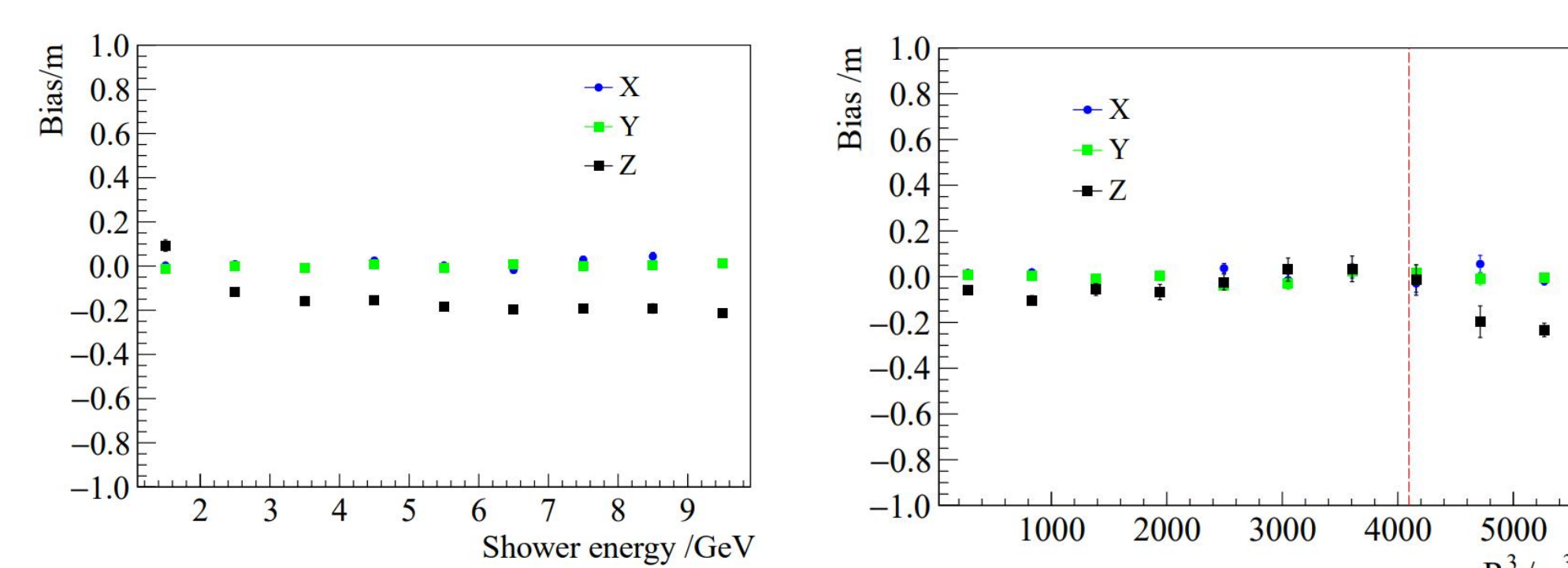


Results

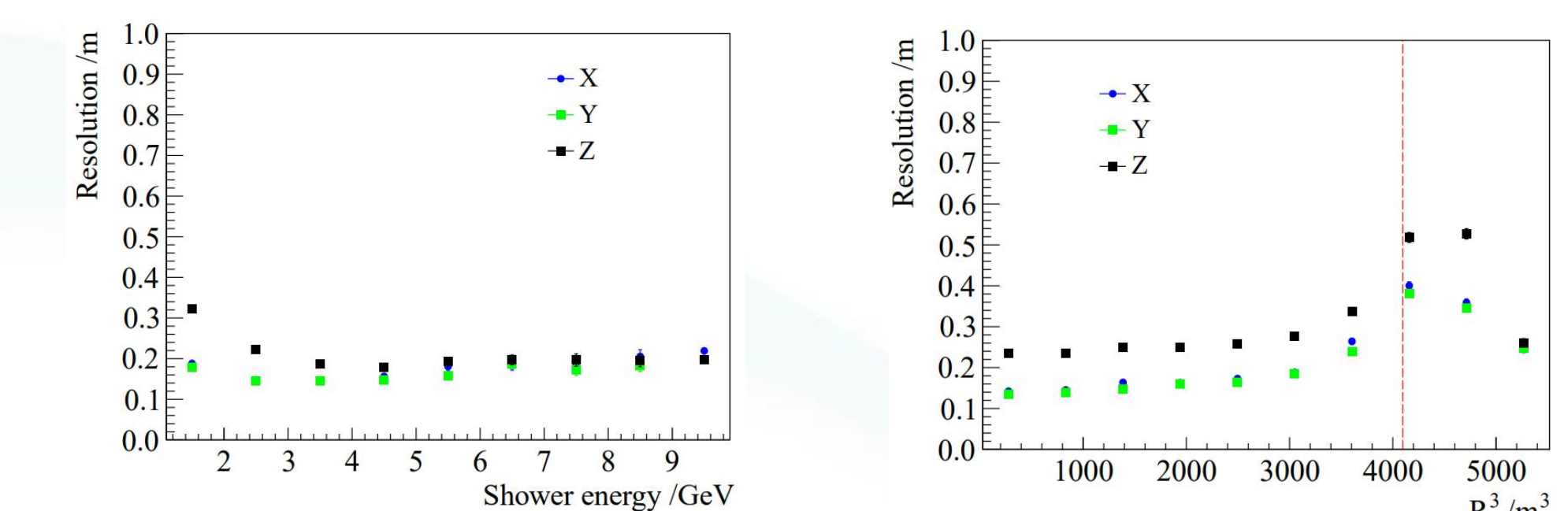
Single shower vertex reconstruction result:



- ✓ The Gaussian fit is significantly less than RMS, while the overall distribution does not conform to a Gaussian distribution, so we use 0.68 quantile as resolution.
- ✓ The reconstruction resolution (68 % quantile) of x and y is about 0.16 m and 0.15 m. The reconstruction resolution of z is poor, and about 0.26 m.



- ✓ The bias of x/y is smaller than that of z. The bias of x/y does not vary significantly with the shower energy and vertex position.
- ✓ The bias of z becomes worse as the shower energy increases. The reason is that when the shower energy is larger, the shower forward-peaking is more pronounced, which makes recon. more difficult.



- ✓ The reconstruction resolution (68 % quantile) of x and y is less than 0.4 m. The reconstruction resolution of z is poor, and less than 0.6 m.
- ✓ The reconstruction performance is relatively poor when the shower energy less than 2 GeV and shower vertex at the edges of the detector.

Conclusions

- ✓ The muon shower characteristic has been presented and study the isotopes yield contributed from shower.
- ✓ Shower vertex was reconstructed based on the waveform information.
 - Eliminate the muon track impact by subtracting the waveform from track component.
- ✓ The reconstruction performance:
 - Single shower vertex within 16 m after considering muon track reconstruction performance:
 - Resolution (68% quantile) < 0.4 m for z and < 0.3 m for x/y of shower vertex.

Bibliography

1. JUNO Collaboration. "Potential to Identify the Neutrino Mass Ordering with Reactor Antineutrinos in JUNO". 2025 Chinese Phys. C 49 033104.
2. J. Collaboration, Juno physics and detector, Progress in Particle and Nuclear Physics 123,103927 (2022).