

The DUNE Experiment: Status and Physics Prospects

Luis Manzanillas

On behalf of the DUNE collaboration

*The XIX International Conference on Topics in Astroparticle and Underground Physics
(TAUP2025)*

Xichang, China – August 29, 2025



NUCLÉAIRE
& PARTICULES



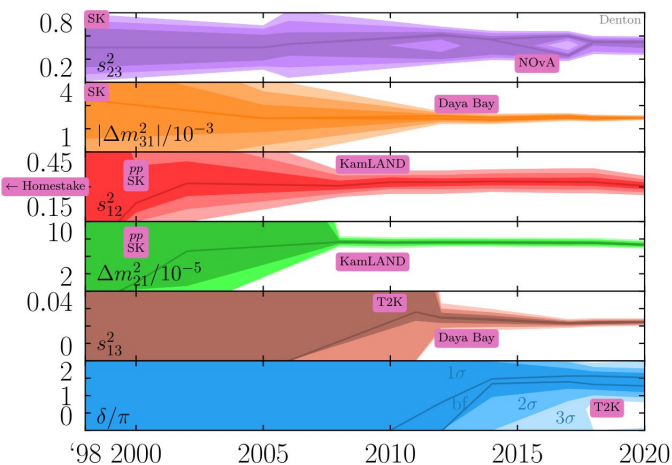
UNIVERSITÉ
SAVOIE
MONT BLANC



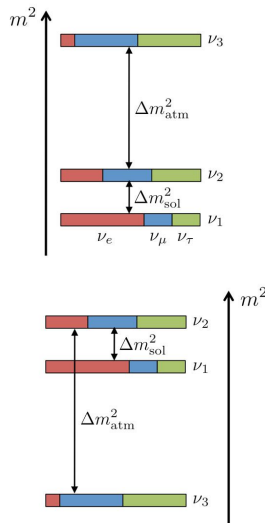
Neutrino physics: The current picture

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U_{PMNS} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$U_{PMNS} = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}}_{\text{Atmospheric Accelerator}} \underbrace{\begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix}}_{\text{Reactor Accelerator}} \underbrace{\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{Solar Reactor}} \underbrace{\begin{pmatrix} i & 0 & 0 \\ 0 & e^{i\alpha} & 0 \\ 0 & 0 & e^{i\beta} \end{pmatrix}}_{\text{O}\nu\beta\beta}$$



arXiv: [2212.00809](https://arxiv.org/abs/2212.00809)



So far data can be “described” in the 3ν framework
Most of the parameters have been determined ✓

But ...

CP violating phase?

Mass ordering?

Octant of θ_{23} ?

Sterile neutrinos?

Is the 3-flavor model correct?

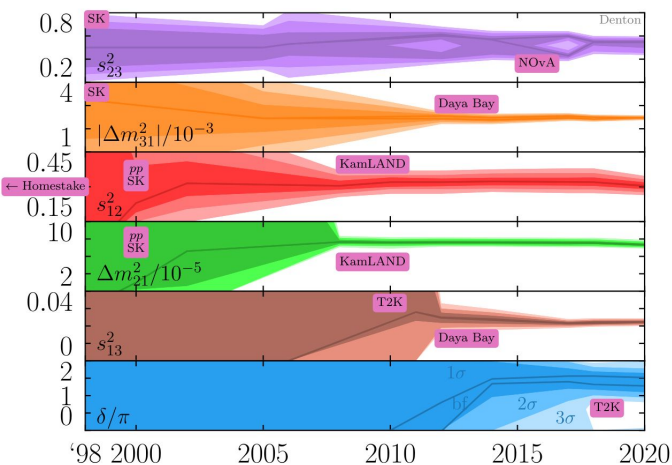
Dirac or Majorana?

Absolute mass scale?

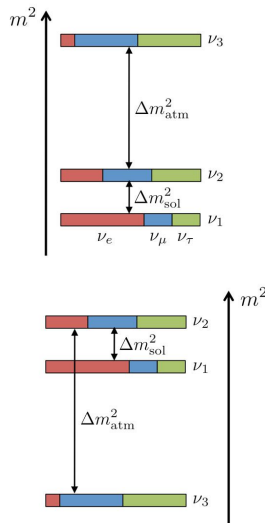
Neutrino physics: The current picture

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U_{PMNS} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$U_{PMNS} = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}}_{\text{Atmospheric Accelerator}} \underbrace{\begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix}}_{\text{Reactor Accelerator}} \underbrace{\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{Solar Reactor}} \underbrace{\begin{pmatrix} i & 0 & 0 \\ 0 & e^{i\alpha} & 0 \\ 0 & 0 & e^{i\beta} \end{pmatrix}}_{\text{O}\nu\beta\beta}$$



arXiv: [2212.00809](https://arxiv.org/abs/2212.00809)



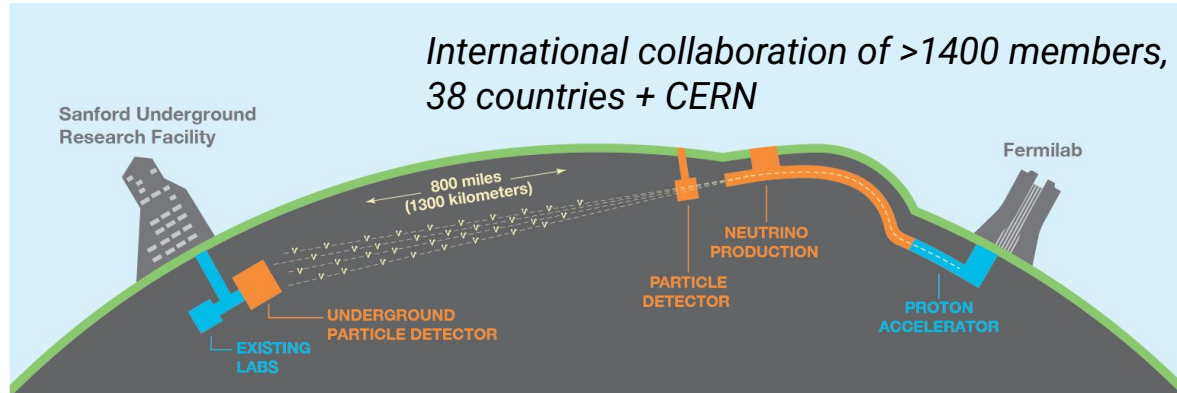
So far data can be “described” in the 3ν framework
Most of the parameters have been determined ✓

But ...

DUNE
scope

- CP violating phase?
- Mass ordering?
- Octant of θ_{23} ?
- Sterile neutrinos?
- Is the 3-flavor model correct?
- Dirac or Majorana?
- Absolute mass scale?

The Deep Underground Neutrino Experiment

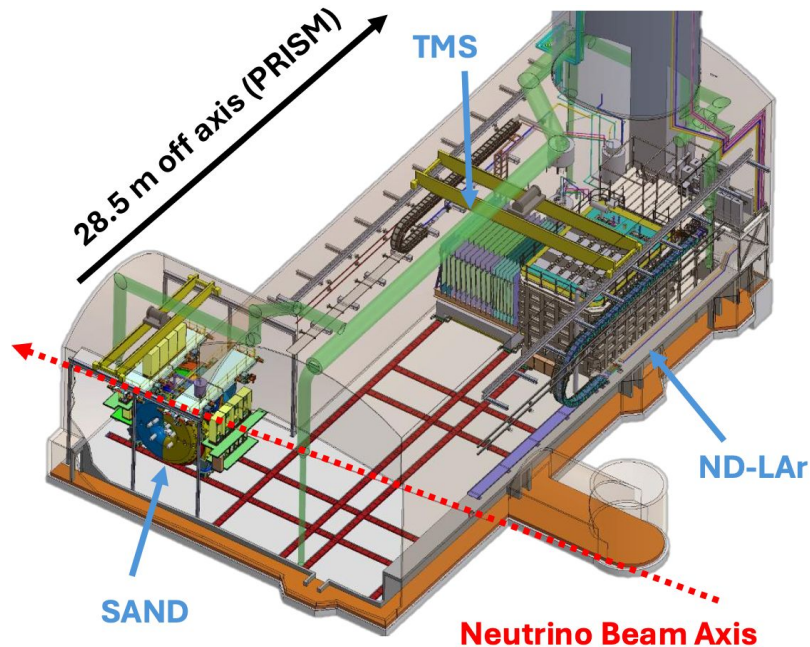


● DUNE – Next-Generation Long-Baseline Neutrino Experiment

- **Baseline:** 1300 km from Fermilab to SURF (South Dakota) → enables *unambiguous* neutrino mass ordering determination
 - **Far Detector Location:** 1.5 km underground at the Sanford Underground Research Facility
 - Project Structure: Developed in **two phases**
 - **Beam Power:** World's most powerful neutrino beam – 2 MW ν beam from Fermilab
 - **Wide-band neutrino energy spectrum** allows **precise measurement of oscillation parameters**
- ## ● Detector systems:
- **Near Detector (ND):** 574 m from source; measures unoscillated neutrino flux
 - **Far Detector (FD):** 3 × 17 kton **Liquid Argon TPC** modules + module of opportunity (LArTPC or other) at SURF

Near detectors

The Near Detector complex will constrain DUNE's flux and cross section systematic uncertainties



Three detectors located 574 m from the beam & 60 m underground

- **ND-LAr:** 50t fiducial volume detector
 - Primary target, FD technology + pixelated readout
- **The muon spectrometer (TMS)**
 - Measure μ 's escaping the first detector
- **SAND:** Tracker surrounded by an electromagnetic calorimeter and magnet
 - Monitor the neutrino beam

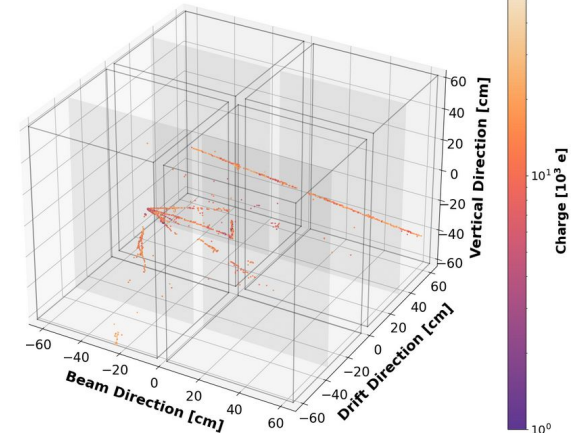
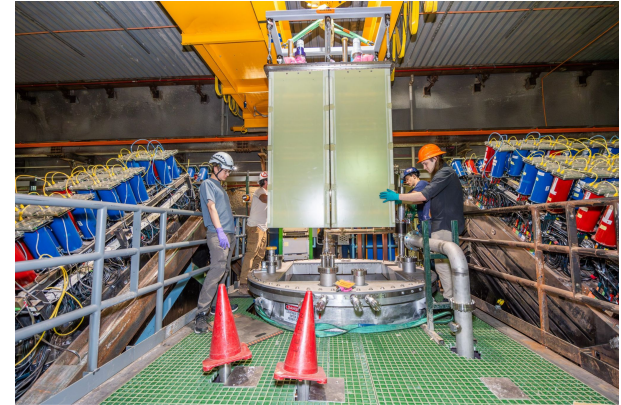
Talk by D. Casazza

ND-LAr and TMS are movable detectors: This is the DUNE-PRISM concept

Near detector prototype: 2x2 ND-LAr

Goals: *Demonstrate modular design with natively 3D pixelated readout*

- Four **LArTPC modules** built and operated in LAr in Bern with a total of $\sim 330\text{k}$ pixel channels
- Operation of 2x2 ND-LAr in NuMI Neutrino Beam at FERMILAB
 - Four TPC modules installed in former location of MINOS-ND
 - Includes upstream/downstream trackers from MINERvA
 - Development of **neutrino analysis pipelines**
 - **Performance demonstrations**

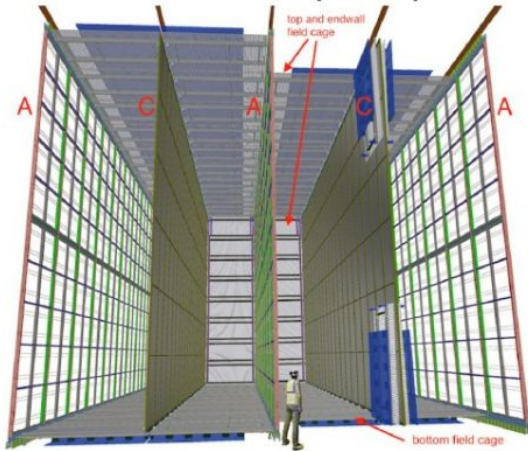


Far detector technologies: HD vs VD

Horizontal drift:

- Wire readout planes (**APAs**)
- **Four drift regions** of ~ 3.5 m
- X-ARAPUCA light readout

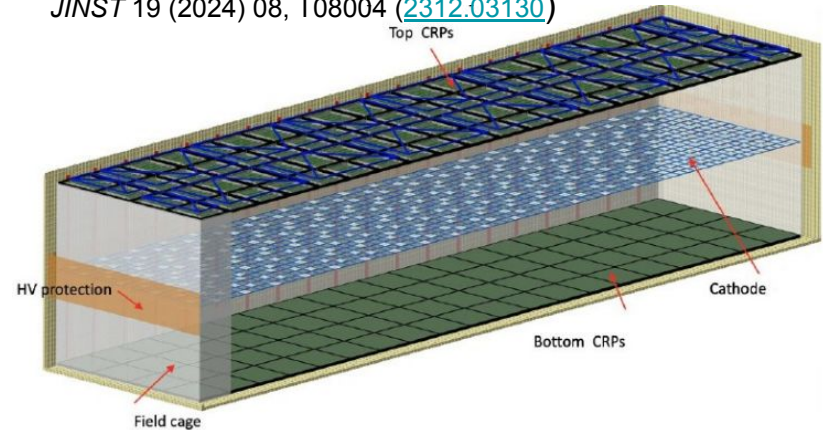
JINST 15 (2020) 08, T08010 ([2002.03010](#))



Vertical drift:

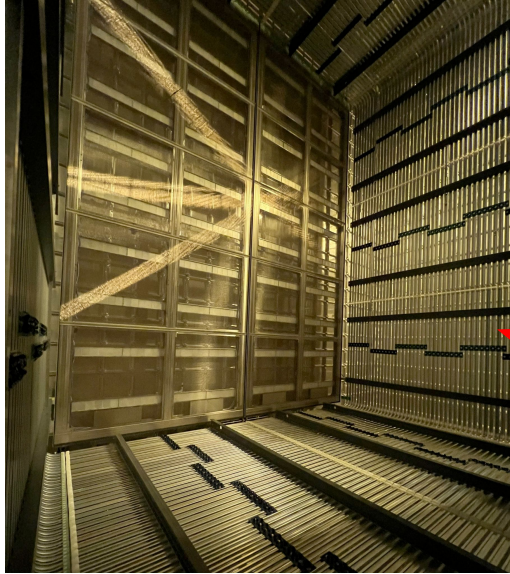
- Charge readout planes (**CRPs**) with PCB-based anodes
- **Two 6.25 m drift regions** and central cathode
- X-ARAPUCA light readout

JINST 19 (2024) 08, T08004 ([2312.03130](#))

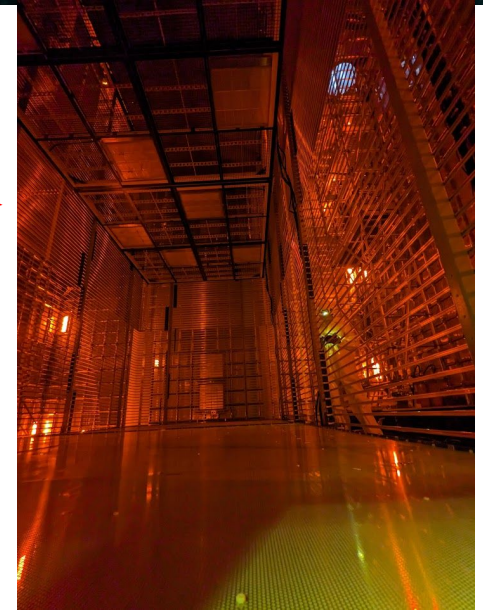
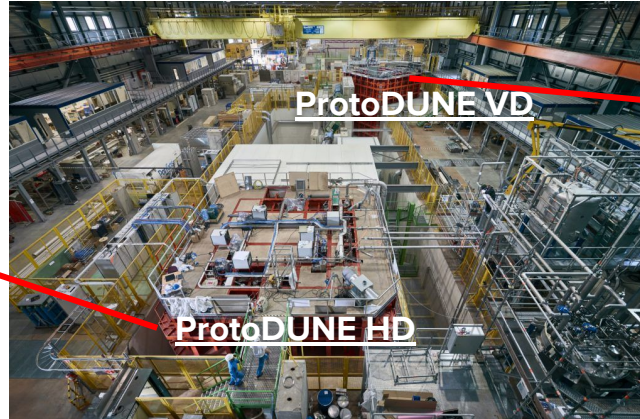


VD simpler to install → first DUNE FD module will use vertical drift

Far detector prototypes: ProtoDUNEs at CERN



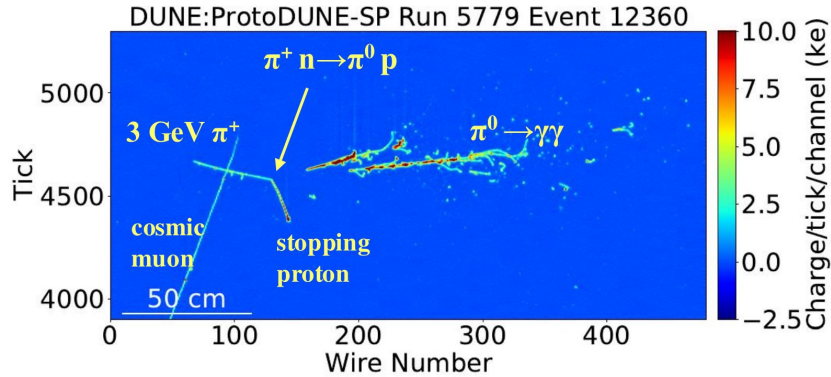
CERN neutrino platform



- Two 750 t LAr detectors with **real-size** readout elements (**APA, CRP, PDS**)
- **Validate the procurement & installation** sequence with 1:1 components
 - Estimate future resources needed and approve **final detector components design**
 - Cosmic and hadron **beam data** to allow **physics studies** and assess the **LArTPC performances**

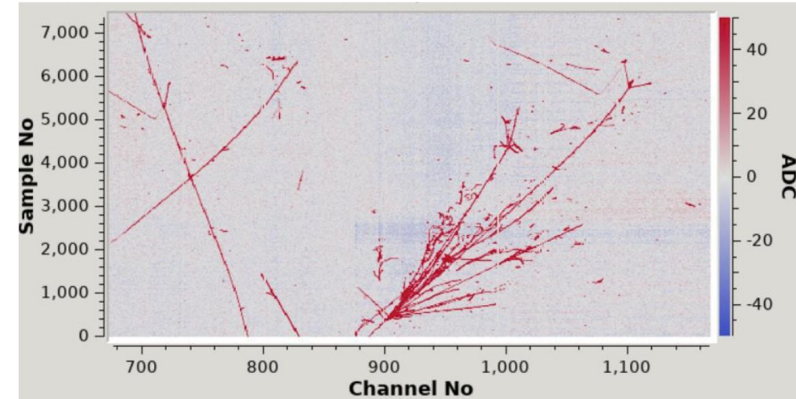
Far detector prototypes: ProtoDUNEs at CERN

ProtoDUNE-HD



- ProtoDUNE-HD completed LAr filling on April 30th, 2024
 - Physics data taking for several months with beam + cosmics + radioactive sources
- Excellent LAr purity: **>30 ms e^- lifetime!**
 - DUNE requirement: 3 ms
- Physics analysis ongoing

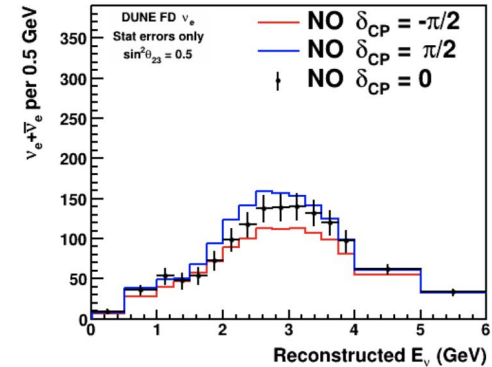
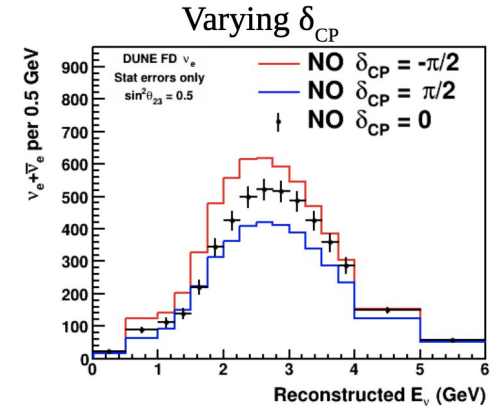
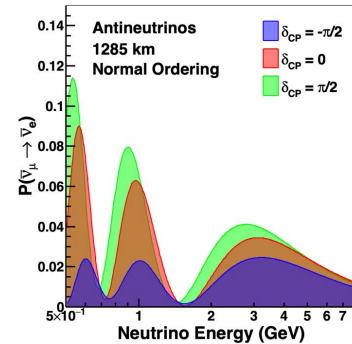
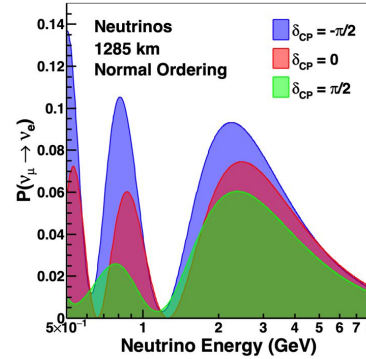
ProtoDUNE-VD



- Test of the **VD concept** for the first time at **large scale**
- LAr transferred from ProtoDUNE-HD completed on December 2025
- Commissioning being completed
 - **First data with nominal operational conditions already taken**
 - Cosmics tracks nicely reconstructed

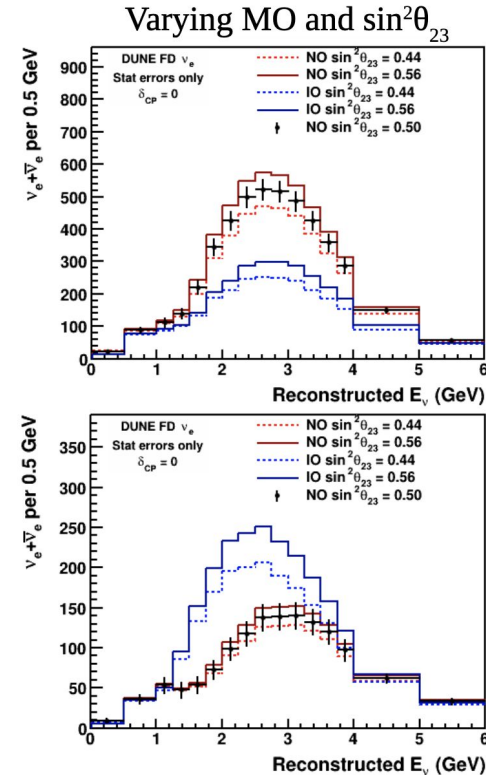
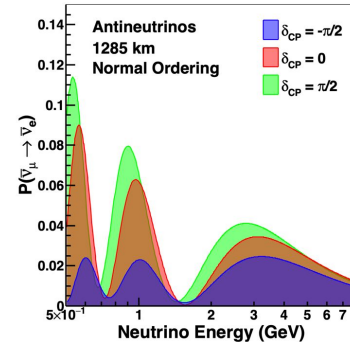
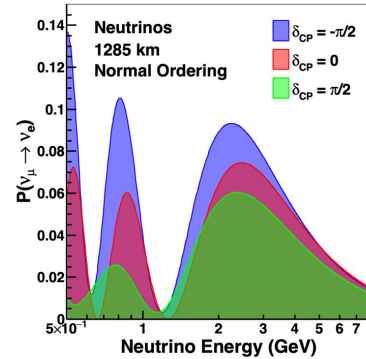
Neutrino energy at the far detectors

- If $\delta_{CP} \sim -\pi/2$
 - DUNE will measure an **enhancement** in electron **neutrino** appearance, and a **reduction** in electron **antineutrino** appearance



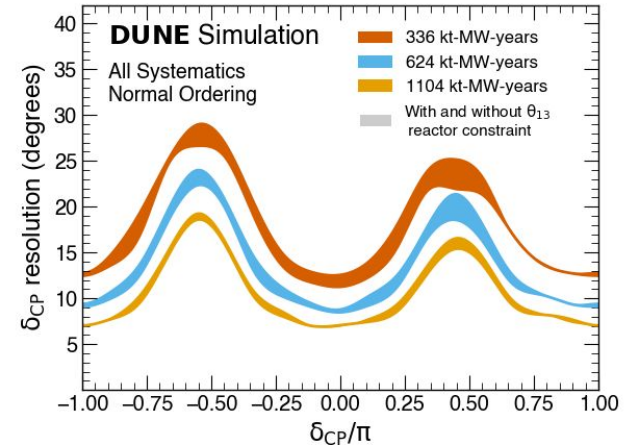
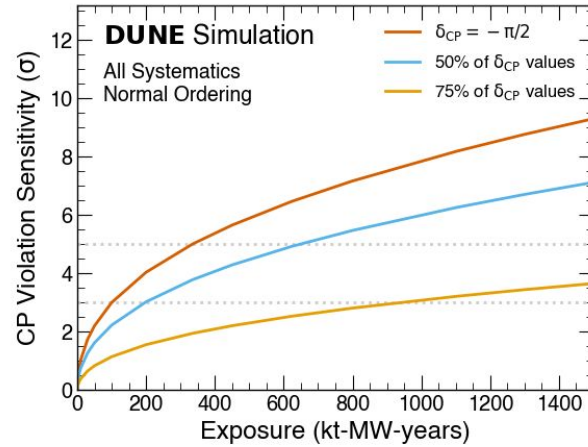
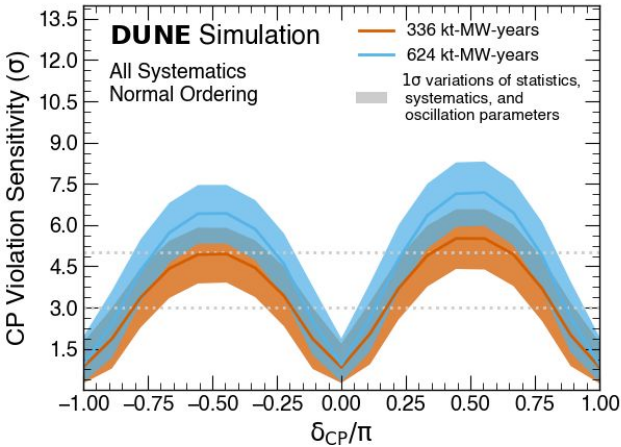
Neutrino energy at the far detectors

- If $\delta_{CP} \sim -\pi/2$
 - DUNE will measure an **enhancement** in electron **neutrino appearance**, and a **reduction** in electron **antineutrino appearance**
- If the **mass ordering** is normal
 - DUNE will measure a much **larger enhancement** in electron neutrino appearance, and a reduction in electron antineutrino appearance
- **MO, δ_{CP} , and θ_{23}** all affect spectra with different shape
 - \rightarrow additional handle on resolving degeneracies
- If **new physics is present**
 - There may be no combination of MO, δ_{CP} , and θ_{23} that fits data



CP sensitivity

- DUNE can establish CPV
 - $>3\sigma$ in 3.5 years for best scenario
 - Over 50% of δ_{CP} values at $>5\sigma$
 - Over 75% of δ_{CP} values at $>3\sigma$ (worst case scenario)
- After 10 years of exposure, strong CP violation discovery potential for both mass ordering options



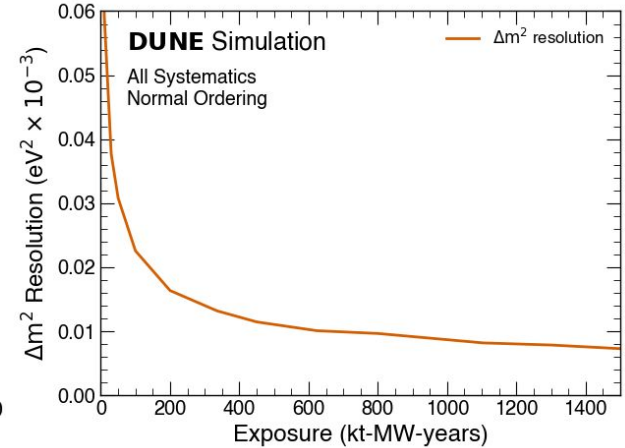
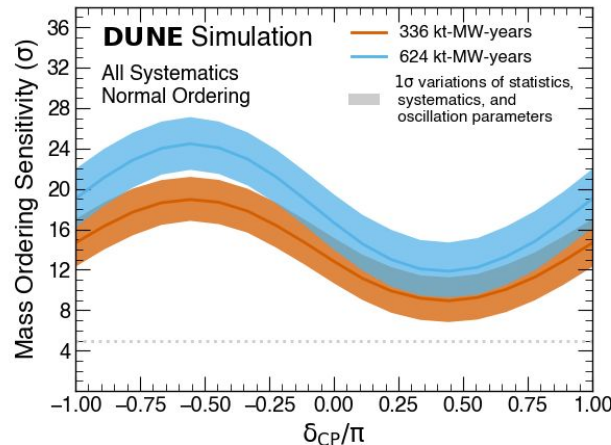
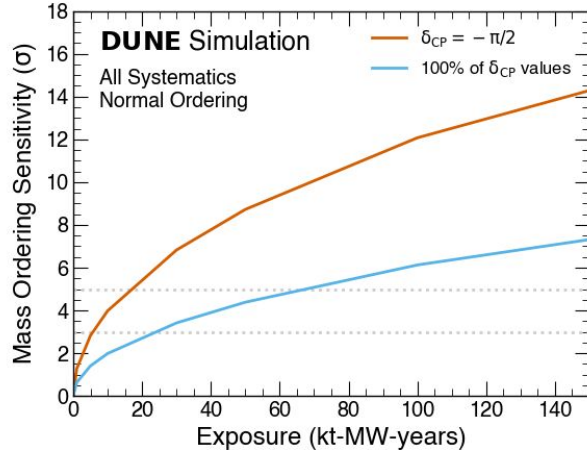
Eur. Phys. J. C 80, 978 (2020)

Mass ordering sensitivity

Regardless of the values of the other oscillation parameters:

- DUNE can establish mass ordering at $> 5\sigma$ in 1 year (best case scenario)
- DUNE can establish mass ordering at $> 5\sigma$ in 3 year (worst case scenario)

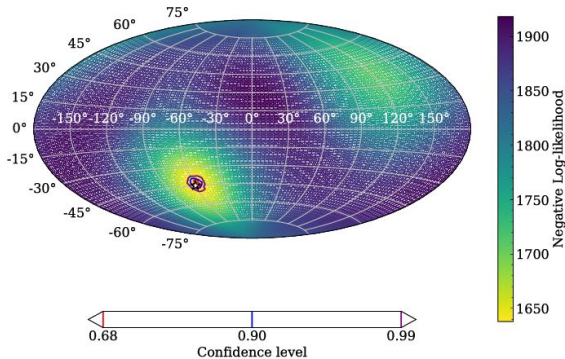
Eur. Phys. J. C 80, 978 (2020)



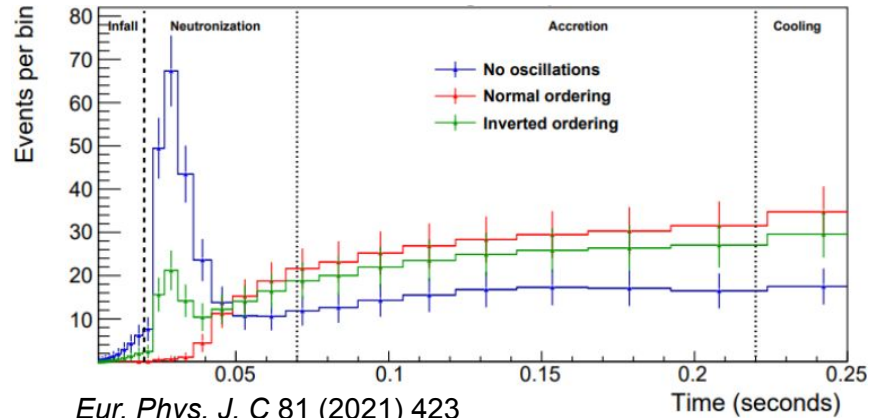
Astrophysical neutrinos

- **Ar target** provides unique sensitivity to MeV electron neutrinos
 - **CC channel:** $\nu_e + {}^{40}\text{Ar} \rightarrow e^- + {}^{40}\text{K}^*$
 - **ES channel:** $\nu_x + e^- \rightarrow \nu_x + e^-$ (pointing)
- Neutrinos from **core-collapse supernovae**
 - Neutronization burst measurements \rightarrow mass ordering measurement
 - ES channel enables $\sim 5^\circ$ pointing resolution (40 kt, 10 kpc)
- **Solar neutrinos**
 - Discovery sensitivity to the **hep solar flux**
 - **Improve upon existing solar oscillation measurements** via day-night asymmetry induced by Earth matter effects

Poster by S. Butchart



Phys.Rev.D 111 (2025) 9, 092006



Eur. Phys. J. C 81 (2021) 423

Construction status



- Cryostat installation starting in 2026
- Inner detector components installation starting in ~2027
- Physics data taking starting in ~2030

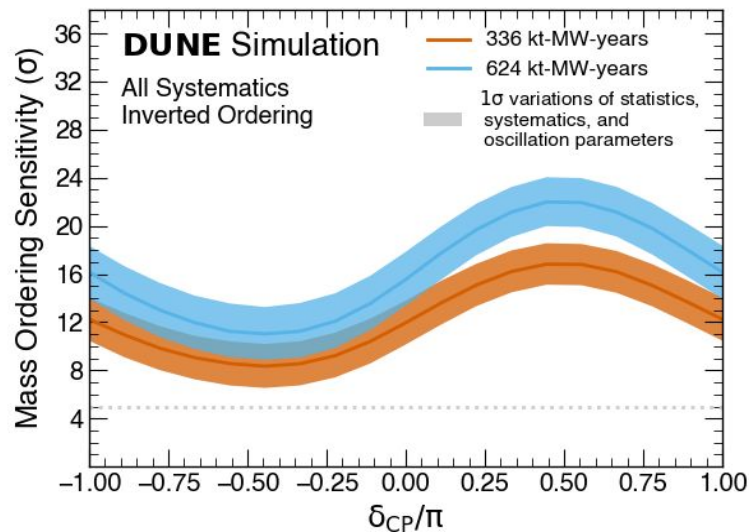
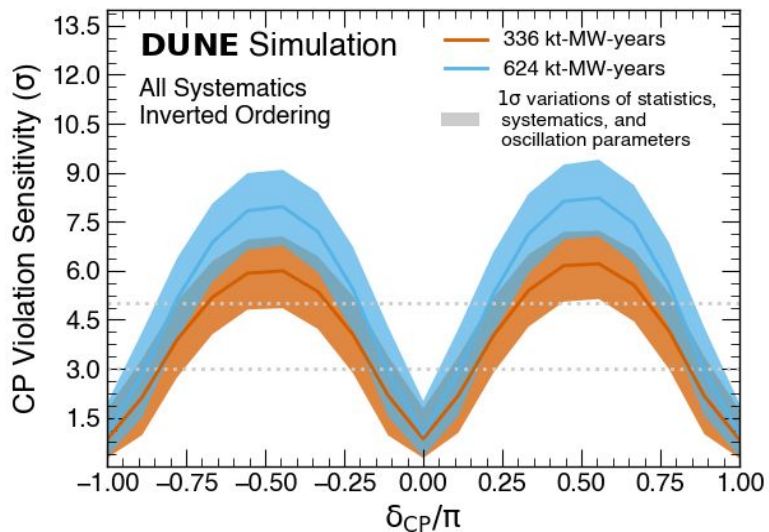
DUNE: Summary & Outlook

- World-leading long-baseline neutrino experiment under construction in the U.S. aimed at major discoveries:
 - **Neutrino mass ordering**
 - **Leptonic CP violation**
- Unique sensitivity to MeV-scale neutrinos:
 - Excellent **detection of supernova ν_e**
 - Potential to observe **hep solar flux**
 - Precision on solar neutrino parameters
- Complementary to the global neutrino effort
- Active prototyping at CERN & Fermilab
- **DUNE science starts this decade!**



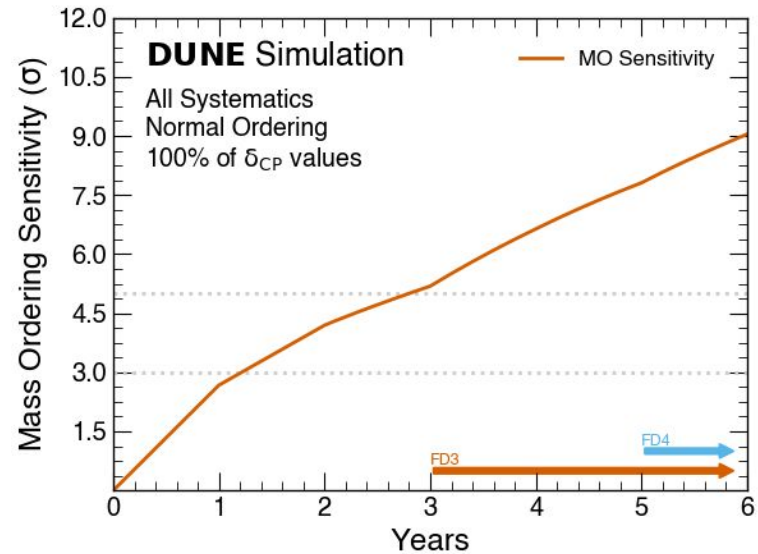
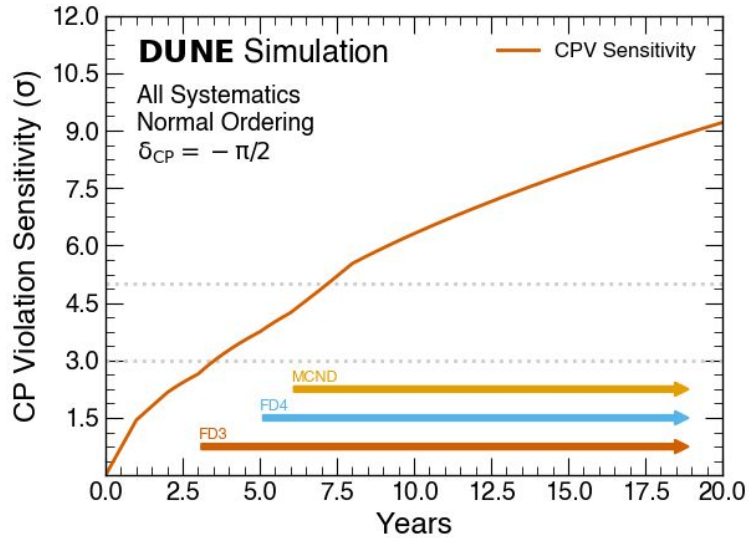
Thank you for your attention!

Inverted ordering



Eur. Phys. J. C 80, 978 (2020)

Sensitivity and FD modules



Eur. Phys. J. C 80, 978 (2020)

MeV neutrinos

