

Coherent Neutrino Scattering, Searches for sterile States

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The XIX International Conference on Topics in Astroparticle and Underground Physics

TAUP2025



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The Motivation for sterile Neutrinos

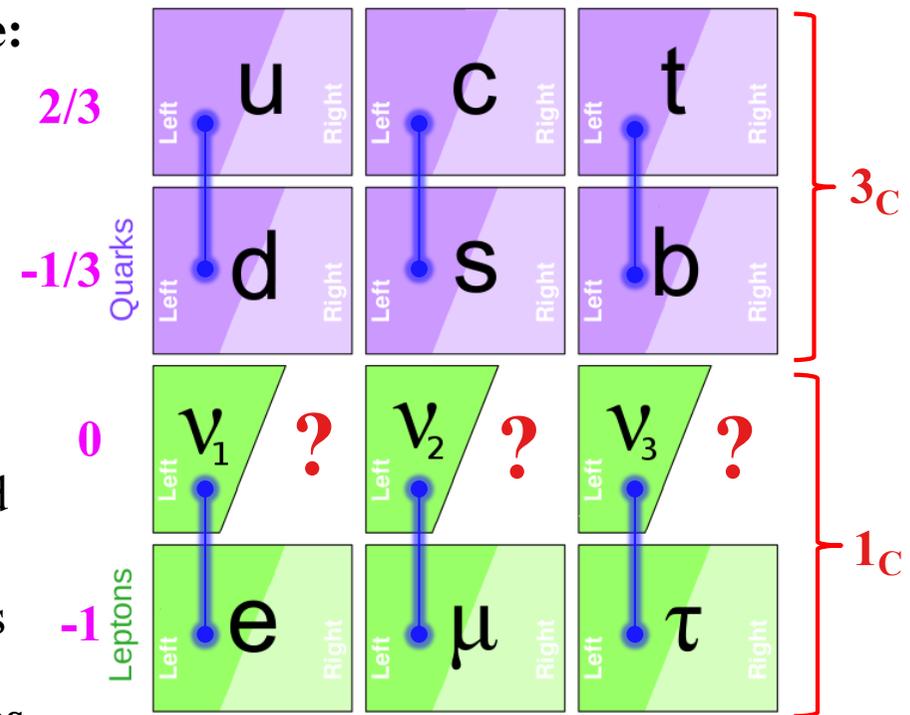
the Standard Model = facts learned from Nature:

- success of renormalizable QFT \rightarrow SM
- gauge group: $SU(3)_c \otimes SU(2)_L \otimes U(1)_Y$
- **chiral** representations: 2_L and 1_L , 3_C and 1_C
- electric charge: $Q = T_3 + Y/2$

questions:

- which representations? why generations?
- ν_R states appear missing? $m_\nu=0 \leftarrow$ not required
- **neutrino masses \rightarrow simplest option: add ν_R**
- beware: image is easily misleading: 1,2,3, GUTs

$\rightarrow \nu_R = (1_C, 1_L, Y=0)$ has important consequences



$$\begin{array}{c}
 Y_N \\
 \hline
 \nu_L \quad \nu_R \\
 \vdots \\
 \times \\
 m_D = Y_N^* \langle \Phi \rangle
 \end{array}$$

$$\begin{array}{c}
 M_R \gg m_D \\
 \hline
 \nu_R \quad \nu_R \\
 \times \\
 \text{Majorana} \\
 \not\sim
 \end{array}$$



6x6: see-saw \rightarrow block diagonalization

$$\begin{pmatrix} \bar{\nu}_L & \bar{\nu}_R^c \end{pmatrix} \begin{pmatrix} 0 & m_D \\ m_D & M_R \end{pmatrix} \begin{pmatrix} \nu_L^c \\ \nu_R \end{pmatrix}$$

- **3 active: $\sim m_D^2/M_R$, 3 sterile $\sim M_R$**
- **expect tiny active/sterile mixings $\sim m_D/M_R$**

Constraints from Neutrino Oscillations

M_R very big \rightarrow ignore 6x6 picture \rightarrow 3x3 PMNS mixing matrix U of active neutrinos

$$\begin{bmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{bmatrix} = U \begin{bmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{bmatrix} = \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{bmatrix} \begin{bmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{bmatrix}$$

3 x 3 unitarity:
 $UU^+ = U^+U = 1$

normalization:

orthogonality:

$$|U_{\alpha 1}|^2 + |U_{\alpha 2}|^2 + |U_{\alpha 3}|^2 = 1 \quad (\alpha = e, \mu, \tau)$$

or

$$|U_{ei}|^2 + |U_{\mu i}|^2 + |U_{\tau i}|^2 = 1 \quad (i = 1, 2, 3)$$

$$U_{\alpha 1}U_{\beta 1}^* + U_{\alpha 2}U_{\beta 2}^* + U_{\alpha 3}U_{\beta 3}^* = 0 \quad (\alpha, \beta = e, \mu, \tau, \alpha \neq \beta)$$

or

$$U_{ei}U_{ej}^* + U_{\mu i}U_{\mu j}^* + U_{\tau i}U_{\tau j}^* = 0 \quad (i, j = 1, 2, 3, i \neq j)$$

\rightarrow oscillations from active to sterile neutrinos would show up as 3x3 unitarity violation

- global 3v oscillation fits work perfectly:
 - **BARI group:** Capozzi et al, PRD 104 (2021) 083031
 - **Valencia group:** globalfit.astroparticles.es
 - **NuFIT:** www.nu-fit.org
- fits without assumed unitarity also work very well \rightarrow **at most ~1-3% deviation**

$\rightarrow U_{\alpha 4}$ expected to be small both from see-saw and 3x3 unitarity

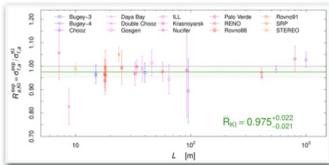
Old and newer Anomalies

→ results which neither fit into:

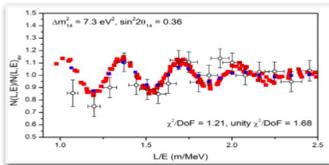
- a consistent 3+1 or 3+2, oscillation picture → other mechanisms?
- cosmology (CMB+LSS, BBN, ...)

reactor anomaly (flux)

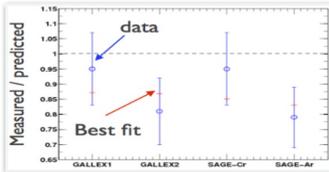
→ probably resolved with improved fluxes
(origin of 5 MeV bump still open)



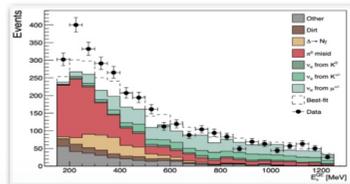
reactor spectra: **NEUTRINO-4**



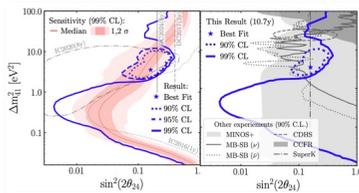
Gallium anomaly: **GALLEX/GNO, SAGE, BEST**



LSND, MiniBooNE



IceCube



L/E oscillation pattern?

oscillation rates with sterile ν's:

- disappearance
- appearance

$$P_{ee} \quad P_{e\mu}$$

$$P_{\mu e} \quad P_{\mu\mu}$$

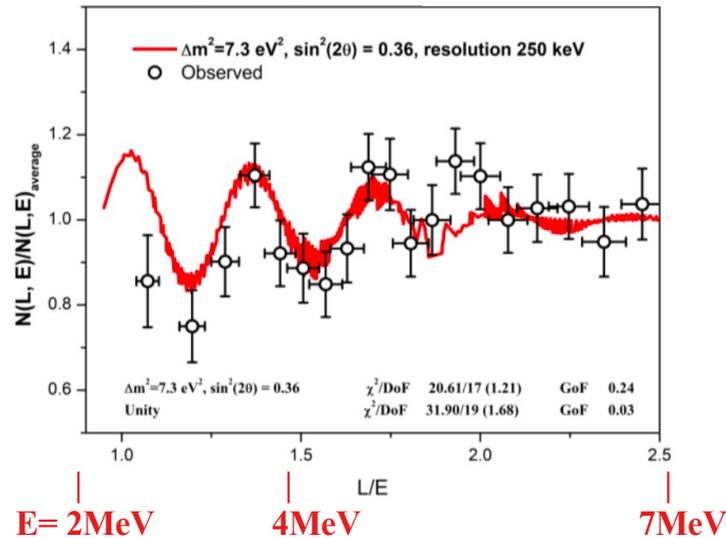
adding a 4th neutrino

→ +2 parameters: (U_{e4} , $U_{\mu4}$) + unitarity

NEUTRINO-4 ee

100 MW_{thermal} @SM-3 reactor @Dimitrovgrad (Russia)
anti-neutrino detector (LS) @5m
ON/OFF measurement

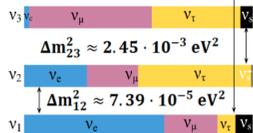
Phys.Rev.D 104, 032003 (2021)



$$\Delta m_{14}^2 = 7.3 \pm 1.17 \text{ eV}^2$$

$$\sin^2 2\theta_{14} = 0.36 \pm 0.12$$

$$2.9\sigma$$



concerns / debates about:

- analysis
- energy resolution
- background
- ...

→ oscillations: do not fit

→ other explanations???

other experiments with spectral information:

– STEREO (France), $L = 9\text{-}11\text{m}$ } on-going combined analysis + Daya Bay ...stay tuned

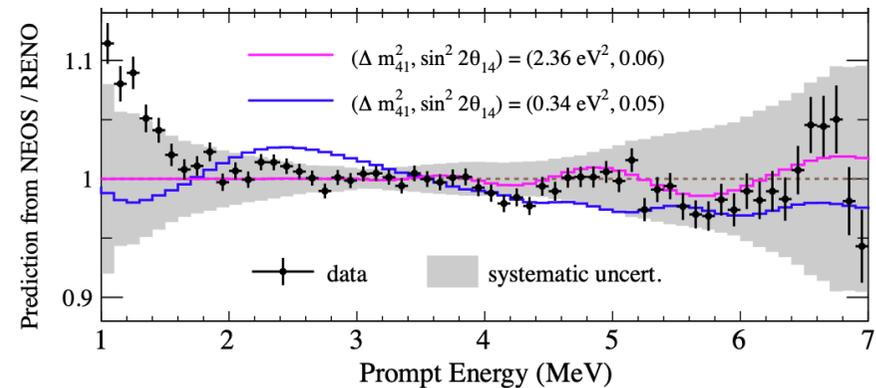
– PROSPECT (USA), $L = 7\text{-}12\text{m}$ }

– DANSS (Russia), $L = 10.9\text{-}12.9\text{m}$

– SoLid (Belgium), PRD 111, 072005 (2025)

→ no evidence

– NEOS + RENO (Korea), commercial, $L = 23.7\text{ m}$



→ wiggles? PRD 105 (2022) L11101

best fit $|\Delta m_{41}^2| = 2.41 \text{ eV}^2$; $\sin^2 2\theta_{14} = 0.08$ (1σ)

→ combined with others... not significant

experiments with near and far detector:

→ less reactor spectral uncertainties (bump, HE tail)

The Gallium Anomaly ee

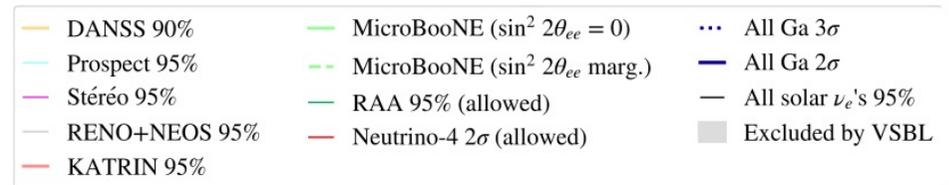
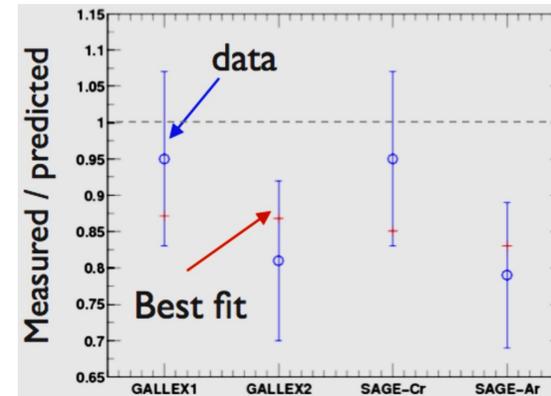


calibrated with intense ${}^{51}\text{Cr}$ and ${}^{37}\text{Ar}$ sources

~20 years ago: **GALLEX/GNO** and **SAGE**

now: **BEST** (2022)

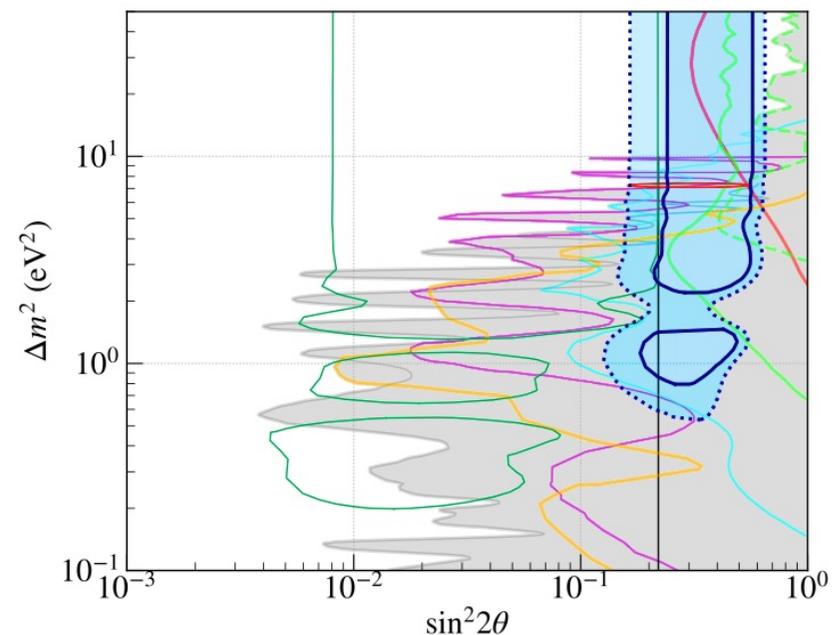
GALLEX	{	$R_1(\text{Cr}) = 0.953 \pm 0.11$	$\rightarrow 0.80 \pm 0.05$		
		$R_2(\text{Cr}) = 0.812 \pm 0.11$			
SAGE	{	$R_3(\text{Cr}) = 0.95 \pm 0.12$		$\sim 4\sigma$	
		$R_4(\text{Ar}) = 0.79 \pm 0.095$			
BEST	{	$R_5(\text{I}) = 0.791 \pm 0.05$			PRC 105, 065502 (2022)
		$R_6(\text{O}) = 0.766 \pm 0.05$			



BEST confirms **GALLEX** and **SAGE** deficits

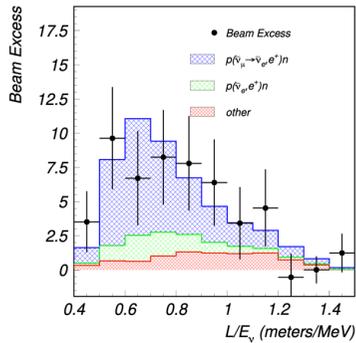
PRL 128, 232501 (2022)

- looks so far like a real effect
- oscillation into ν_s explanation excluded
- other physics? no known explanation



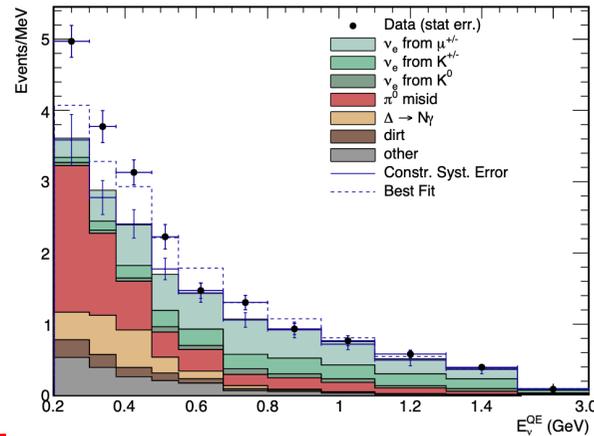
LSND - MiniBooNE - MicroBooNE μe

PRD 64:112007 (2001)



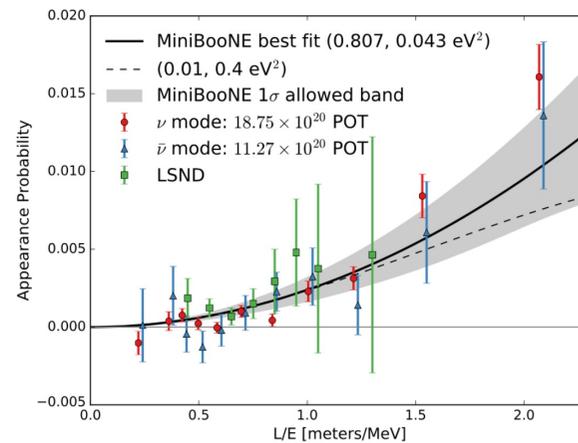
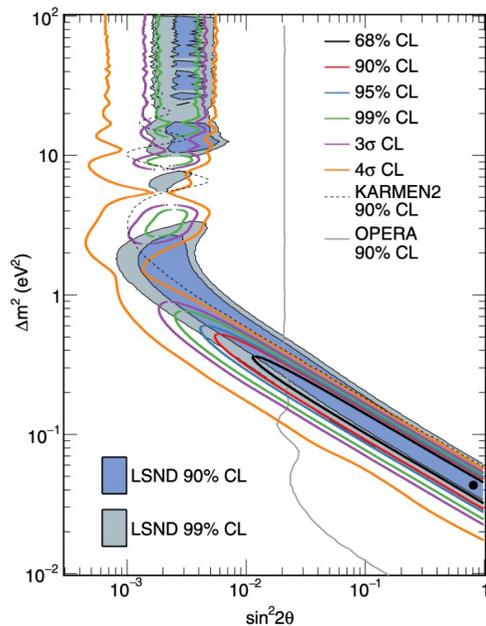
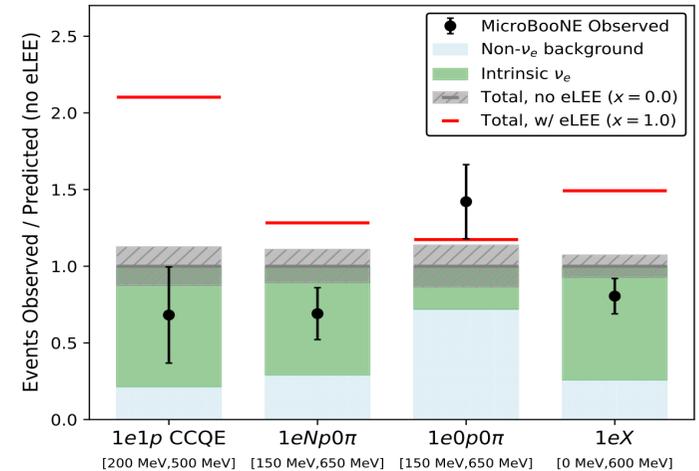
$\nu_\mu \rightarrow \nu_e$ appearance, $\sim 30m$
oscillation into ν_s ? $\rightarrow 3.8\sigma$

PRL 121, 221801 (2018), PRD 103, 052002 (2021)



PRL 128.241801 (2022)

inclusive and exclusive searches for ν_e CC



$\nu_\mu \rightarrow \nu_e$ appearance: 4.7σ
LSND + MiniBooNE: 6.1σ

red = expected from MiniBooNE
 \rightarrow no excess of low-energy ν_e

S. Martynenko @ NuDM24
MICROBOONE- Note-1127-PUB

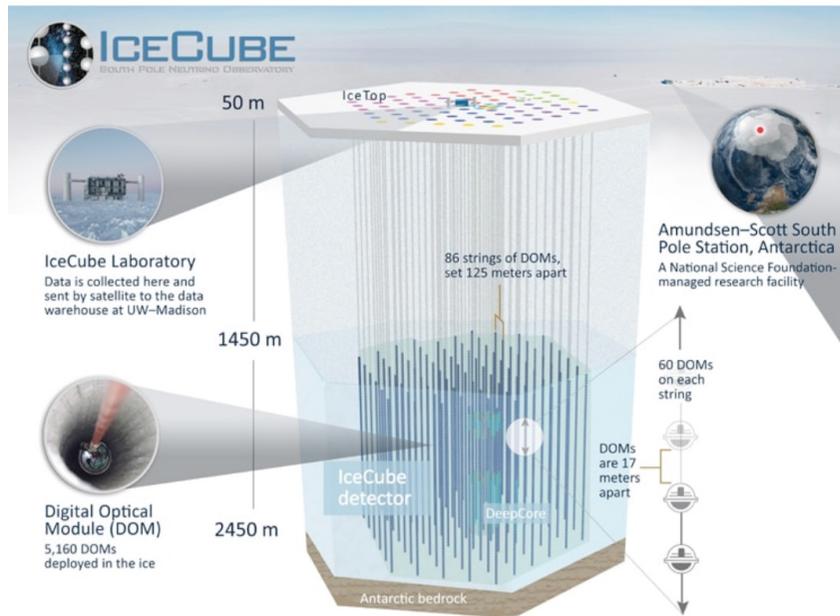
- $\sim 65\%$ more statistics
- new sideband constraints

\rightarrow no excess at $> 99\%$ CL
oscillations: tension app. /disapp.

\rightarrow looking forward to
FNAL SBN program

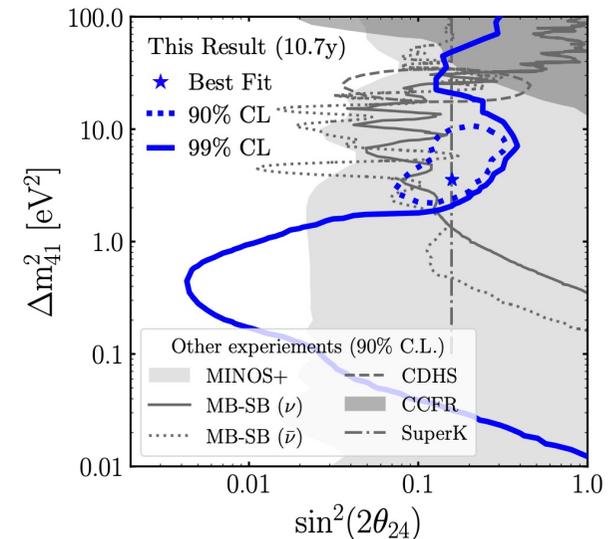
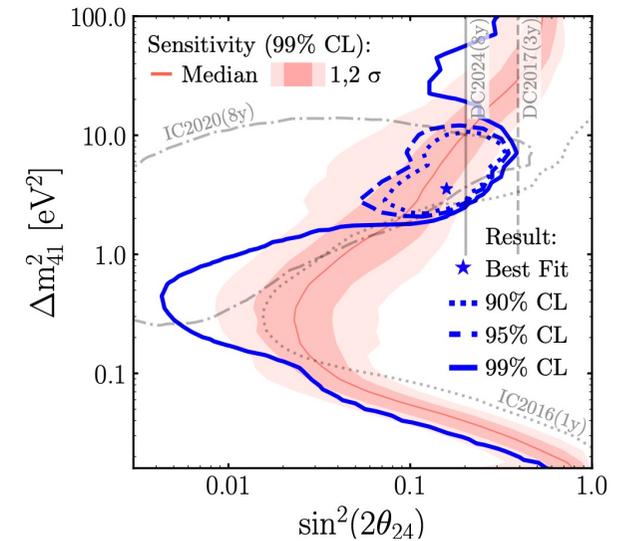
Atmospheric Neutrinos in ICECUBE $\mu\mu$

- km³ neutrino detector 1.5–2.5 km beneath the surface in the ice at the South Pole
- powerful sterile neutrino searches at both high (≥ 400 GeV) and low (≤ 60 GeV) energies
- sensitive to ν_μ disappearance at eV²



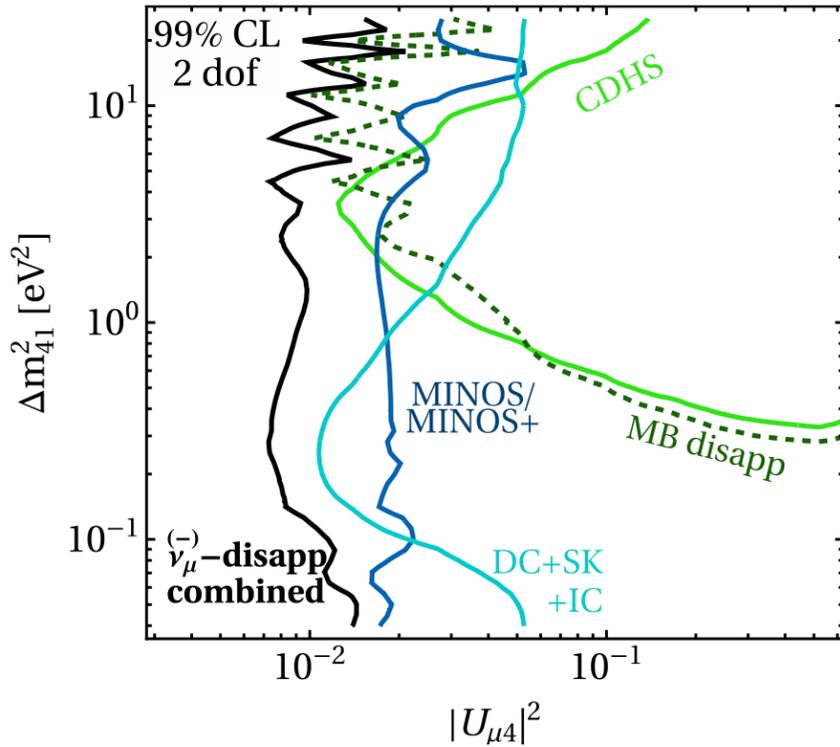
PRL 133 (2024) 20, 201804

- 3+1 sterile neutrino search using 10.7 years of IceCube data
- atmospheric ν_μ traversing Earth with 0.5-100 TeV
- best-fit for a 3+1 model: $\sin^2(2\theta_{24}) = 0.16$ and $\Delta m_{41}^2 = 3.5$ eV²
- comparison to MINOS+ and others \rightarrow tension/unlikley



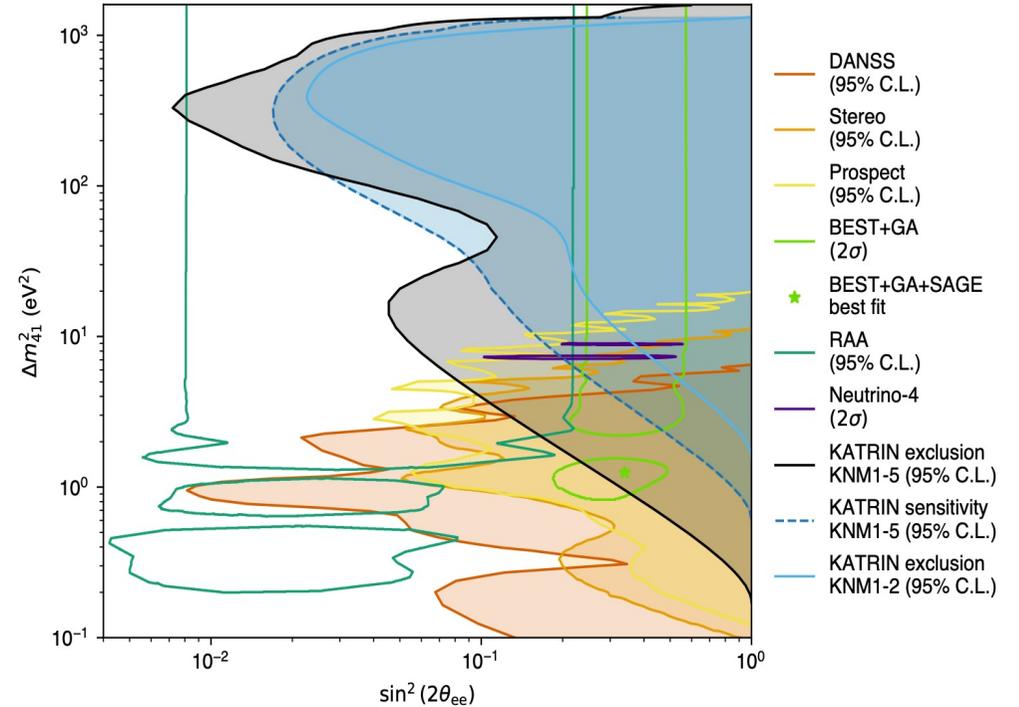
$\mu\mu$ Exclusion Limits ee

e.g, Dentler et al., JHEP 08 (2018) 010



a number of experiments with no hint for ν_μ disappearance

new result from KATRIN arXiv: 2503.18667



• 95% C.L. exclusion limits by KATRIN

shown also:

- DANSS
- STEREO
- Prospect

anomalies do not fit to an oscillations interpretation

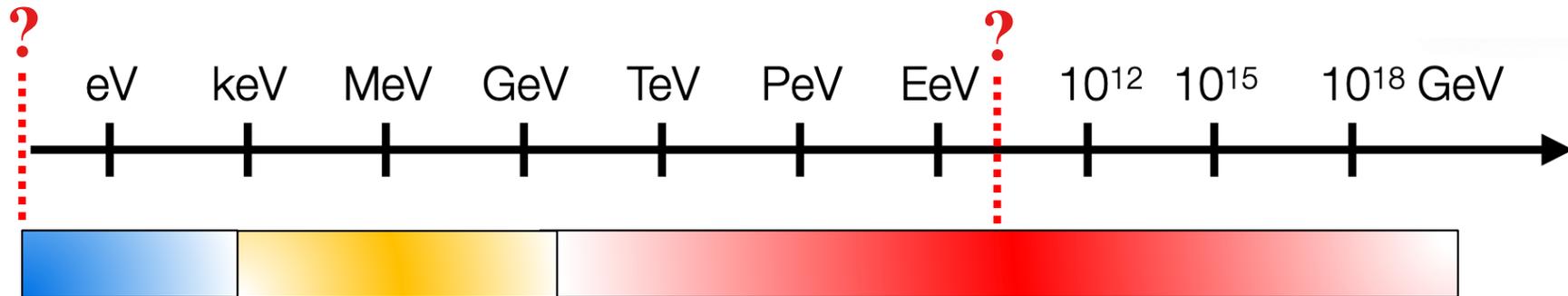
* reactor anomaly (RAA)

* gallium anomaly

* green star = best-fit BEST, GALLEX, SAGE

* Neutrino-4 (purple)

Sterile Neutrinos: A Bigger Perspective



modified oscillations
why small mixings?

long lived
heavy ν_s (HNL)

generic see-saw territory
tiny mixings naturally expected

anomalies:

3+1 \rightarrow tension in data samples

3+2, 3+n ...???

persist? \rightarrow other mechanisms?

oscillation is an assumption

\rightarrow alternatives...

- decay
- NSI = short distance transition
- time dependent
- dark sectors
- ...

+ **see-saw \leftrightarrow BAU**

\leftrightarrow many other BAU mechanisms

- 3rd hierarchy problem if M_R too heavy

\leftarrow improved EW fits with TeV-ish sterile ν_s ,

$\leftarrow M_R = g_Y \cdot \langle \phi \rangle$ with heavy $\langle \phi \rangle$ and tiny g_Y

• many connections to BSM, dark sectors,...

• cosmology

- keV sterile neutrinos as dark matter

- DESI+Planck+CMB

$$\Sigma m_\nu < 0.064 \text{ eV}, \Delta N_{\text{eff}} = 3.23 \pm 0.34 \text{ (SM=3.04)}$$

\rightarrow very important to clarify anomalies

sterile ν 's: no clear sign, but still well motivated!

Coherent Elastic Neutrino-Nucleus Scattering (CEvNS)

The simple picture:

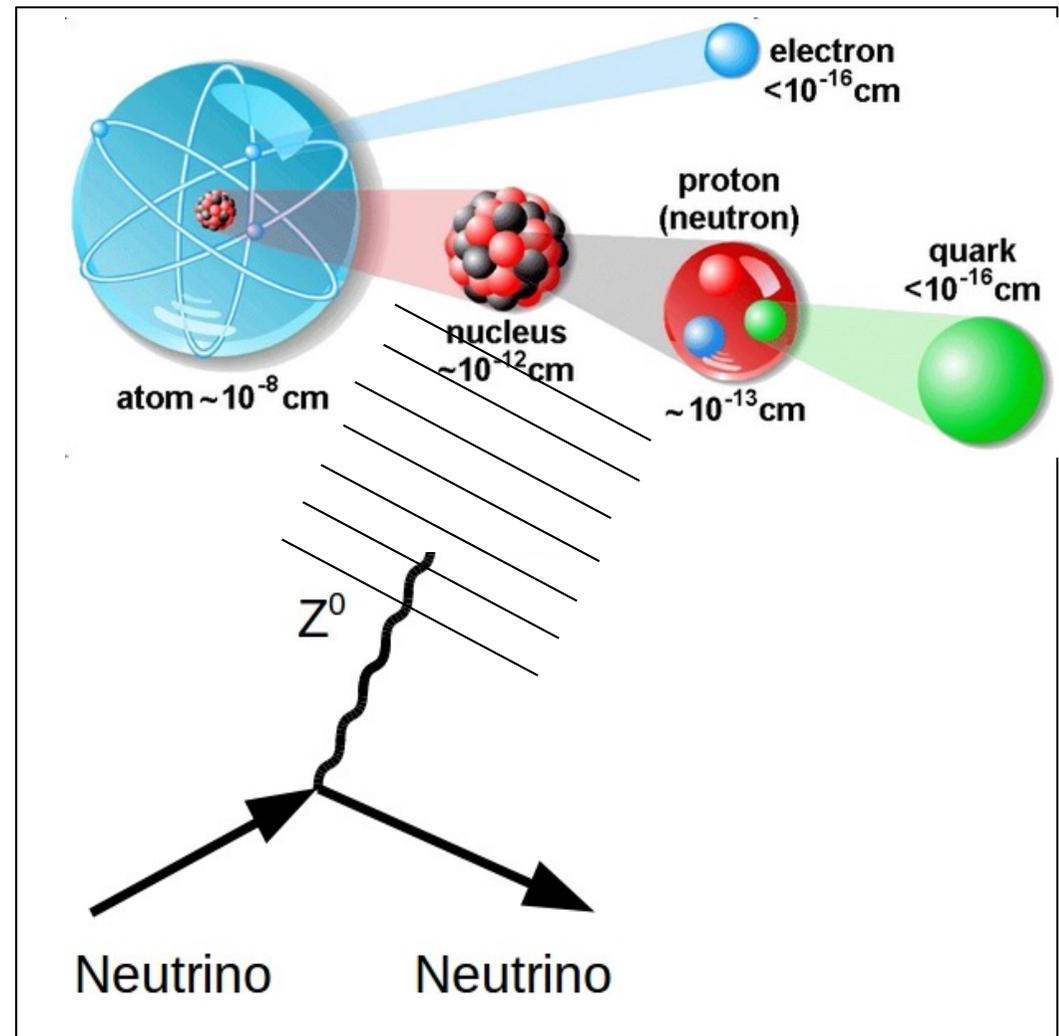
1. ν -wavelength big enough
→ Z scattering @ whole nucleus
2. small momentum transfer

→ coherent scattering

- Coherence length $\sim 1/E$
 - E_ν below $O(50)$ MeV
 - lower for full coherence
- cross section $\sim (E_\nu)^2$
 - overcompensate:
 - very high ν flux
 - coherence $\sim N^2$

E.g. Germanium target:

$N \simeq 40-44 \rightarrow N^2 \simeq 1600-1900 \rightarrow$ conventional detector mass 10t \leftrightarrow few kg

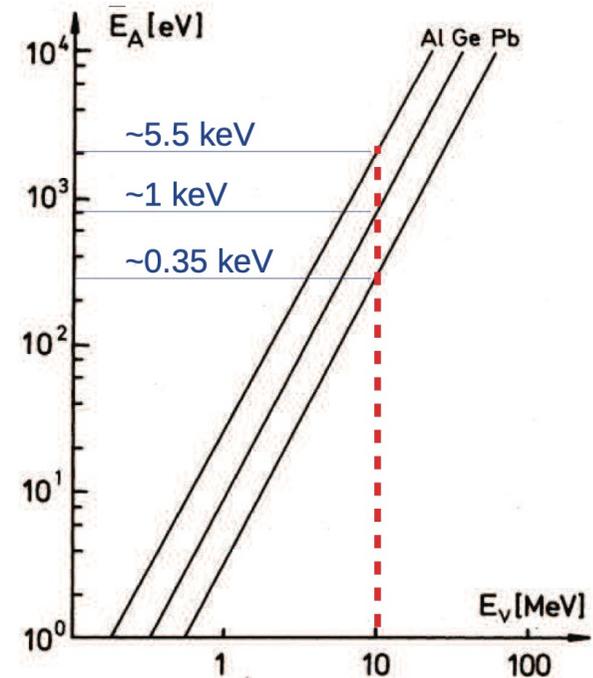


Z-scattering of ν with whole nucleus

$$\frac{d\sigma(E_\nu, T)}{dT} = \frac{G_f^2}{4\pi} Q_w^2 M \left(1 - \frac{MT}{2E_\nu^2}\right) F(Q^2)^2$$

$$Q_w = N - \underbrace{(1 - 4 \sin^2 \theta_w)}_{\simeq 0} Z \sim N$$

- neutrons dominate: $\sigma \sim N^2 \rightarrow$ enhancement
- form factor $F(Q^2) \leftrightarrow$ size of nucleus
- form factor matters for $E_\nu = 20\text{-}50$ MeV
- $F(Q^2) \rightarrow 1$ for $E_\nu \ll 50$ MeV
- kinematics:
 M: mass of nucleus
 E_ν : neutrino energy
 T: nuclear recoil energy
 θ : scattering angle

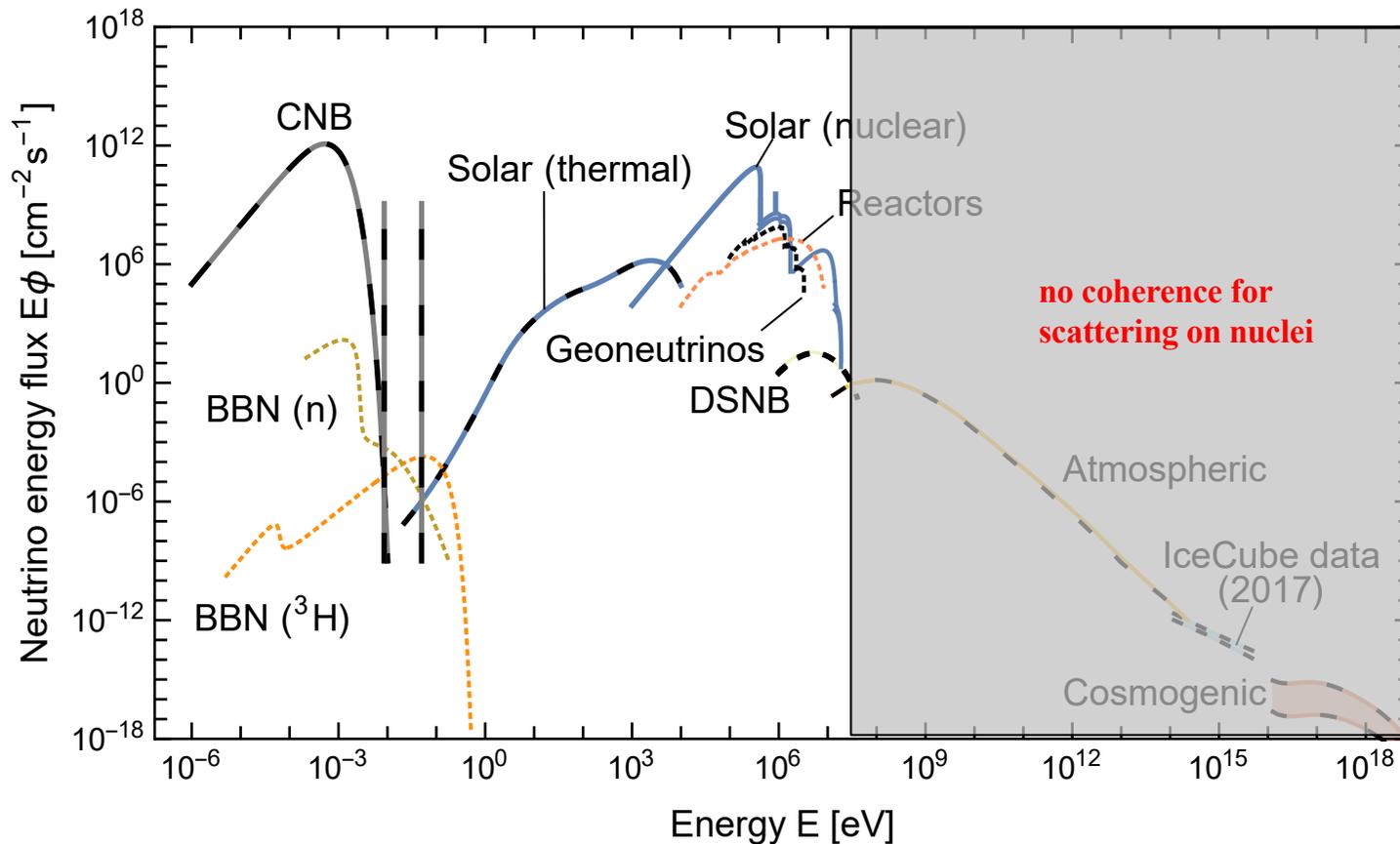


A.Drukier, L.Stodolsky, Phys.Rev.D 30 (1984) 11

$$|\vec{q}| = \sqrt{2MT} = \sqrt{2E_\nu^2(1 - \cos \theta)}$$

\rightarrow very small momentum transfer

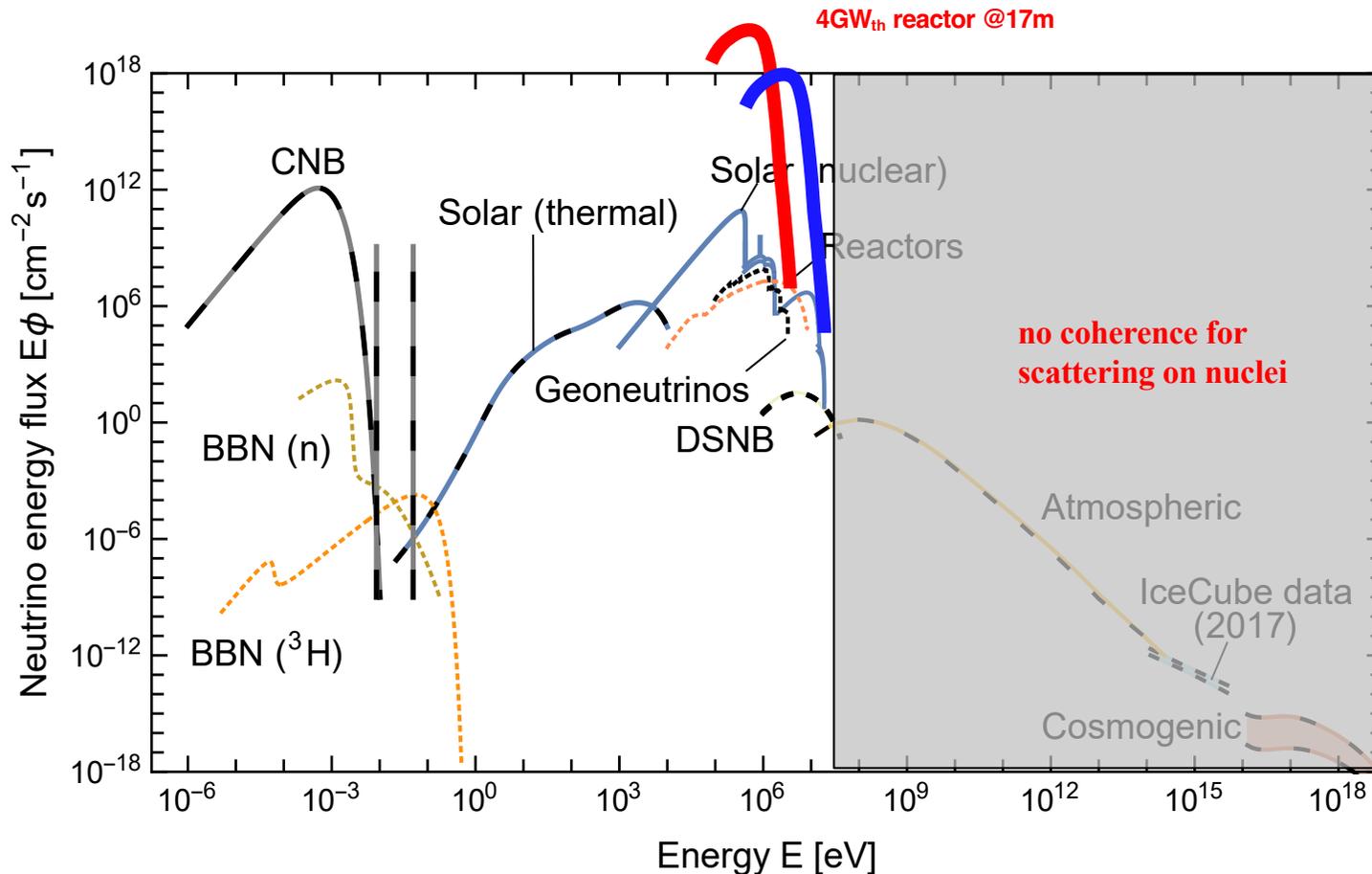
Neutrino Sources: Flux \otimes Energy



Vitagliano, Tamborra, Raffelt
Rev.Mod.Phys. 92 (2020) 45006
arXiv:1910.11878

Neutrino Sources: Flux \otimes Energy

→ very different close to a nuclear power reactor and in a stopped π -beam or a supernova

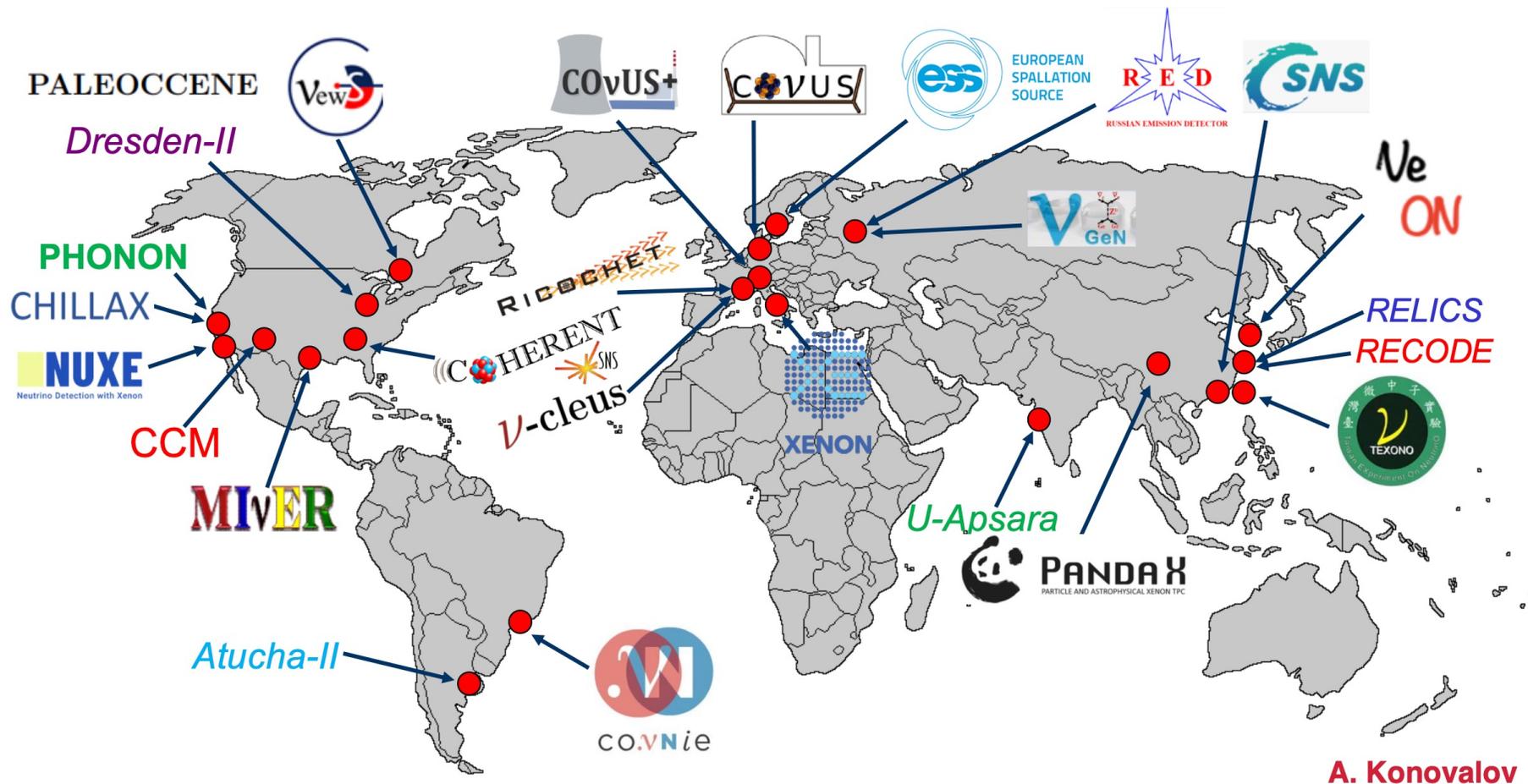


Vitagliano, Tamborra, Raffelt
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→ event rates: flux \otimes detector size \otimes time \leftrightarrow backgrounds

→ sources: beams, reactors, Sun, supernovae

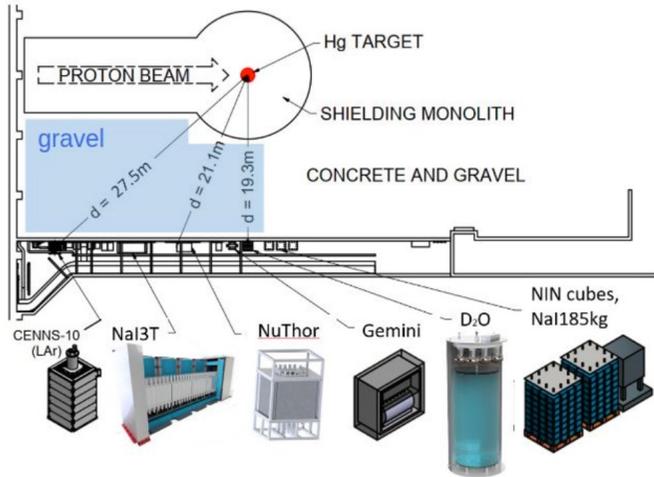
CEvNS Projects in the World



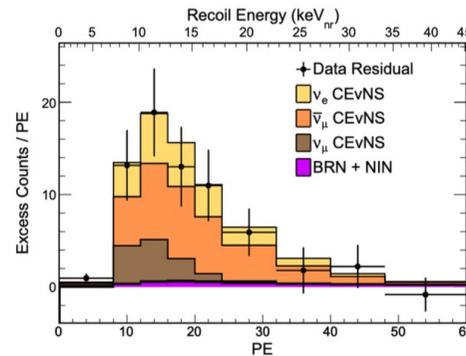
- 1974 – D. Freedman: coherent scattering of neutrinos on the whole nucleus (CEvNS)
- 2017 – observation by COHERENT in stopped π -beams \leftrightarrow $F(q^2)$ - not fully coherent
- 2024 – solar ^8B neutrinos by XENONnT and PandaX
- 2025 – CONUS+ with reactor neutrinos

COHERENT

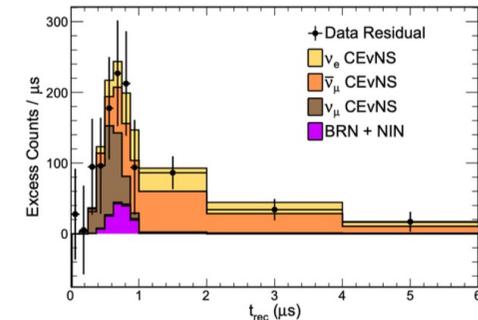
Spallation Neutron Spource (SNS) @Oak Ridge National Laboratory, USA



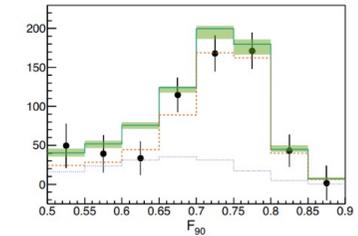
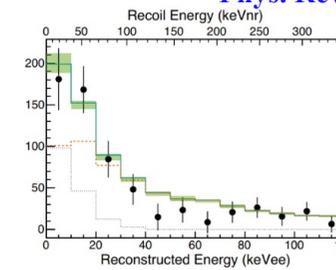
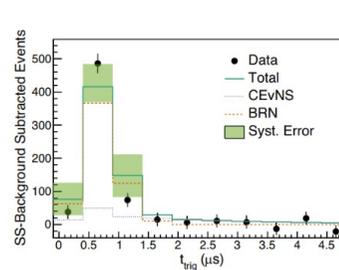
CsI



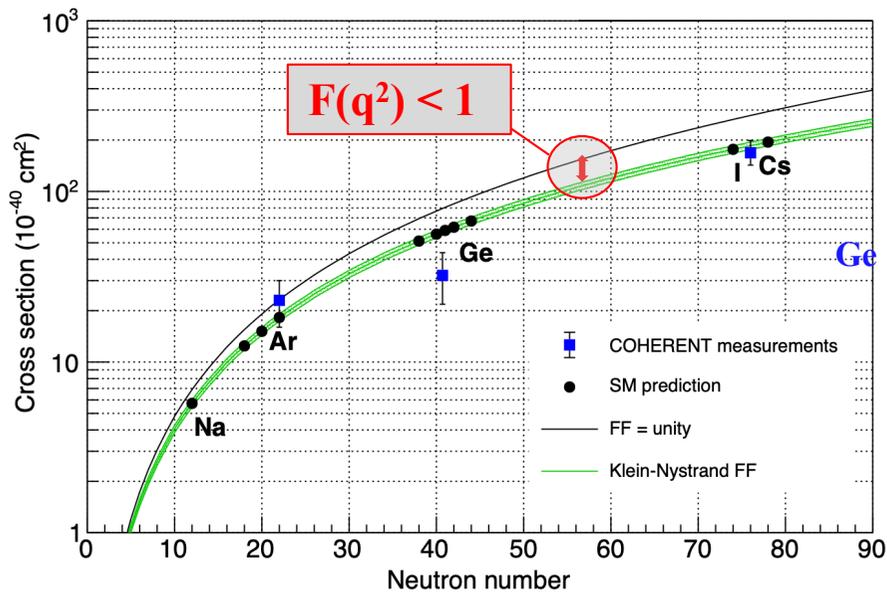
Phys. Rev. Lett. 129, 081801, (2022)



Phys. Rev. Lett. 126, 012002 (2021)

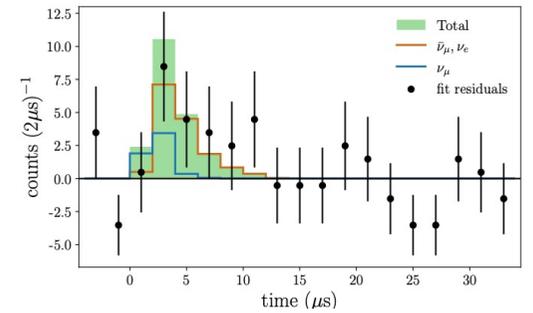
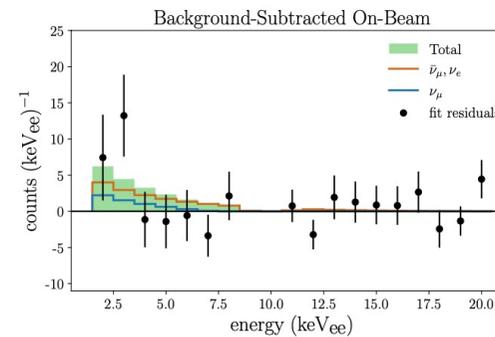


Ar



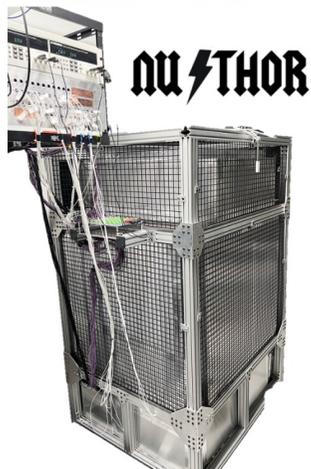
Ge

Phys. Rev. Lett. 134, 231801 (2025)



Future COHERENT

- continue data taking → improvements to come
- on-going upgrades

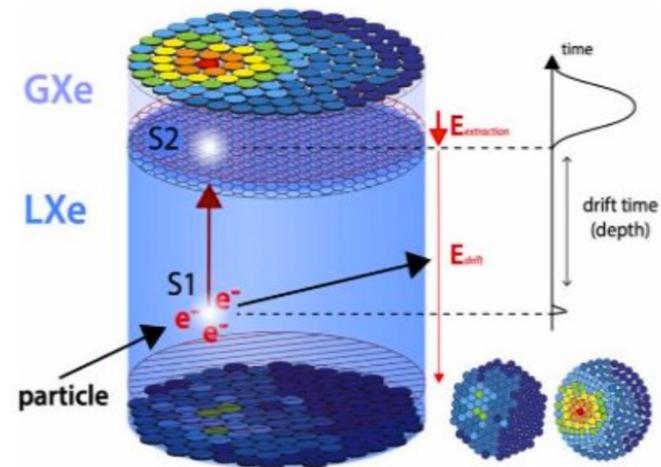
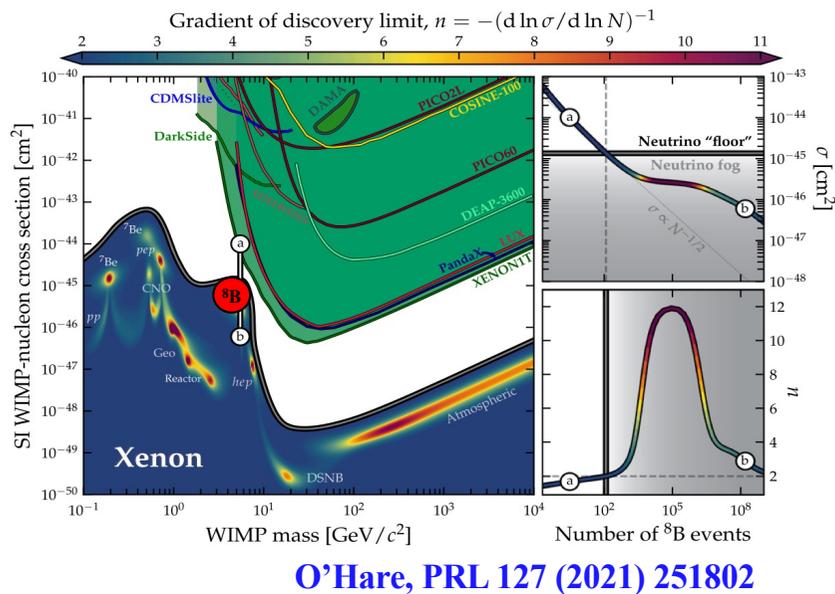


- SNS Proton Power Upgrade complete
- Ge acquiring 6x statistics in 2025 as PRL release, lower thresholds and backgrounds
- D₂O addressing leading syst. uncertainty
- 750-kg LAr target commissioning at ORNL now; will see ~5000 events/year
- 2.4 tn NalvETe (NaI) taking data in 2025
- Cryo-CsI R&D in progress
- NuThor observed ν -induced fission
- Pb glass 22 tn detector for inelastic in the works

→ towards precision CE ν NS

XENONnT and PandaX

- solar neutrinos (^8B)
 - atmospheric neutrinos
- “neutrino fog” in DM experiments



XENONnT - 2.7σ
PRL 133, 191002 (2024)

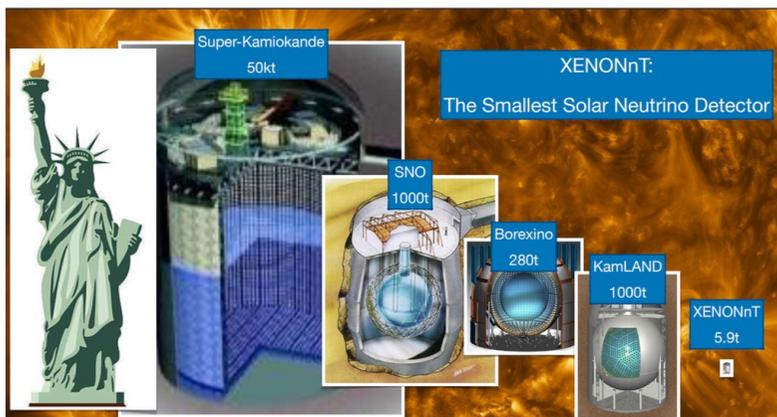
PandaX-4T - 2.6σ
PRL 133, 191001 (2024)

lowered S1,S2 thresholds
NR signals as low as 0.5 keV
exposure $3.51 \text{ t} \times \text{yr}$
best-fit: $10.7+3.7-4.2$ ^8B events

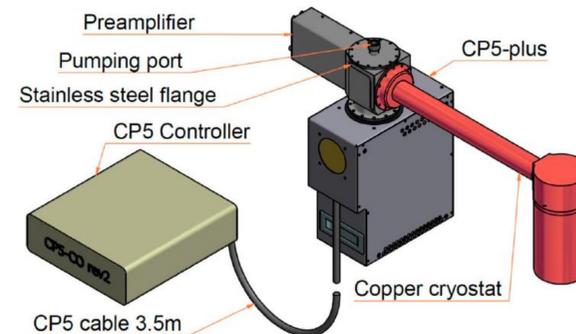
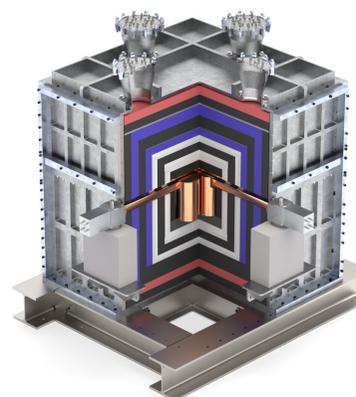
2 channels (paired S1-S2, US2)
thresholds of 1.1 and 0.33 keV
1.20 and 1.04 $\text{t} \times \text{yr}$
 3.5 ± 1.3 (paired) ; 75 ± 28 (US2)

future: +LZ, more statistics

long term: XLZD, PandaxT → precision CEvNS



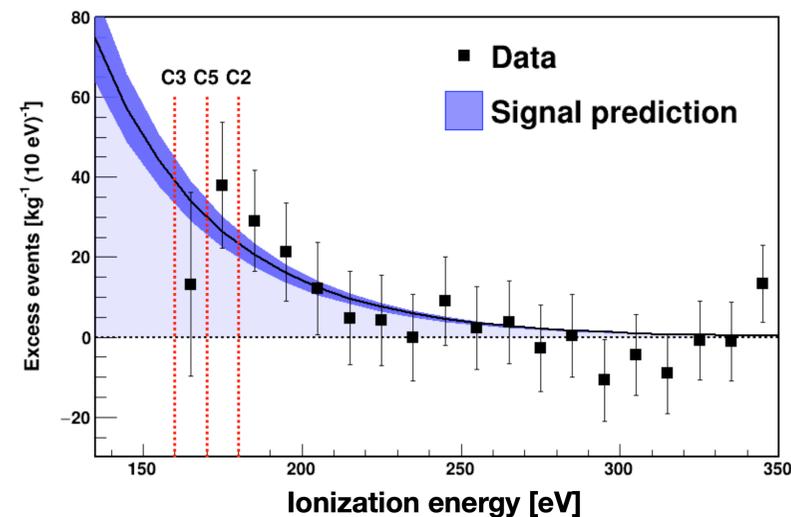
CONUS+



Leibstadt nuclear power plant, CH (KKL)

3.6 GW_{th}, 20.7m → flux: $1.45 \cdot 10^{13}$ v/cm²/s

- sophisticated shield → “virtual depth”
- improved veto systems
- extensive bckg measurements (n, μ’s, materials,...)
- 4 × 1kg low bckg point-contact HPGe detectors
- pulse shape discrimination (slow pulses)
- ASIC electronics (improved low E trigger efficiency)
- reduced point contact size & bonding technology
- lower threshold: $\sim 250 \rightarrow 160$ eV
- very sophisticated & well tested bckg model
- $F(Q^2) = 1 \rightarrow$ deviations \rightarrow BSM



Nature 643, 1229 (2025)

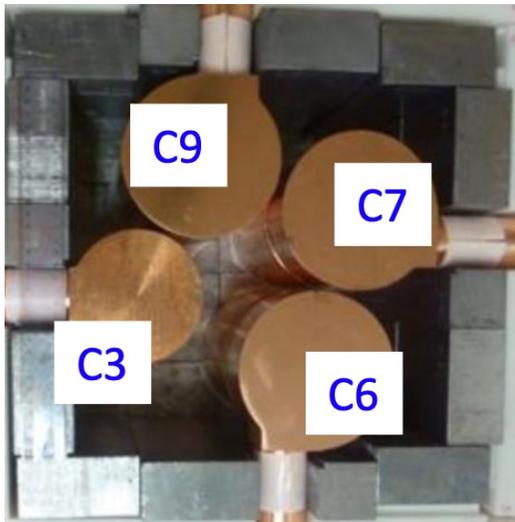
395 ± 106 CEvNS events $\rightarrow 3.7\sigma$
1st detection of CEvNS with reactor ν’s!

Comparison of CEνNS Results

Experiment	Target	Source	Neutrino energy	Flux [cm ⁻² s ⁻¹]	Data [counts]	Data/ SM prediction	Significance of null hypothesis rejection
COHERENT	Cs	Accelerator	10–50 MeV	$5 * 10^7$	306^{+20}_{-20}	$0.90^{+0.14}_{-0.14}$	11.6σ
COHERENT	Ar	Accelerator	10–50 MeV	$5 * 10^7$	140^{+40}_{-40}	$1.22^{+0.49}_{-0.49}$	3.5σ
COHERENT	Ge	Accelerator	10–50 MeV	$5 * 10^7$	21^{+7}_{-6}	$0.59^{+0.26}_{-0.24}$	3.9σ
XENONnT	Xe	Sun (⁸ B)	< 15 MeV	$5 * 10^6$	11^{+4}_{-2}	$0.90^{+0.65}_{-0.67}$	2.73σ
PandaX-4T	Xe	Sun (⁸ B)	< 15 MeV	$5 * 10^6$	4^{+1}_{-1}	$1.25^{+0.69}_{-0.69}$	2.64σ
CONUS+	Ge	Reactor	< 10 MeV	$1.5 * 10^{13}$	395^{+106}_{-106}	$1.14^{+0.36}_{-0.36}$	3.7σ

→ all will collect more statistics ; longer term → expect precision

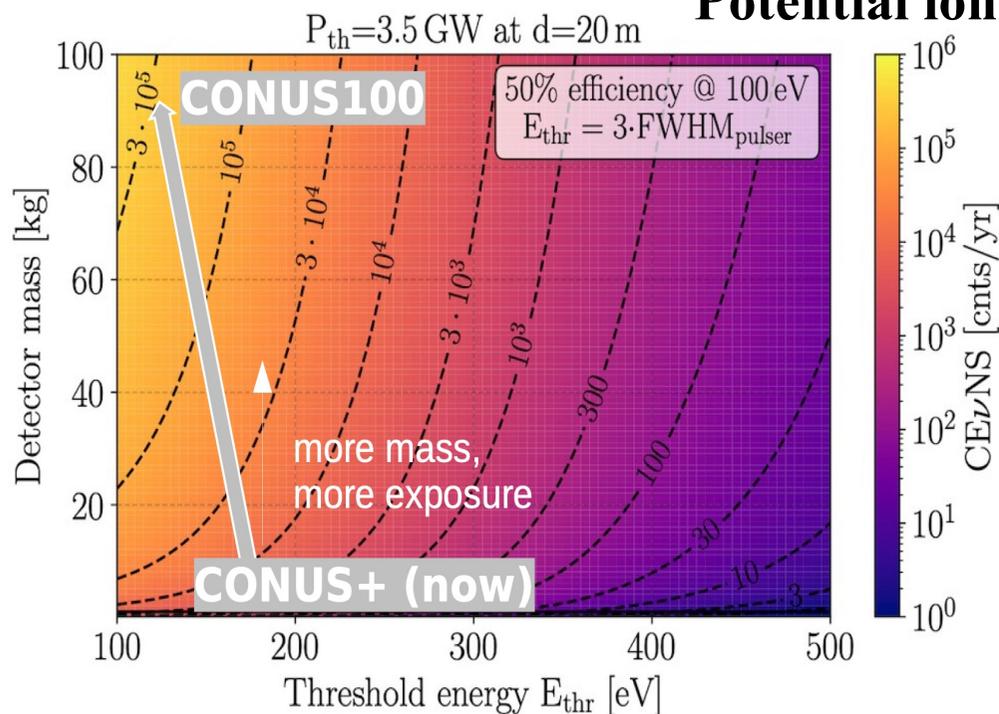
Future of CONUS+



- new 2.4 kg PPC Ge detectors $3 \text{ kg} \rightarrow 1+3*2.4 = 8.2 \text{ kg}$
- various improvements (background, threshold, ...)
- installed in 11/2024
- taking data \rightarrow improved results $O(1\text{yr})$

\rightarrow improved CEvNS results on shorter time scales

Potential long term potential of established technology



ML, T. Rink, M. Sen, JHEP 08 (2024) 171

- scaling to $O(100\text{kg})$ Germanium
- even lower threshold seems feasible
- 5 years of operation
- \rightarrow 500.000 events in 5y conceivable

\rightarrow precision CEvNS physics

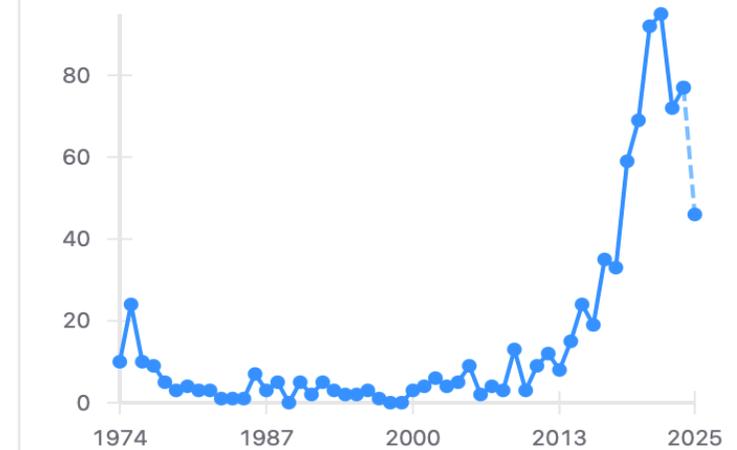
Many interesting CEvNS Topics

- clean SM cross sections
- $\sin^2\theta_W$ at low energies
- Sterile neutrino searches
- ν magnetic moment, $\langle r_\nu^2 \rangle$
- ν millicharges
- NSI operators
- Tests of lepton universality
- Connections to models of dark matter
- Measurements of neutron formfactors (nuclear structure) → **unique**
- Nuclear reactor monitoring (non-proliferation) → **applications**
- Flux & spectral measurements → **synergy w. oscillation experiments**
- CEvNS & energy transport in supernovae → **important for next SN**
- SN neutrino detection → **SNEWS, pointing, ...**
- Connection to dark matter direct detection (ν floor) → **solar ν 's**

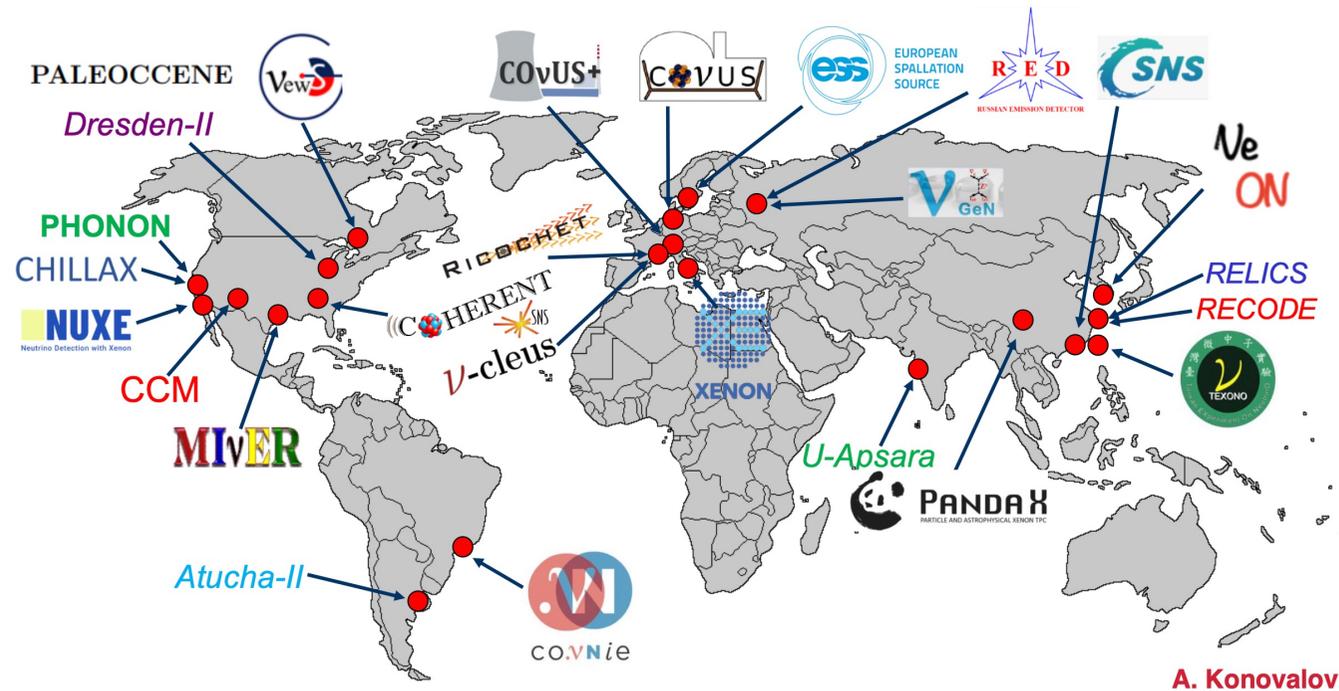
unique
BSM
sensitivity

D.Z. Freedman, PRD9, 1389 (1974)

Citations per year



Other CEνNS Experiments



various very promising other technologies & experiments

→ very interesting, healthy & motivating competition

→ see parallel sessions for status & details:

→ moving to precision CEνNS physics

very interesting potential to test / discover BSM physics

Conclusions

Sterile neutrinos: many reasons - very well motivated: simple see-saw + many variants

- so far limits assuming oscillations ($e\bar{e}$, $\mu\bar{\mu}$, μe)
- anomalies which do not fit into an oscillation picture
 - might hint some other new physics or might go away
 - very important to clarify
- constraints from cosmology and astrophysics

CEvNS: finally observed O(50) years after invention

- COHERENT (stopped π beam)

- XENONnT and PandaX-4T (solar ${}^8\text{B}$ ν 's)

- CONUS+ (reactor neutrinos)

} improved results to come
+ other experiments coming online

- so far consistent with SM expectations → already very interesting BSM limits
- all aim at upgrades which will lead to significantly increased statistics
 - moving towards an era of precision neutrino physics

ν masses: 1st BSM physics → much more than masses & mixings

More Details in Parallel Talks

sterile neutrinos

Gallium anomaly:	CIUFFOLI, Emilio BENATO, Giovanni JI, Xiangpan
MicroBooNE:	LUO, Xiao JI, Xiangpan
KATRIN:	MOHANTY, Shailaja NAVA, Andre
DANSS:	ALEKSEEV, Igor SVIRIDA, Dmitry
Daya Bay:	LI, Ruhui
SBND:	WESTER, Thomas
LiF crystals:	WOO, KyungRae
Jinping Neutrino Experiment (JNE):	ZHU, Yutao

CE ν NS

nuGeN:	LUBASHEVSKIY, Alexey
NUCLEUS:	BOSSIO, Elisabetta
CONNIE:	AGUILAR-AREVALO, Alexis
XENON:	LI, Shengchao LIU, Kexin
COHERENT:	LIU, Jing
RED-100:	KOZLOVA, Ekaterina
Ga-anomaly:	BENATO, Giovanni
RELICS:	XIE, Lingfeng
NR tracking:	CARATELLI, David
CONUS+:	LINDNER, Manfred