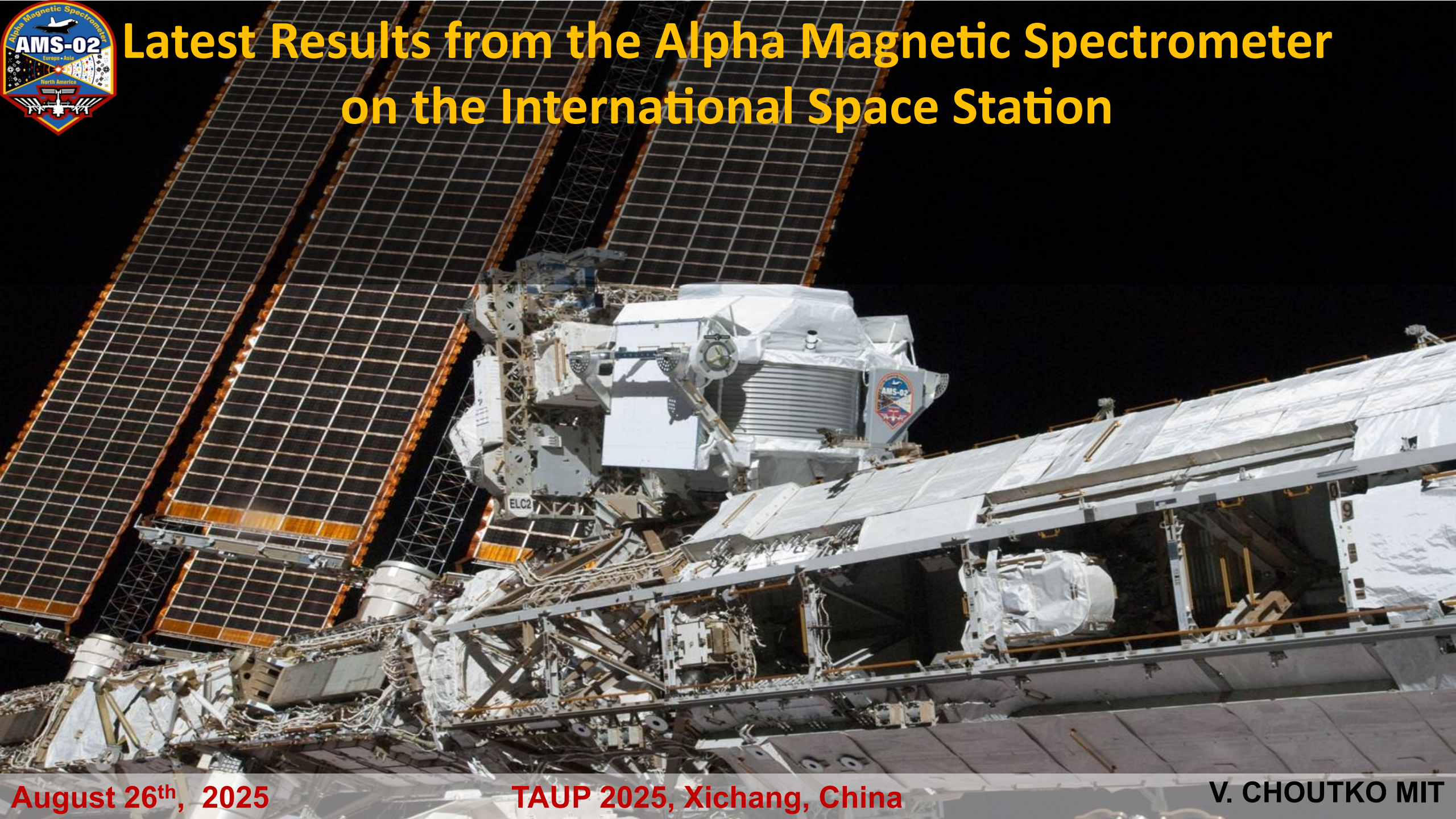




Latest Results from the Alpha Magnetic Spectrometer on the International Space Station



August 26th, 2025

TAUP 2025, Xichang, China

V. CHOUTKO MIT

AMS Launched May 2011

Space Shuttle Endeavour

Mission STS-134

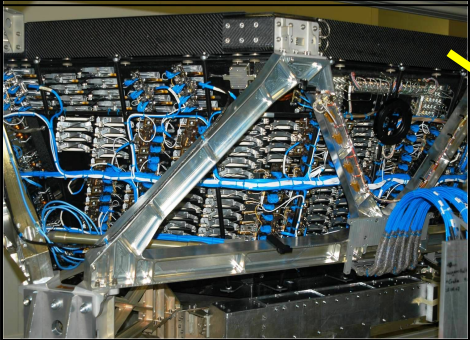
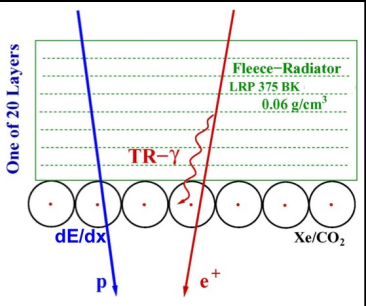


AMS installed on the ISS
Near Earth Orbit:
altitude 400 Km
inclination 52°

AMS-02: A TeV precision magnetic spectrometer in space

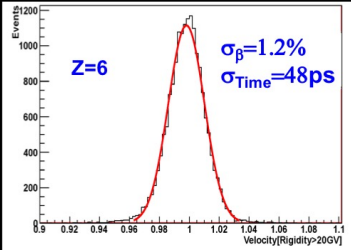
Transition Radiation Detector

Identifies e^+ , e^-



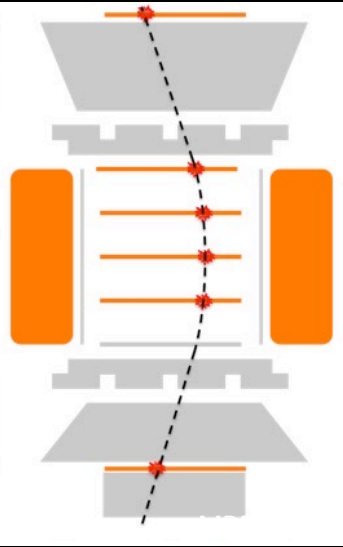
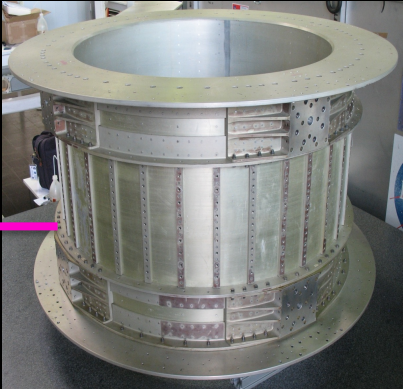
Time Of Flight

Z, β



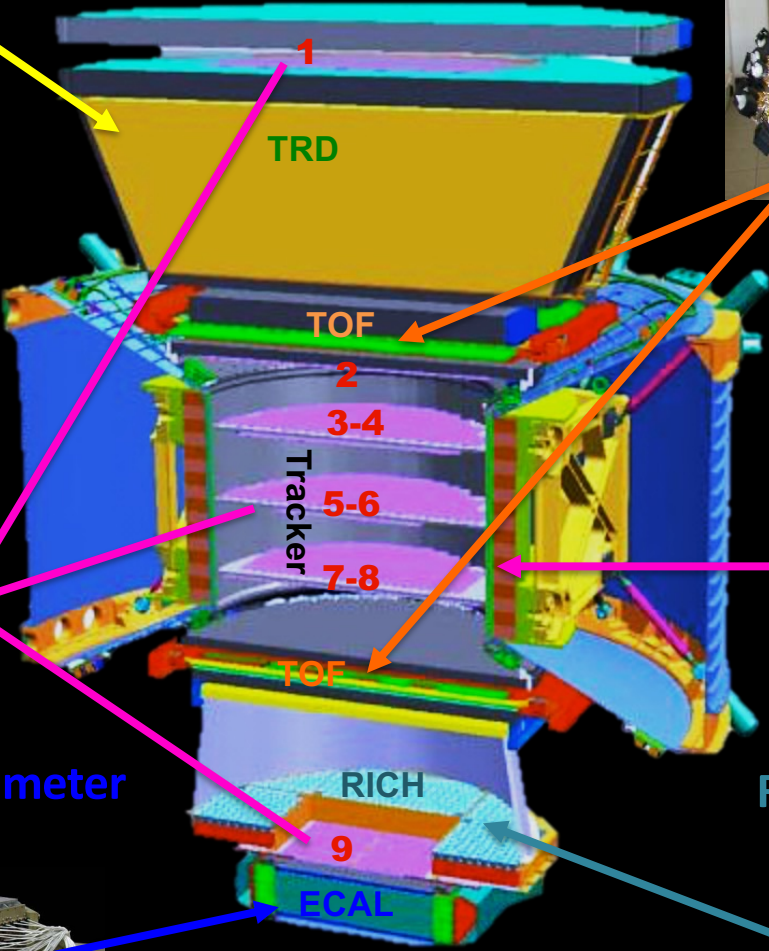
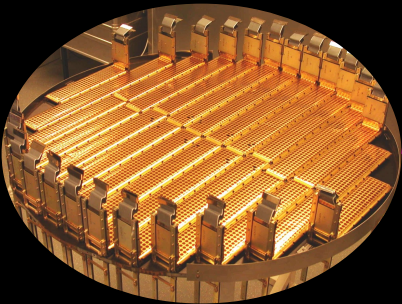
Magnet

$\pm Z$



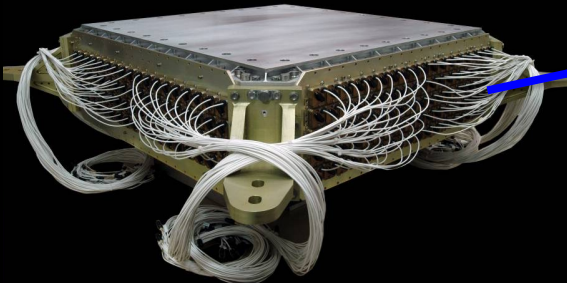
Silicon Tracker

$Z, \text{Rigidity}=p/Zc$



Electromagnetic Calorimeter

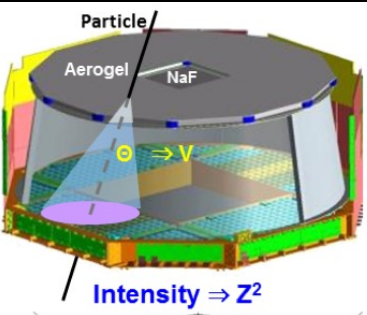
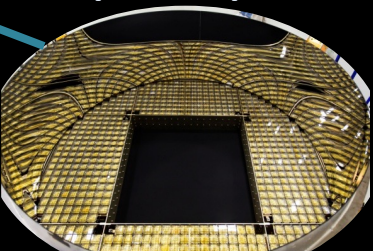
Energy of e^+ , e^-



Ring Imaging Cherenkov

Z, β

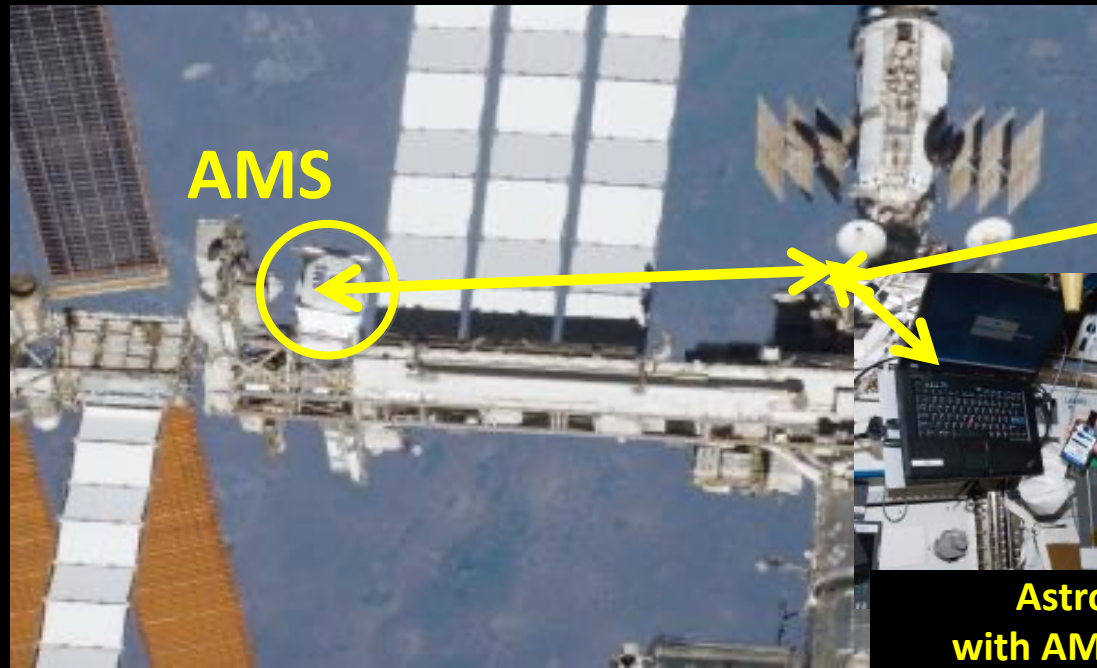
Isotopic composition



AMS Continious Data Flow ISS to POCC

To date 253 billion particles have been measured by AMS: e^+ , e^- , p , \bar{p} , nuclei, γ ,...

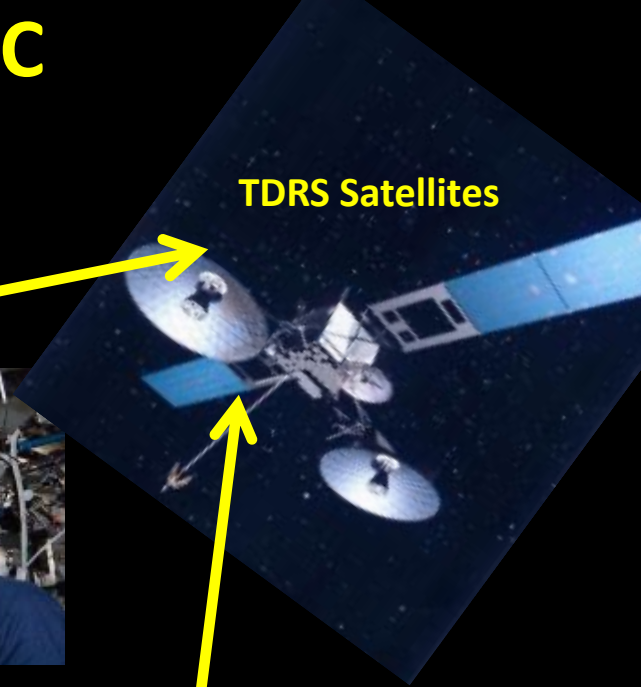
360 billion events expected to 2030



AMS



Astronaut with AMS Laptop



TDRS Satellites



Payload Operations Control Centers (POCC) at CERN, JSC, Asia



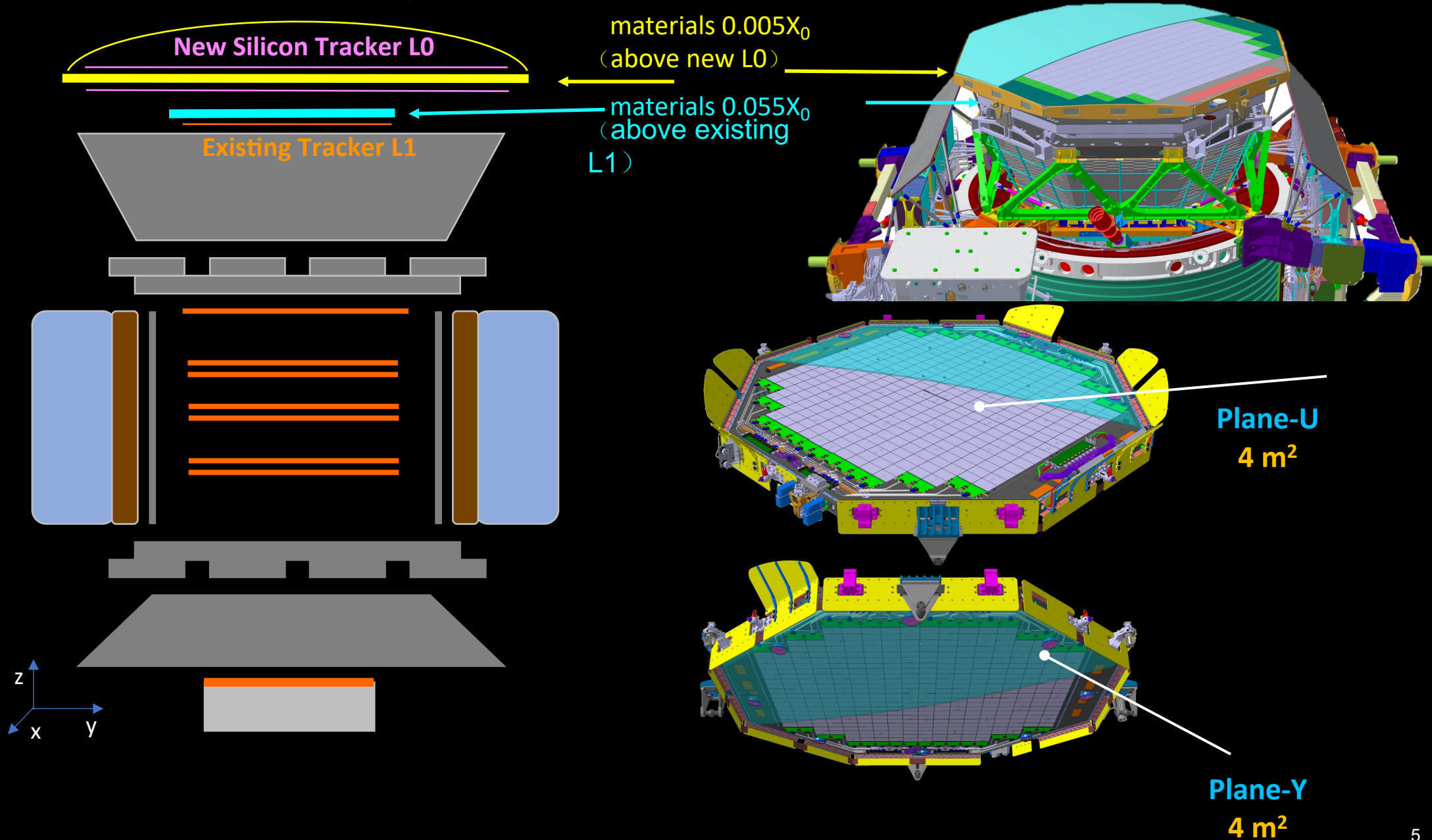
AMS Computers Marshall, Alabama



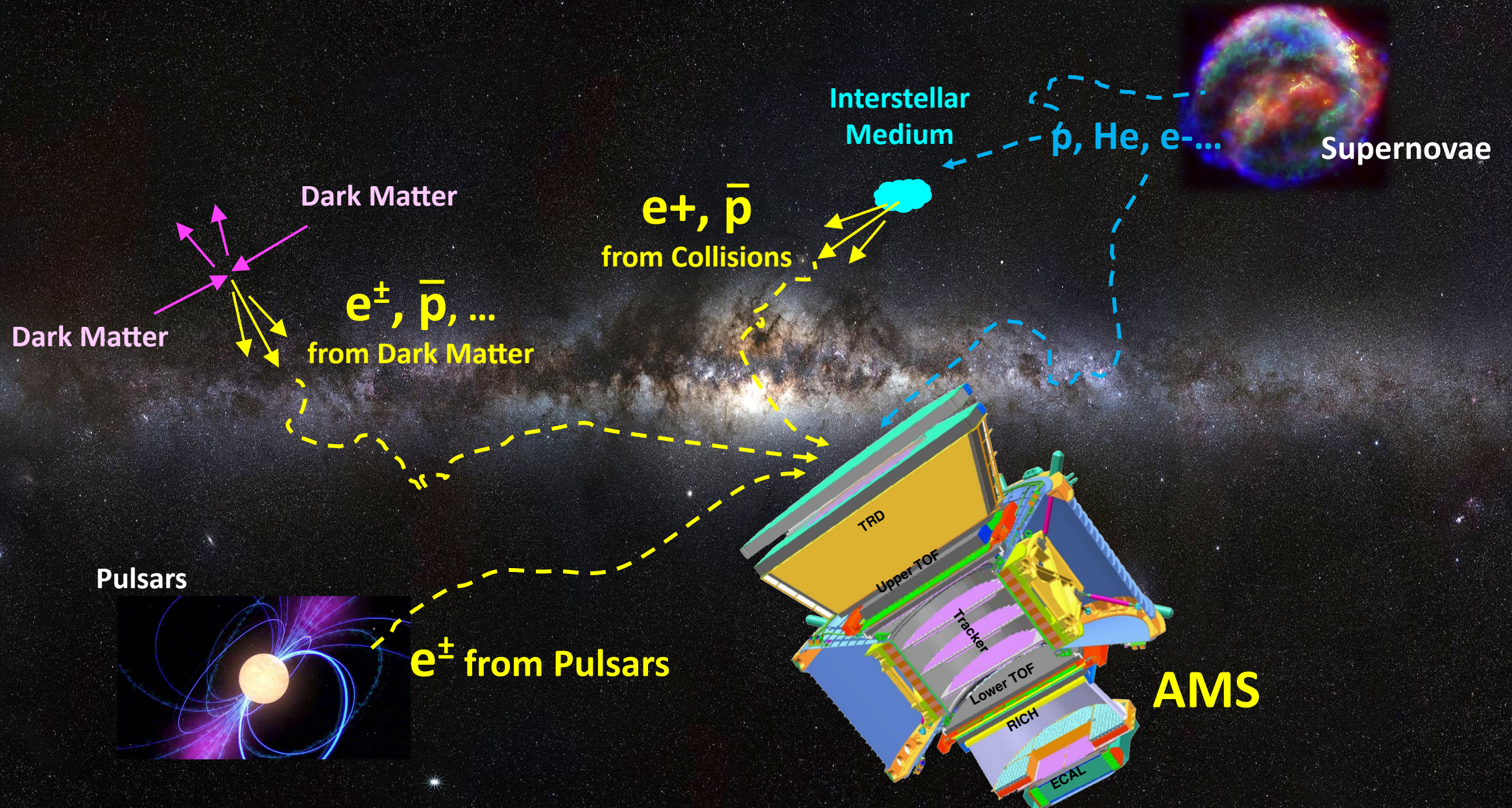
White Sands Ground Terminal, New Mexico

AMS on ISS: 2026-2030+

With Completion of L0 Upgrade AMS Acceptance will be Increased by 300% with Minimal Materials Above



Latest Results on e^+ , e^- , and \bar{p}

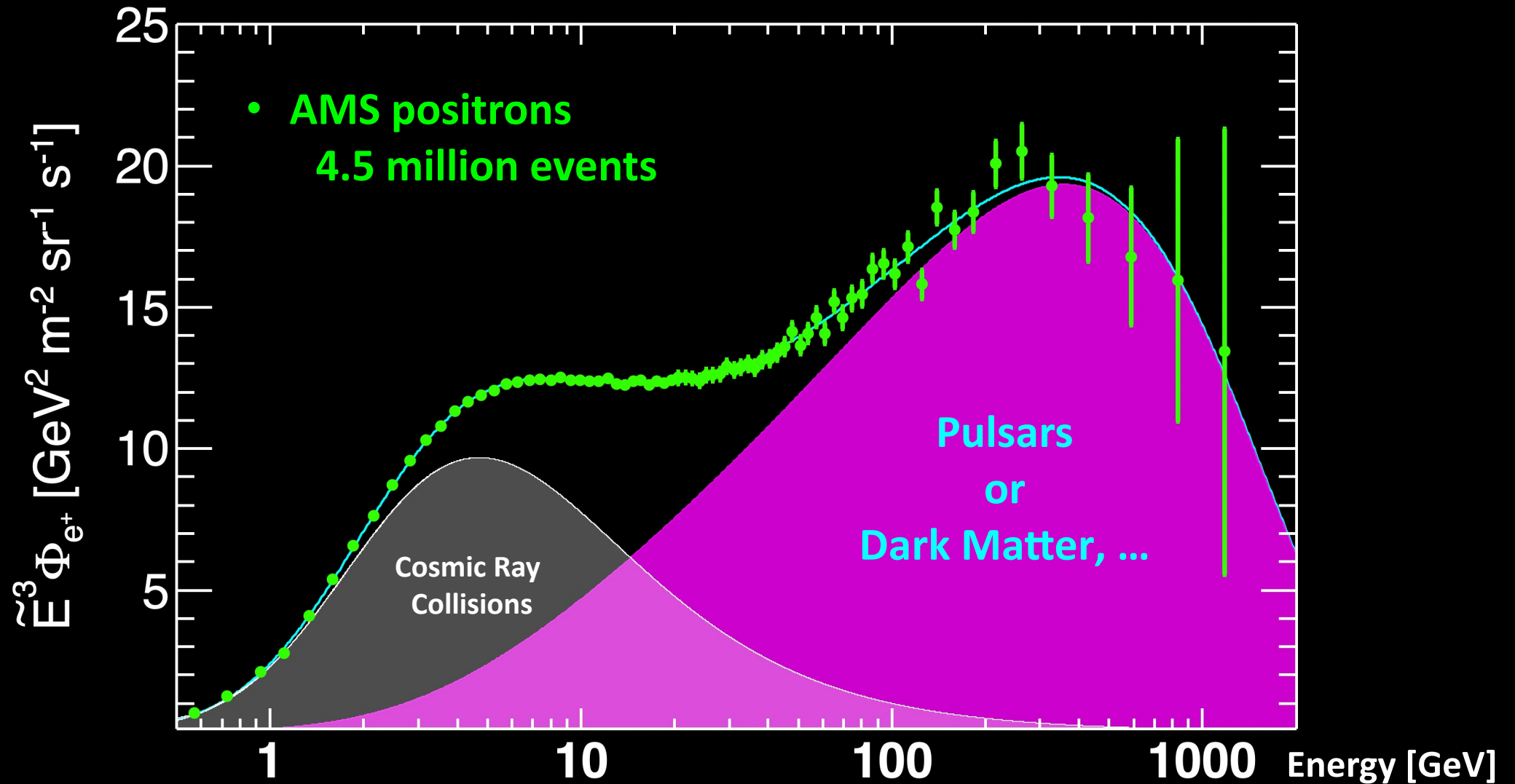


The positron flux is the sum of low-energy part from cosmic ray collisions plus a high-energy part from pulsars or dark matter with a cutoff energy

Empirical model: $\Phi_{e^+}(E) = \frac{E^2}{\hat{E}^2} \left[C_d (\hat{E}/E_1)^{\gamma_d} + C_s (\hat{E}/E_2)^{\gamma_s} \exp(-\hat{E}/E_s) \right]$

$\chi^2/\text{dof} = 40/66$

Solar Collisions Pulsars or Dark Matter

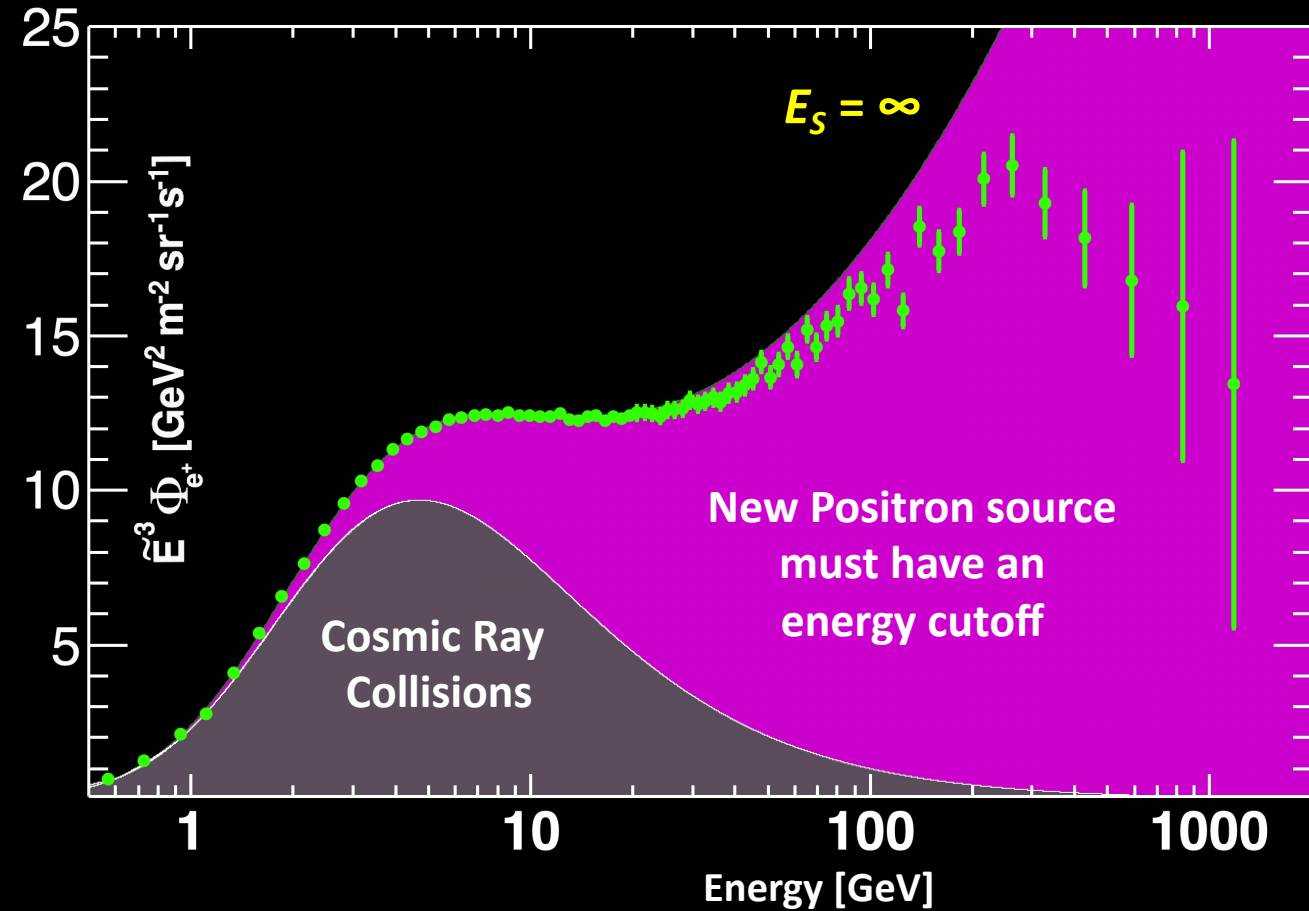
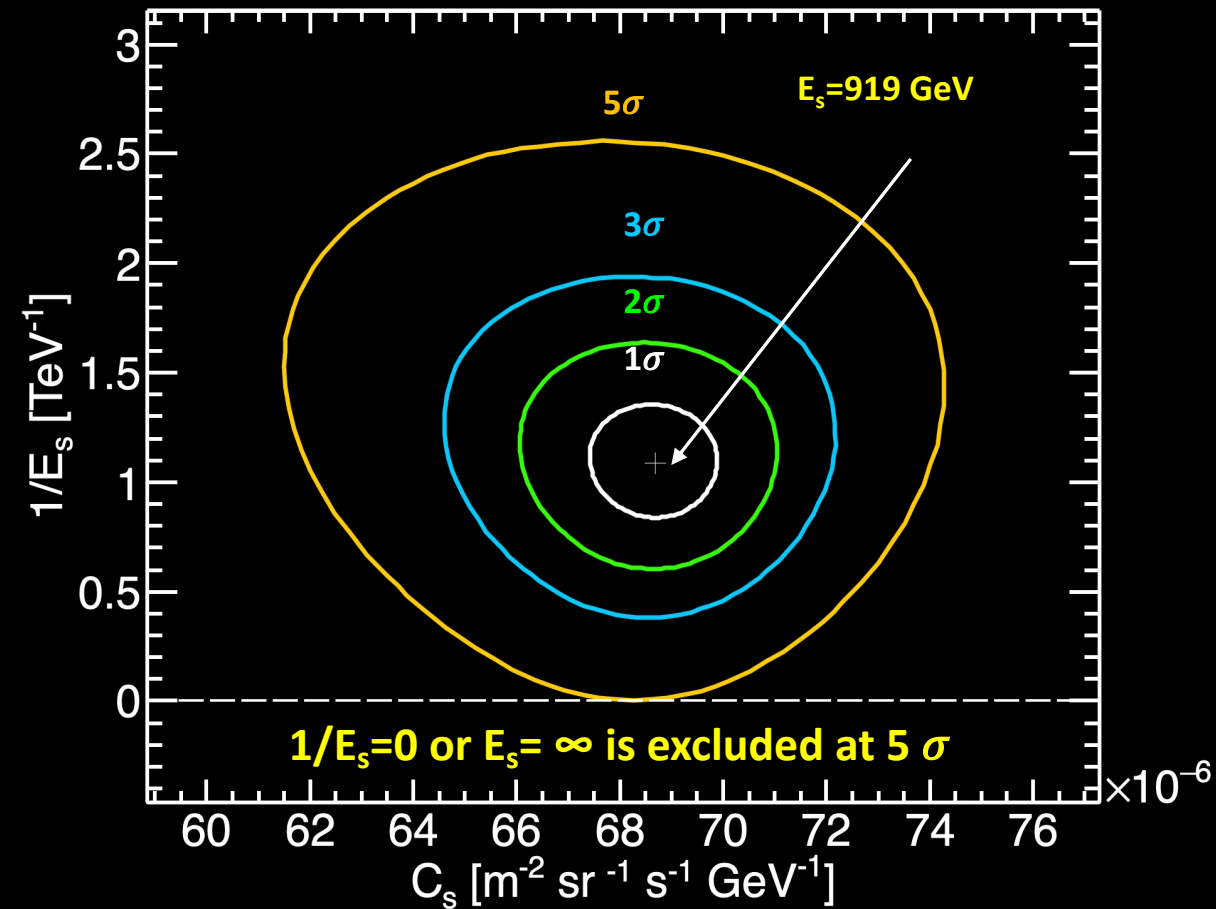


Surprising Observation: The existence of a finite cutoff energy E_s

$$\Phi_{e^+}(E) = \frac{E^2}{\hat{E}^2} \left[C_d (\hat{E}/E_1)^{\gamma_d} + C_s (\hat{E}/E_2)^{\gamma_s} \exp(-\hat{E}/E_s) \right]$$

Collisions

Pulsars or Dark Matter Collisions



$1/E_s = 0$ or $E_s = \infty$ is excluded at 5σ

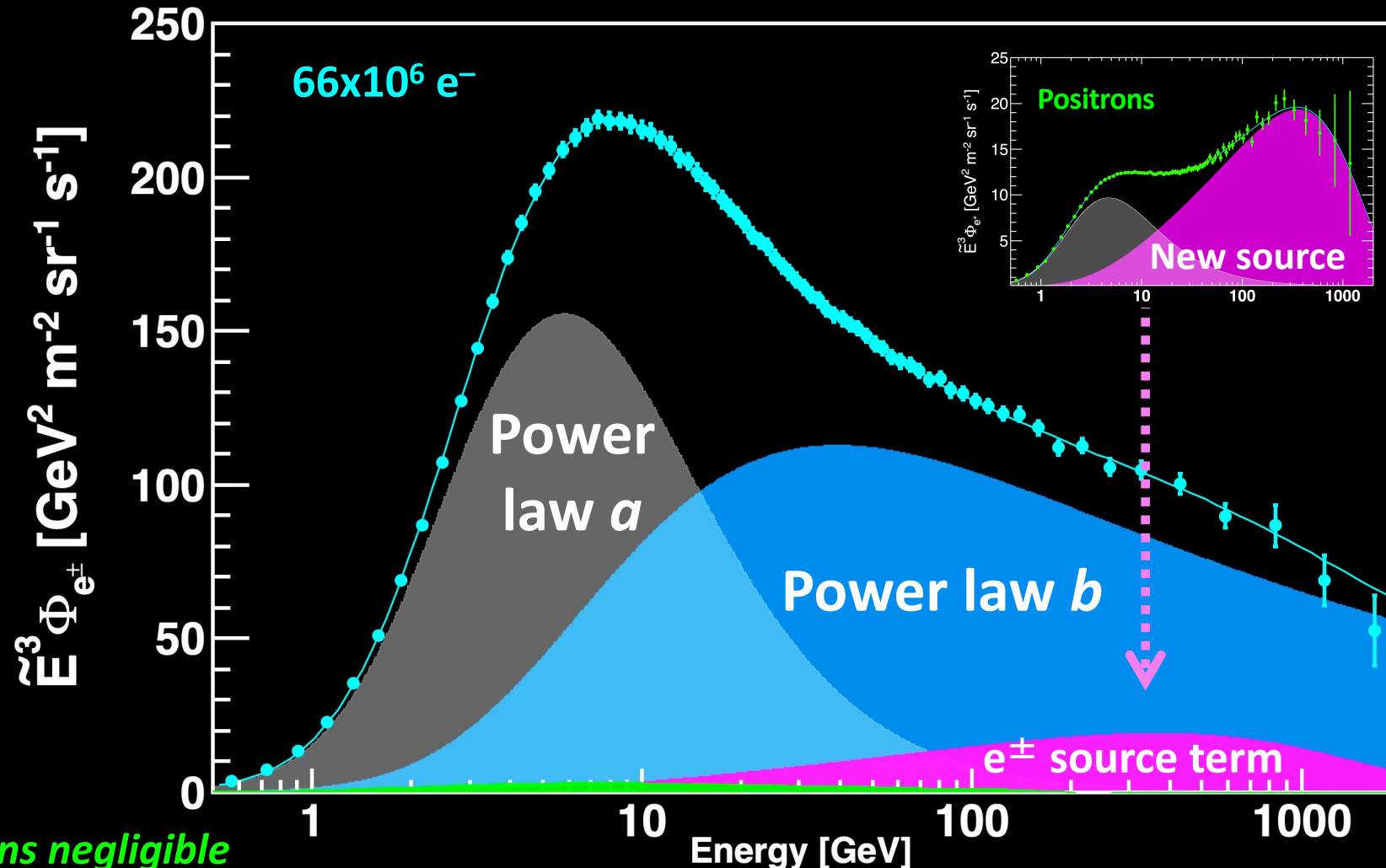
Latest Result on the electron spectrum

The spectrum fits well with two power laws (a , b) and the measured positron source term

Empirical model:
 $\chi^2/\text{dof} = 25/67$

$$\Phi_{e^-}(E) = \frac{E^2}{\hat{E}^2} (C_a \hat{E}^{\gamma_a} + C_b \hat{E}^{\gamma_b} + \text{Positron Source Term})$$

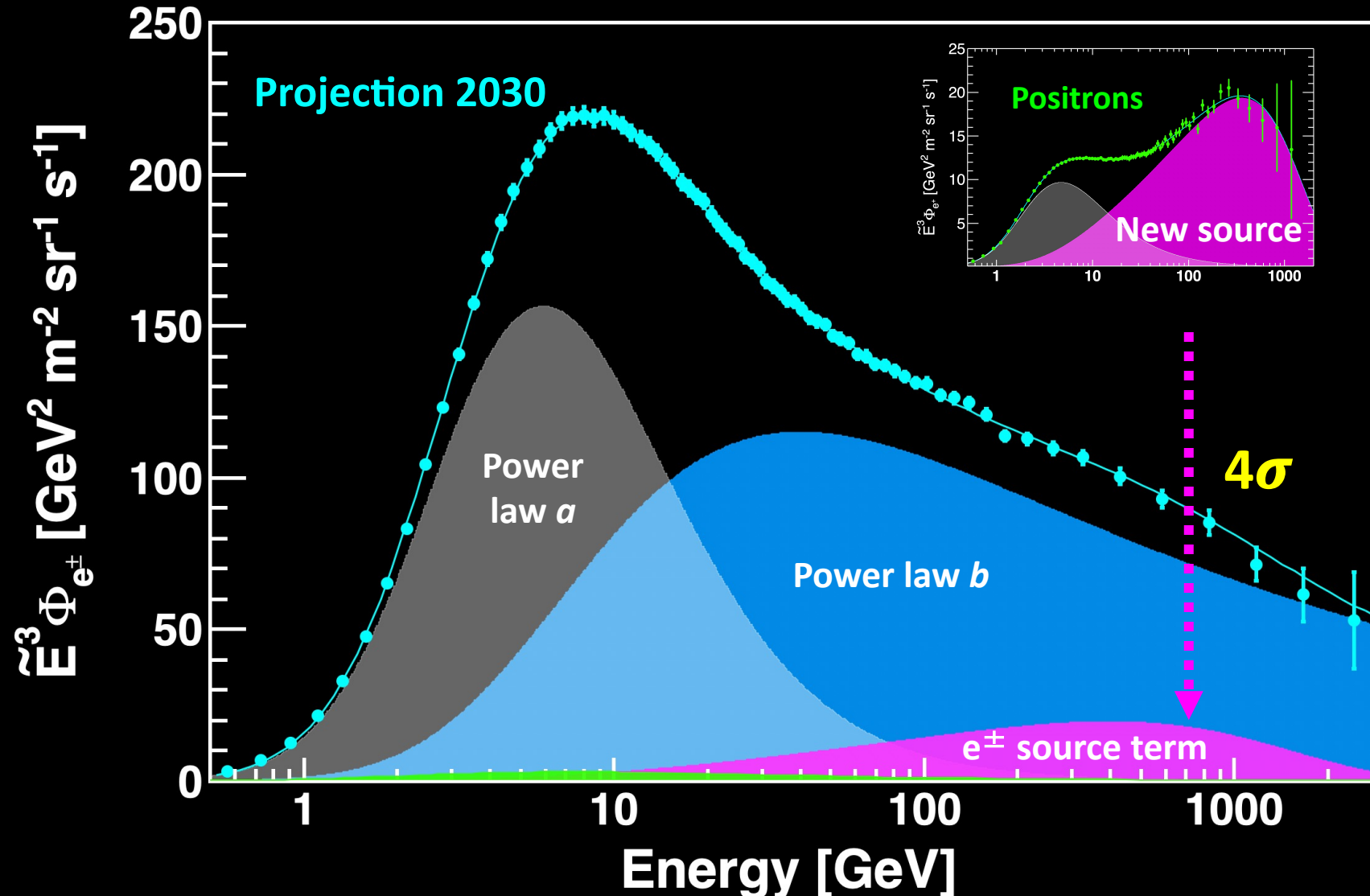
Solar Power law a Power law b



e^- from collisions negligible

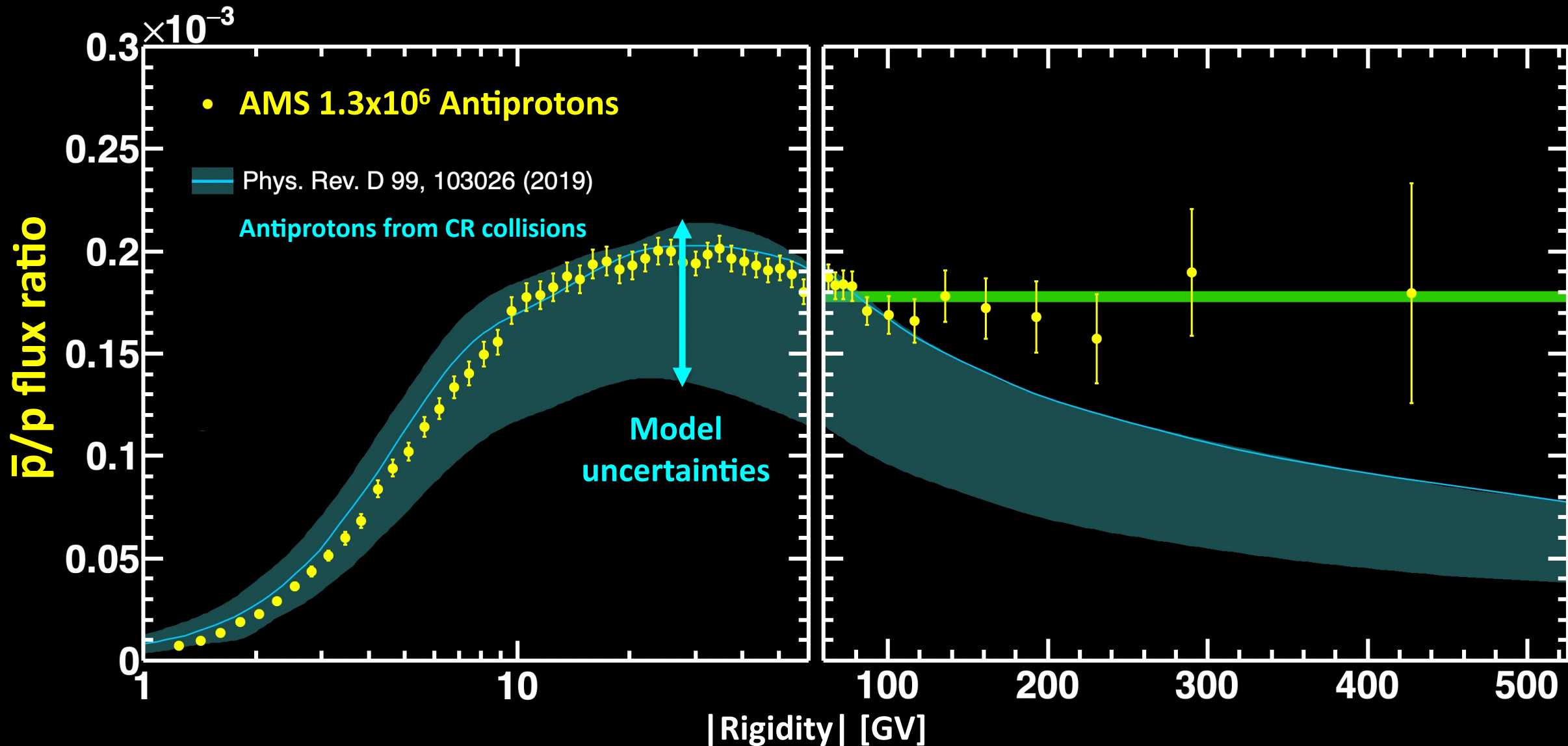
New sources, like Dark Matter or Pulsars, produce equal amounts of e^+ and e^-

By 2030, the charge-symmetric nature of the high energy source will be established at the 4σ level, due to increase of statistics and energy range



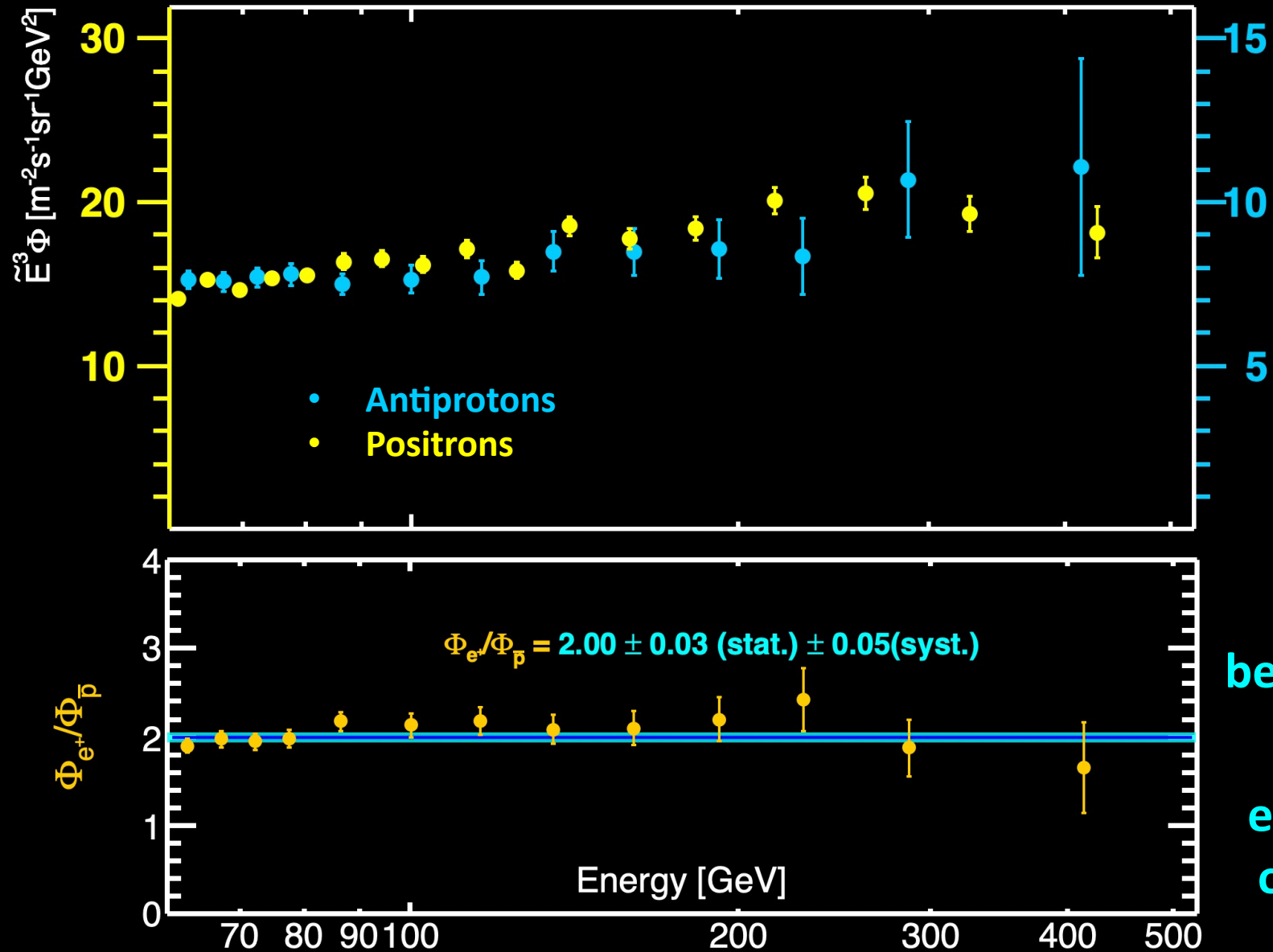
Properties of Cosmic Antiprotons

Above 60 GV, the antiproton-to-proton flux ratio is energy independent.

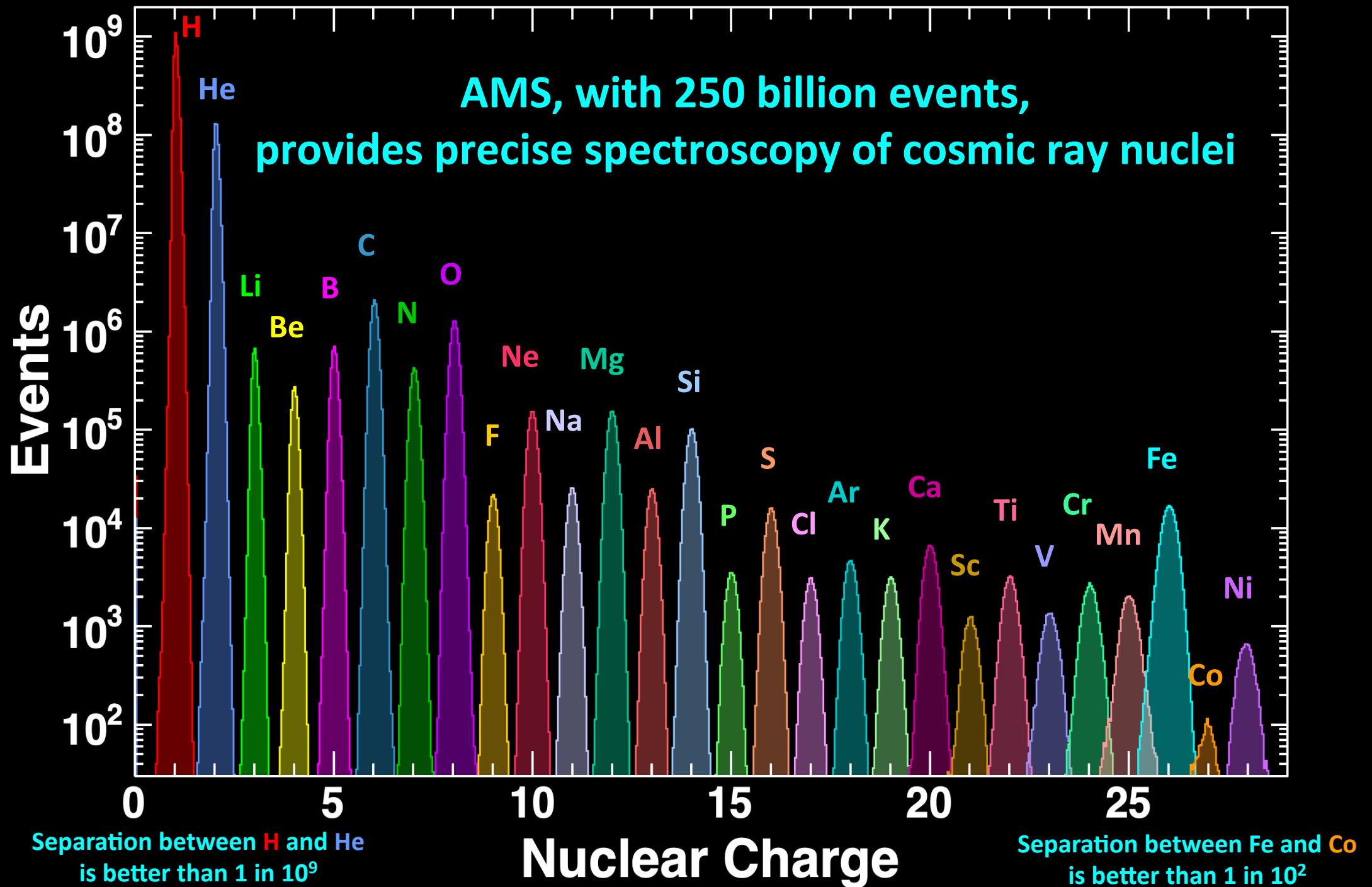


Cosmic Antiprotons and Positrons

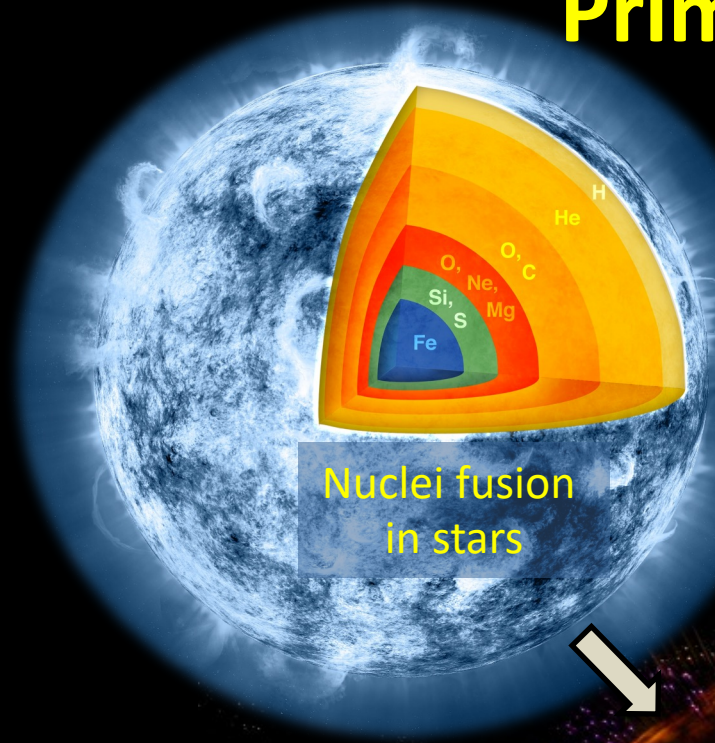
Above 60 GeV, the \bar{p} and e^+ fluxes have identical rigidity dependence



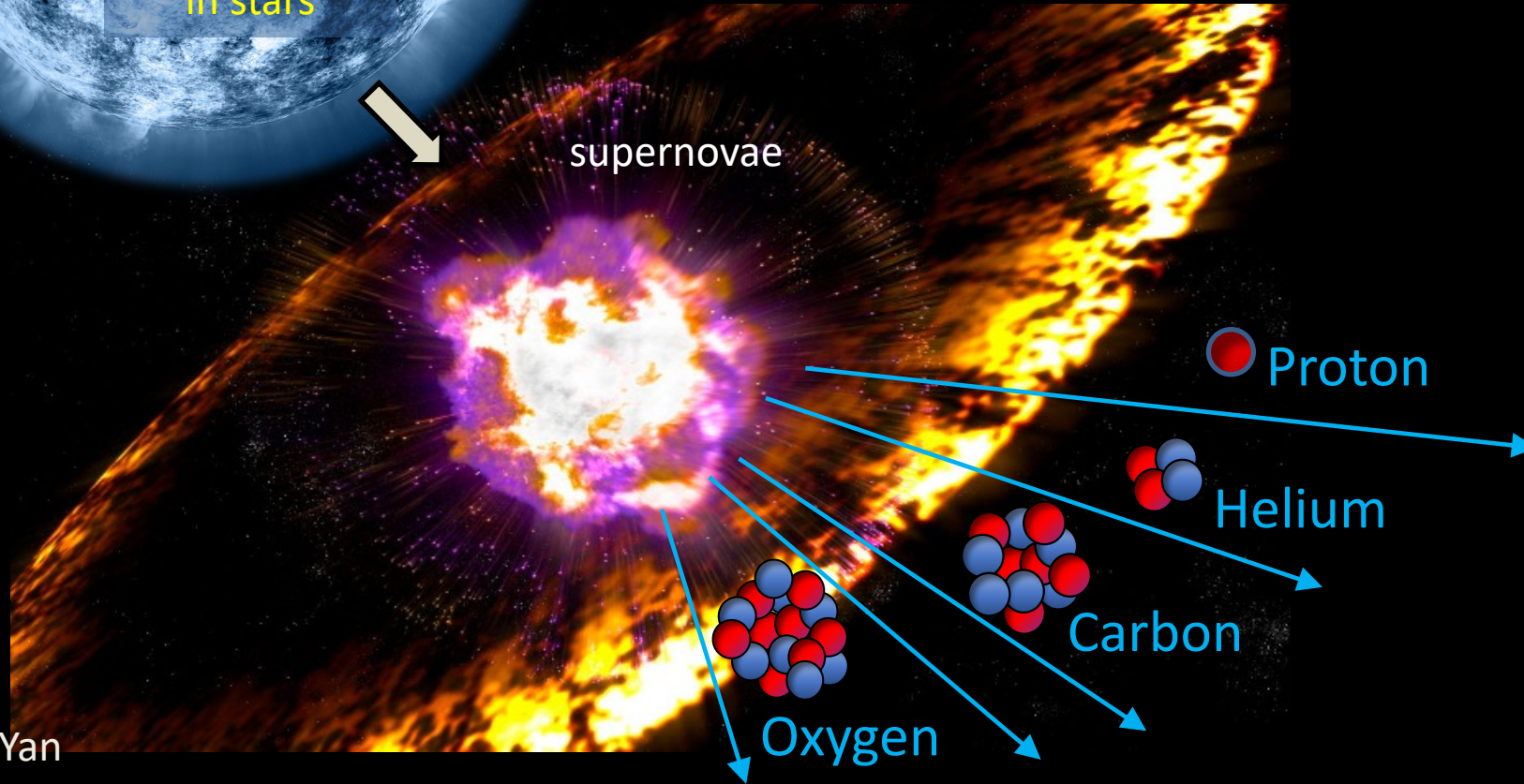
The identical behavior of positrons and antiprotons excludes the pulsar origin of positrons



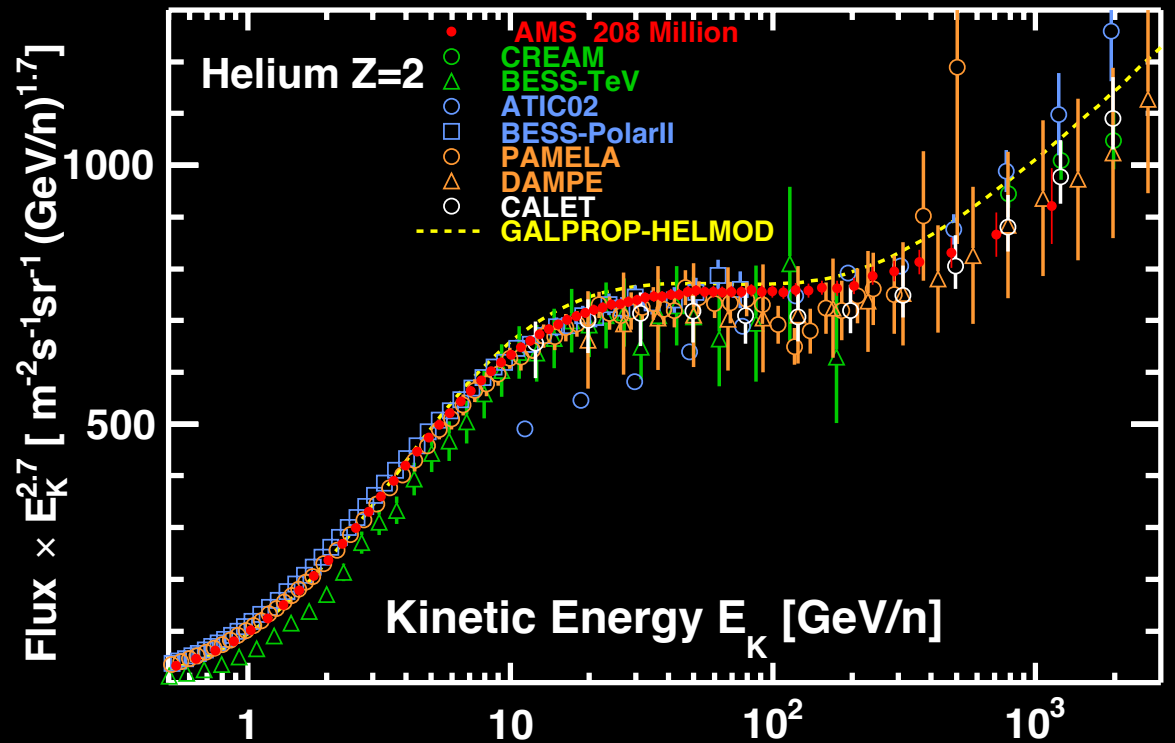
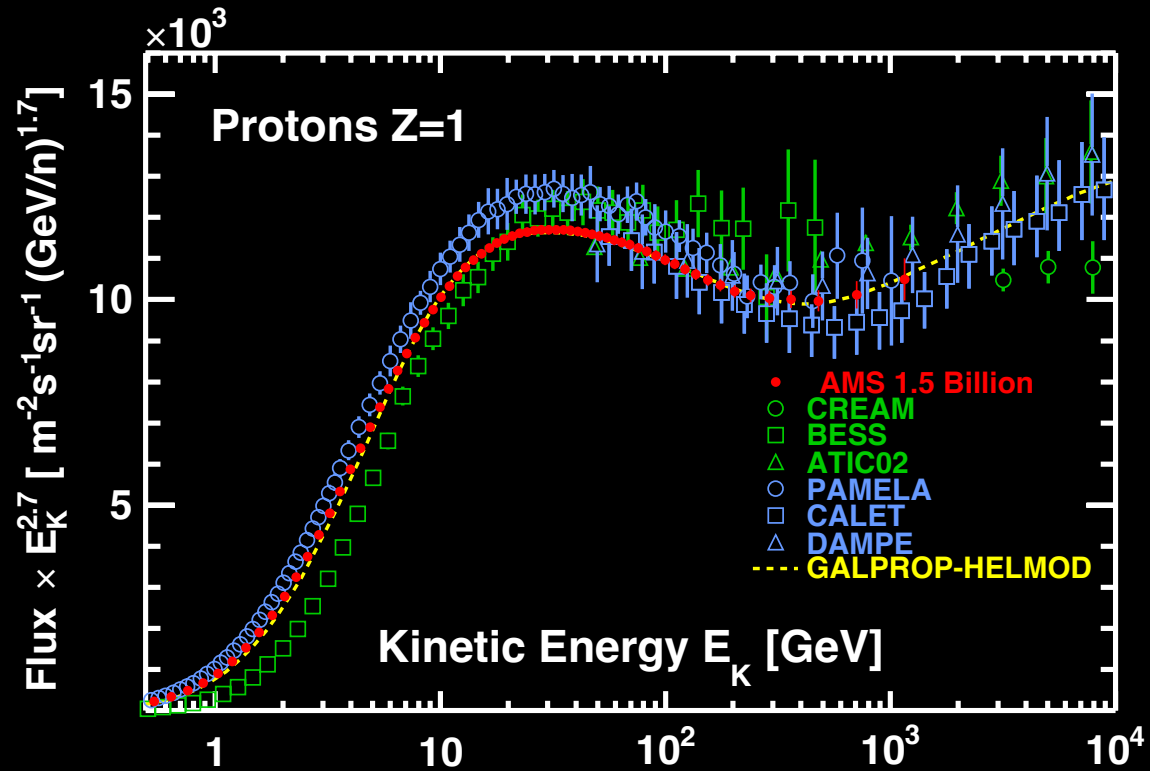
Primary Cosmic Rays



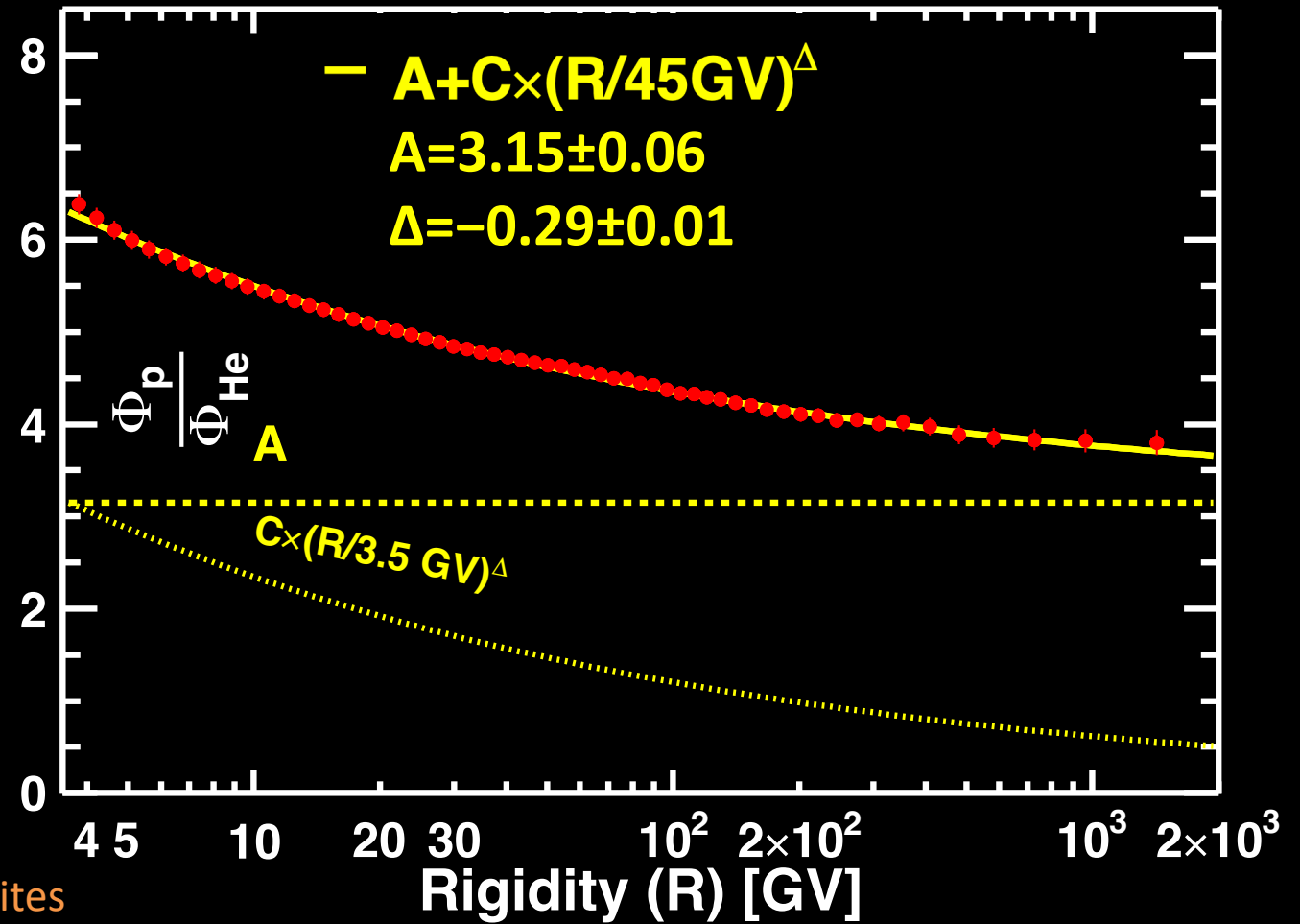
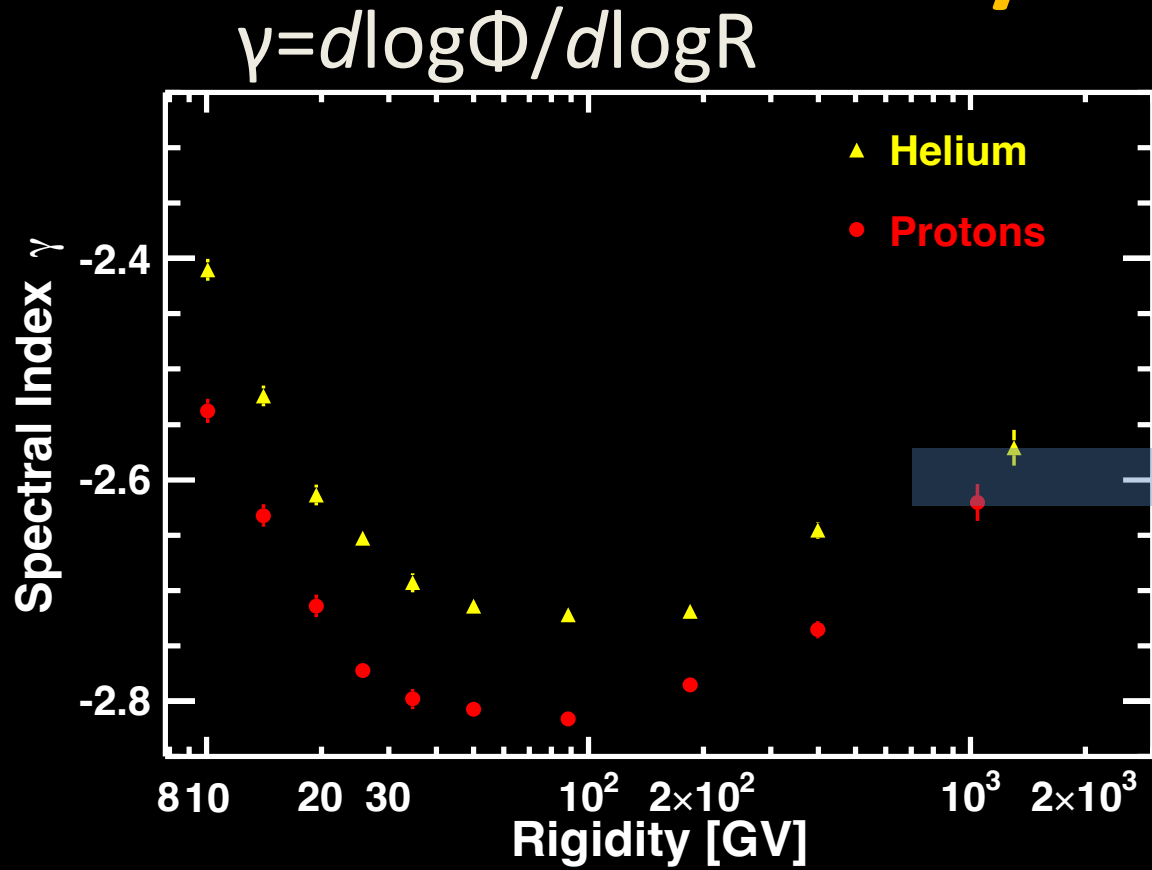
Primary cosmic rays (p, He, C, O, Ne, Mg, S ..., Fe) are mostly produced during the lifetime of stars and are accelerated in supernovae shocks, whose explosion rate is about 2-3 per century in our Galaxy.



Latest AMS proton and He flux measurement



Proton/Helium Flux Ratio

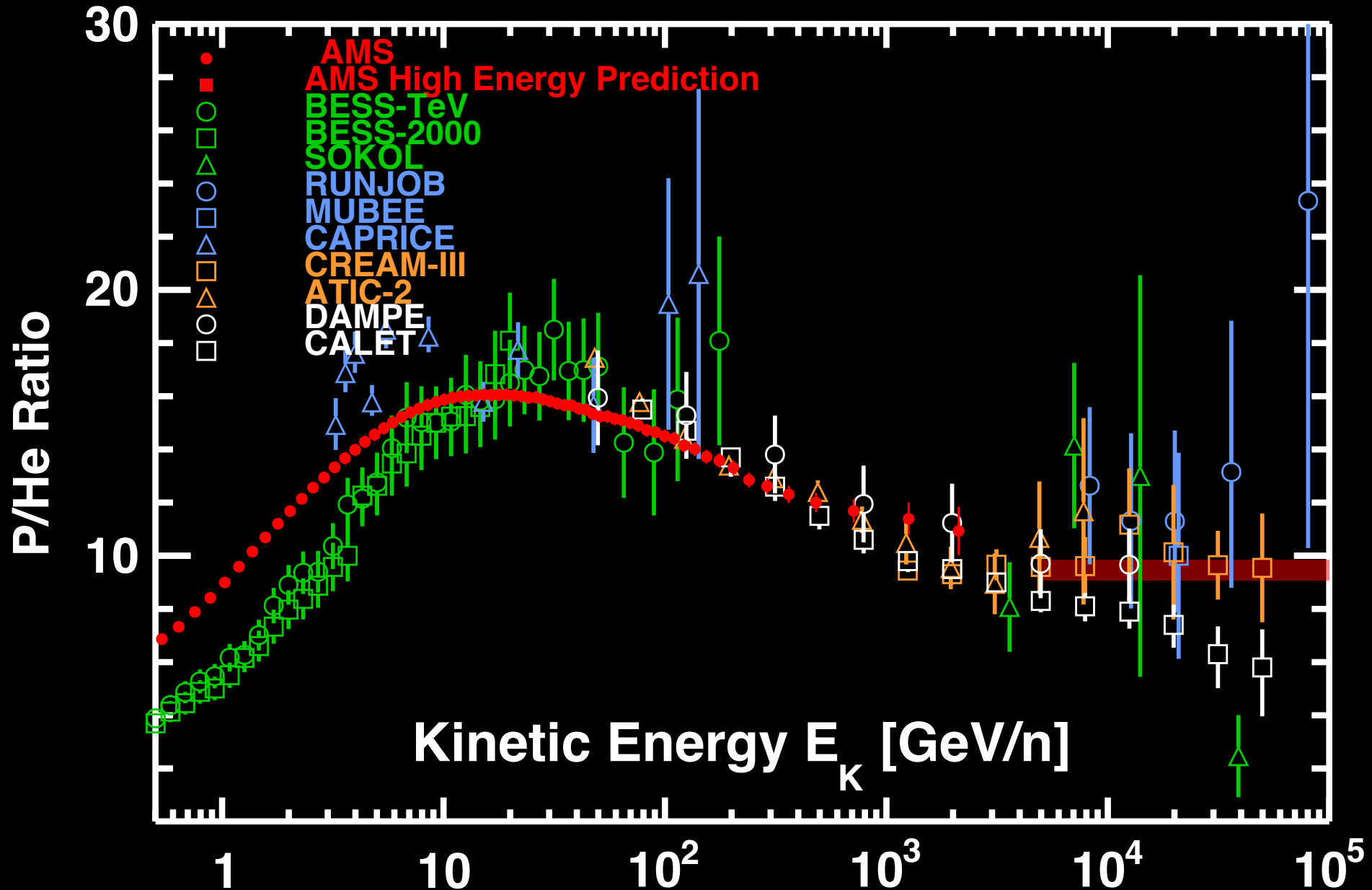


P and He may have same spectral index at highest rigidities

Physics Reports, 894, 1 (2021) :

AMS found that proton flux have two components,
one is like Helium and another is unique to proton flux.

Proton/Helium Flux Ratio

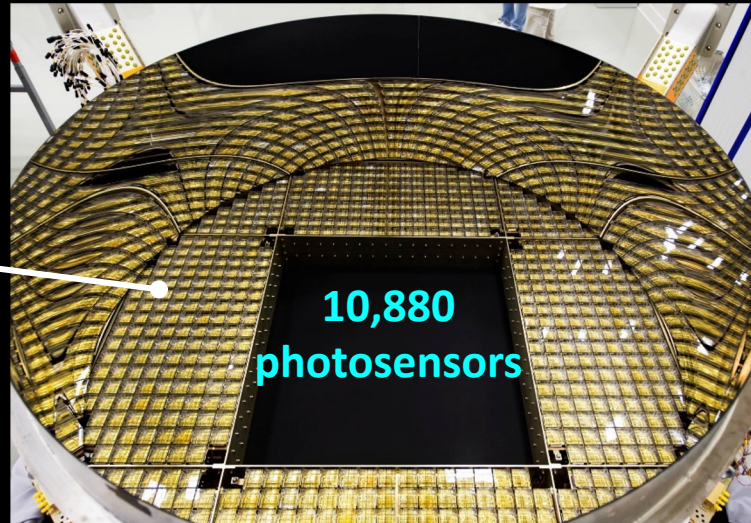
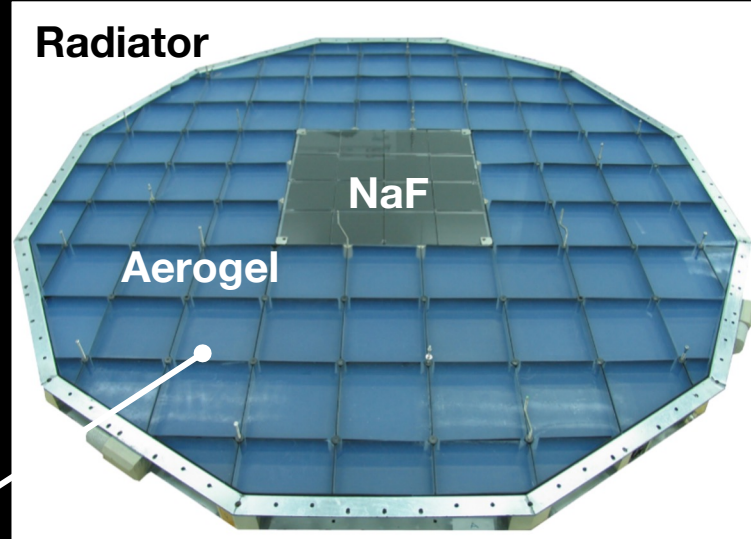
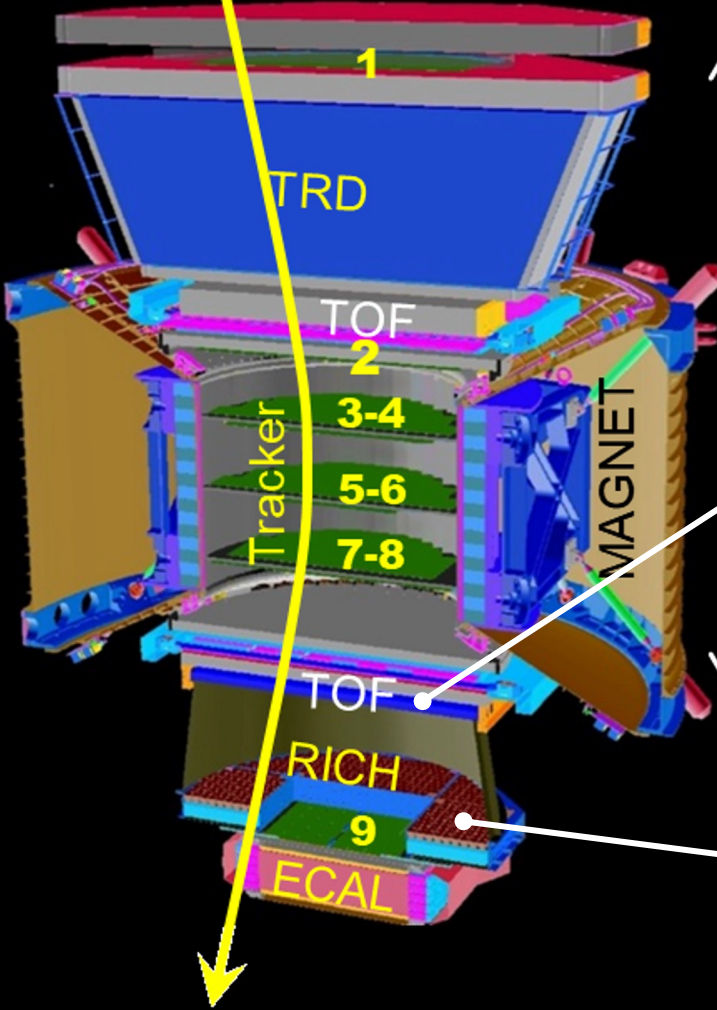


What about other $Z=1$ heavy cosmic rays? Let's look on Deuterons.

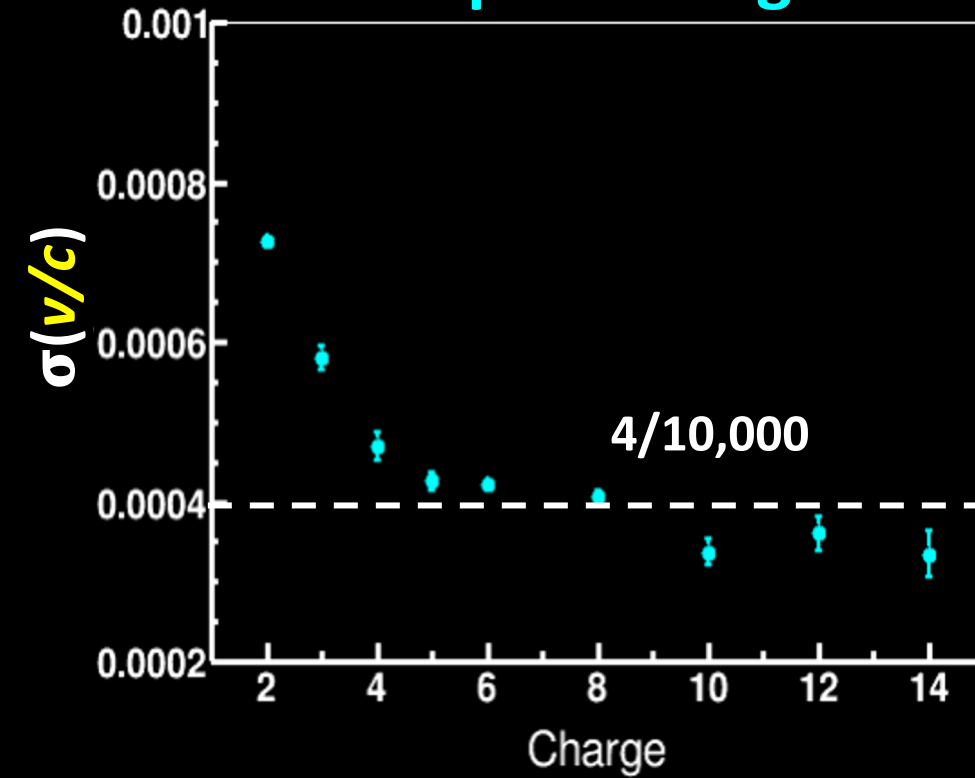
Precision measurement of isotopes

Isotopes: Same Z , different m

Ring Imaging CHerenkov (RICH)



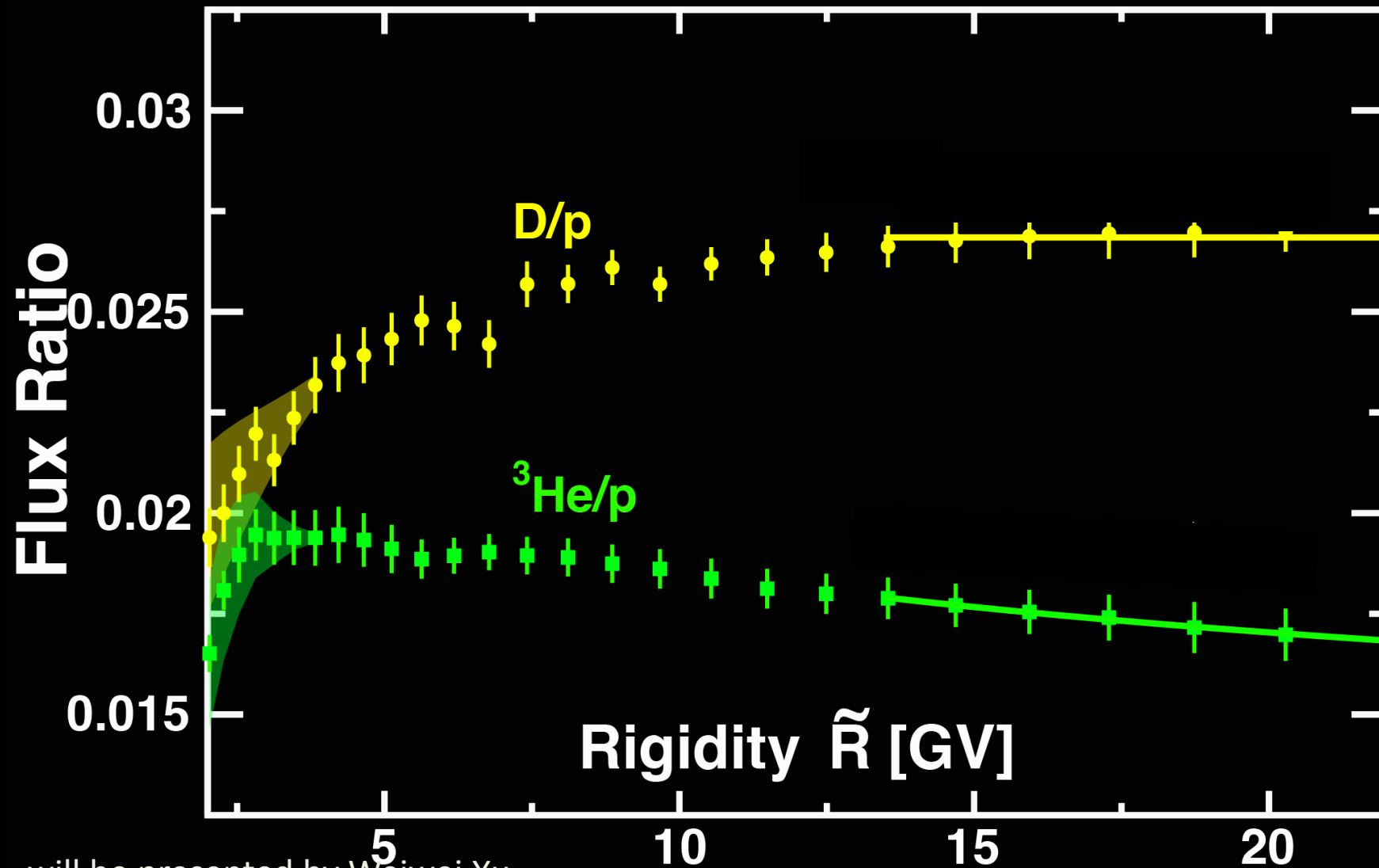
**RICH velocity v resolution
at the speed of light c**



Origin of Cosmic Deuterons D

He, C, O, ... + Interstellar Media \rightarrow (D, ^3He) + X

D and ^3He are both known to be secondary cosmic rays



The flux ratio of $^3\text{He}/p$ decreases with rigidity above 4 GV.

If D is pure secondary, the flux ratio of D/p must also decrease with rigidity above ~ 4 GV

The flux ratio of D/p increases with rigidity and is about to be constant above 13 GV.

D must have an additional primary source

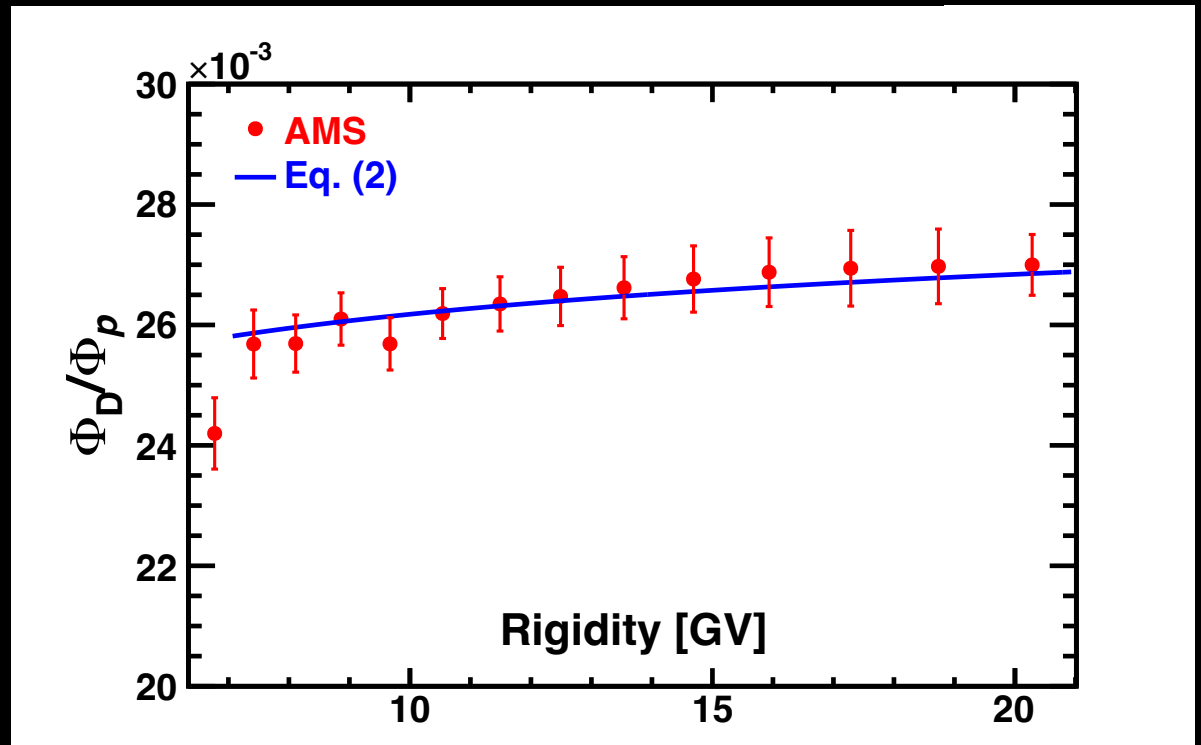
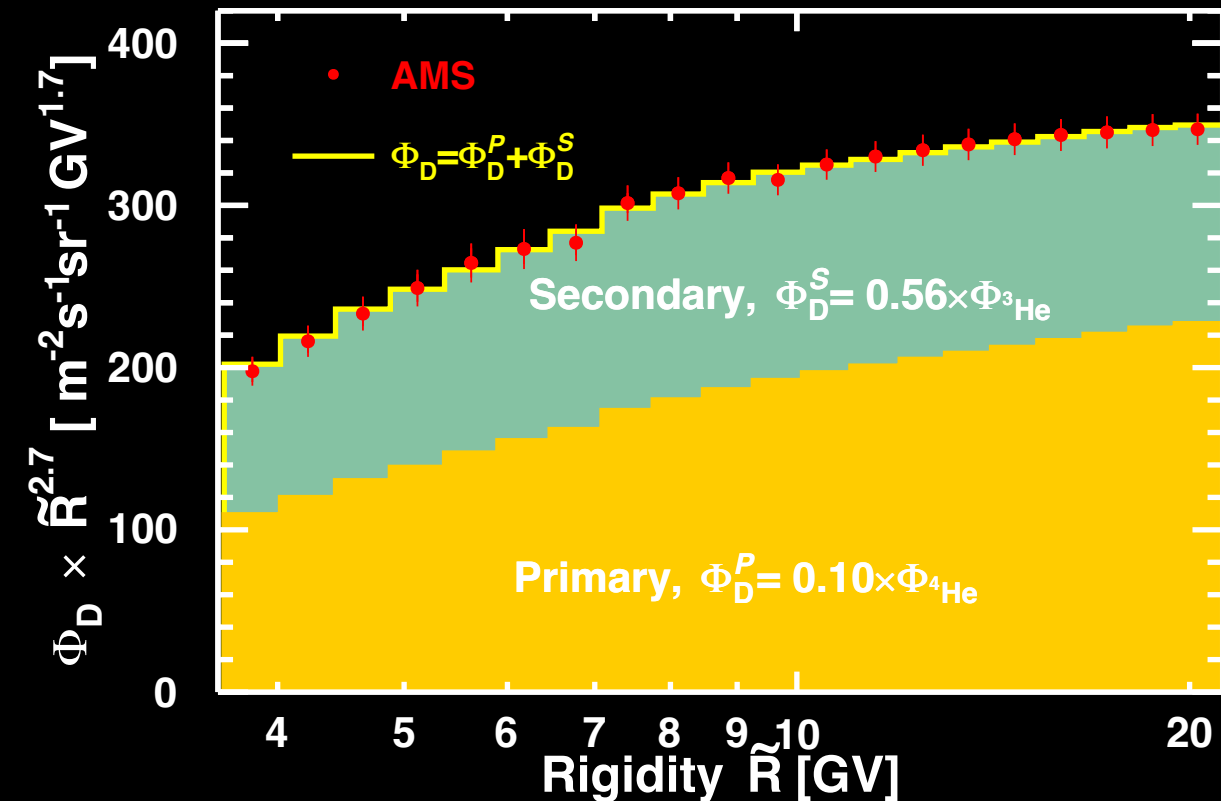
To find the D/p flux ratio rigidity dependence and to predict its value at high rigidities we note that

$$\frac{D/p}{1 + D/p} \equiv D/({}^3\text{He} + {}^4\text{He}) \times \text{He}/(p + D) \quad (1)$$

We then evaluate the D flux rigidity dependence as a weighted sum of primary and secondary fluxes $D = 0.56 \times {}^3\text{He} + 0.1 \times {}^4\text{He}$, ${}^3\text{He}/{}^4\text{He} = 0.153 \times (R/3.5\text{GV})^{-0.29}$ and the $\text{He}/(p+D) = 3.15 + 3.18 \times (R/3.5\text{GV})^{-0.29}$

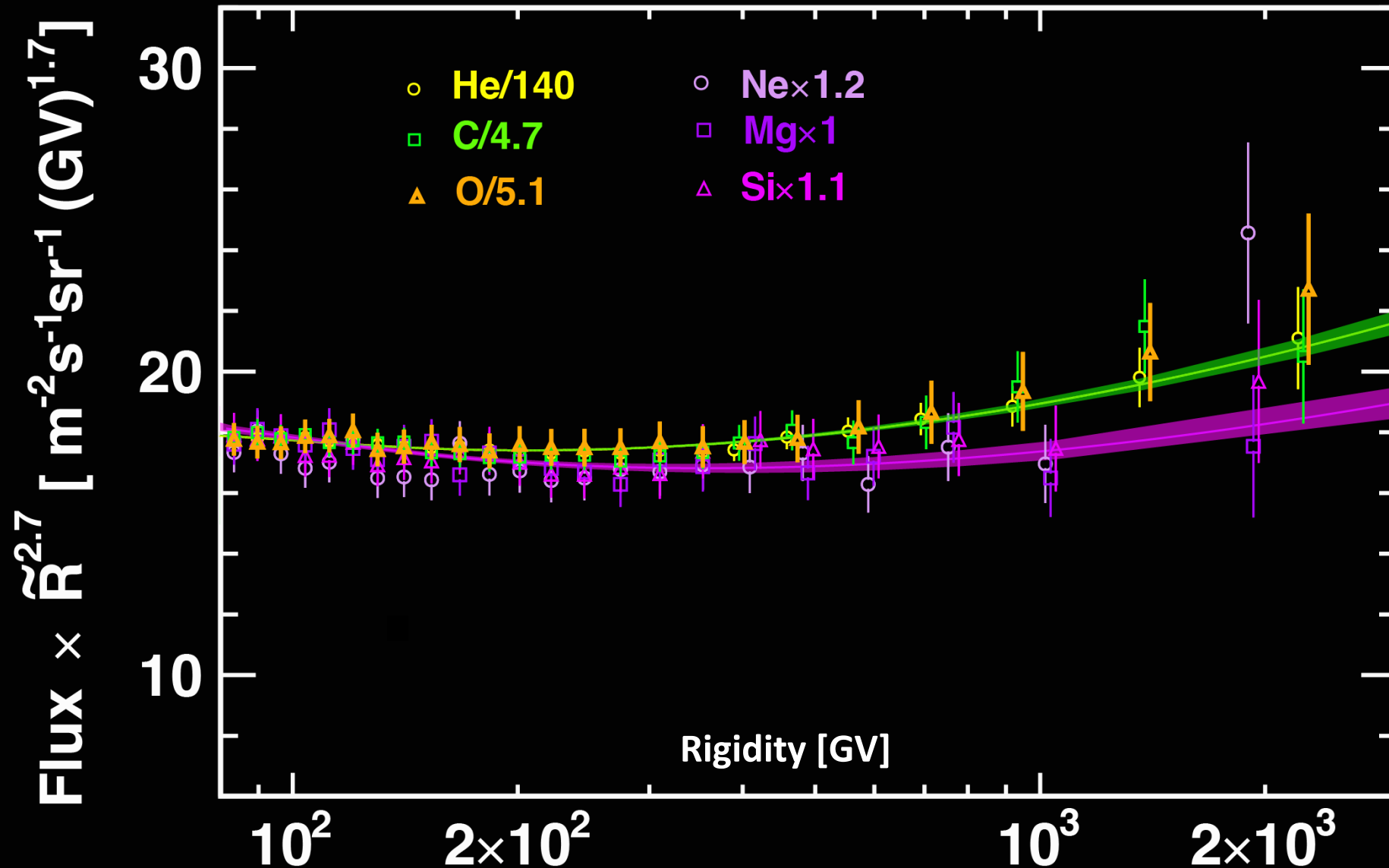
$$\frac{D/p}{1 + D/p} = \frac{1 + 0.90(R/3.5\text{GV})^{-0.29}}{1 + 0.153(R/3.5\text{GV})^{-0.29}} \times \frac{0.030}{(1 + 1.01 \times (R/3.5\text{GV})^{-0.29})} \quad (2)$$

From the equation, we got for the D/p flux ratio at high rigidities, $D/p^{R \rightarrow \infty} = 0.031 \pm 0.001$.

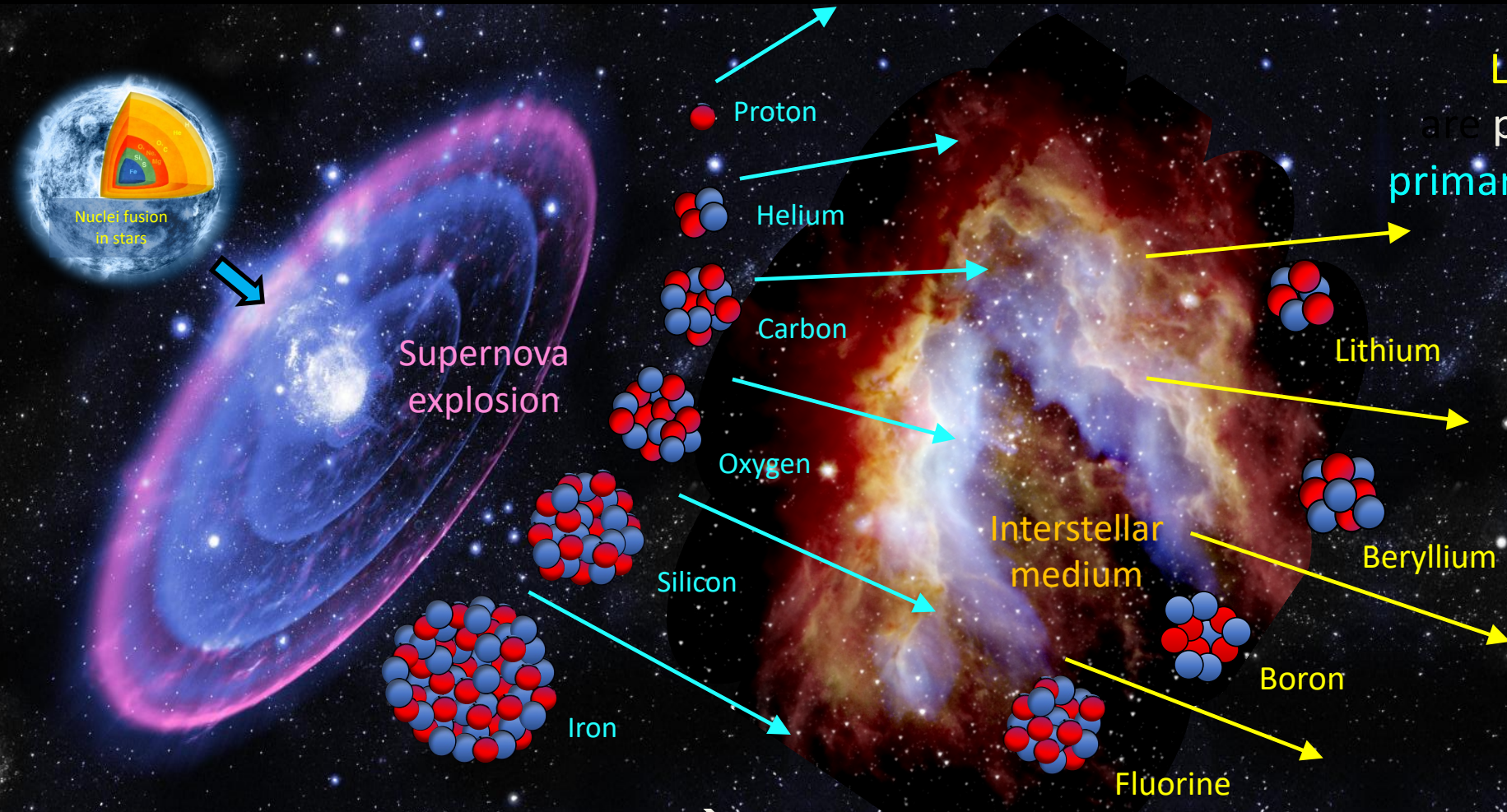


AMS Results: Primary cosmic rays have two classes

Light elements He-C-O and Heavier elements Ne-Mg-Si
each have their own rigidity dependence



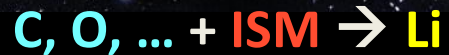
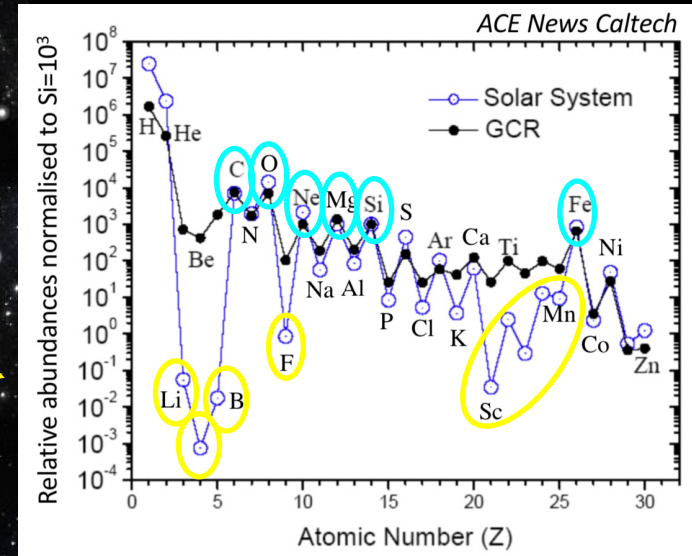
Secondary Cosmic Rays



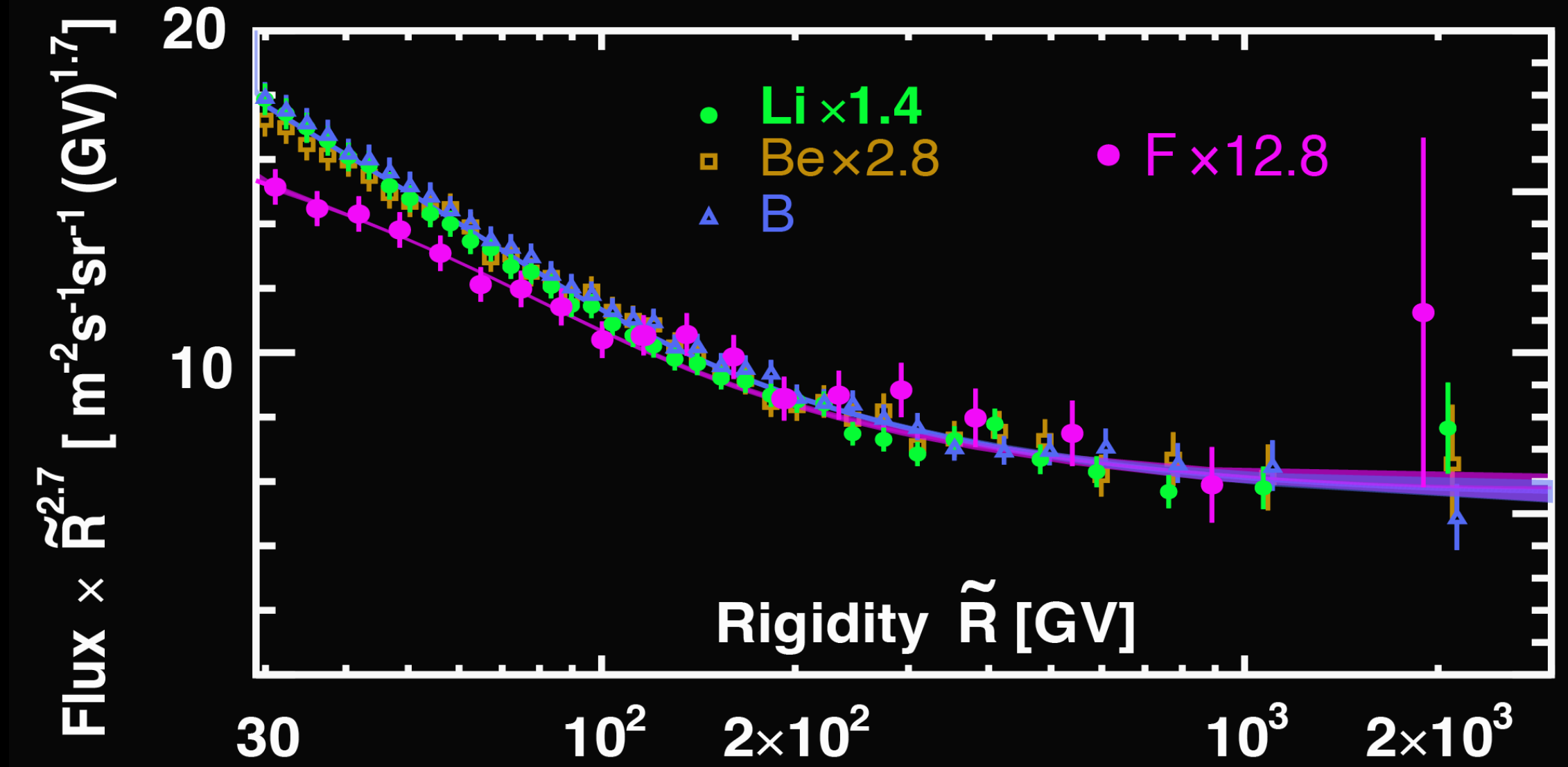
Secondary

Li, Be, B, F, sub-Fe nuclei

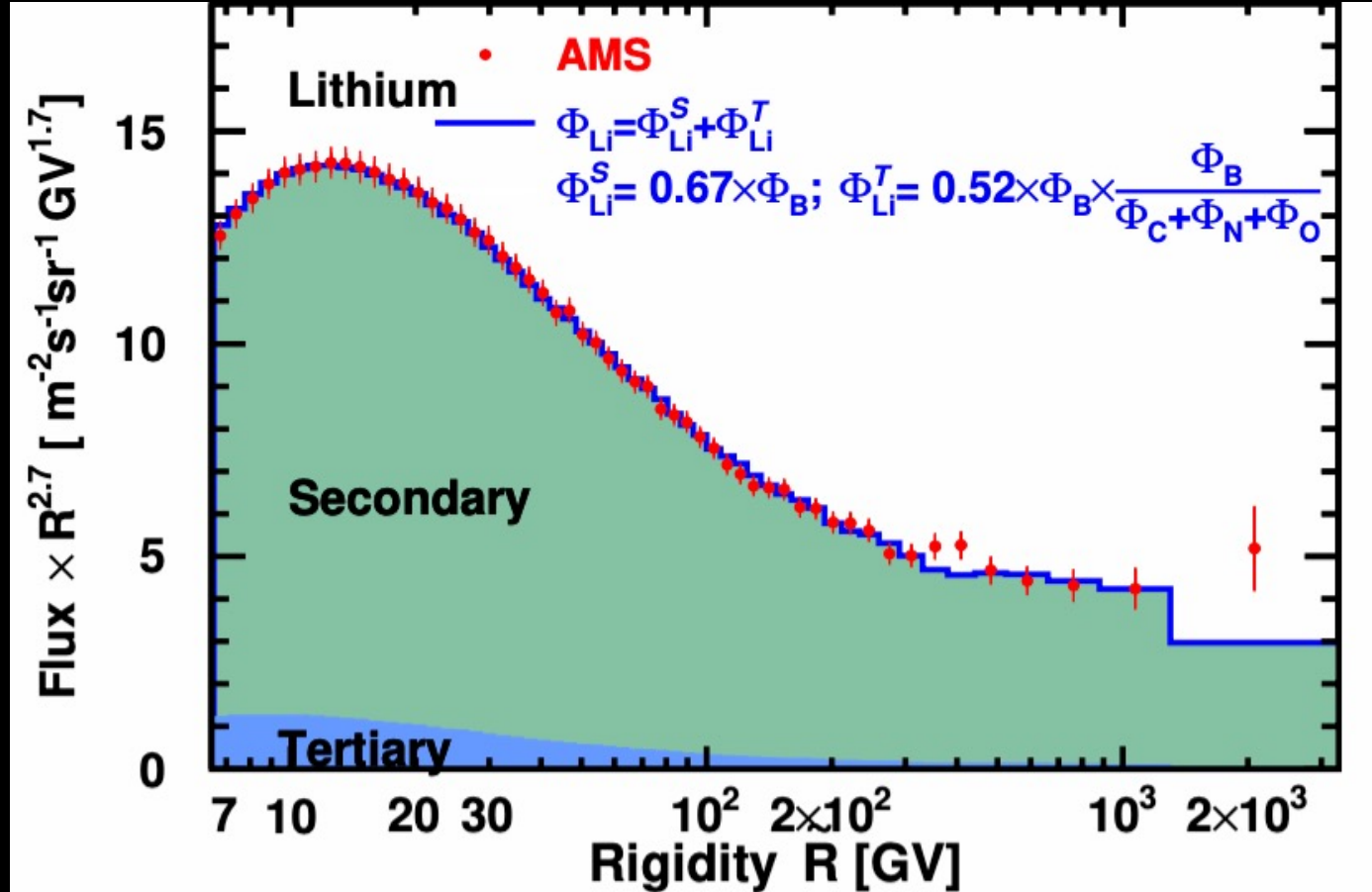
are produced by the collision of primary cosmic rays C, O, Si, ... , Fe with the interstellar medium



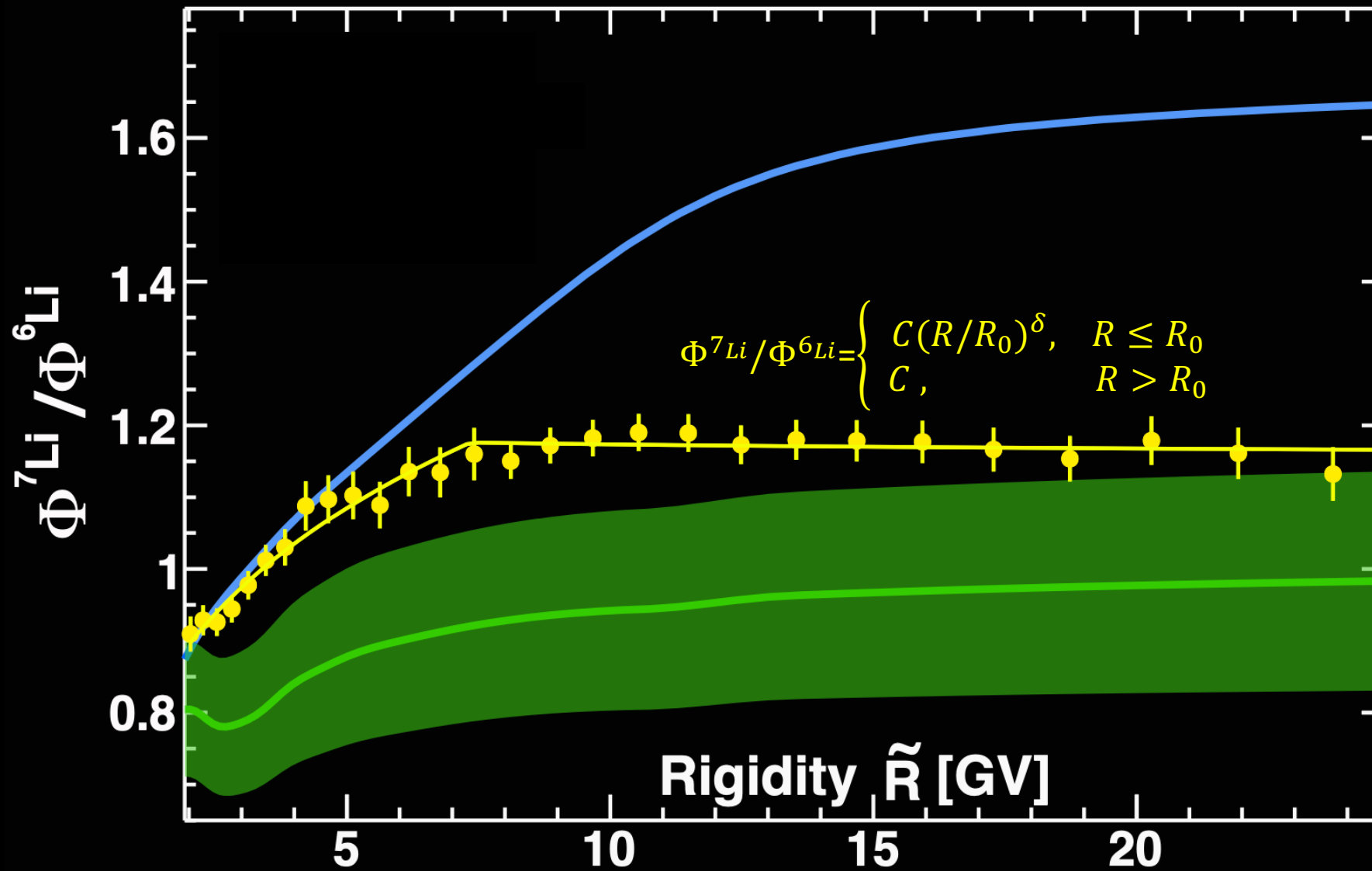
Secondary cosmic rays have two classes **Li-Be-B** and **F**
have their own rigidity dependence



With more statistics we can study the differences between Li and B fluxes rigidity dependence at low rigidities. One can expect that some part of Li flux has a tertiary nature, i.e. produced by the interaction of B and Be with interstellar medium. To obtain the secondary and the tertiary components of the Li flux we fit $\Phi_{\text{Li}} = \Phi_{\text{Li}}^S + \Phi_{\text{Li}}^T$ to the weighted sum of a secondary cosmic ray flux, namely boron Φ_B , and of a tertiary cosmic ray flux, namely $\Phi_B \times \frac{\Phi_B}{\Phi_C + \Phi_N + \Phi_O}$ above 6 GV. The fit yields $\Phi_{\text{Li}}^S = (0.67 \pm 0.02) \times \Phi_B$ and $\Phi_{\text{Li}}^T = (0.52 \pm 0.08) \times \Phi_B \times \frac{\Phi_B}{\Phi_C + \Phi_N + \Phi_O}$ with a $\chi^2/\text{d.o.f.} = 28/53$.



Results on Lithium Isotopes



GALPROP Model
assuming a
primary component
in the ^7Li flux

AMS data and fit, $R_0 = 7\text{GV}$

USINE Model
assuming secondary
origin of ^6Li and ^7Li

Above $\sim 7\text{ GV}$, the rigidity dependence of ^6Li and ^7Li fluxes are identical

Excludes the existence of a sizable primary component in the ^7Li flux

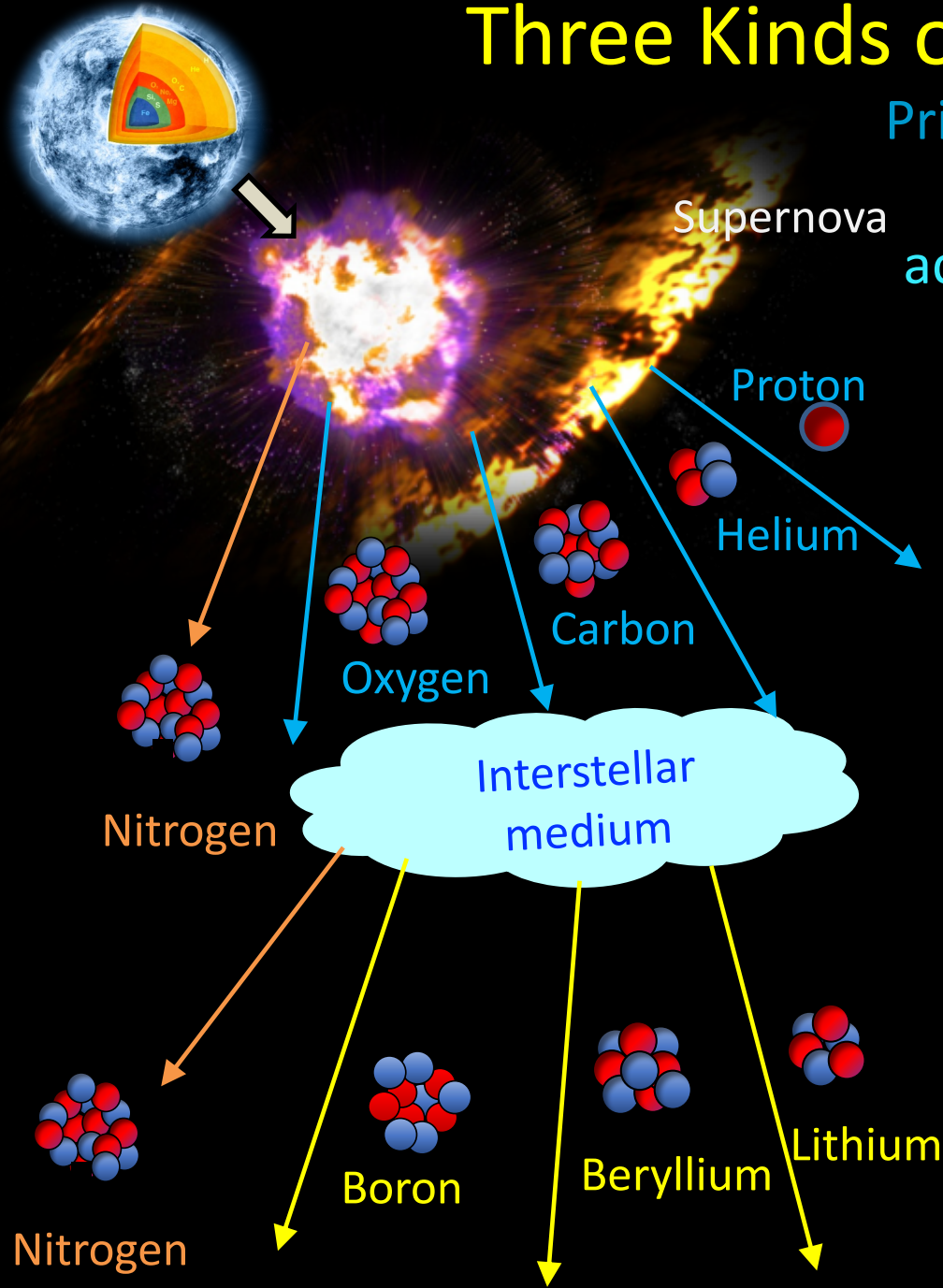
Three Kinds of Charged Cosmic Rays

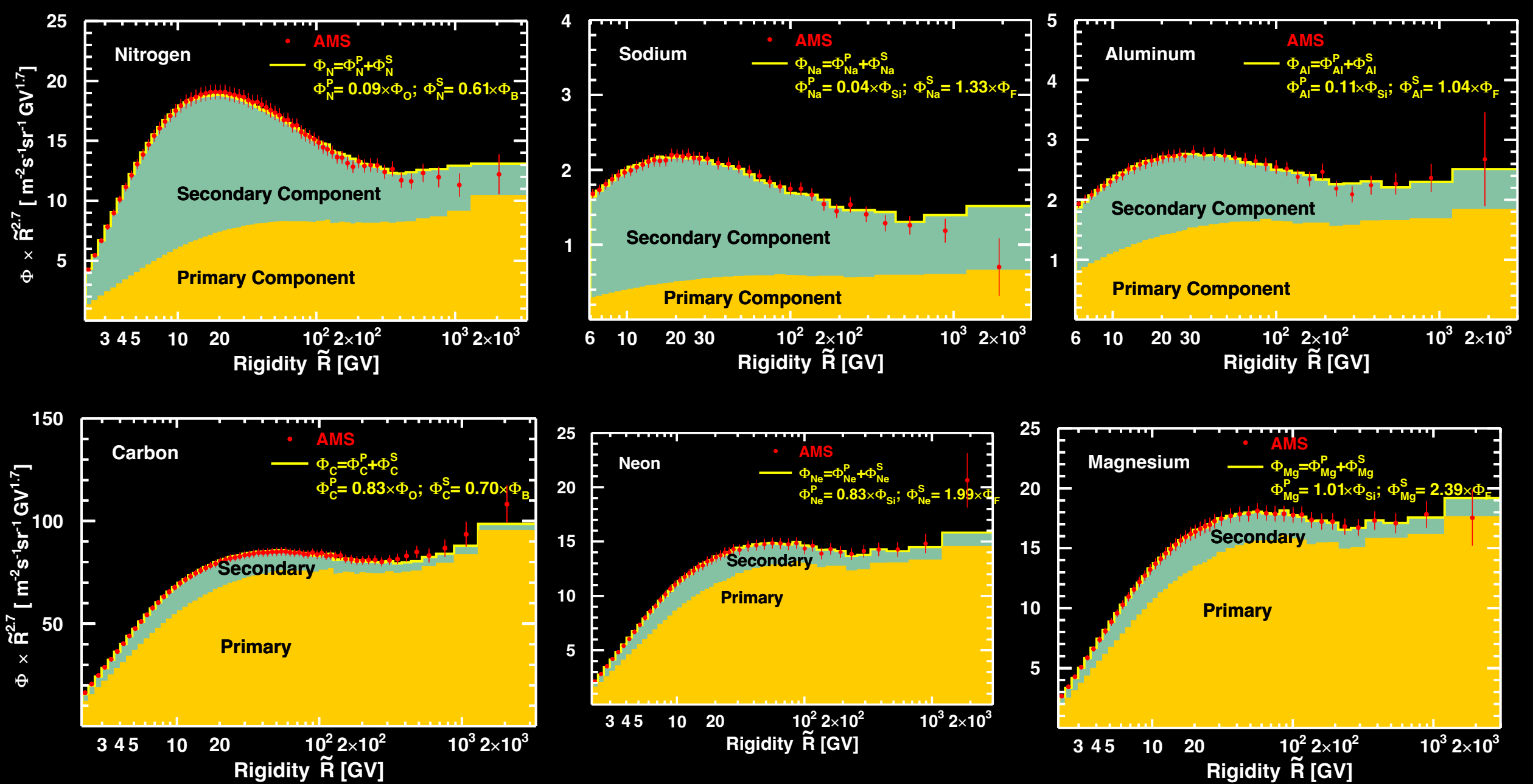
Primary cosmic rays (p, He, C, O, Ne, Mg, S, Ar, ..., Fe) are mostly produced during the lifetime of stars and are accelerated in supernovae shocks, whose explosion rate is about 2-3 per century in our Galaxy.

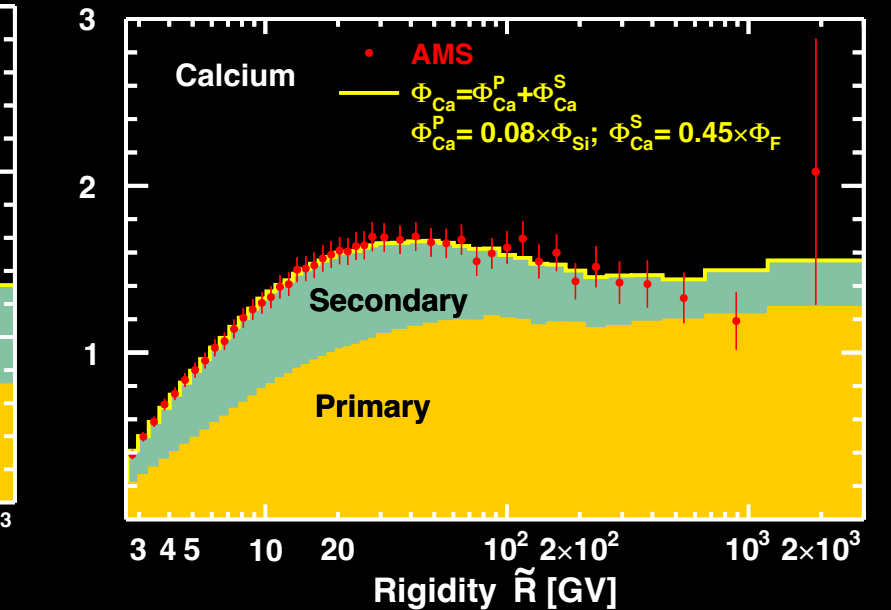
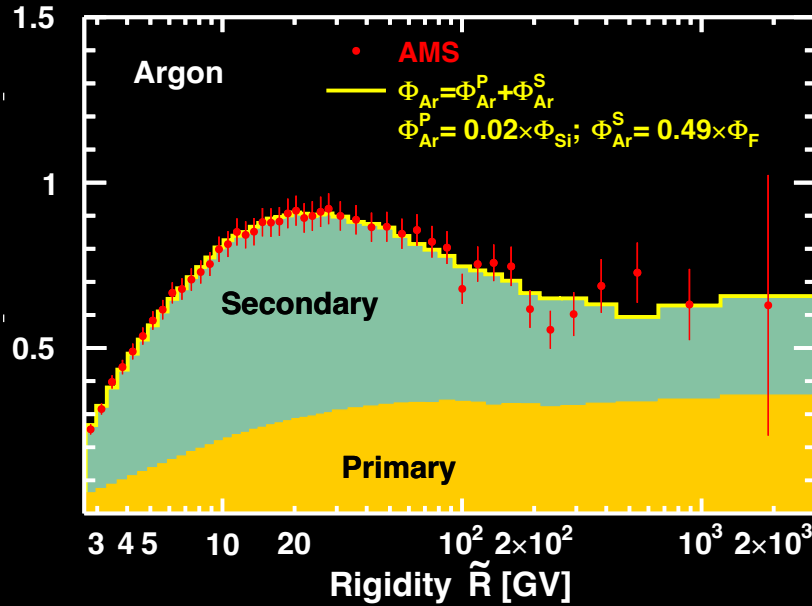
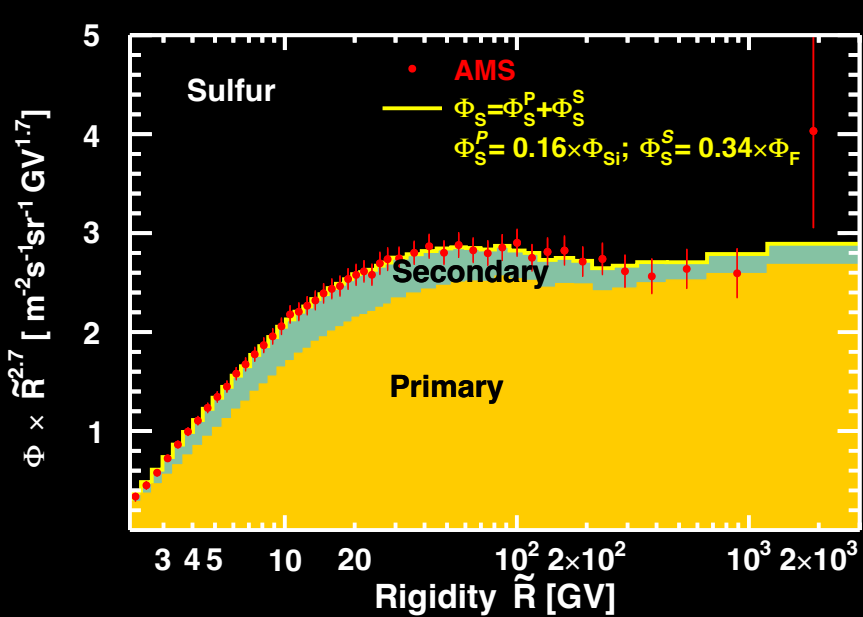
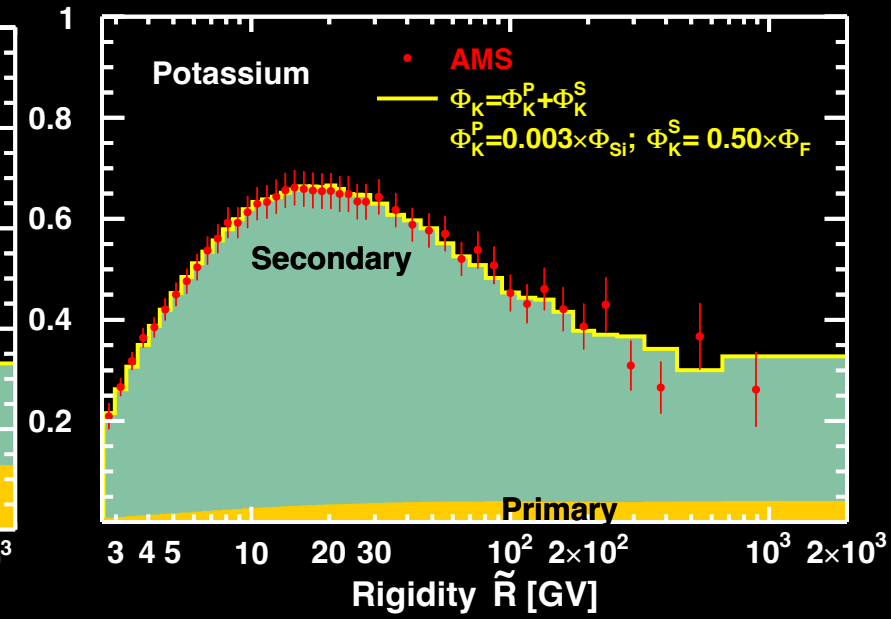
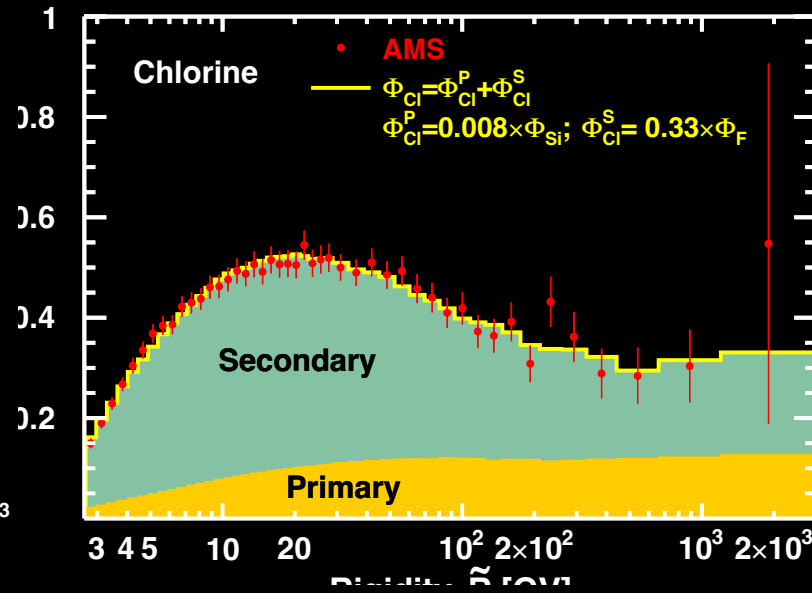
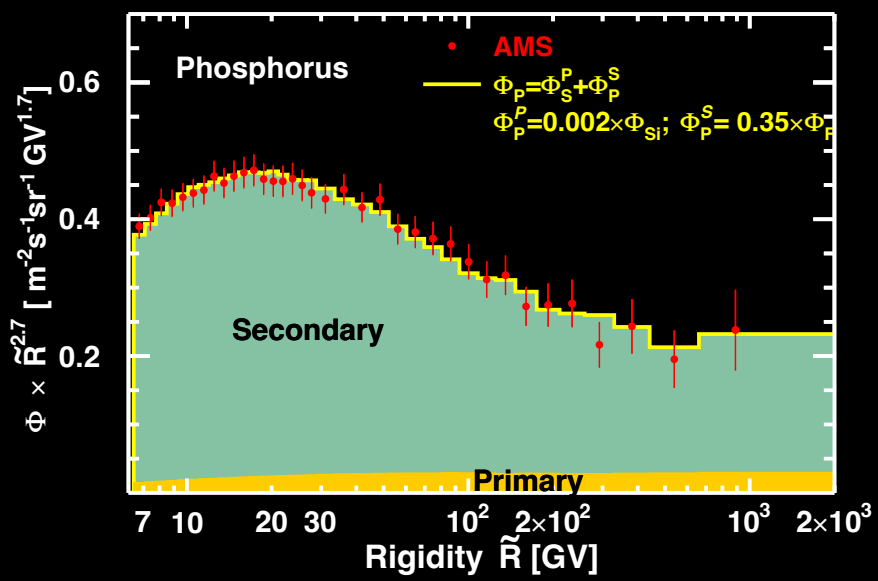
Secondary cosmic nuclei (Li, Be, B, F, ...) are produced by the collisions of primary cosmic rays and interstellar medium.

Cosmic nuclei with both Primary and Secondary Components (N, Na, Al, Cl, ...) . Many primary cosmic rays C, Ne, Mg, S are also expected to have sizeable secondary component.

AMS found that all such fluxes can be presented as a weighted sums of the characteristic primary flux (O, Si) and a characteristic secondary flux (B, F)







Preliminary results, please refer to fututre AMS publication

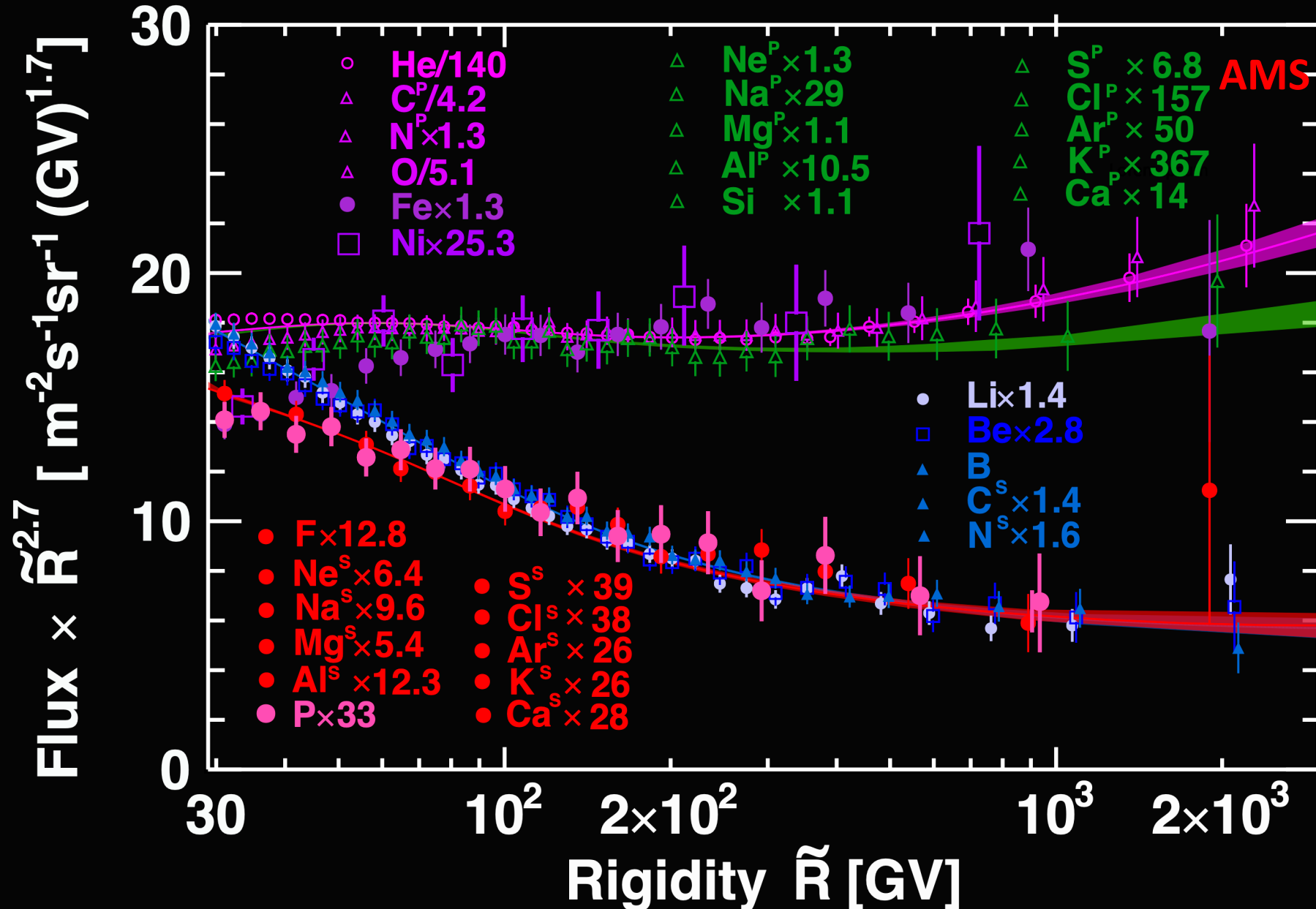
Summary Primary and Secondary Components

Nuclei Flux	Primary	Secondary	Secondary Fr, % 6 GV	Secondary Fr, % 2 TV
Φ_C	$(0.835 \pm 0.025) \times \Phi_O$	$(0.70 \pm 0.03) \times \Phi_B$	20 \pm 1	4 \pm 0.5
Φ_N	$(0.091 \pm 0.002) \times \Phi_O$	$(0.61 \pm 0.02) \times \Phi_B$	69 \pm 1	23 \pm 2
Φ_{Ne}	$(0.831 \pm 0.025) \times \Phi_{Si}$	$(1.99 \pm 0.14) \times \Phi_F$	24 \pm 1	5 \pm 0.5
Φ_{Na}	$(0.038 \pm 0.003) \times \Phi_{Si}$	$(1.33 \pm 0.04) \times \Phi_F$	81 \pm 2	38 \pm 12
Φ_{Mg}	$(1.008 \pm 0.025) \times \Phi_{Si}$	$(2.39 \pm 0.17) \times \Phi_F$	25 \pm 1	5 \pm 0.5
Φ_{Al}	$(0.105 \pm 0.004) \times \Phi_{Si}$	$(1.04 \pm 0.03) \times \Phi_F$	57 \pm 2	22 \pm 8
Φ_P	$(0.002^{+0.002}_{-0.001}) \times \Phi_{Si}$	$(0.35 \pm 0.02) \times \Phi_F$	98 \pm 2	92 \pm 2
Φ_S	$(0.162 \pm 0.005) \times \Phi_{Si}$	$(0.33 \pm 0.04) \times \Phi_F$	18 \pm 3	3 \pm 1
Φ_{Cl}	$(0.008 \pm 0.001) \times \Phi_{Si}$	$(0.33 \pm 0.01) \times \Phi_F$	86 \pm 3	63 \pm 4
Φ_{Ar}	$(0.021 \pm 0.002) \times \Phi_{Si}$	$(0.49 \pm 0.02) \times \Phi_F$	74 \pm 4	44 \pm 4
Φ_K	$(0.003 \pm 0.001) \times \Phi_{Si}$	$(0.50 \pm 0.02) \times \Phi_F$	96 \pm 3	87 \pm 3
Φ_{Ca}	$(0.076 \pm 0.003) \times \Phi_{Si}$	$(0.45 \pm 0.03) \times \Phi_F$	44 \pm 2	18 \pm 1

The C (Z=6) to Ca (Z=20) cosmic ray nuclei primary and secondary components derived as fractions of O(Si) and B(F) fluxes, respectively, and their secondary fractions at 6 GV and 2 TV.

This allows to measure relative cosmic ray abundances of C/O, N/O, Ne/Si, Na/Si, Mg/Si, Al/Si, S/Si, Cl/Si, Ar/Si, K/Si, and Ca/Si at the source independently of cosmic ray propagation.

Primary and Secondary Decomposition for All Cosmic Ray Fluxes of $2 \leq Z \leq 20$, $Z=26$, $Z=28$

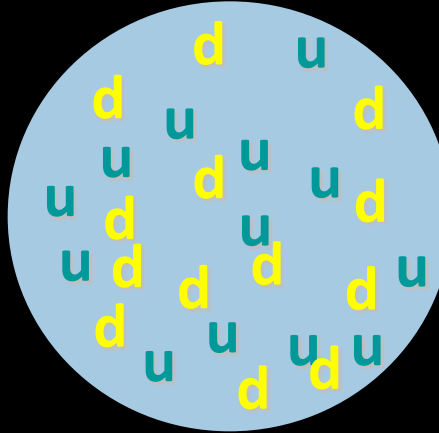


Strange Quark Matter – “Strangelets”

E. Witten, Phys. Rev. D, 272-285 (1984)

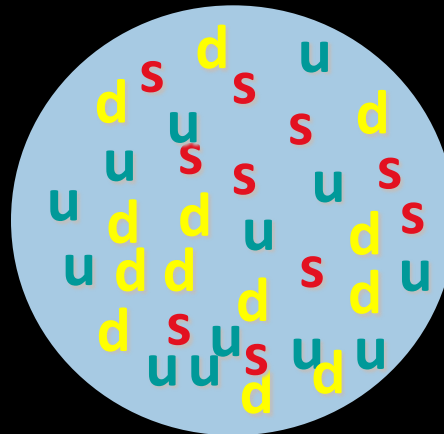
There are six quarks – u, d, s, c, b, and t.

All the material on Earth is made out of u and d quarks



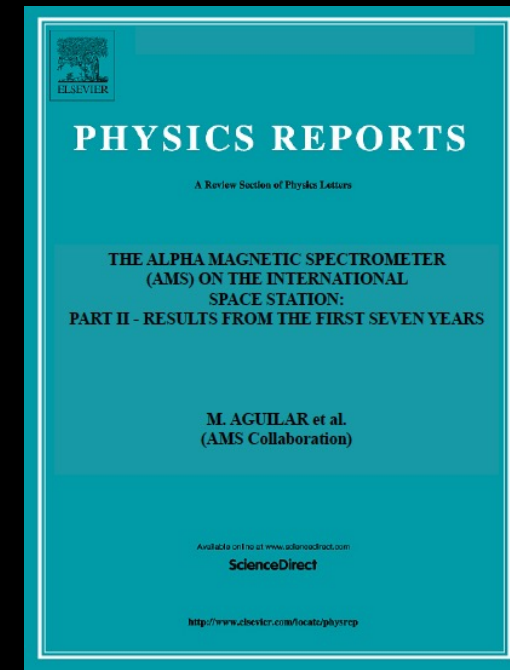
Diamond ($Z/A \sim 0.5$)

Is there material in the universe made up of u, d, & s quarks?

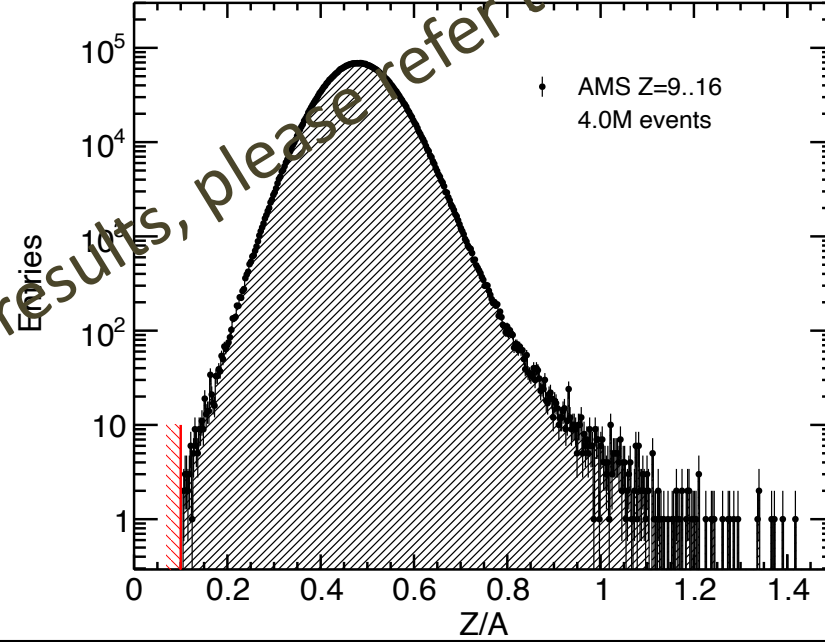
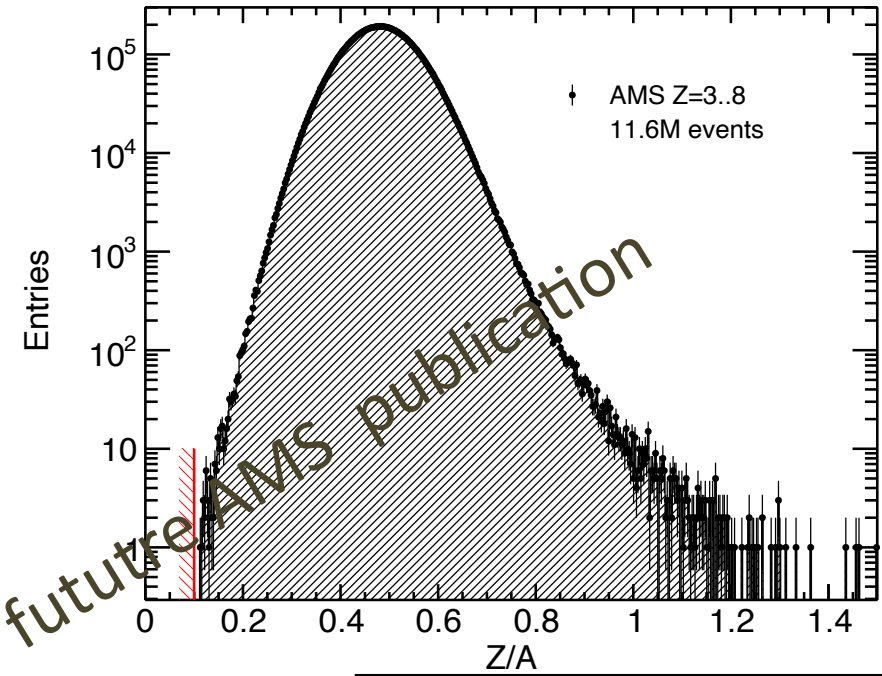
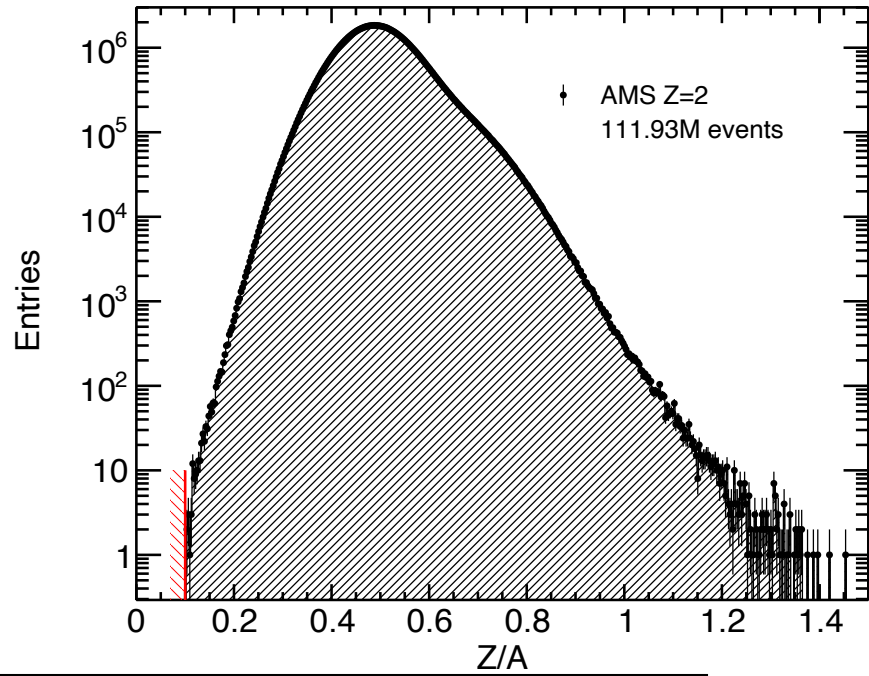


Strangelet ($Z/A < 0.1$)

This is being answered by AMS.

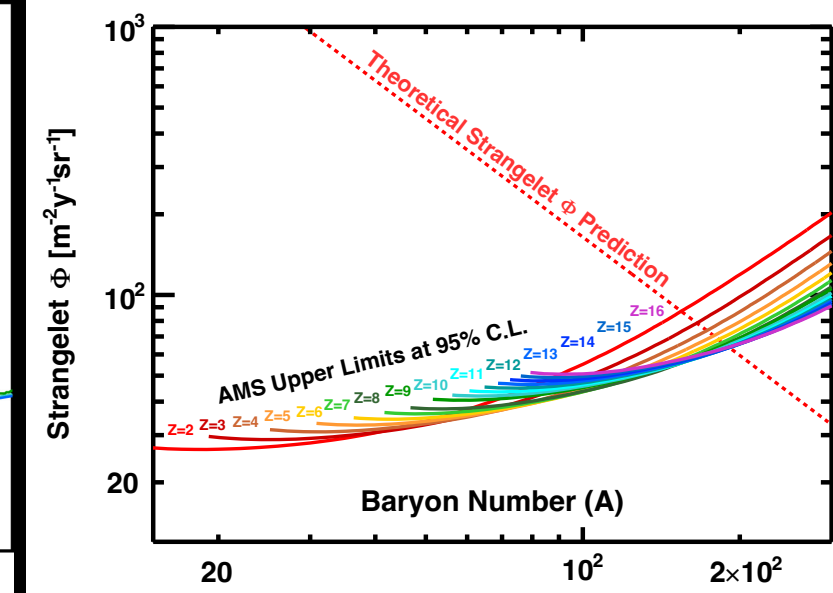
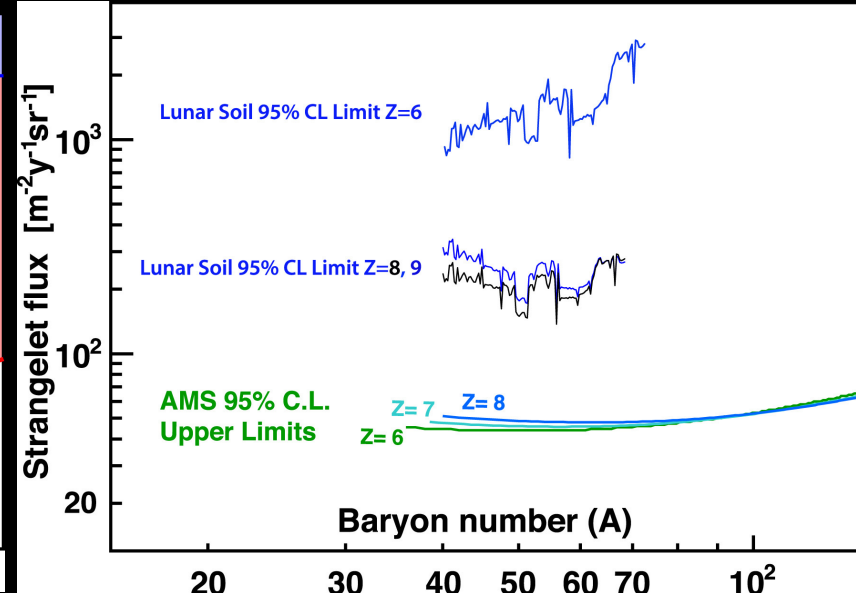
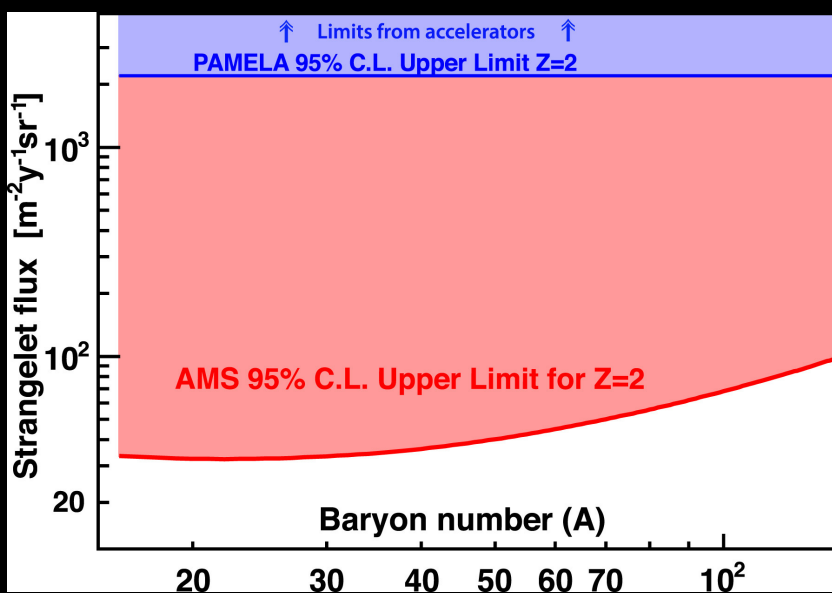


Search for Strangelets $Z=2\dots 16$

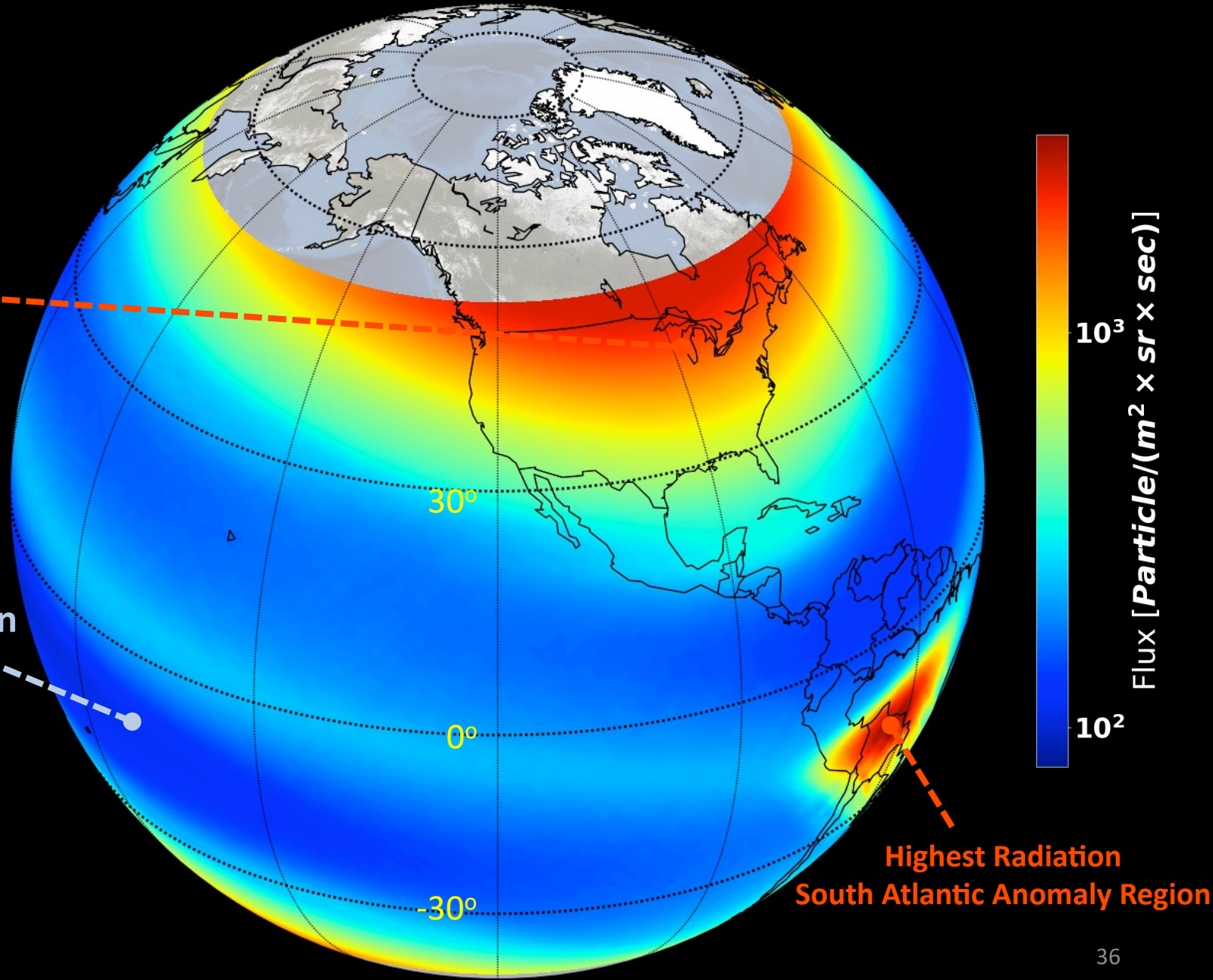
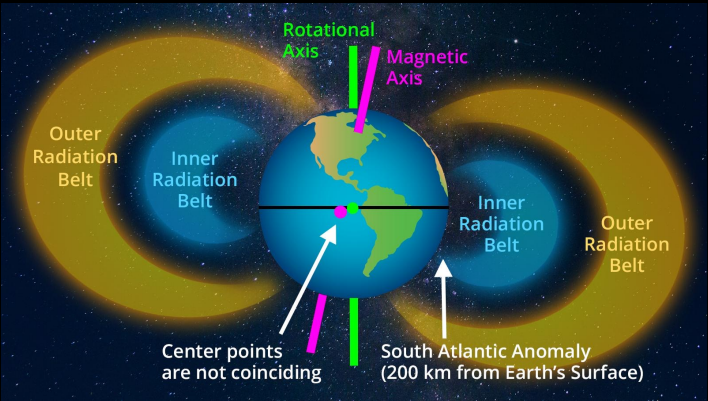
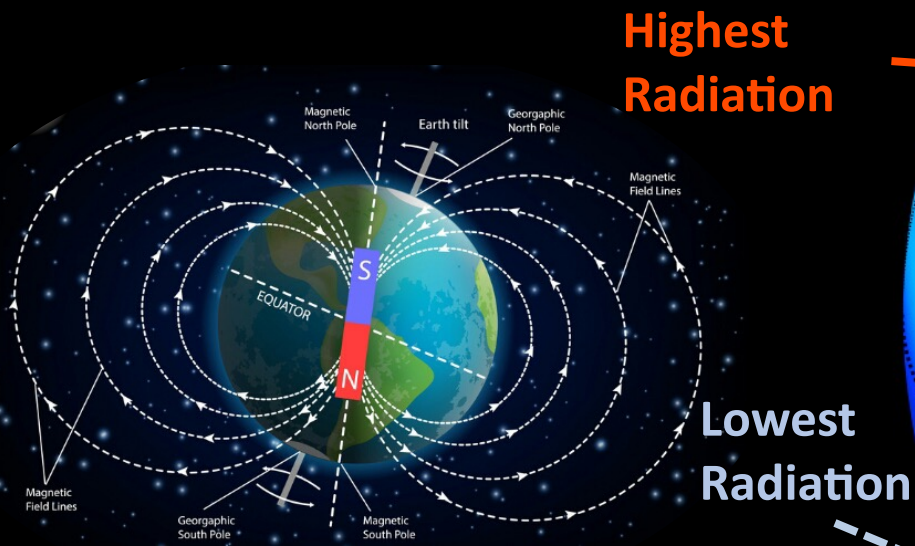


Preliminary results, please refer to future AMS publication

Strangelets with $Z = 2$ to 16 are ruled out by AMS



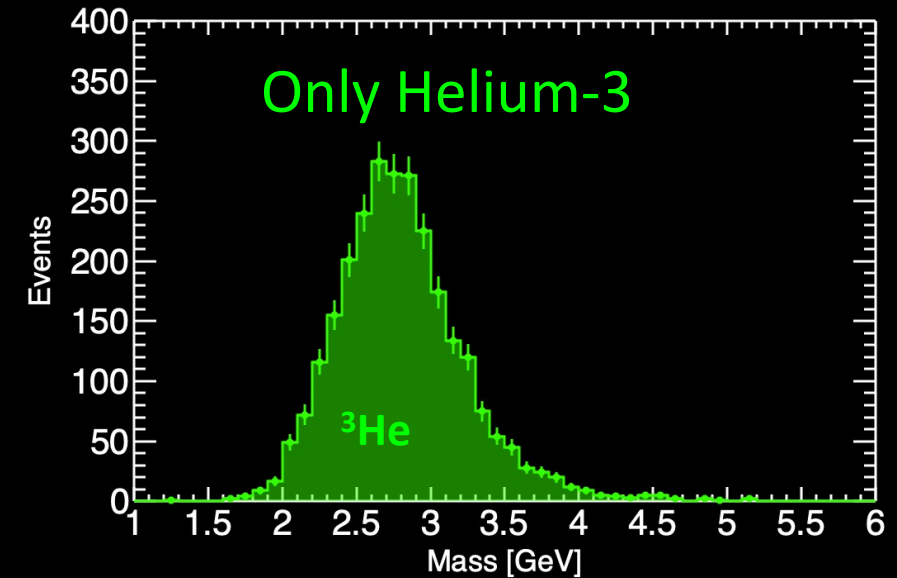
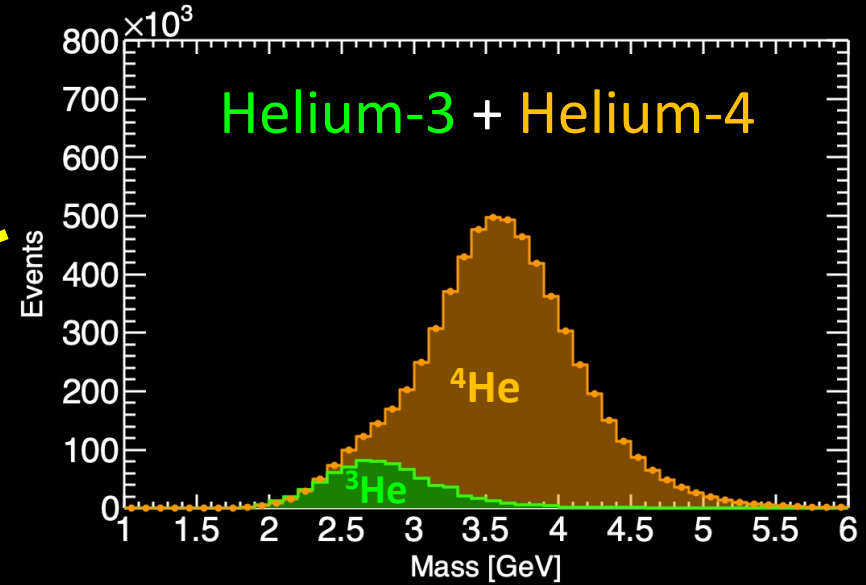
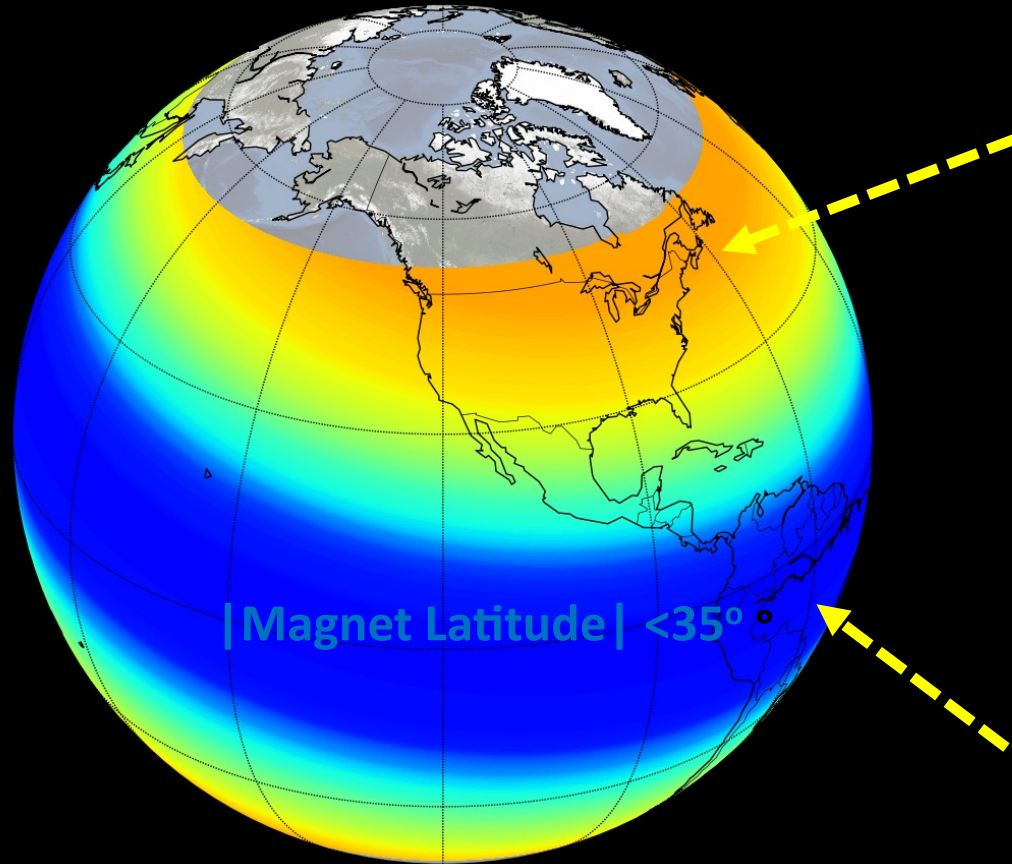
Measurement of near-Earth Radiation



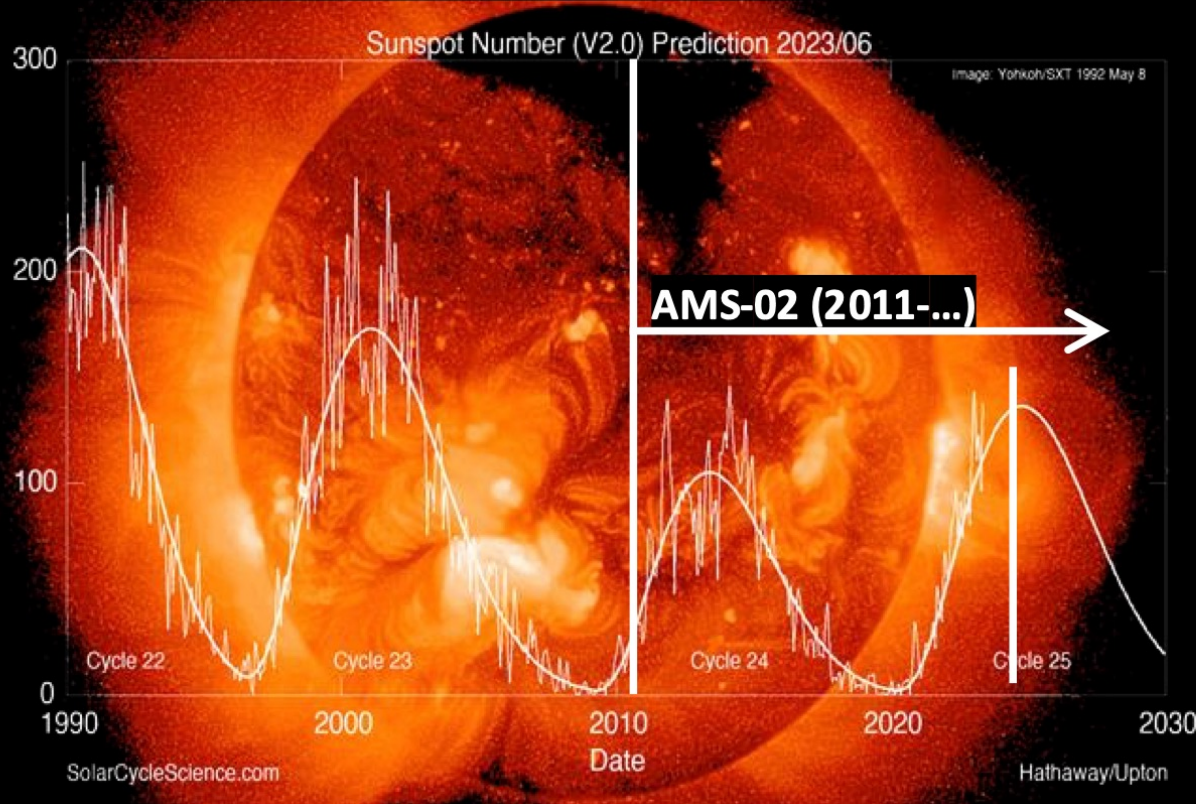
Measurement of near-Earth radiation

In the energy range 100 MeV to 4000 MeV over a large region, $-35^\circ < \Theta_M < +35^\circ$, only ^3He isotope exists.

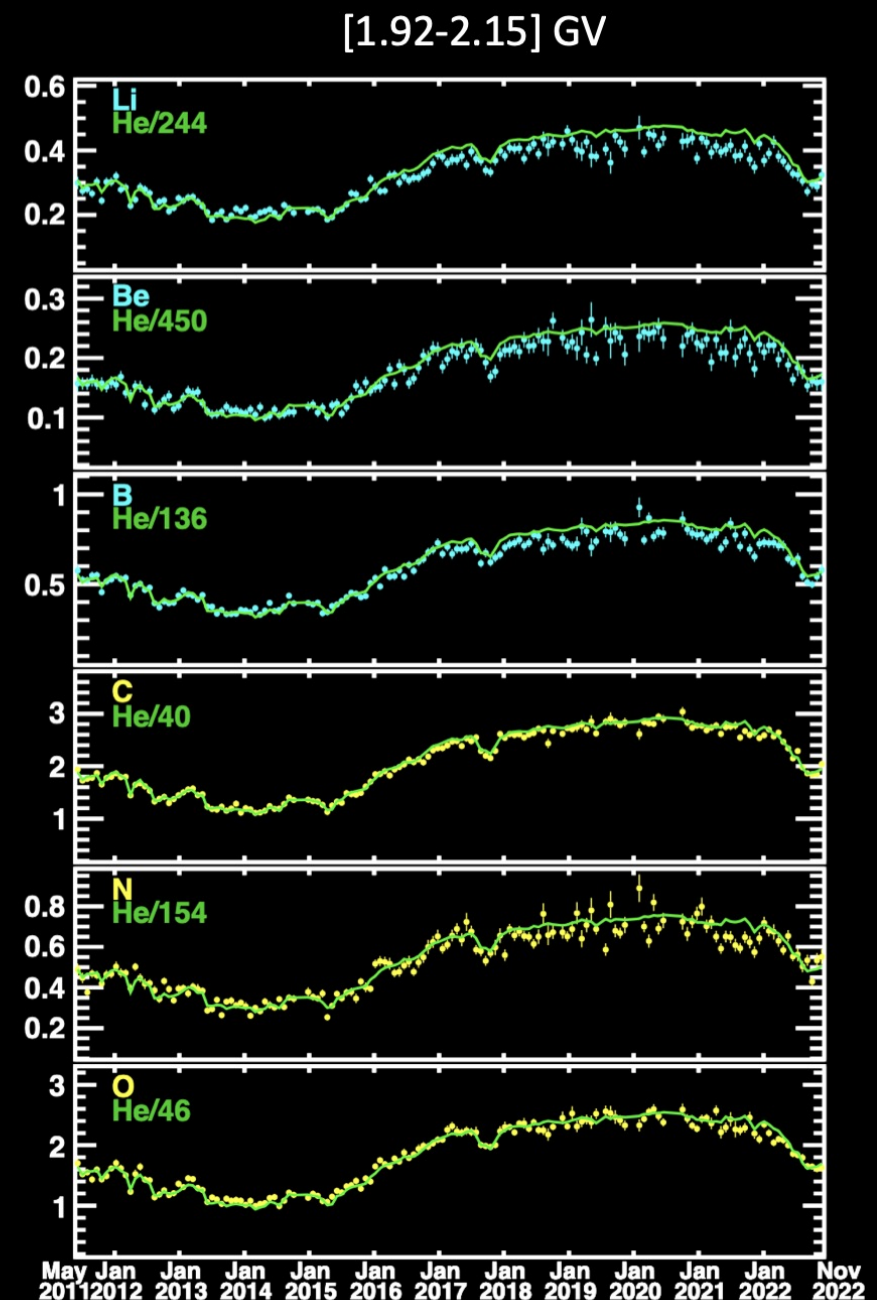
AMS Helium nuclei flux



Time Variation of Cosmic Light Nuclei



Φ_X [$\text{m}^{-2} \text{sr}^{-1} \text{s}^{-1} \text{GV}^{-1}$]



All nuclei exhibit similar long-term and short-term time dependences.

Fluxes are anti-correlated with solar activity, (a) being lower during epoch of high solar activity and (b) higher during epoch of low solar activity.

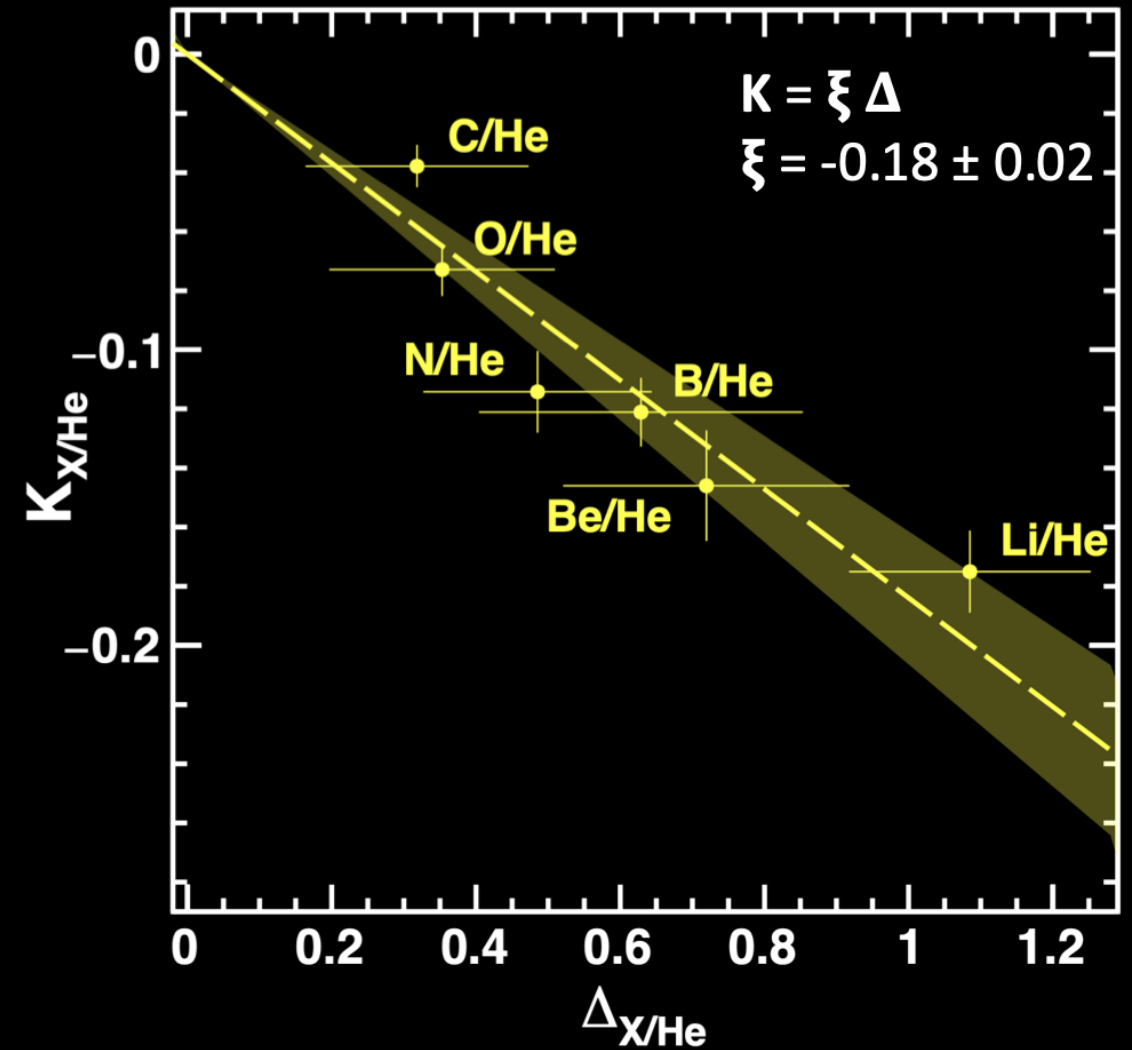
Phys. Rev. Lett. **134**, 051001 (2025)

K-parameter as function of rigidity

[1.92-2.15] GV

$$\frac{\Phi_X^i / \Phi_{\text{He}}^i - \langle \Phi_X^i / \Phi_{\text{He}}^i \rangle}{\langle \Phi_X^i / \Phi_{\text{He}}^i \rangle} = K_{X/\text{He}}^i \frac{\Phi_{\text{He}}^i - \langle \Phi_{\text{He}}^i \rangle}{\langle \Phi_{\text{He}}^i \rangle}$$

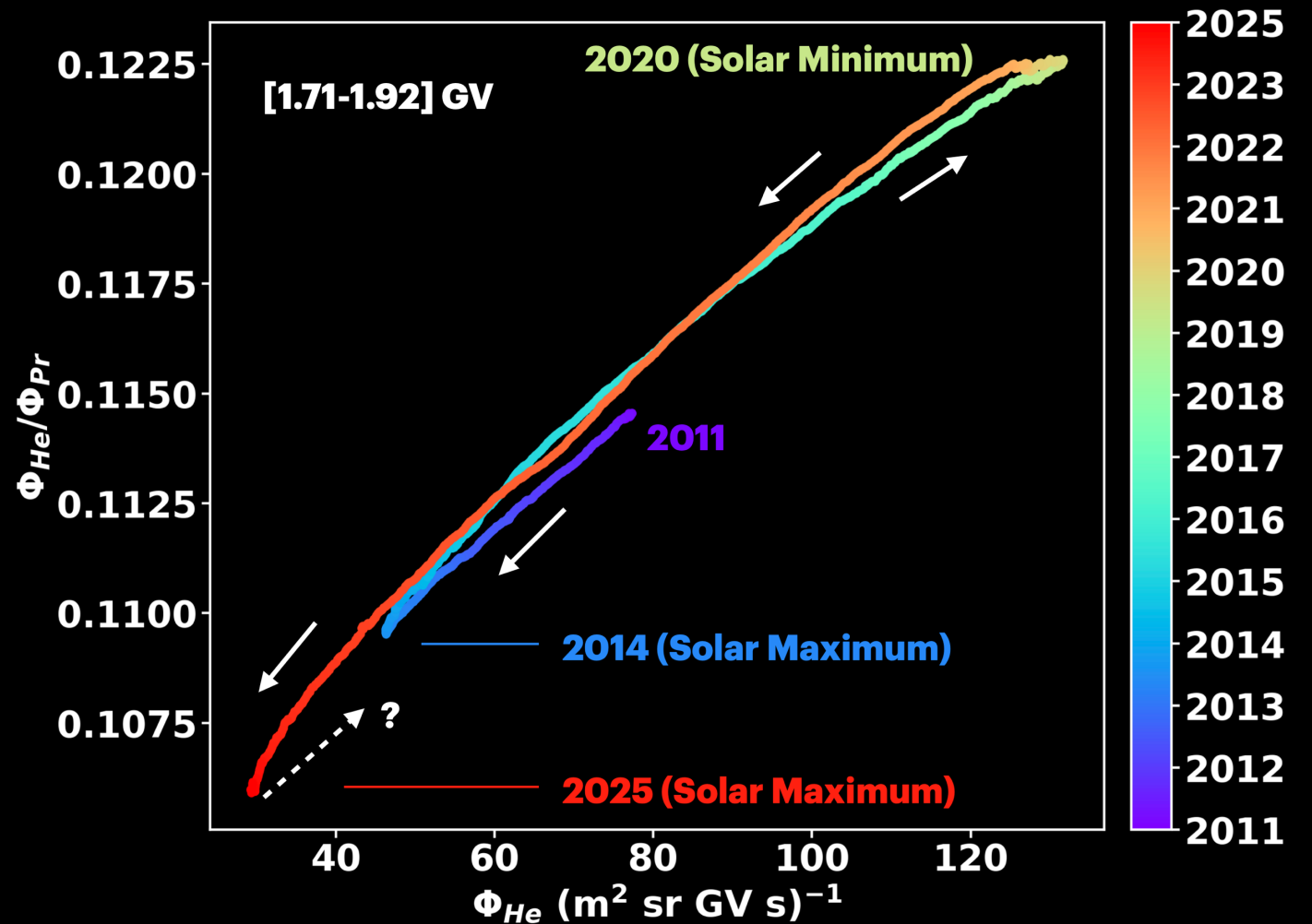
- Li, Be and B are significantly less modulated than He up to 3.6 GV.
- C, N and O are significantly less modulated than He up to 2.15 GV.



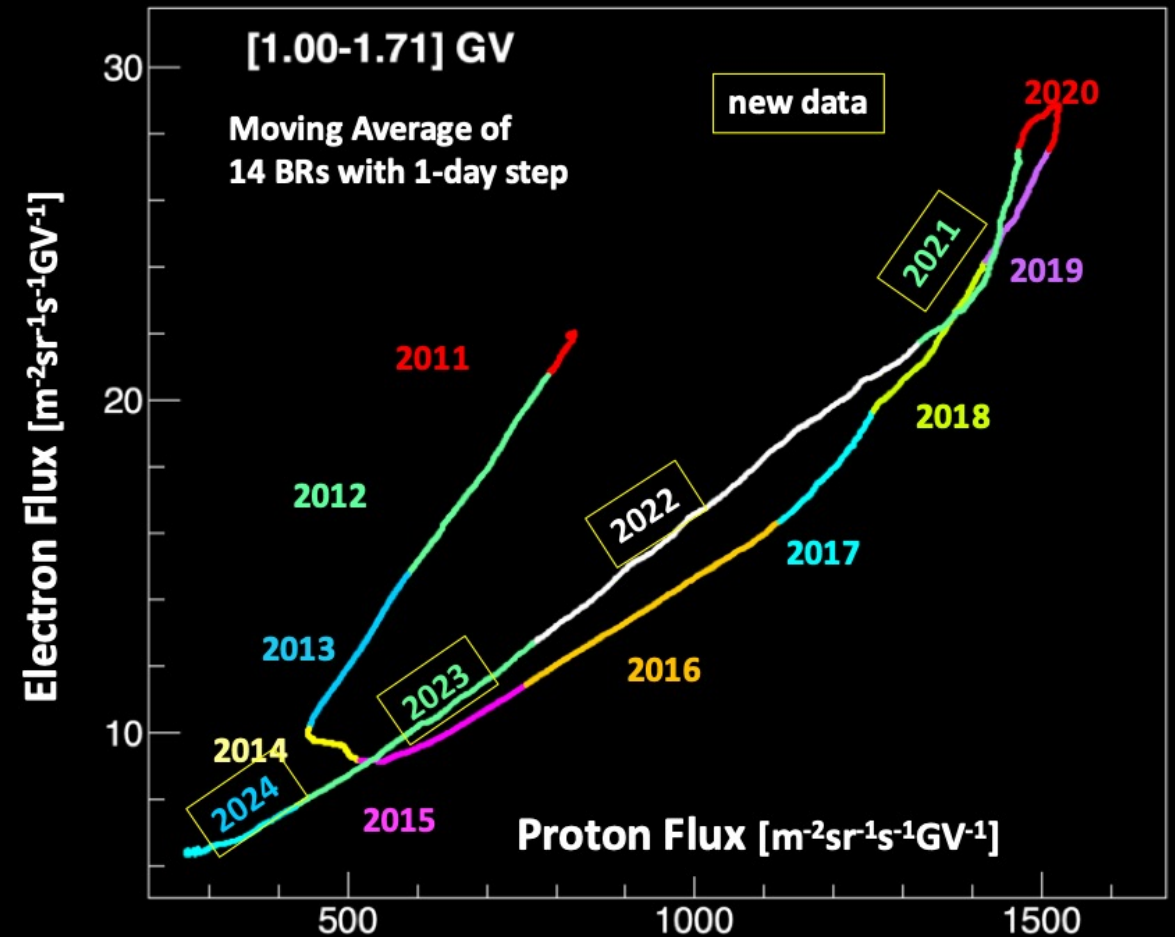
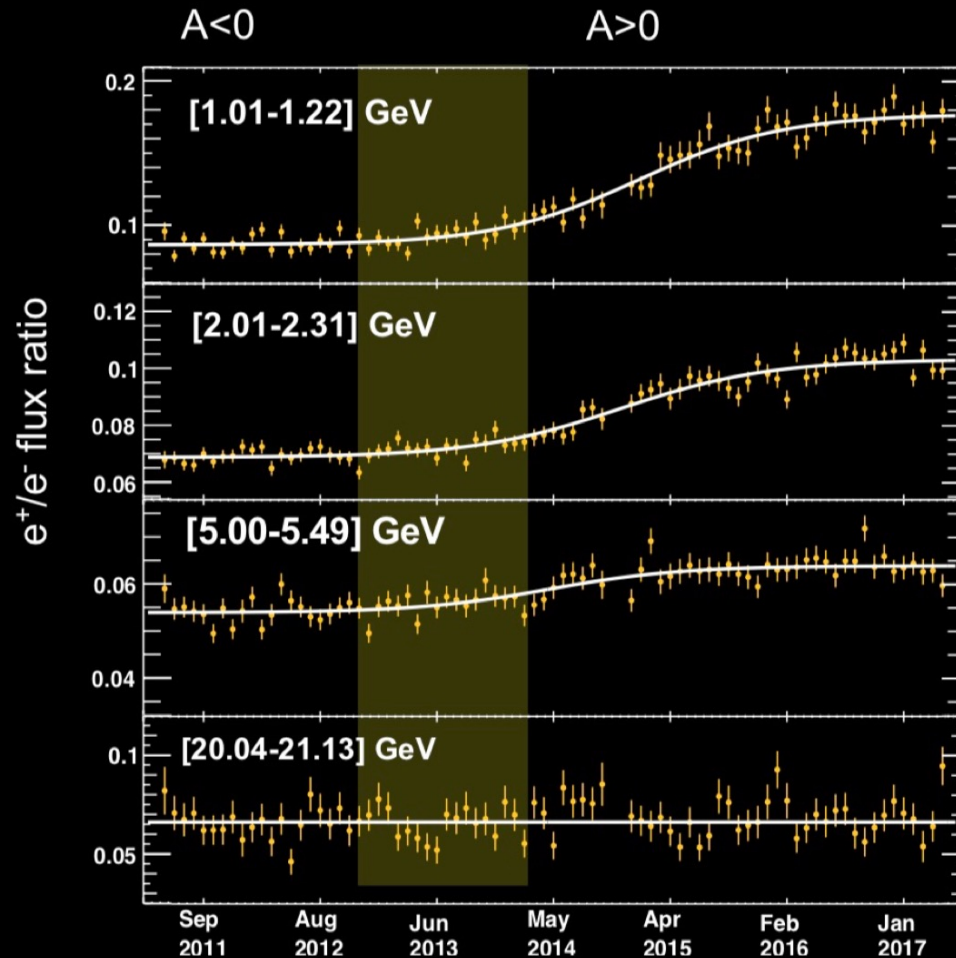
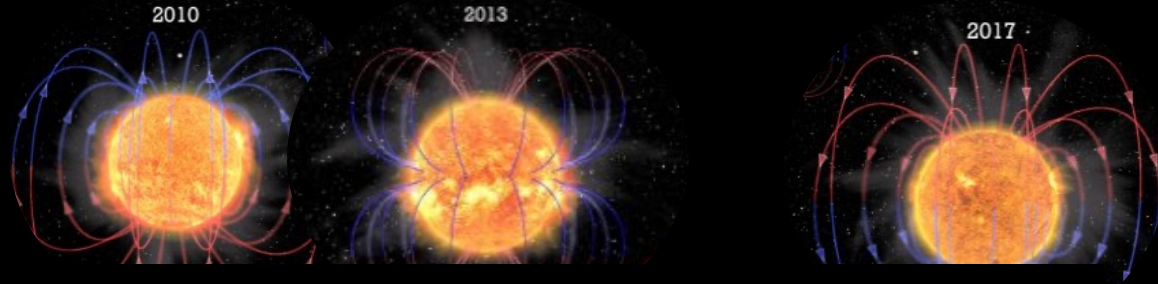
AMS observed that these differences in solar modulation are linearly correlated with the differences in the spectral indices Δ of the cosmic nuclei fluxes. This shows, in a model-independent way, that solar modulation of galactic cosmic nuclei depends on their spectral shape.

He/Pr Hysteresis

- **Correlation** between He flux and He/Pr ratio (moving average).
- At low rigidity, an **hysteresis** between proton and helium fluxes is observed around the **2014** solar maximum.
- During **2025** solar maximum, both **fluxes** are showing a **lower minimum** compared to 2014.

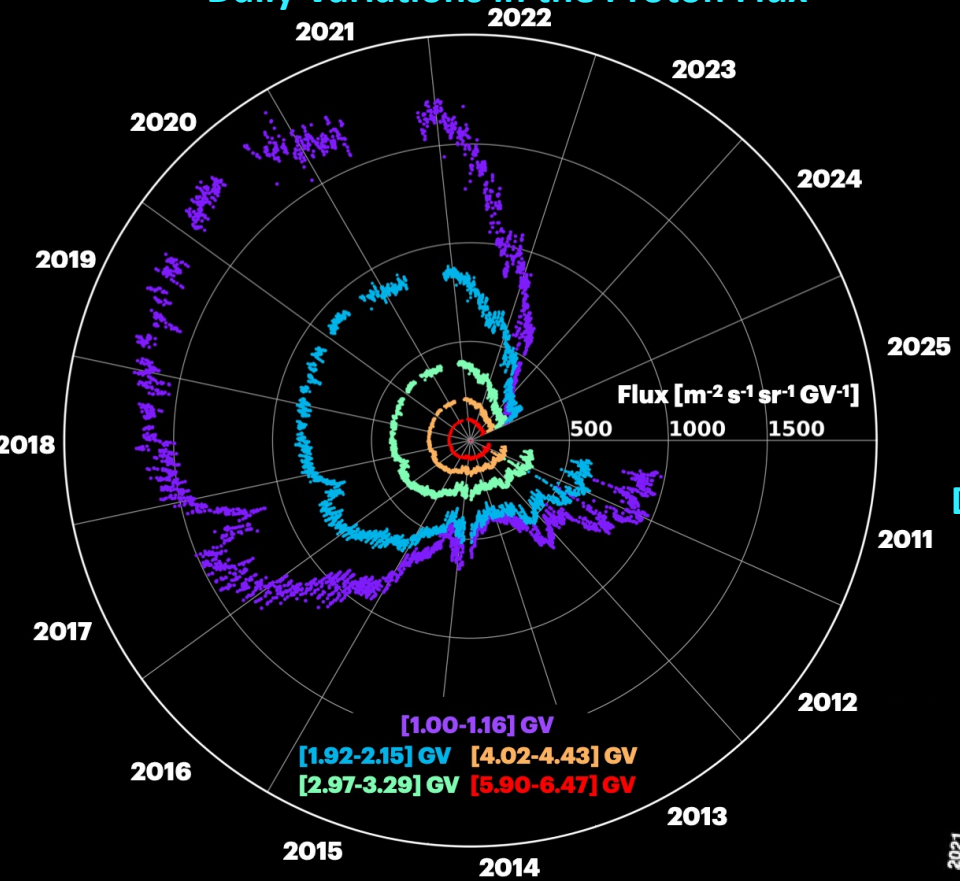


Long Term(> 1 Solar Cycle) Studies of Charge Sign Effect

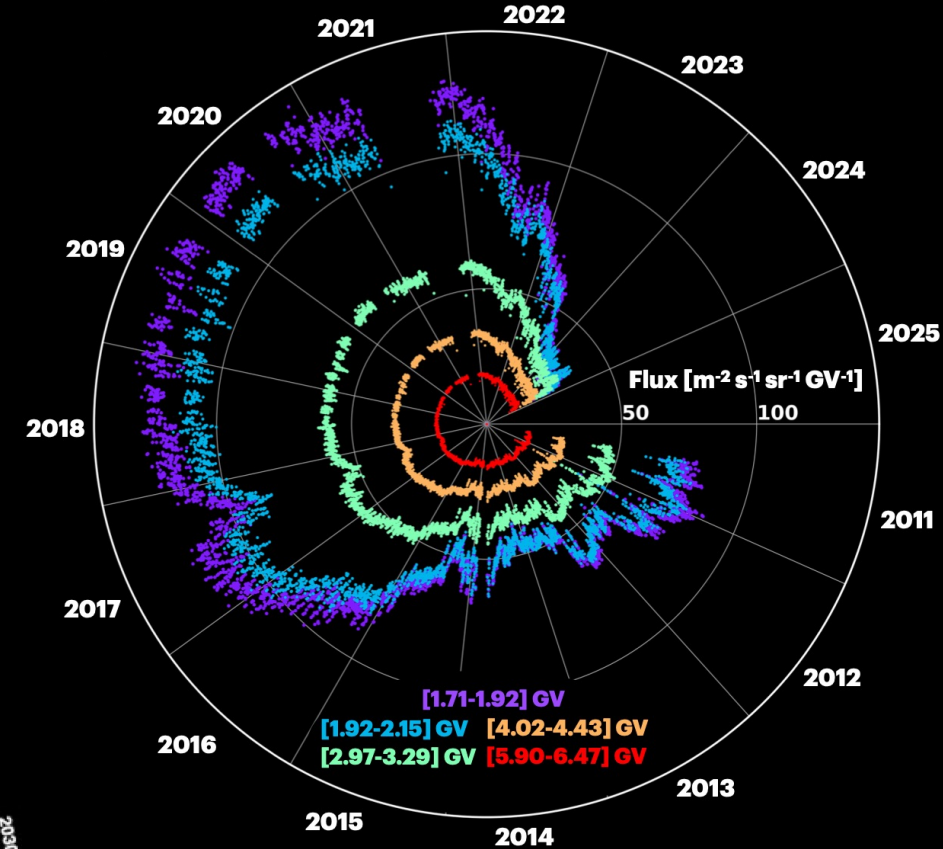


Long Term Studies (> 1 Solar Cycle) Daily Variations of Elementary Particles and light Nuclei

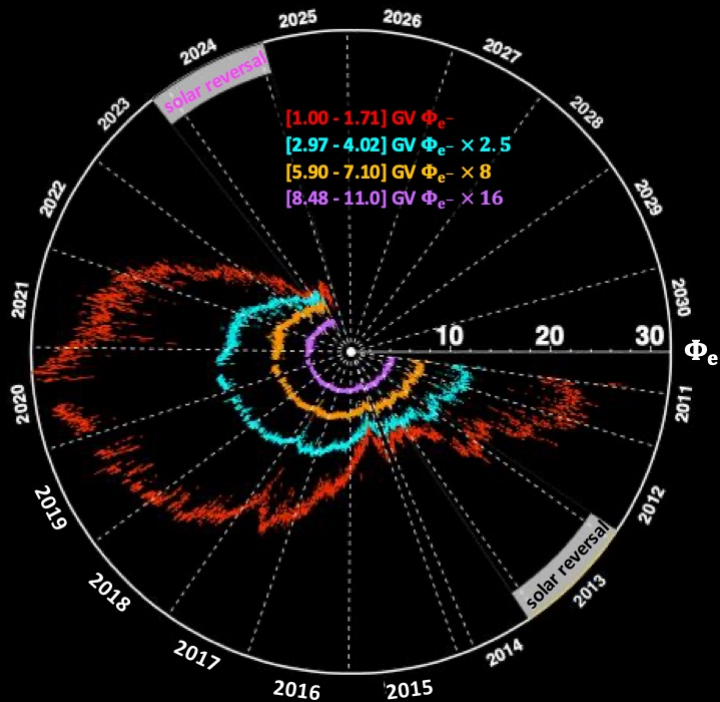
Daily Variations in the Proton Flux



Daily Variations in the Helium Flux



Daily Variations in the Electron Flux



**AMS is the only magnetic spectrometer in space.
The results from AMS are unlocking the secrets of the cosmos.
AMS will continue to take data for the ISS lifetime.**

