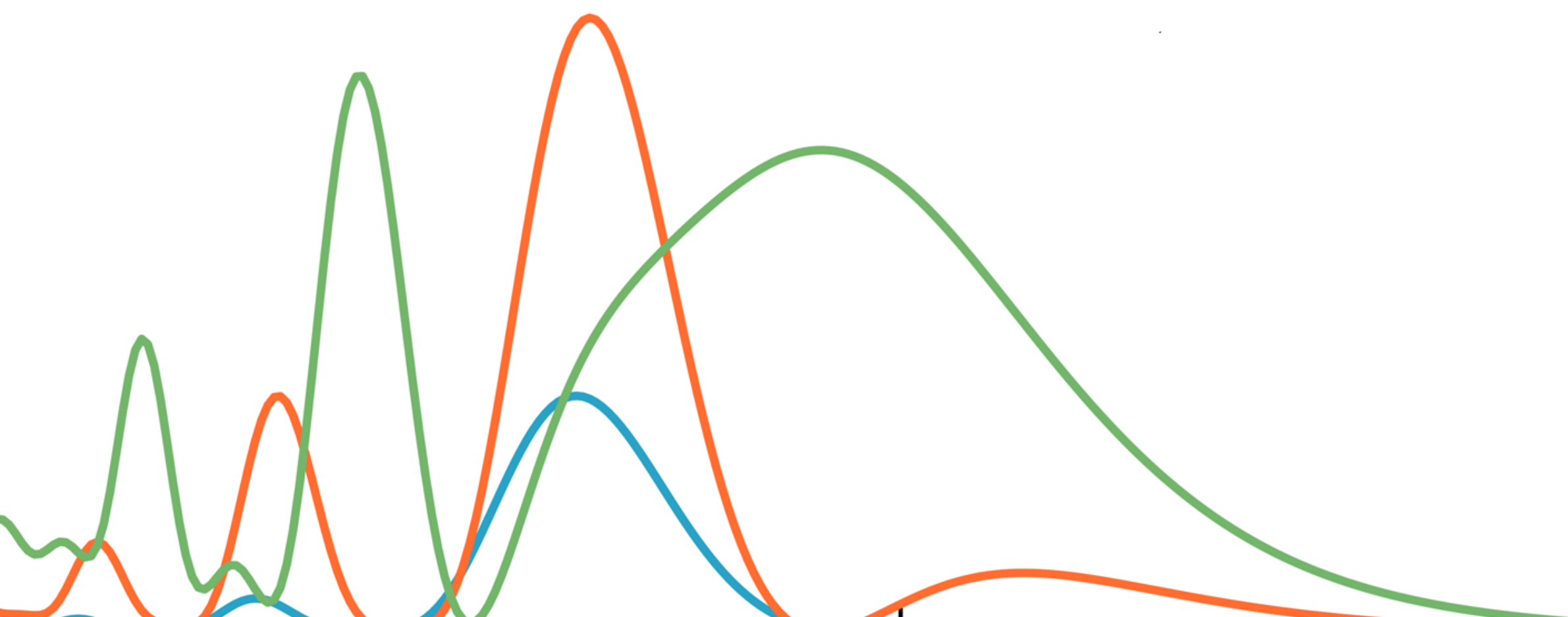


Review of neutrino oscillation experiments

Reactor / Accelerator / Atmospheric

TAUP 2025, Lukas Berns, Tohoku U



Lower left oscillogram from PRX 13, 041055 (2023)

ν -oscillation

For neutrinos **flavor basis** (interaction)
 \neq **Hamiltonian basis**. (propagation)

↓
 Flavor ($\nu_e \mid \nu_\mu \mid \nu_\tau$)
 oscillates over $L \times \Delta m^2 / E$,
 amplitude controlled by
 (PMNS) mixing matrix U :

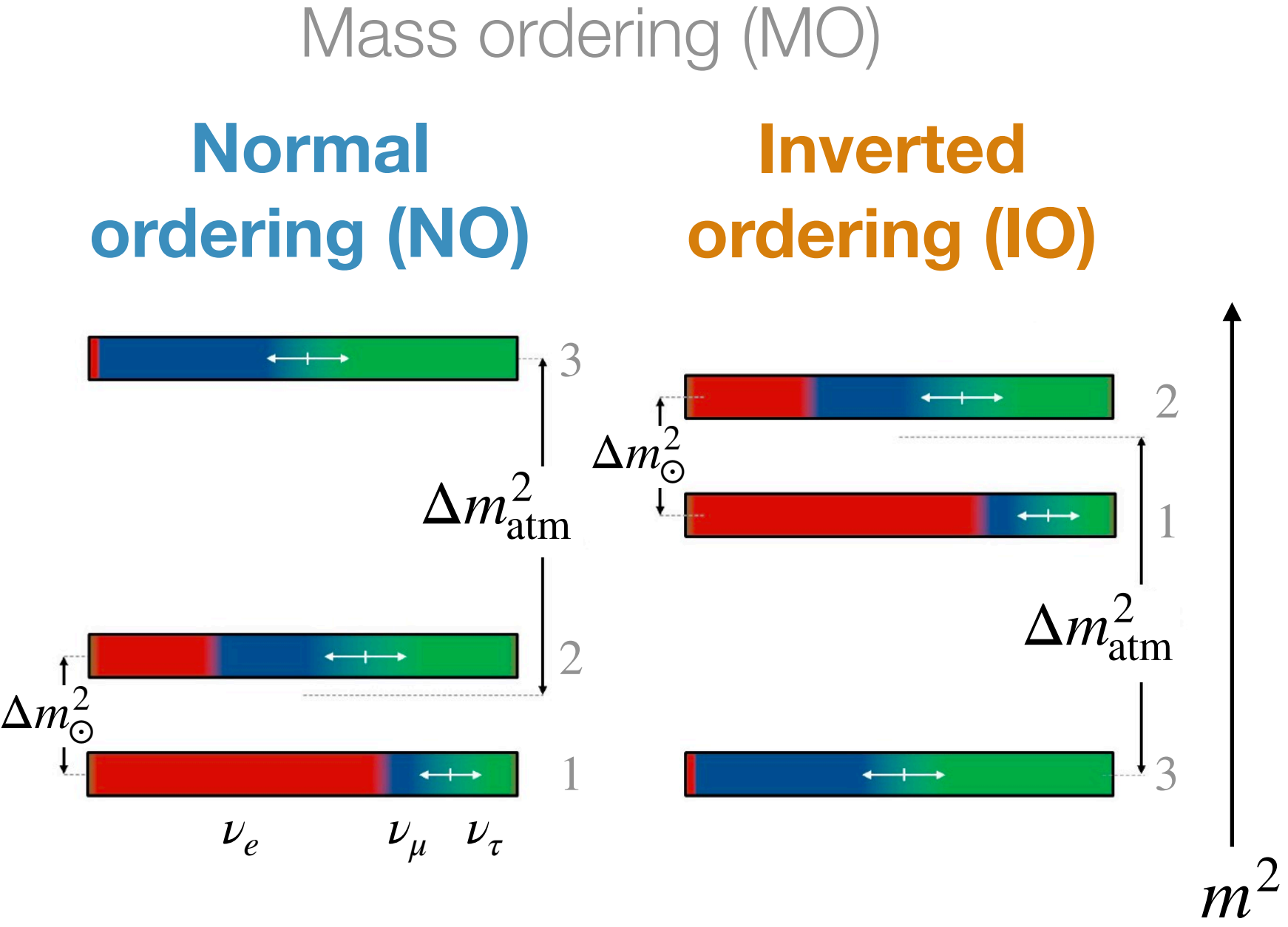
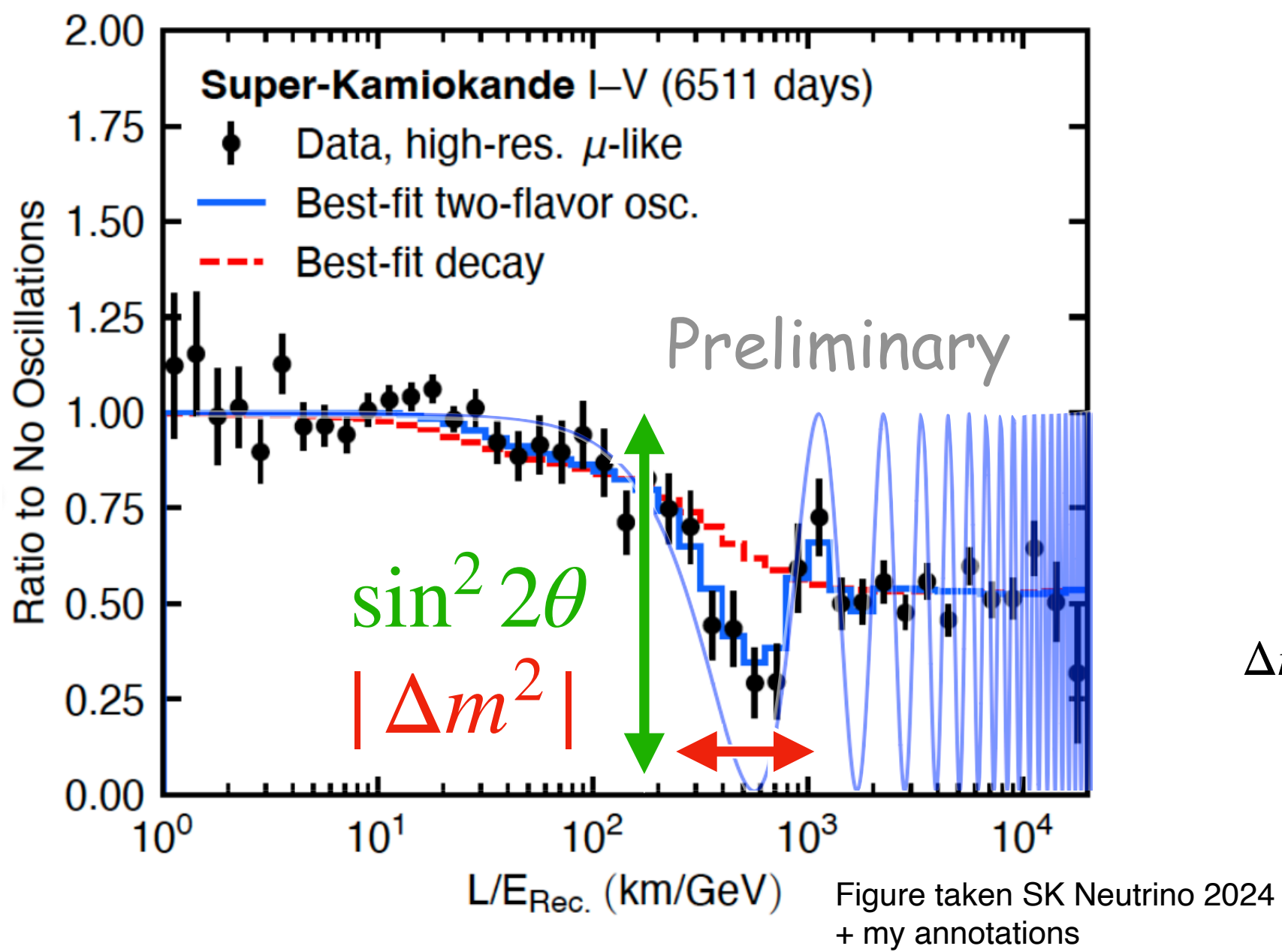
θ_{23}
atmospheric

θ_{13}, δ_{CP}
reactor, accelerator

θ_{12}
solar

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$c_{ij} \equiv \cos \theta_{ij}$
 $s_{ij} \equiv \sin \theta_{ij}$



ν -oscillation

$$U = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}}_{\substack{\theta_{23} \\ \text{atmospheric}}} \underbrace{\begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix}}_{\substack{\theta_{13}, \delta_{CP} \\ \text{reactor, accelerator}}} \underbrace{\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\substack{\theta_{12} \\ \text{solar}}}$$

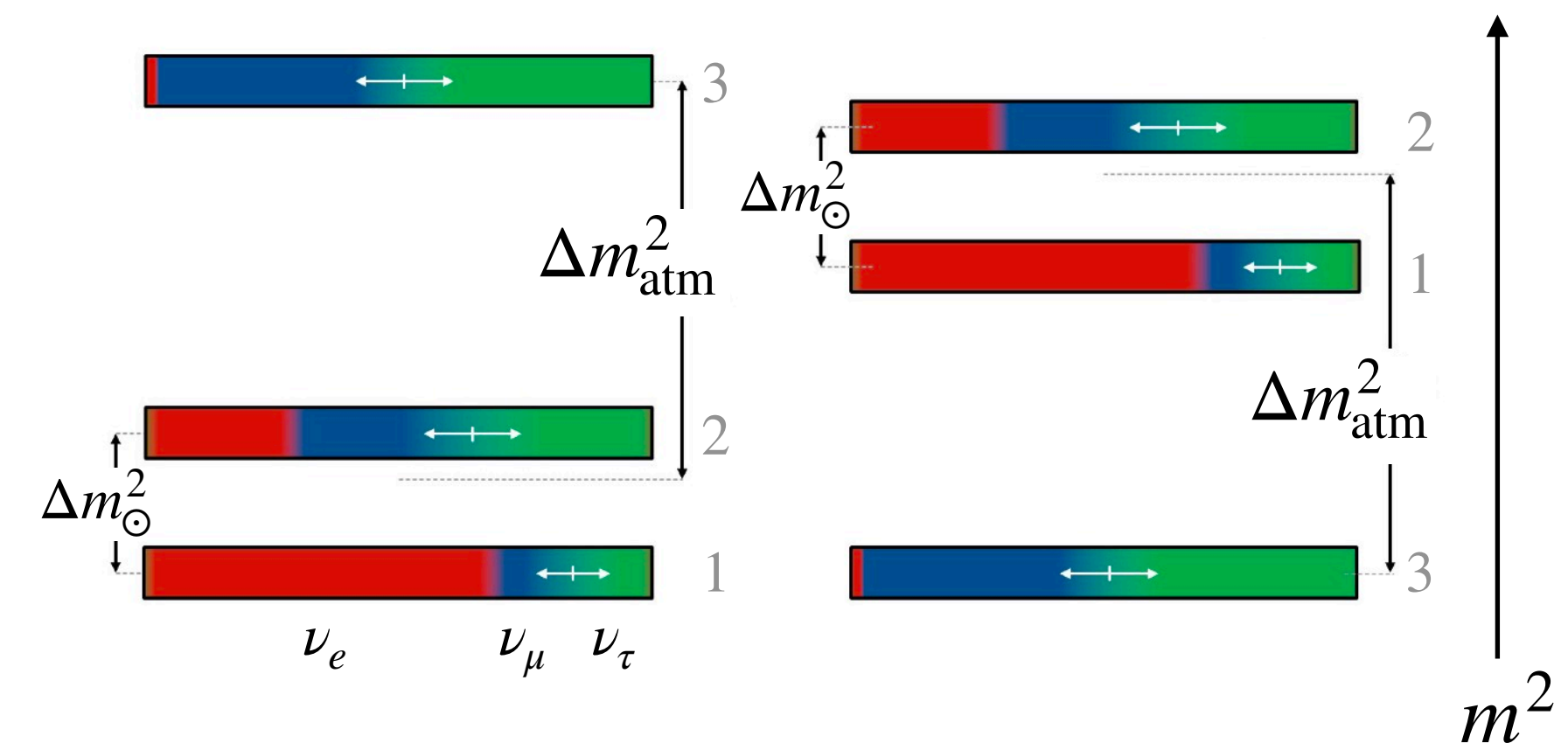
$c_{ij} \equiv \cos \theta_{ij}$
 $s_{ij} \equiv \sin \theta_{ij}$

?

Mass ordering (MO)

**Normal
ordering (NO)**

**Inverted
ordering (IO)**



Open questions:

- value of $\delta_{CP} \rightarrow$ if $\sin \delta_{CP} \neq 0$, CP violation

ν -oscillation

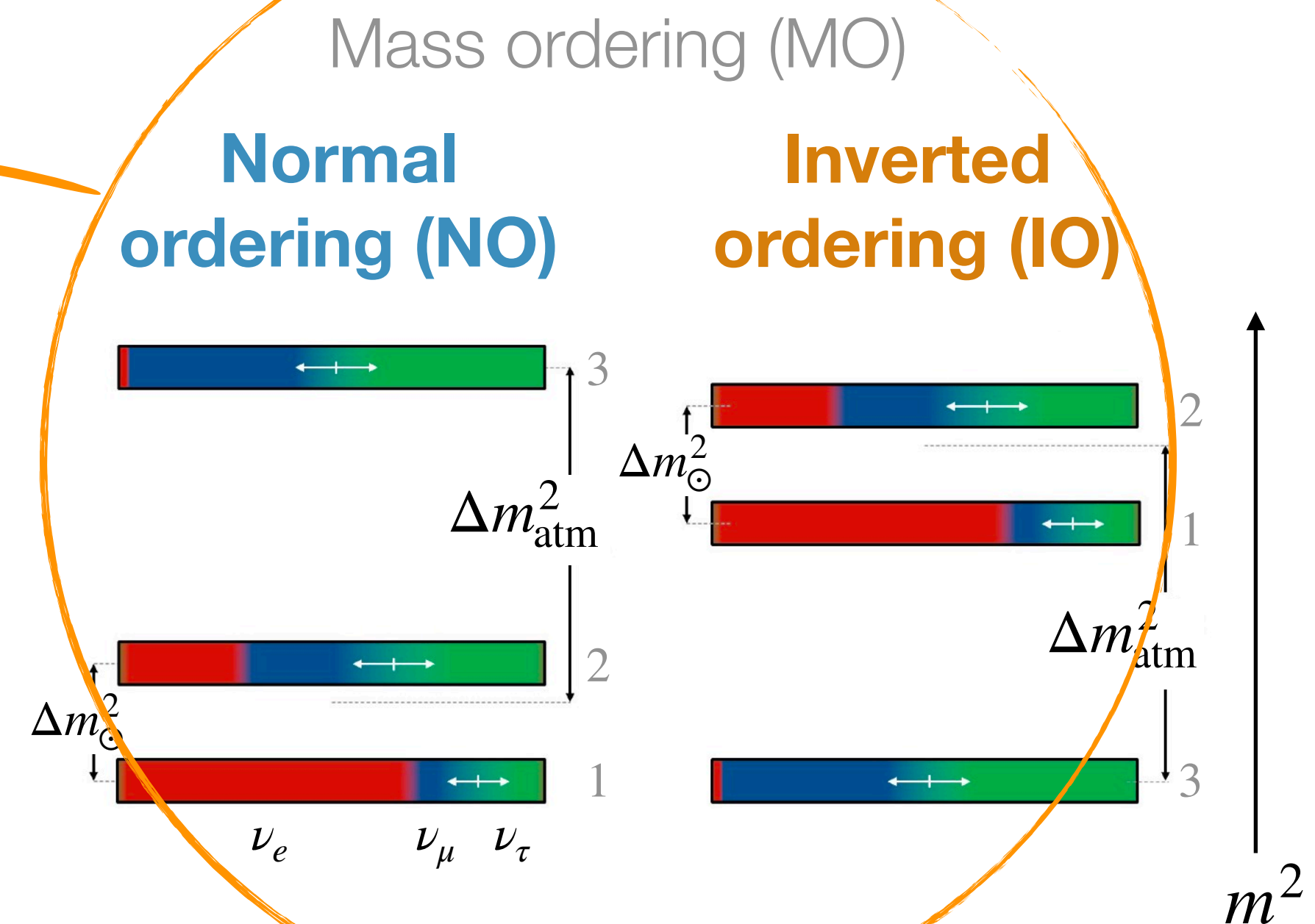
$$U = \begin{array}{c} \theta_{23} \\ \text{atmospheric} \end{array} \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{array}{c} \theta_{13}, \delta_{CP} \\ \text{reactor, accelerator} \end{array} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix} \begin{array}{c} \theta_{12} \\ \text{solar} \end{array} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$c_{ij} \equiv \cos \theta_{ij}$
 $s_{ij} \equiv \sin \theta_{ij}$

Open questions:

- value of δ_{CP} \rightarrow if $\sin \delta_{CP} \neq 0$, CP violation
- sign of Δm_{32}^2 (mass ordering)

Important for
cosmology (leptogenesis...),
0νββ searches



ν -oscillation

$$U = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}}_{\text{atmospheric } \theta_{23}} \underbrace{\begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix}}_{\text{reactor, accelerator } \theta_{13}, \delta_{CP}} \underbrace{\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{solar } \theta_{12}}$$

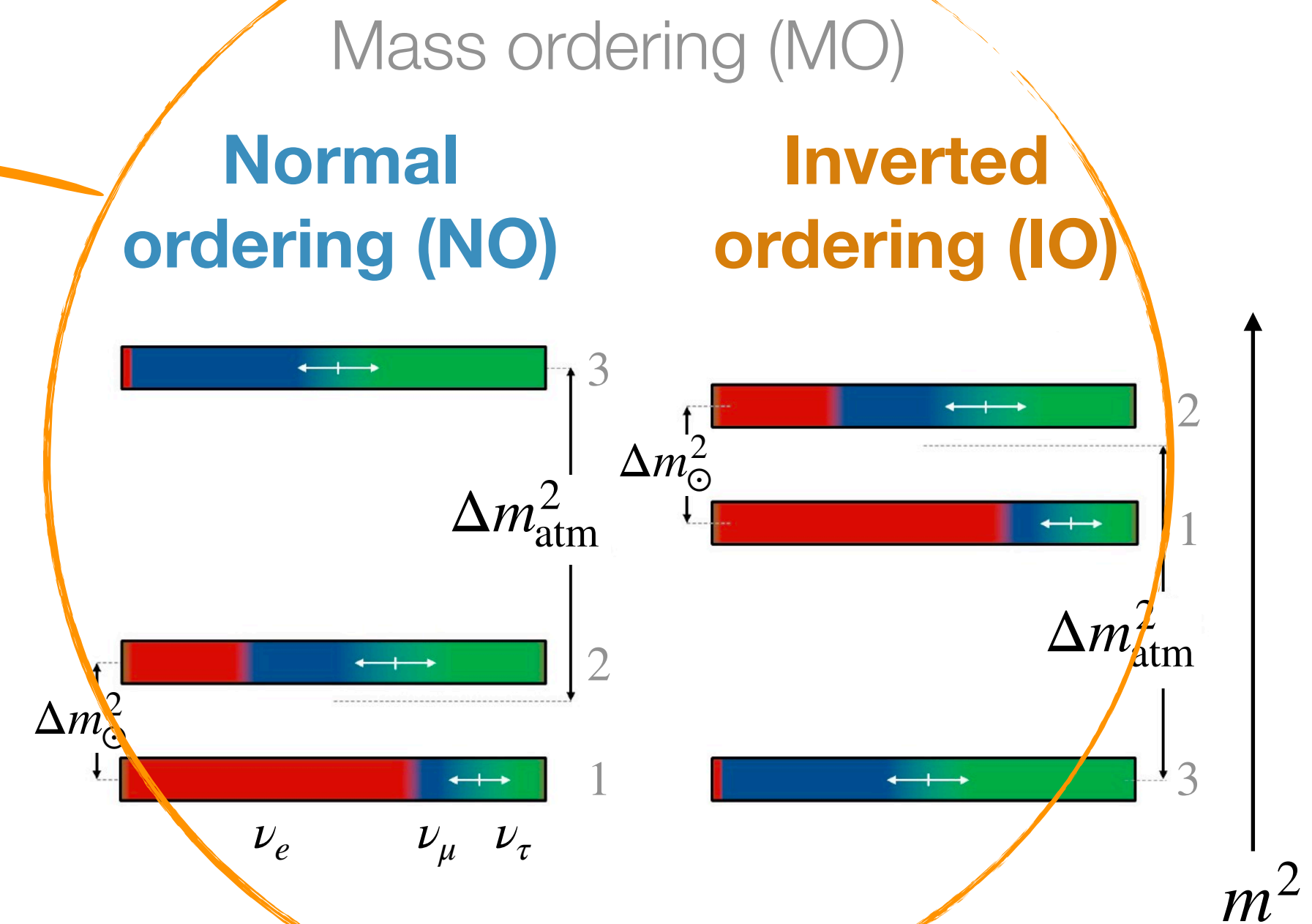
$c_{ij} \equiv \cos \theta_{ij}$
 $s_{ij} \equiv \sin \theta_{ij}$

Open questions:

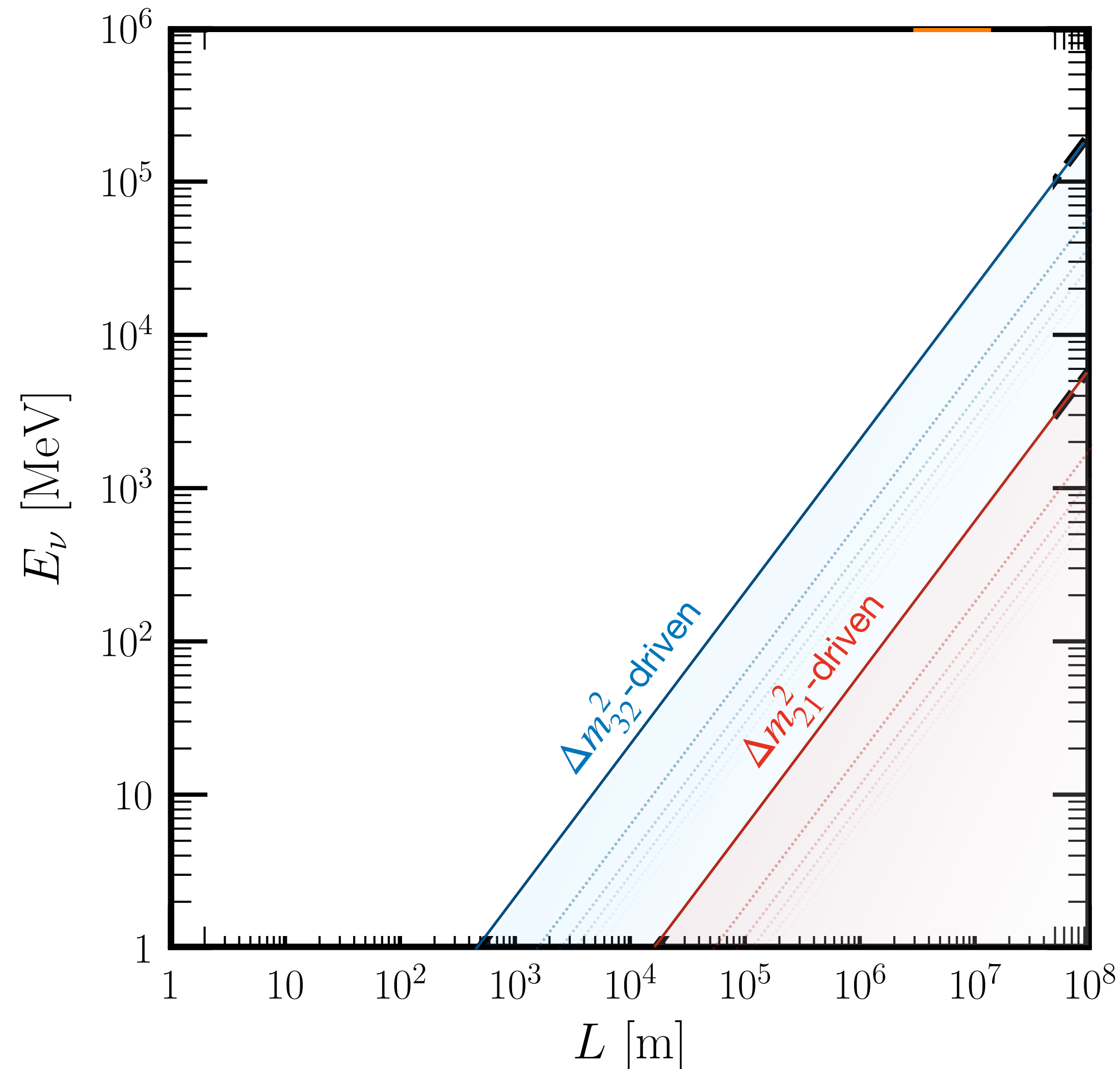
- value of $\delta_{CP} \rightarrow$ if $\sin \delta_{CP} \neq 0$, CP violation
- sign of Δm_{32}^2 (mass ordering)
- Is $\theta_{23} = \frac{\pi}{4}$? $< \frac{\pi}{4}$? $> \frac{\pi}{4}$? (octant)

Important for
cosmology (leptogenesis...),
Ov2 β searches

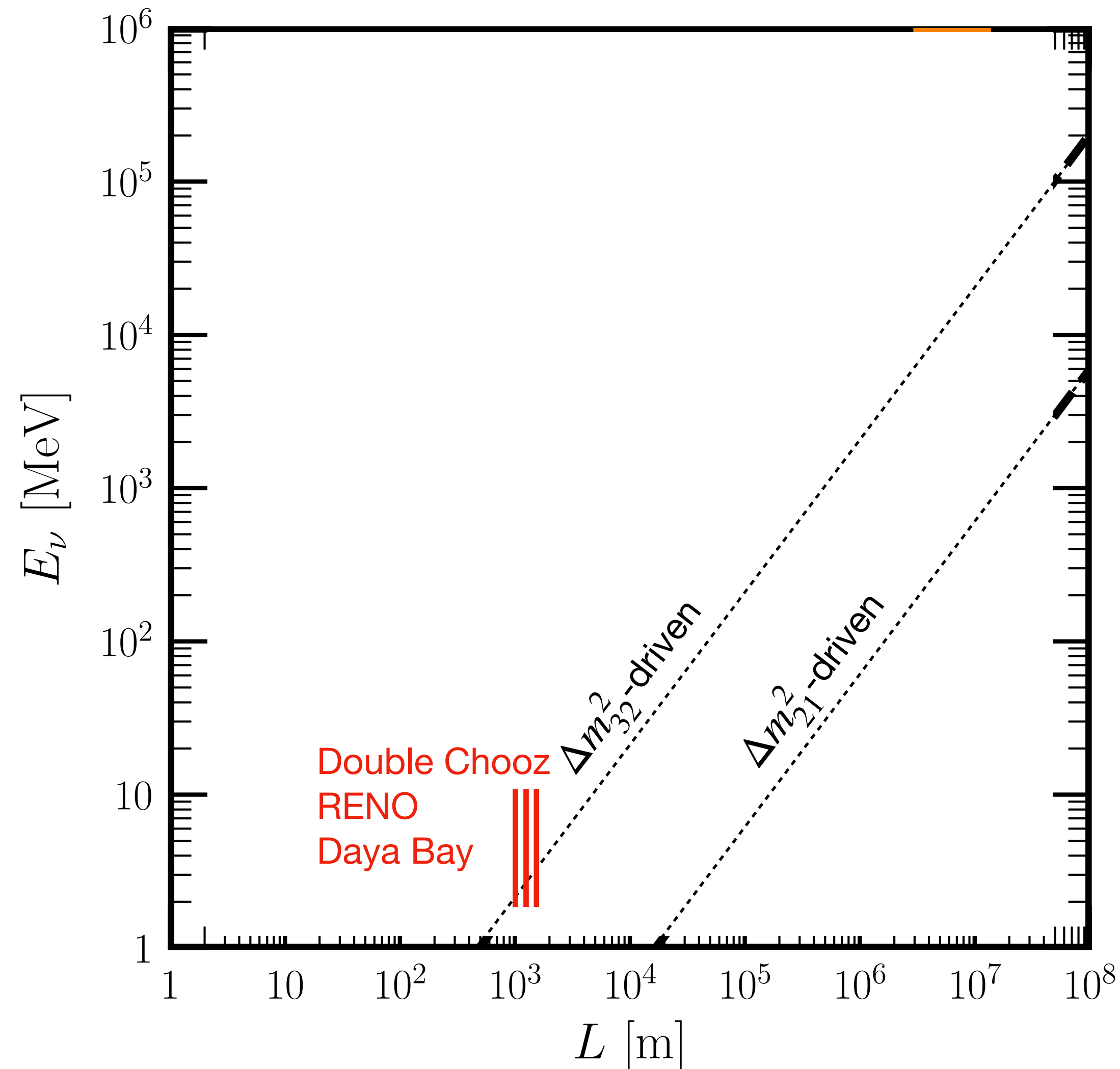
Flavor symmetries?



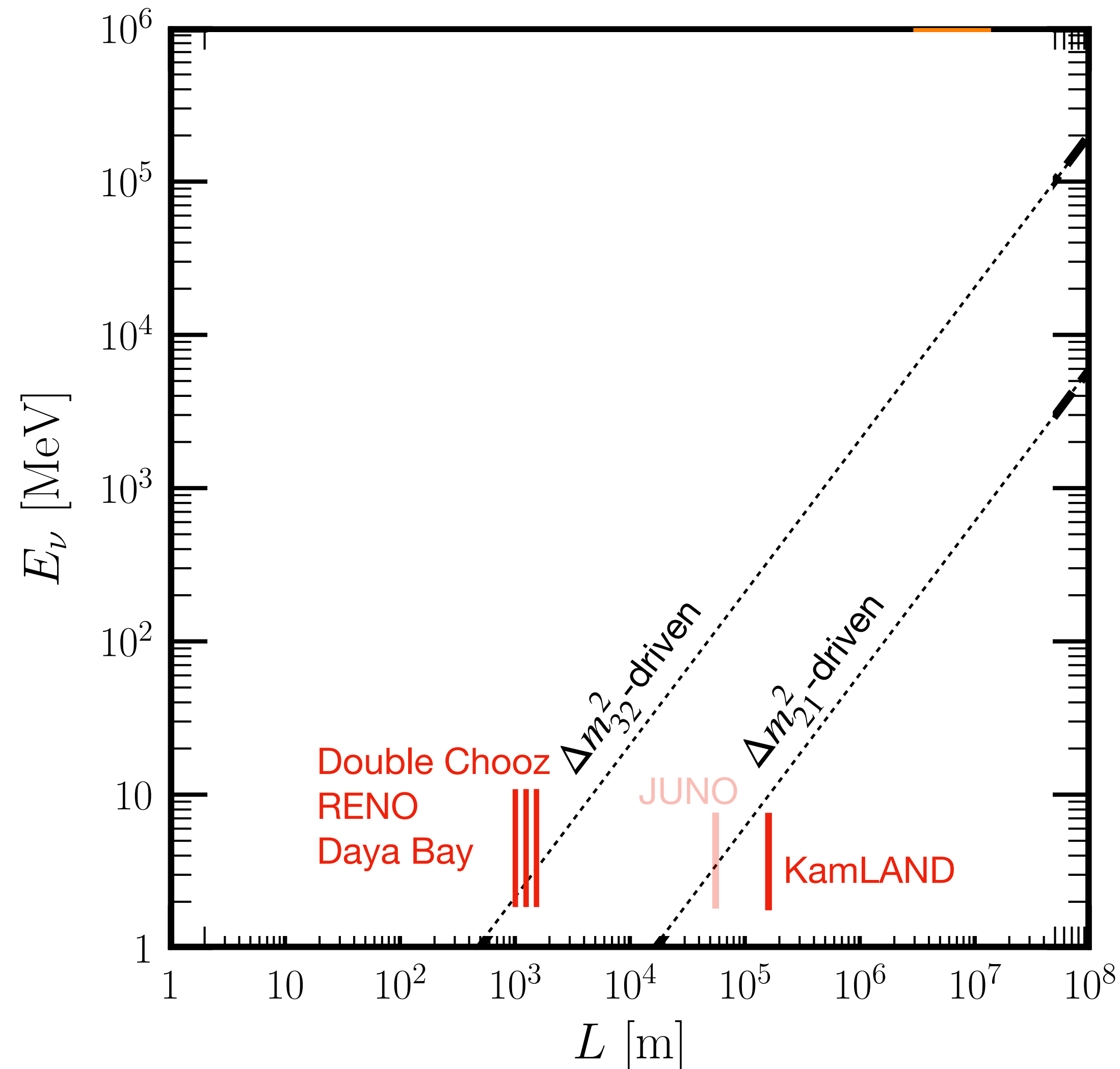
Current (3-flavor) oscillation experiments



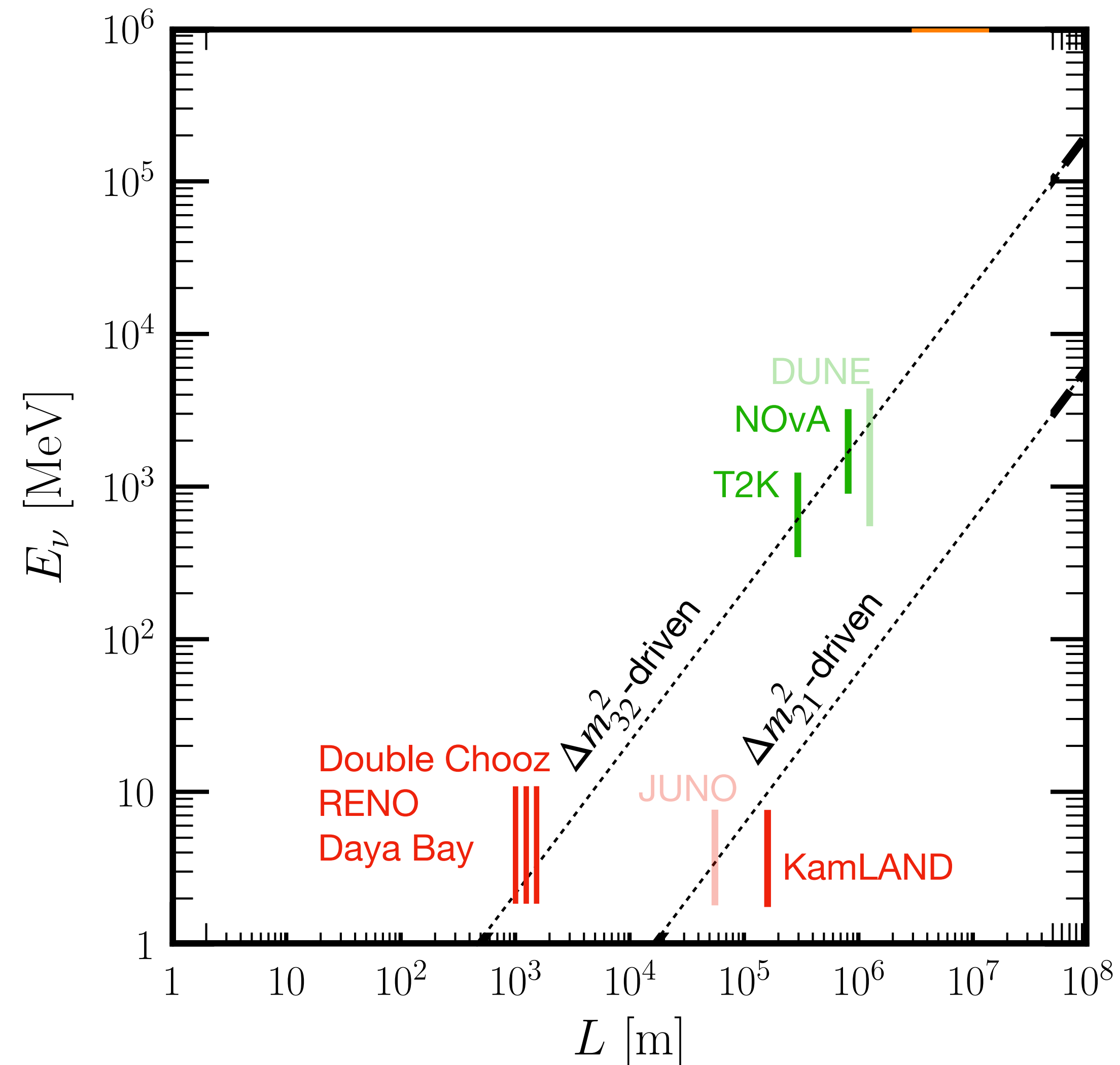
Current (3-flavor) oscillation experiments



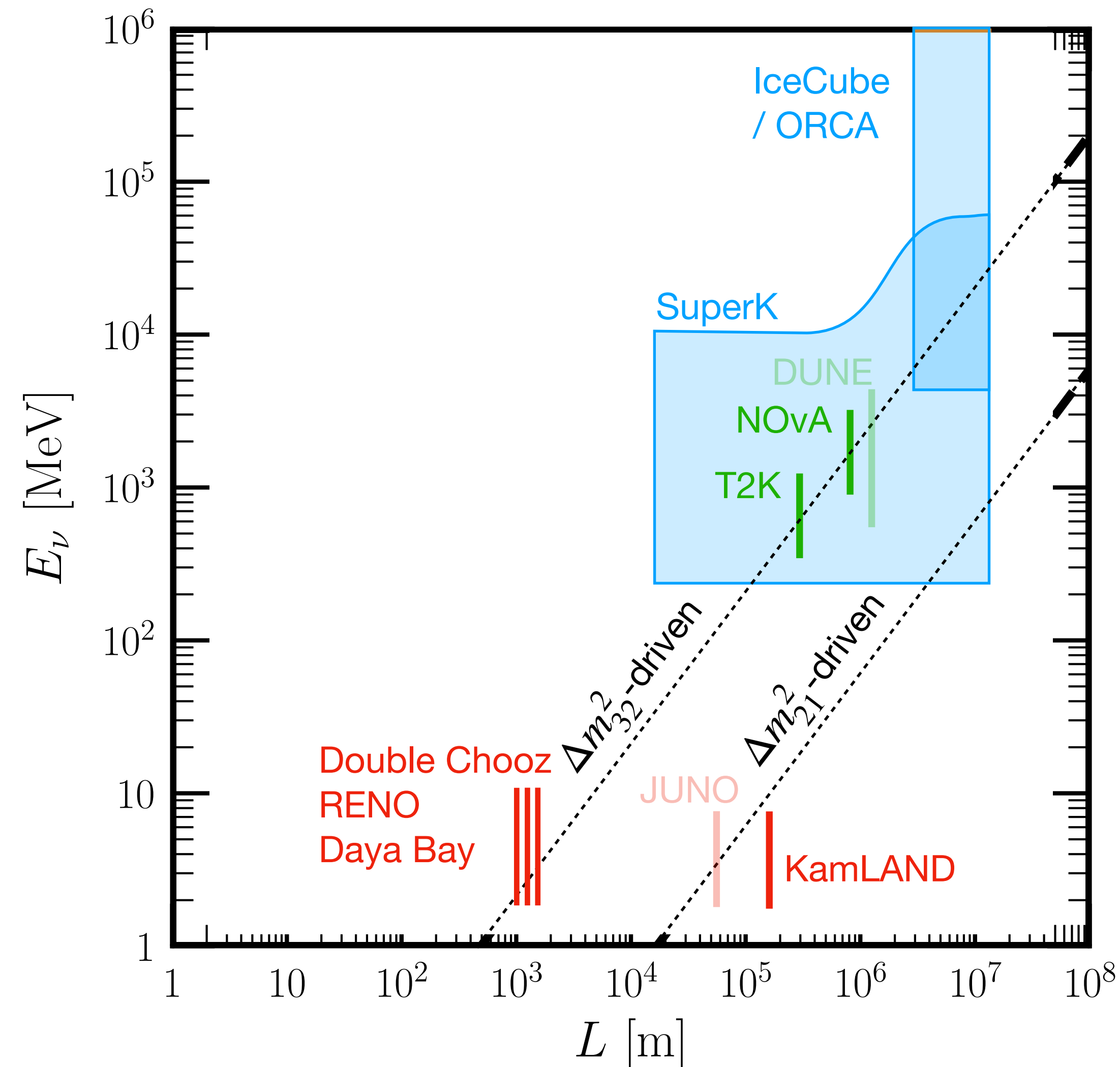
Current (3-flavor) oscillation experiments



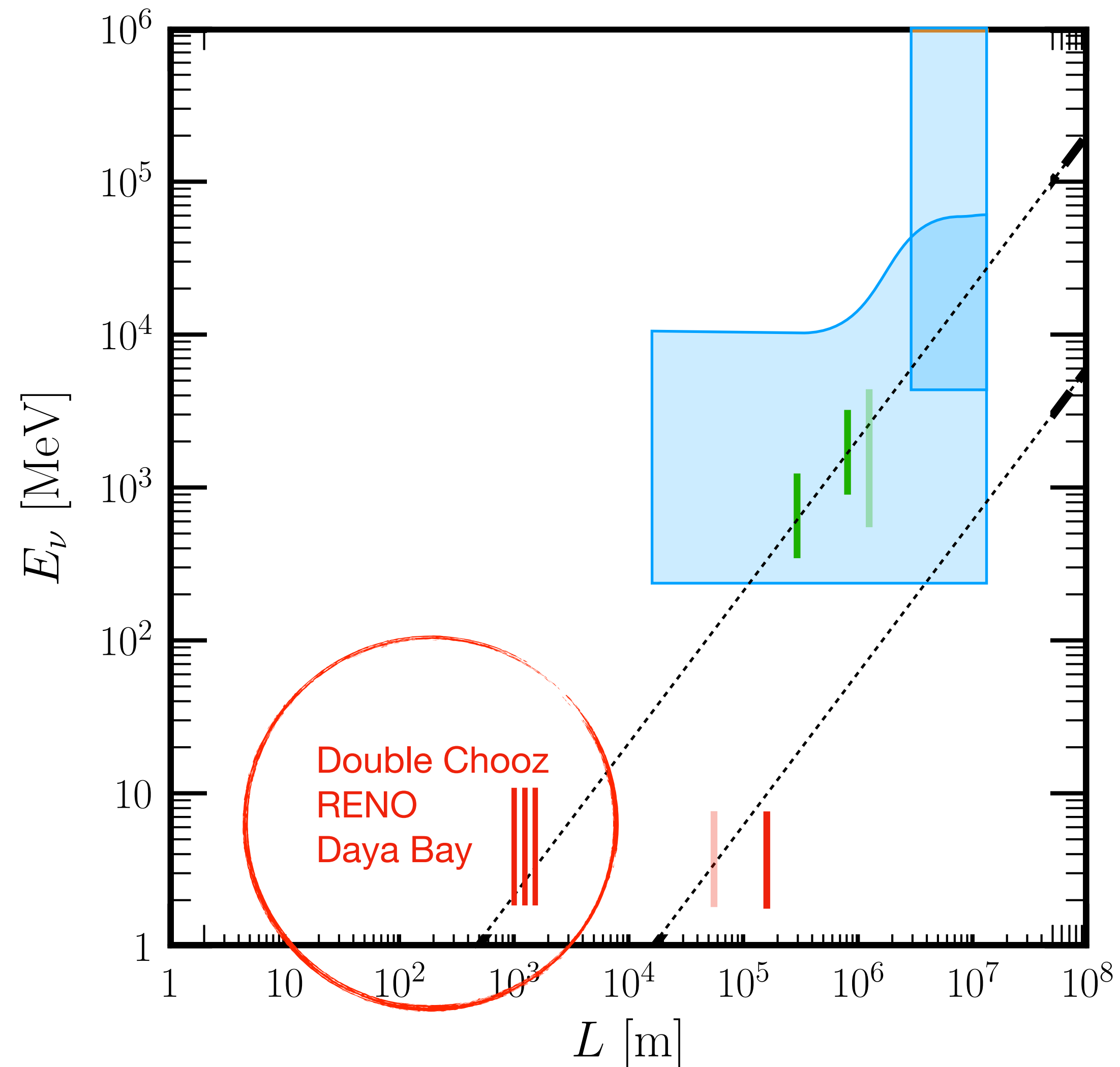
Current (3-flavor) oscillation experiments



Current (3-flavor) oscillation experiments

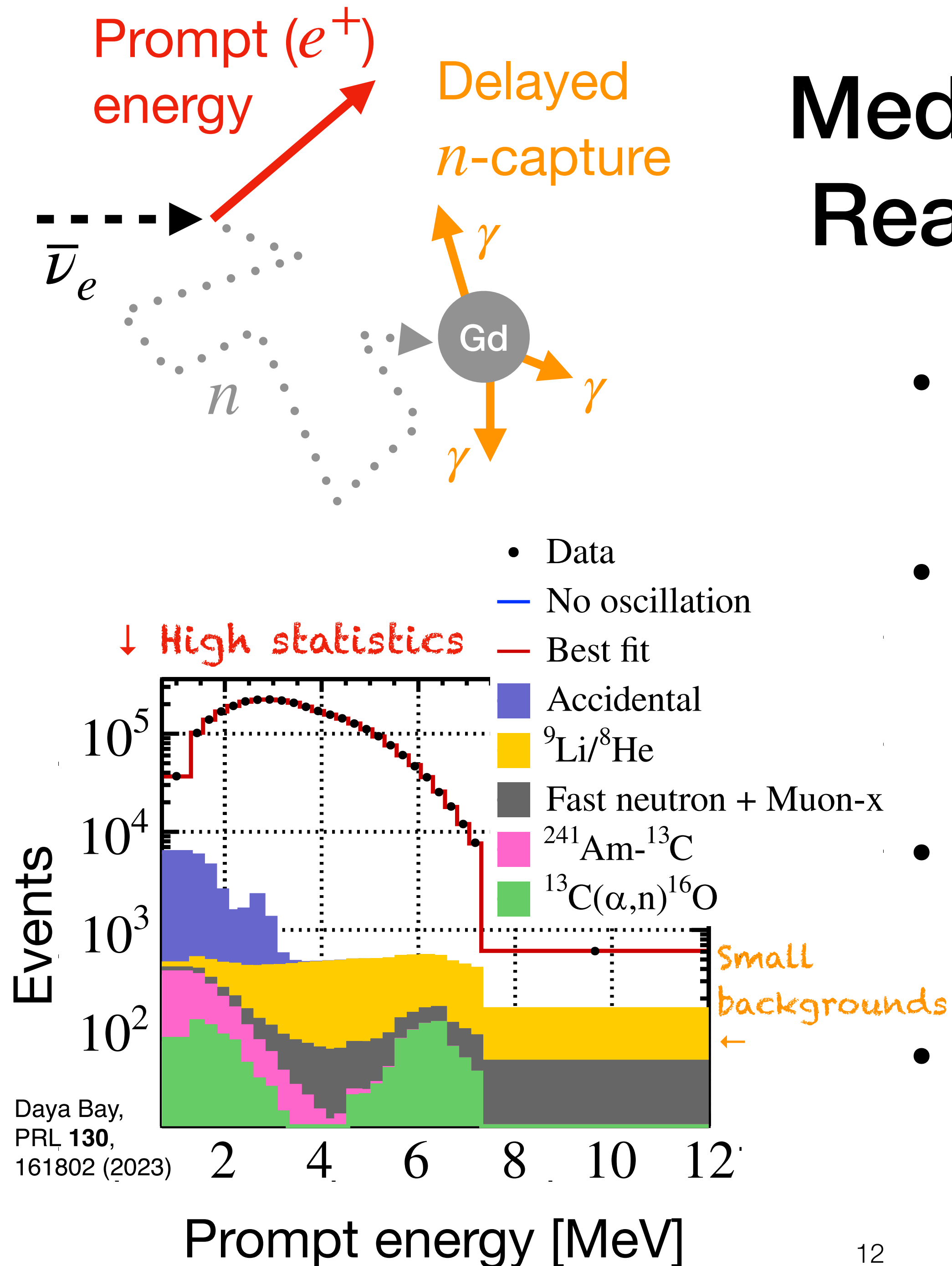


Medium Baseline (km) Reactor Experiments



Assuming typical densities for terrestrial exp.
Plot inspired by PRD 101, 033008 (2020)

Medium Baseline (km) Reactor Experiments



- Intense anti-neutrino flux from reactor beta decay chain
- Detection in liquid scintillator via inverse-beta decay with delayed coincidence $\bar{\nu}_e + p \rightarrow e^+ + n$
- Neutrino energy inferred from e^+ **energy** deposit in LS
- **Delayed γ** from neutron capture on Gd (main) or H (sub) for significant background reduction

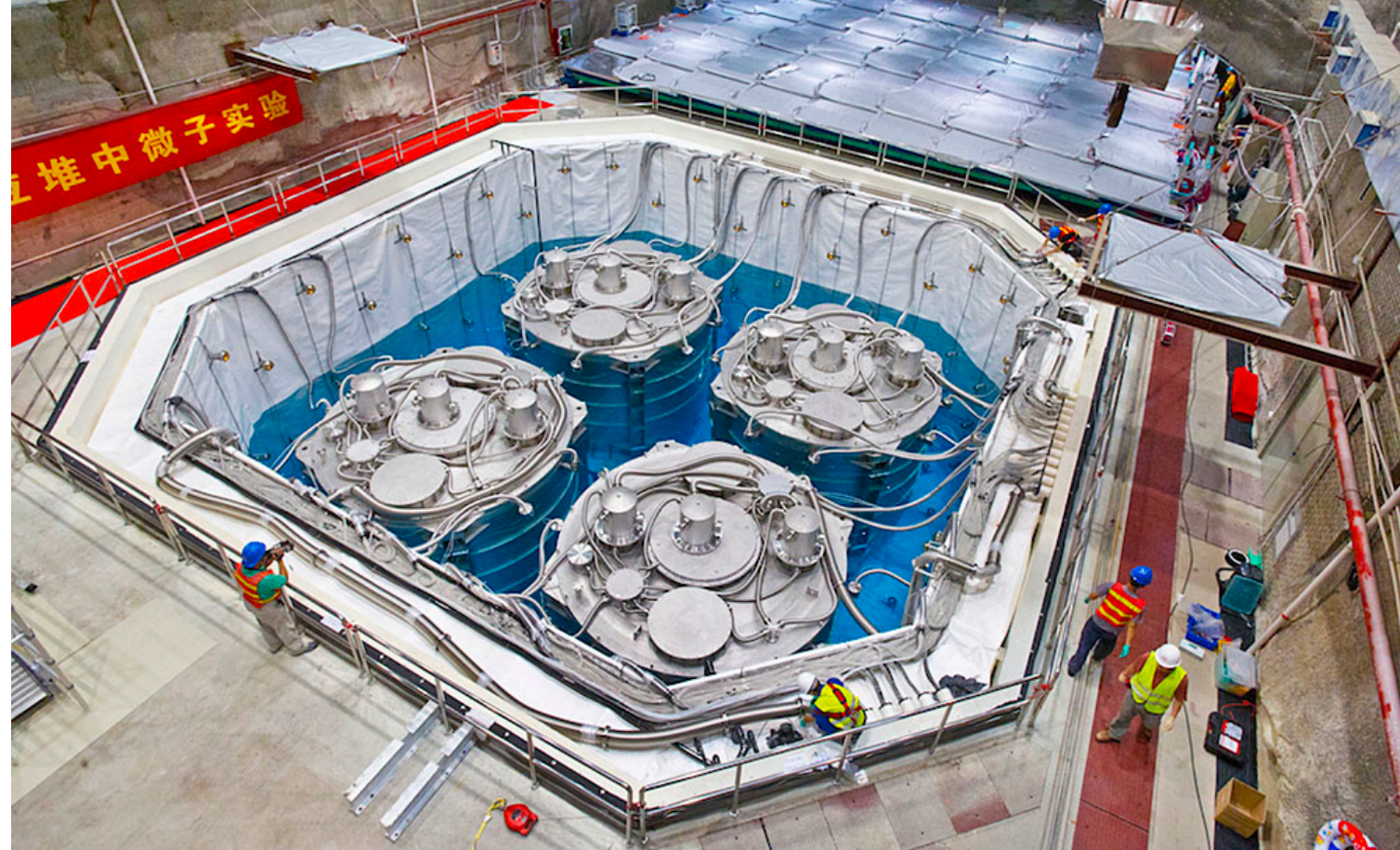
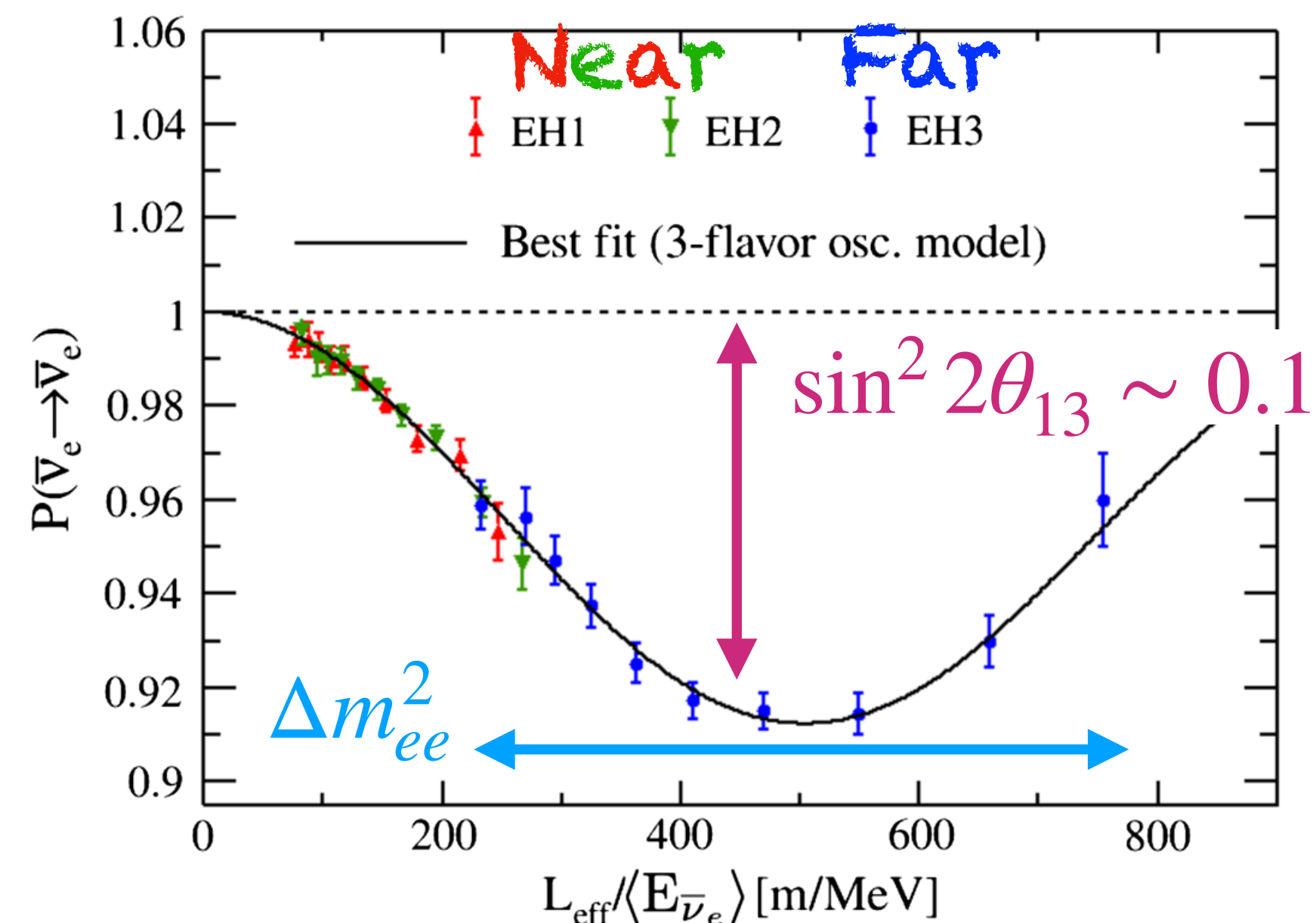


Photo: Roy Kaltschmidt, Berkeley Lab

Medium Baseline (km) Reactor Experiments

- 3 experiments with similar design
Double Chooz (France),
Daya Bay (China),
RENO (Korea)

DayaBay (PRL.130.161802)



Very clear oscillation signature

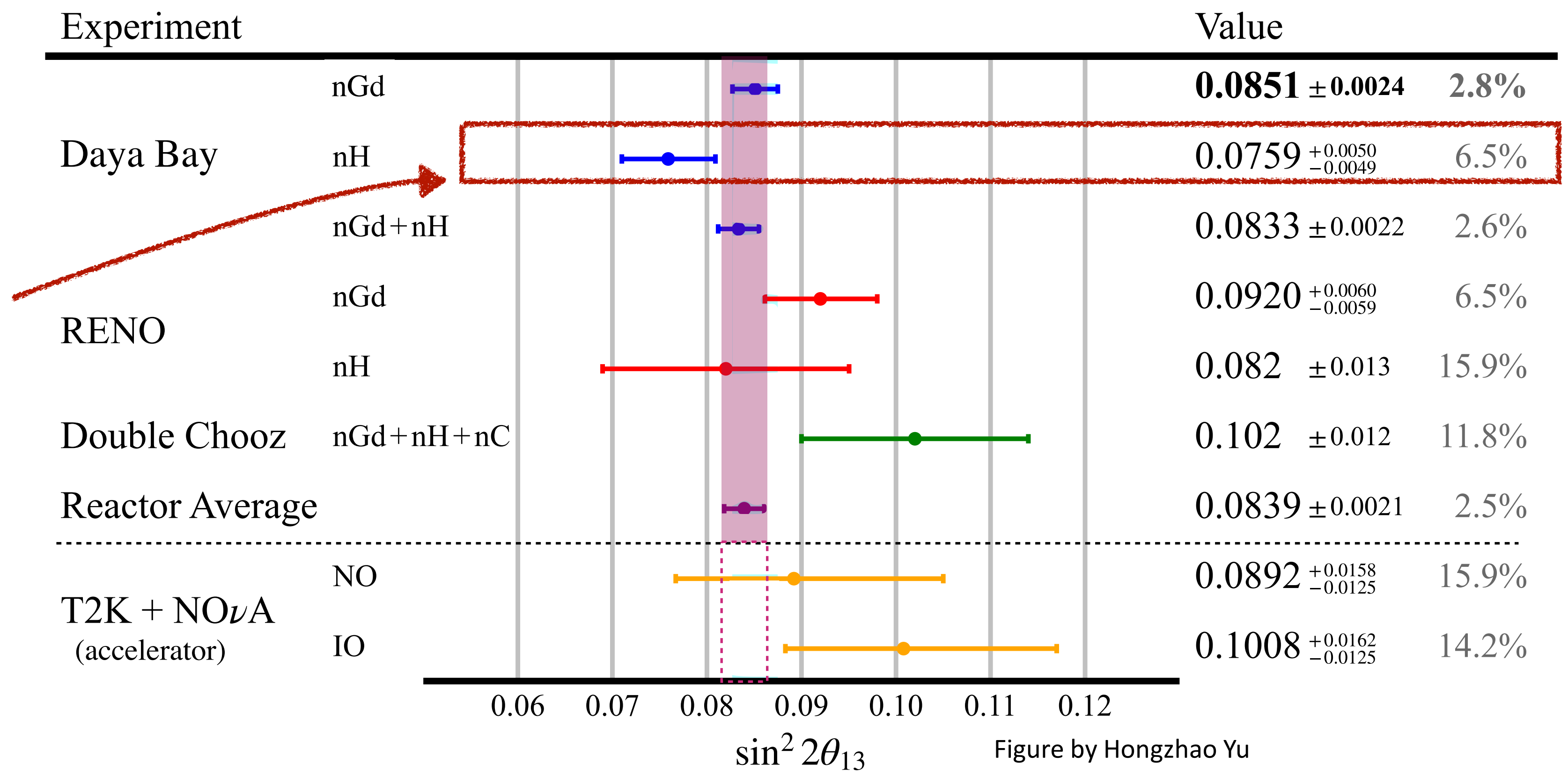
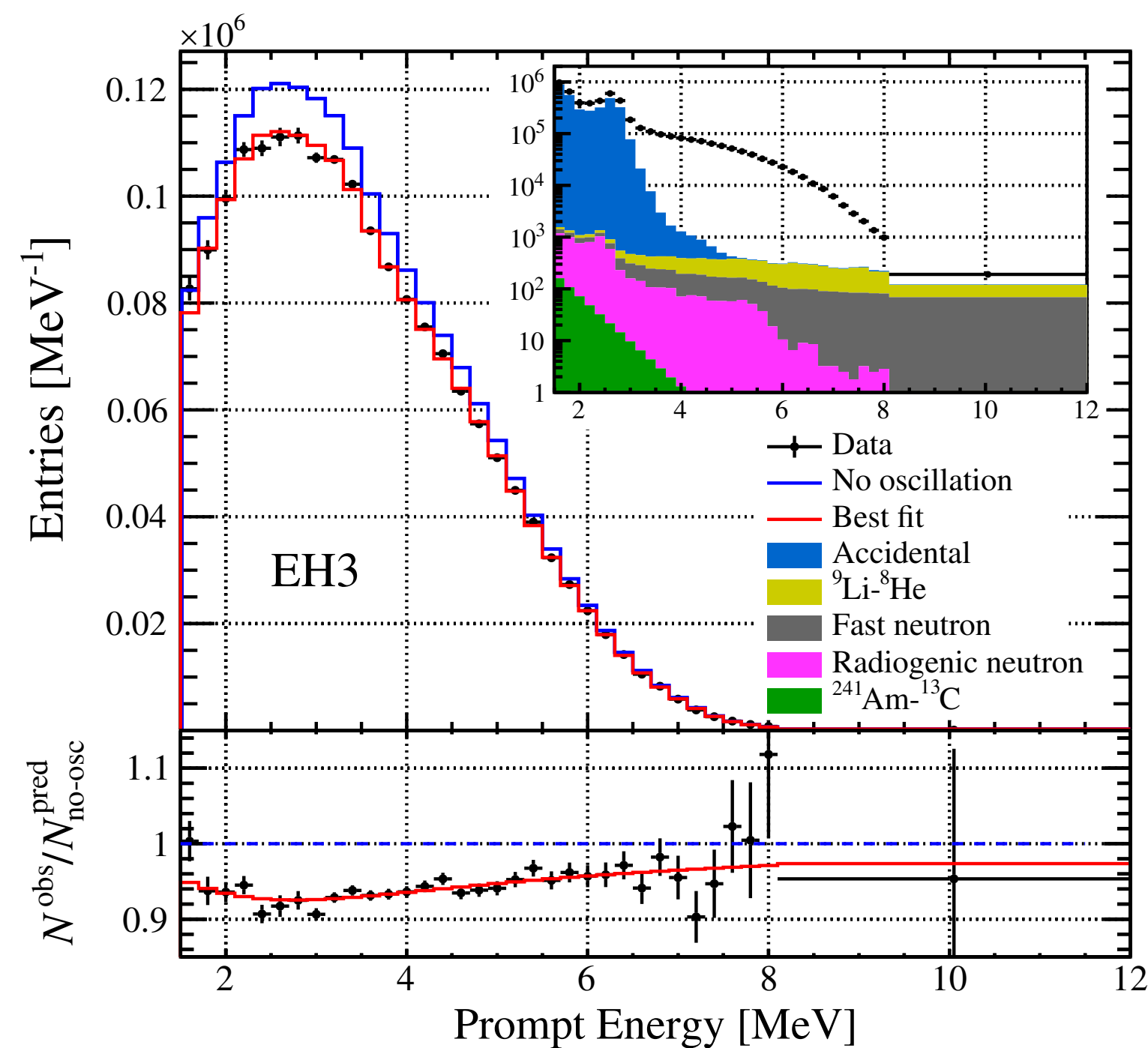
All have completed data taking

- Identical detectors **near** and **far** (~ 1 km) from reactor allows cancellation of many systematic effects
- Oscillation of $\bar{\nu}_e \rightarrow \bar{\nu}_e$ at first osc. max.

Approximately 2-flavor osc. with amplitude $\sin^2 2\theta_{13}$ and frequency $|\Delta m_{ee}^2| \approx |\Delta m_{31}^2|$

θ_{13} constraints

Daya Bay nH analysis
PRL **133**, 151801 (2024)



Note: average is error weighted
average assuming no correlation

θ_{13} constraints

RENO final
PRD 111, 112006 (2025)

$\chi^2/\text{ndf} = 64.0/66$

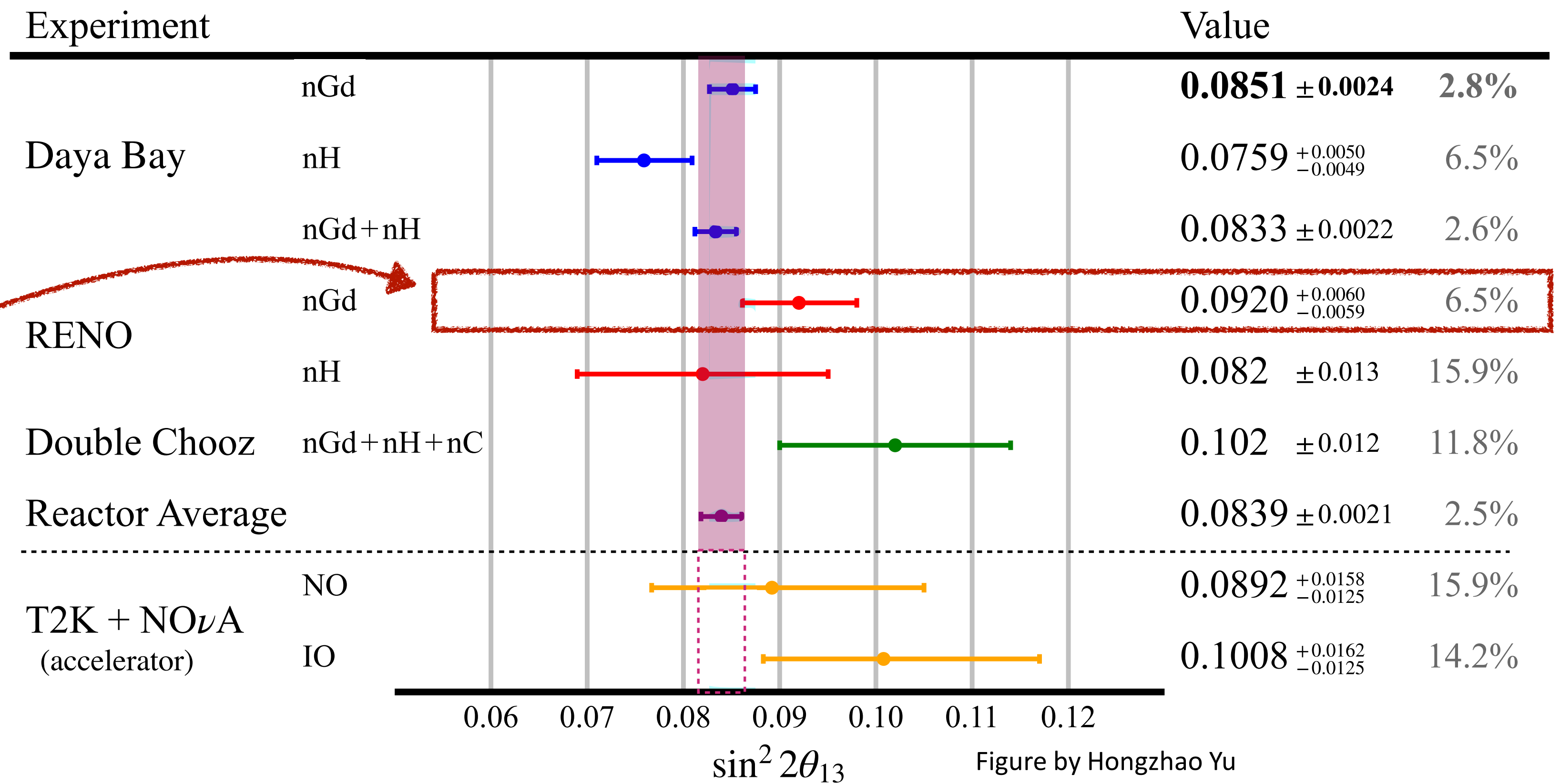
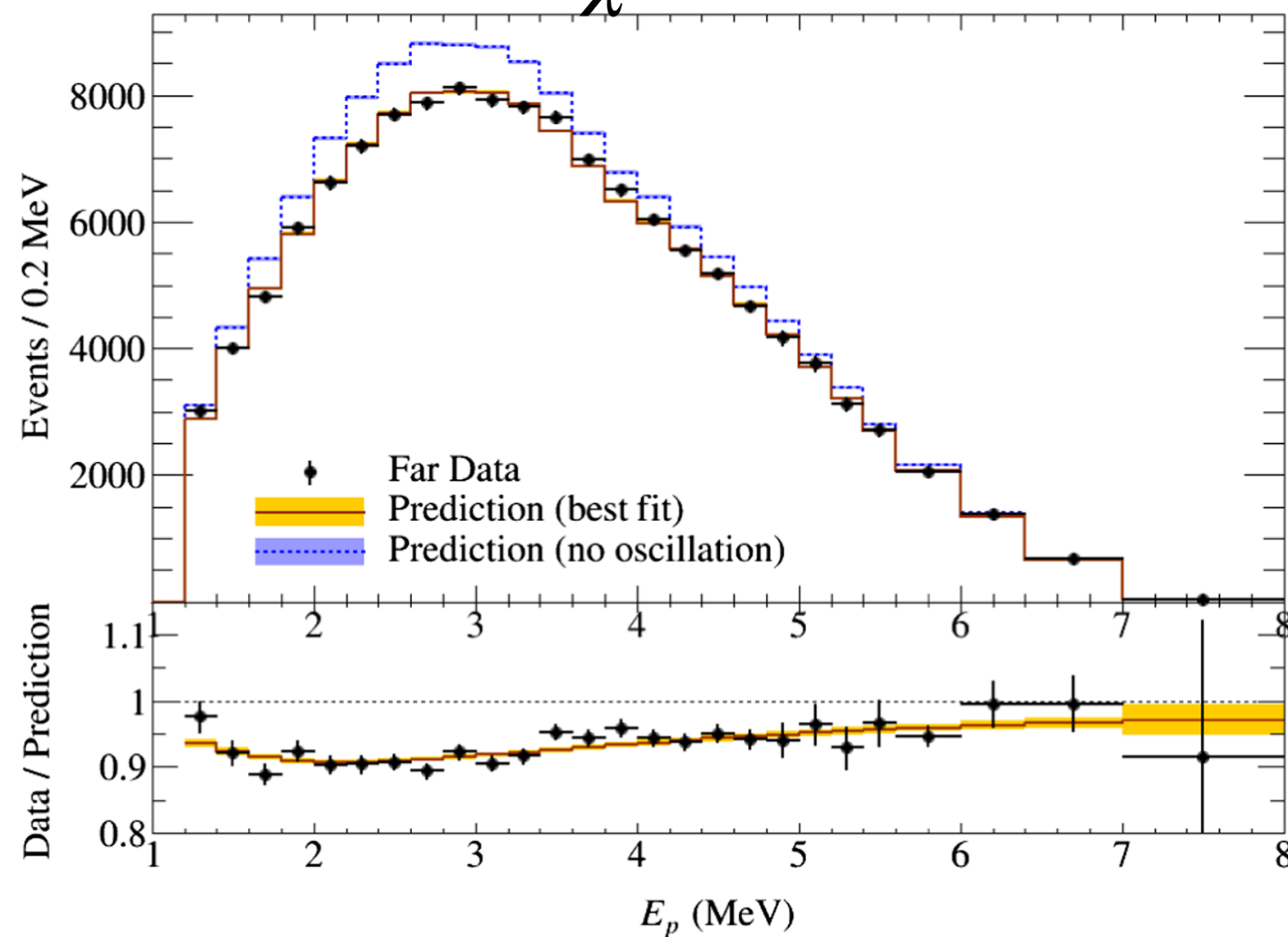


Figure by Hongzhao Yu

Slightly modified emphasis

Note: average is error weighted
average assuming no correlation

θ_{13} constraints

Double Chooz
Final results will be
presented this afternoon
by Philipp Soldin

Neutrino Physics and
Astrophysics Parallel 1B

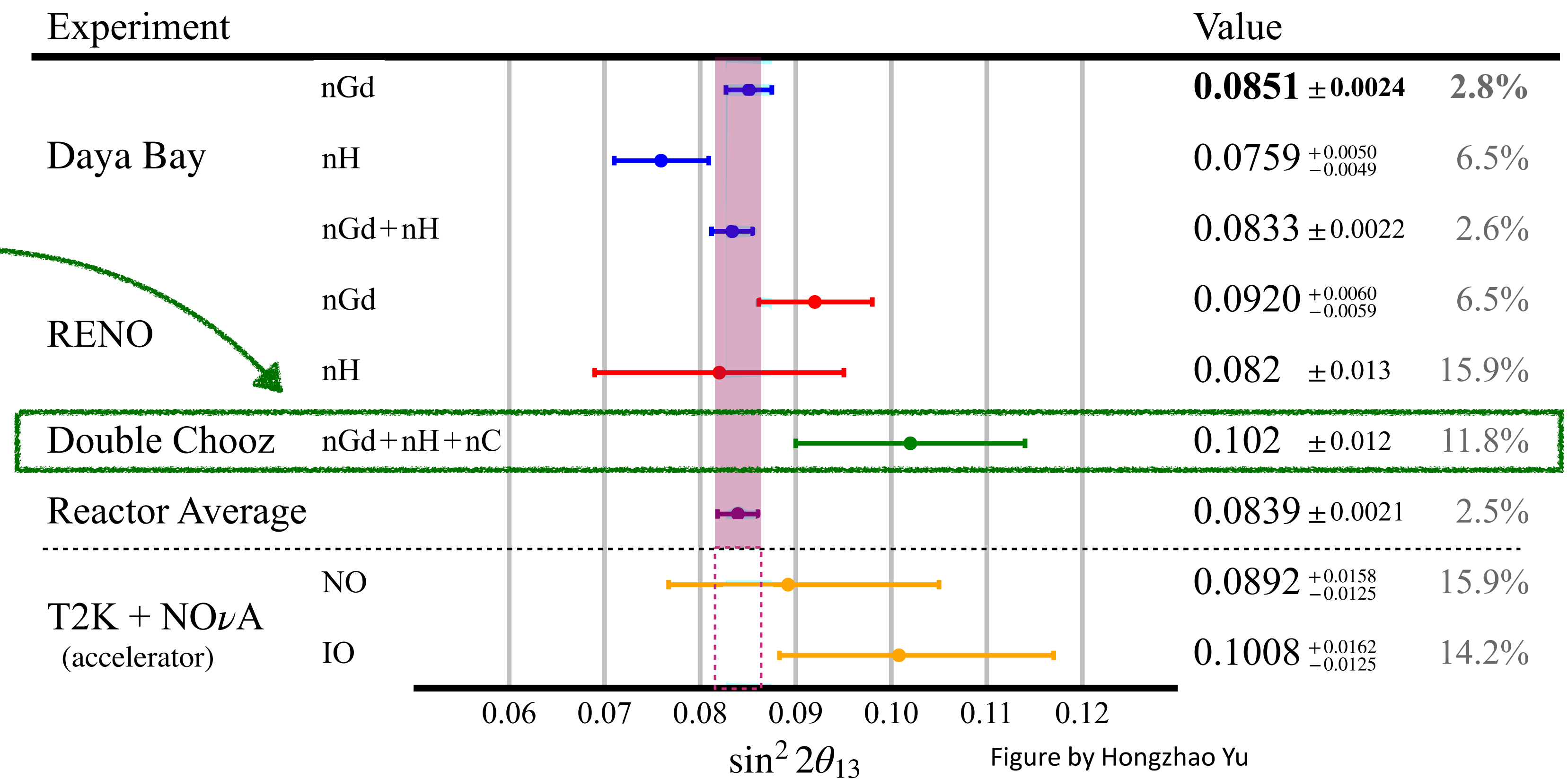


Figure by Hongzhao Yu

Slightly modified emphasis

Note: average is error weighted
average assuming no correlation

θ_{13} constraints

Consistent results also from
 $\nu_{\mu} \rightarrow \nu_e$ appearance at
accelerators
(some implication on MO)

Nice check of 3ν paradigm

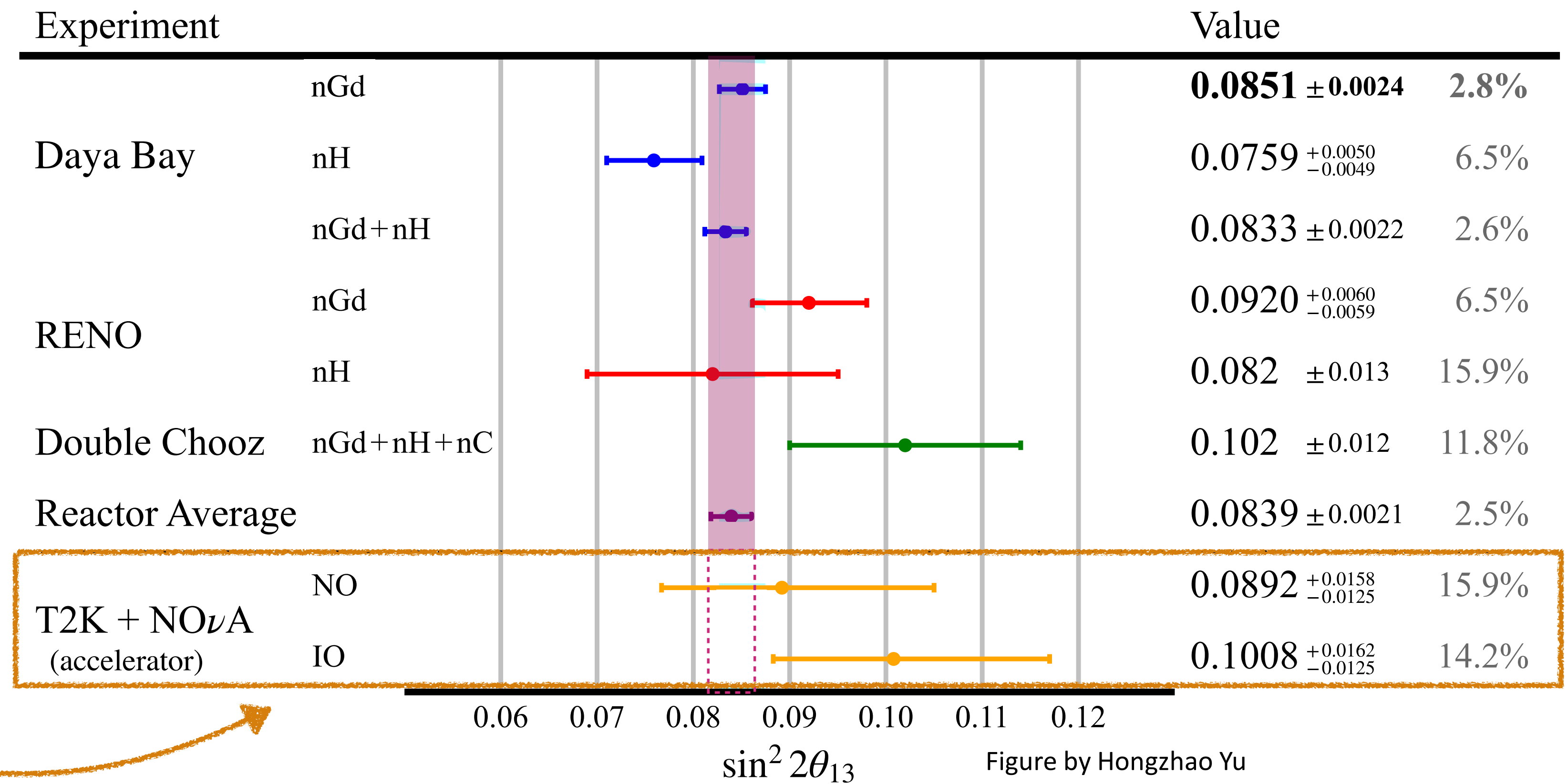
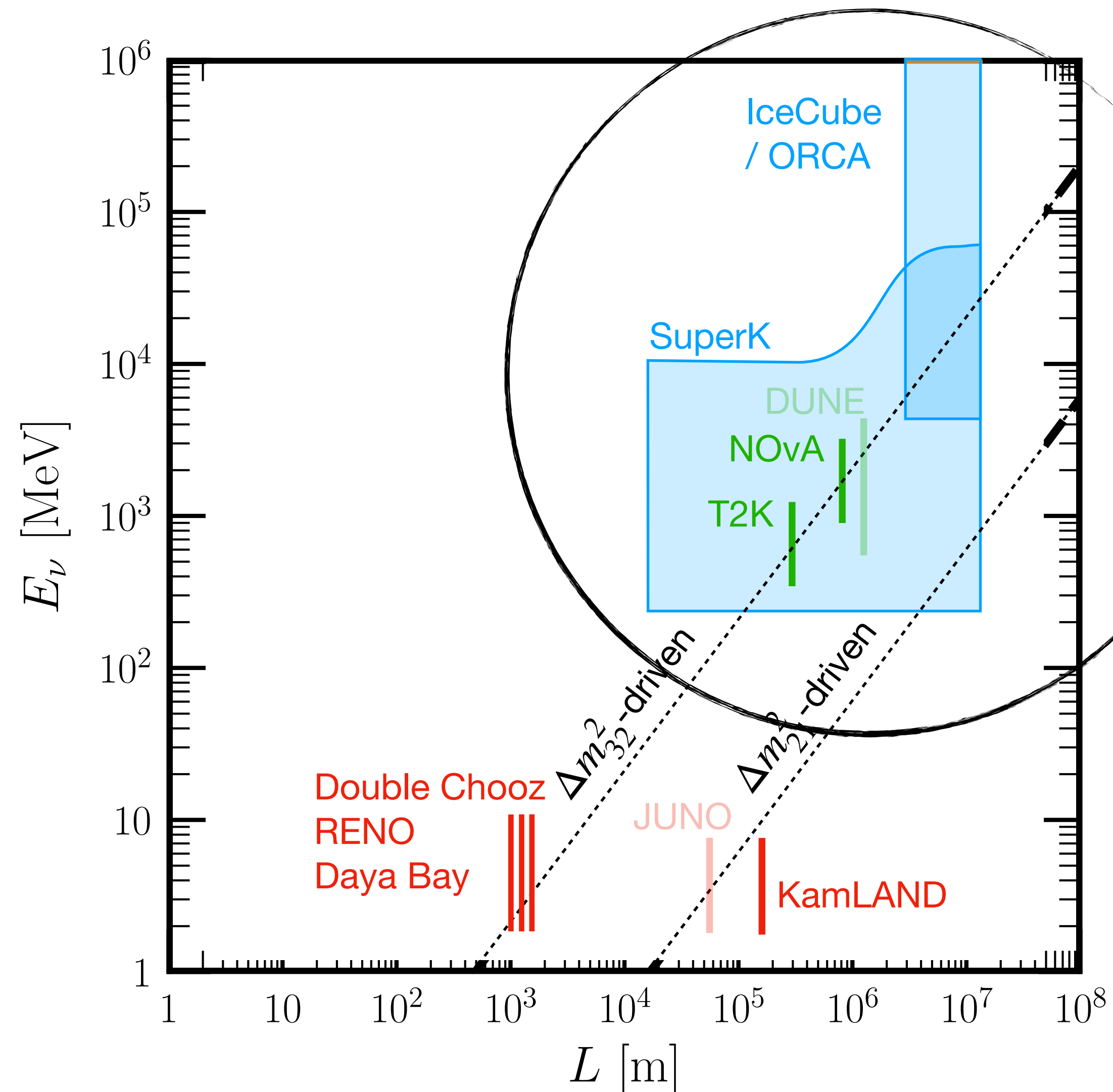


Figure by Hongzhao Yu

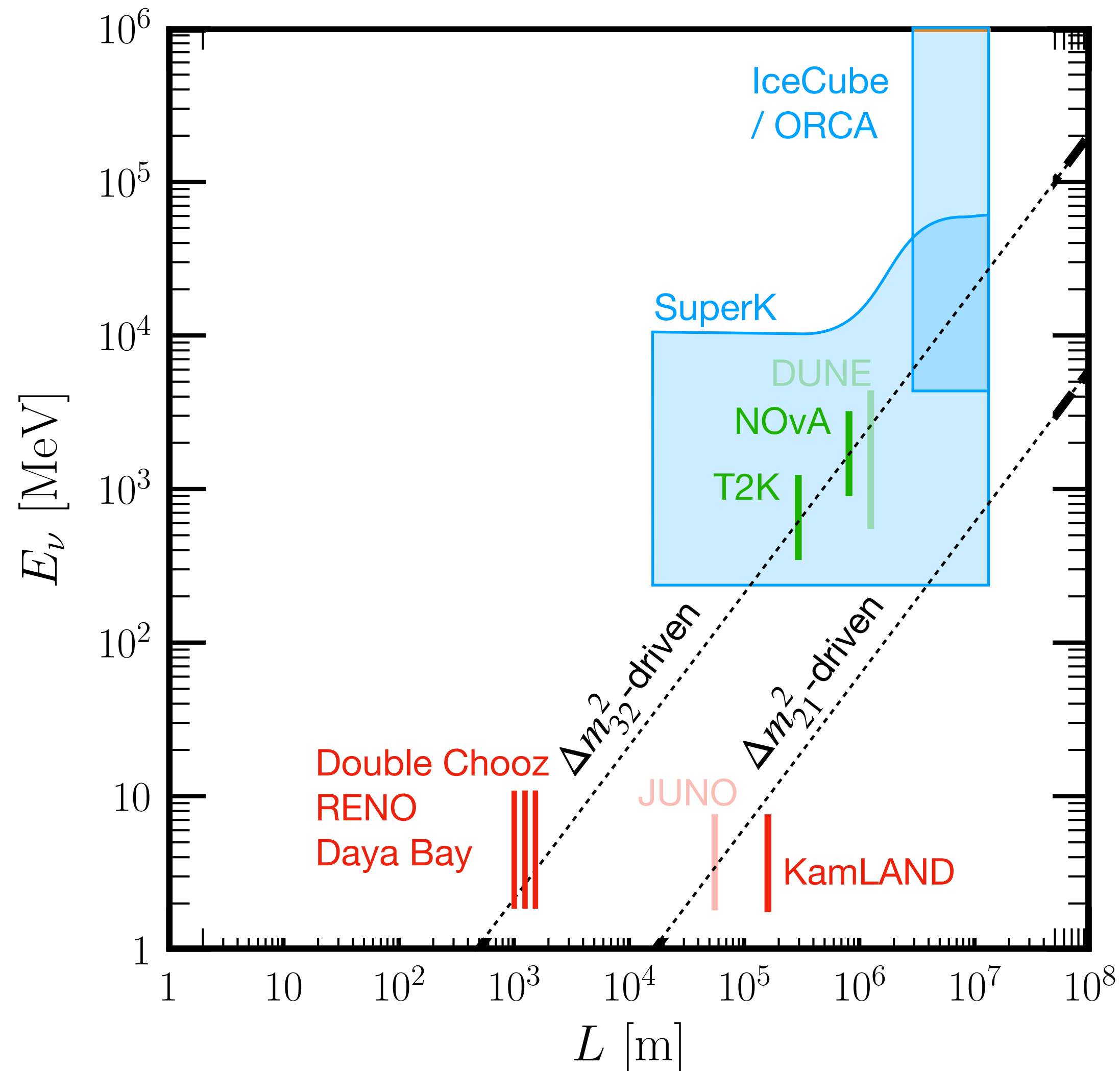
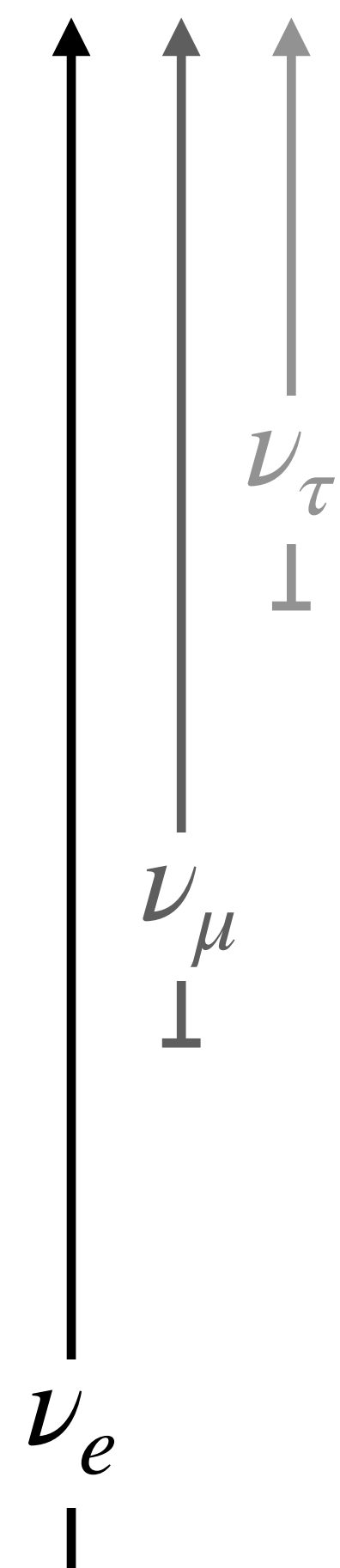
Slightly modified emphasis

Note: average is error weighted
average assuming no correlation

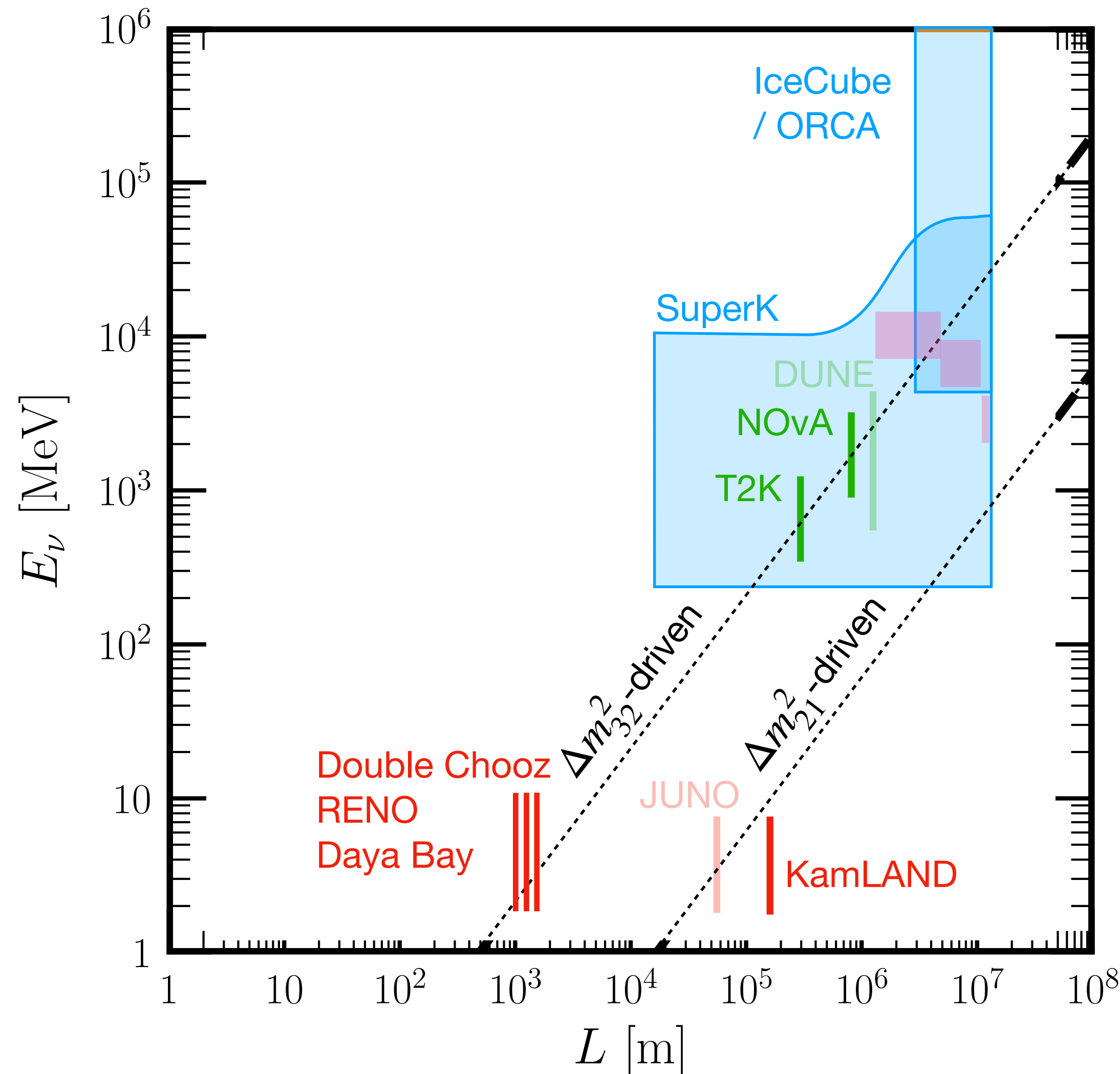
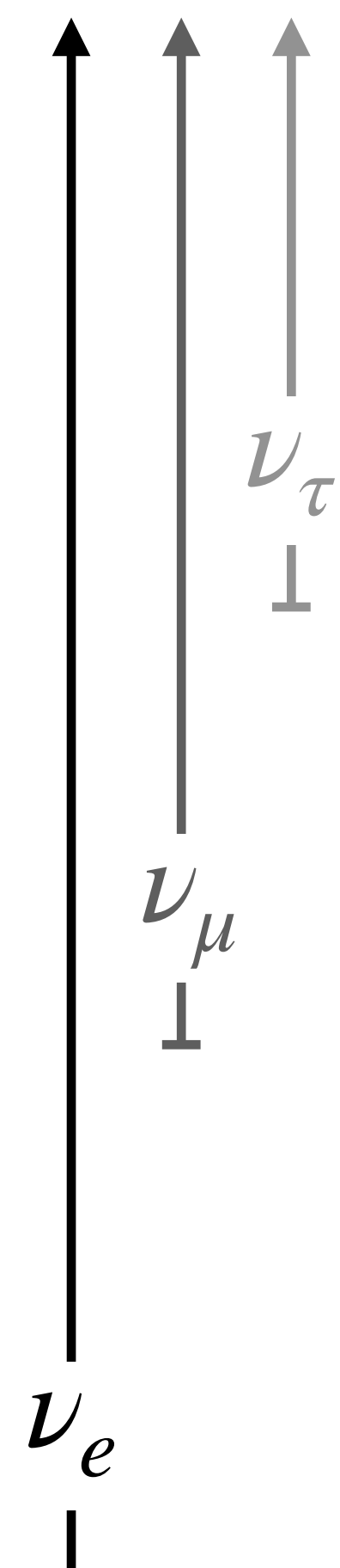
Long-baseline experiments



Long-baseline experiments



Long-baseline experiments



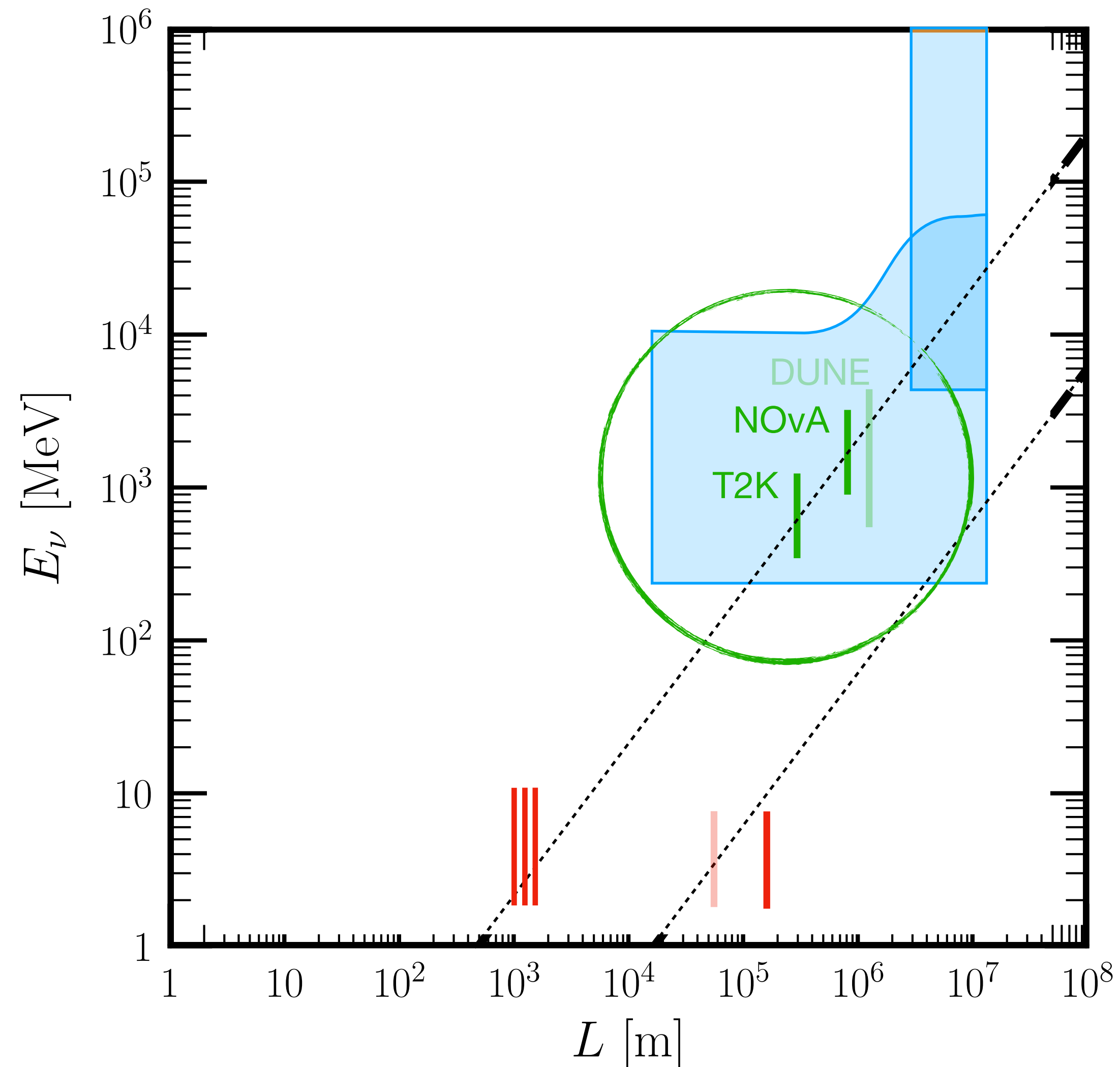
In matter, coherent forward-scattering with e^- modifies oscillations

θ_{13} resonance
 ν for NO
 $\bar{\nu}$ for IO

$\sin^2 2\theta_{13,m}$

Assuming typical densities for terrestrial exp.
 Plot inspired by PRD 101, 033008 (2020)

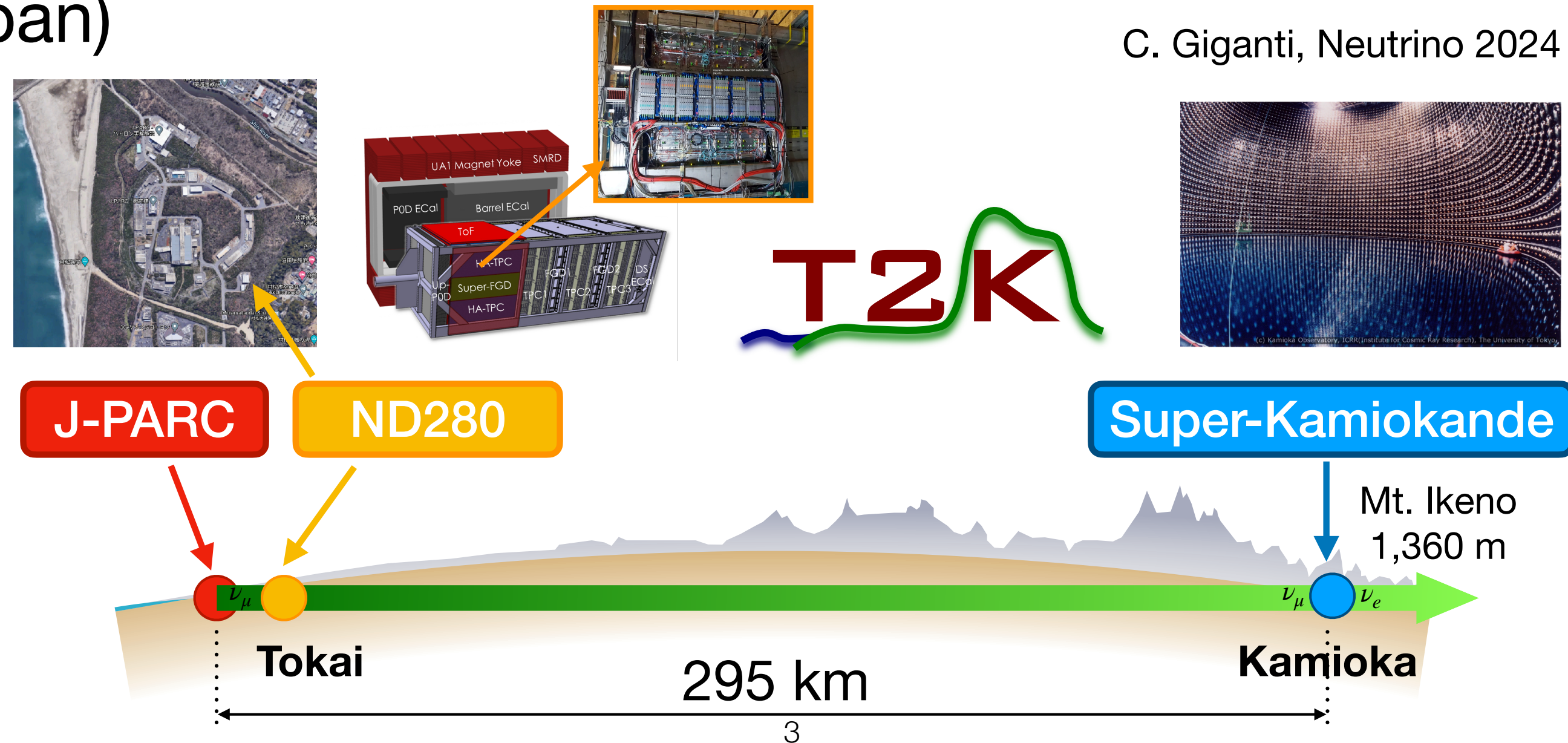
Accelerator experiments



Assuming typical densities for terrestrial exp.
Plot inspired by PRD 101, 033008 (2020)

Accelerator neutrinos

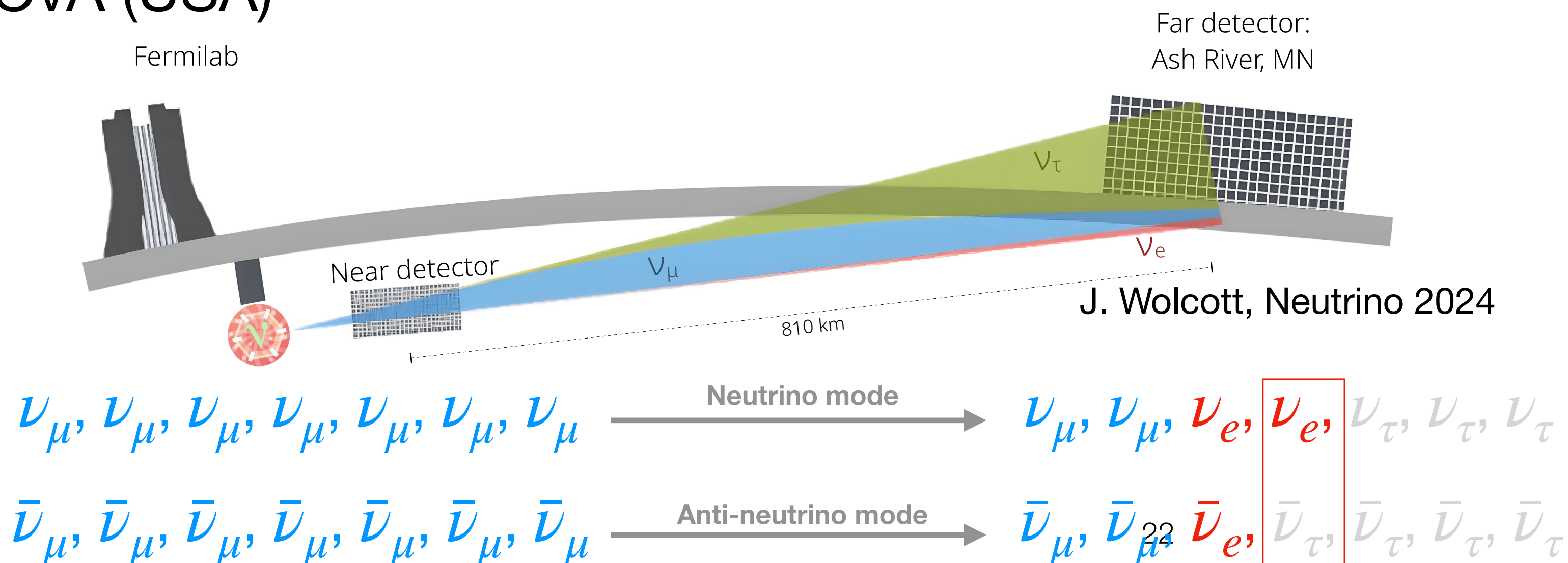
T2K (Japan)



C. Giganti, Neutrino 2024

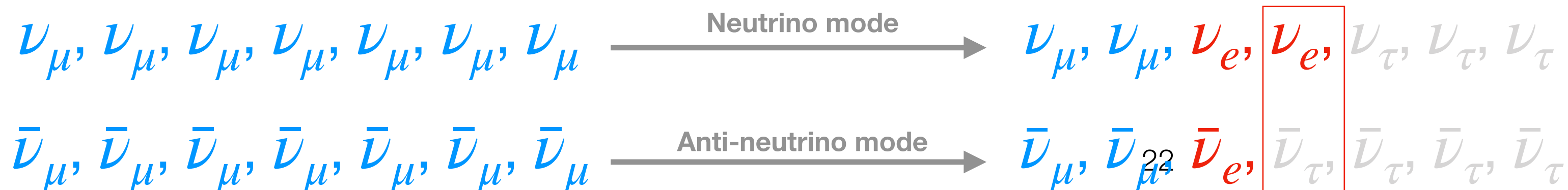
NOvA (USA)

Fermilab



J. Wolcott, Neutrino 2024

- Selectable ν_μ or $\bar{\nu}_\mu$ beams by focusing π^\pm produced by p beam on fixed target
- Precision study of $\nu_\mu \rightarrow \nu_{e,\mu}$ oscillations near first oscillation maximum
- Low $\nu_e/\bar{\nu}_e$ contamination allows study of $\nu_\mu \rightarrow \nu_e$ oscillations for both $\nu/\bar{\nu}$
- Near detectors to measure neutrinos before oscillations
↓
constrain flux \times interaction systematics



ν_μ disappearance

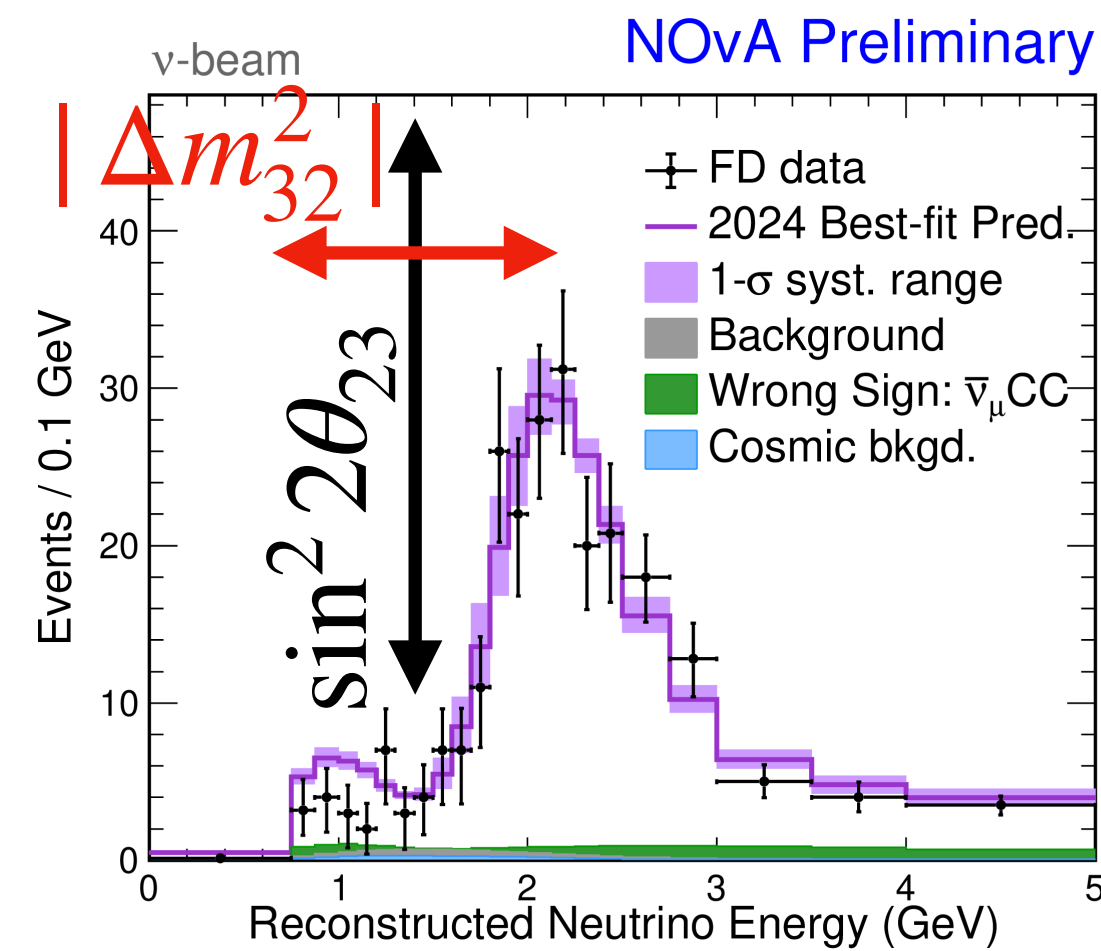
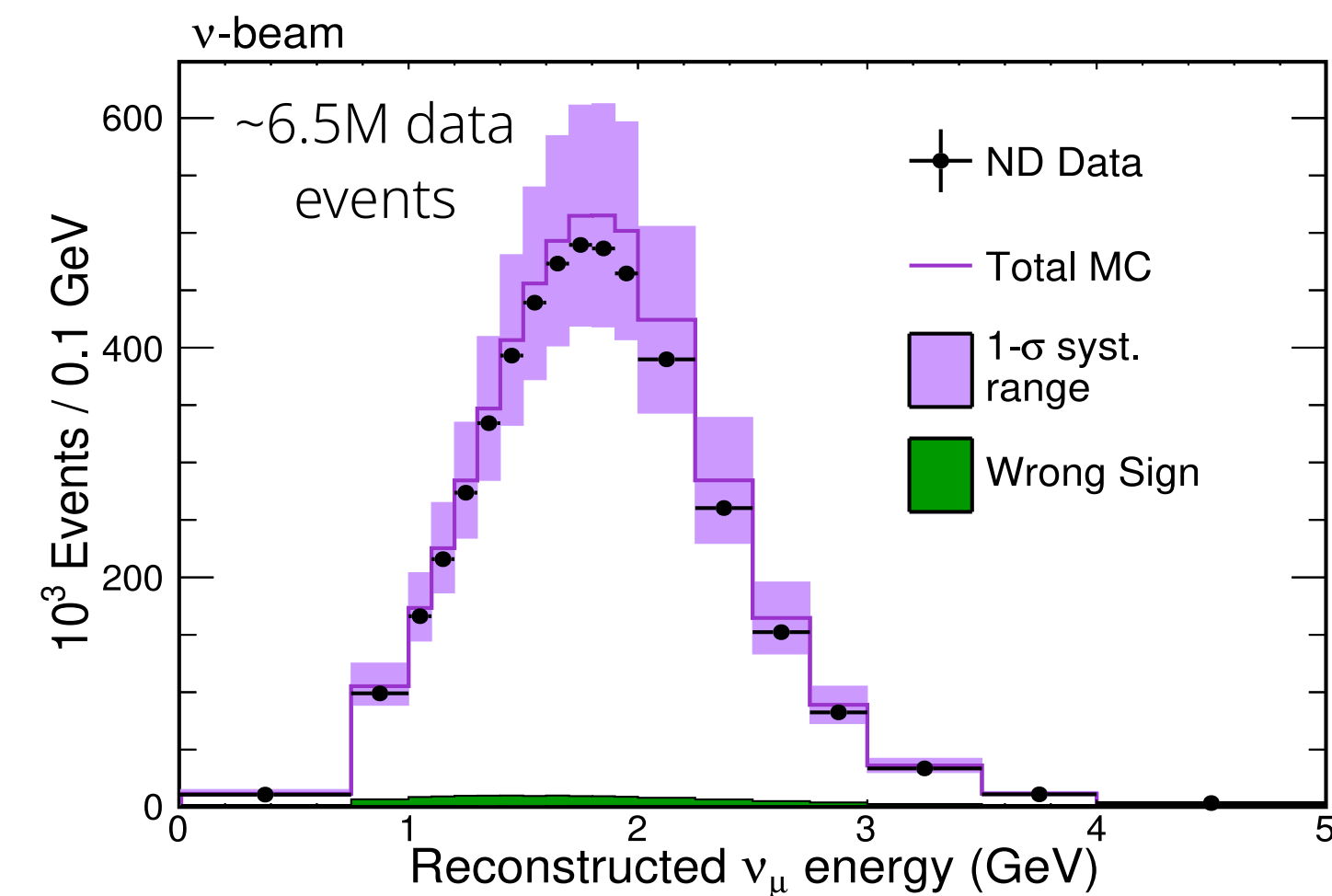
Neutrino 2024 Preliminary

ND

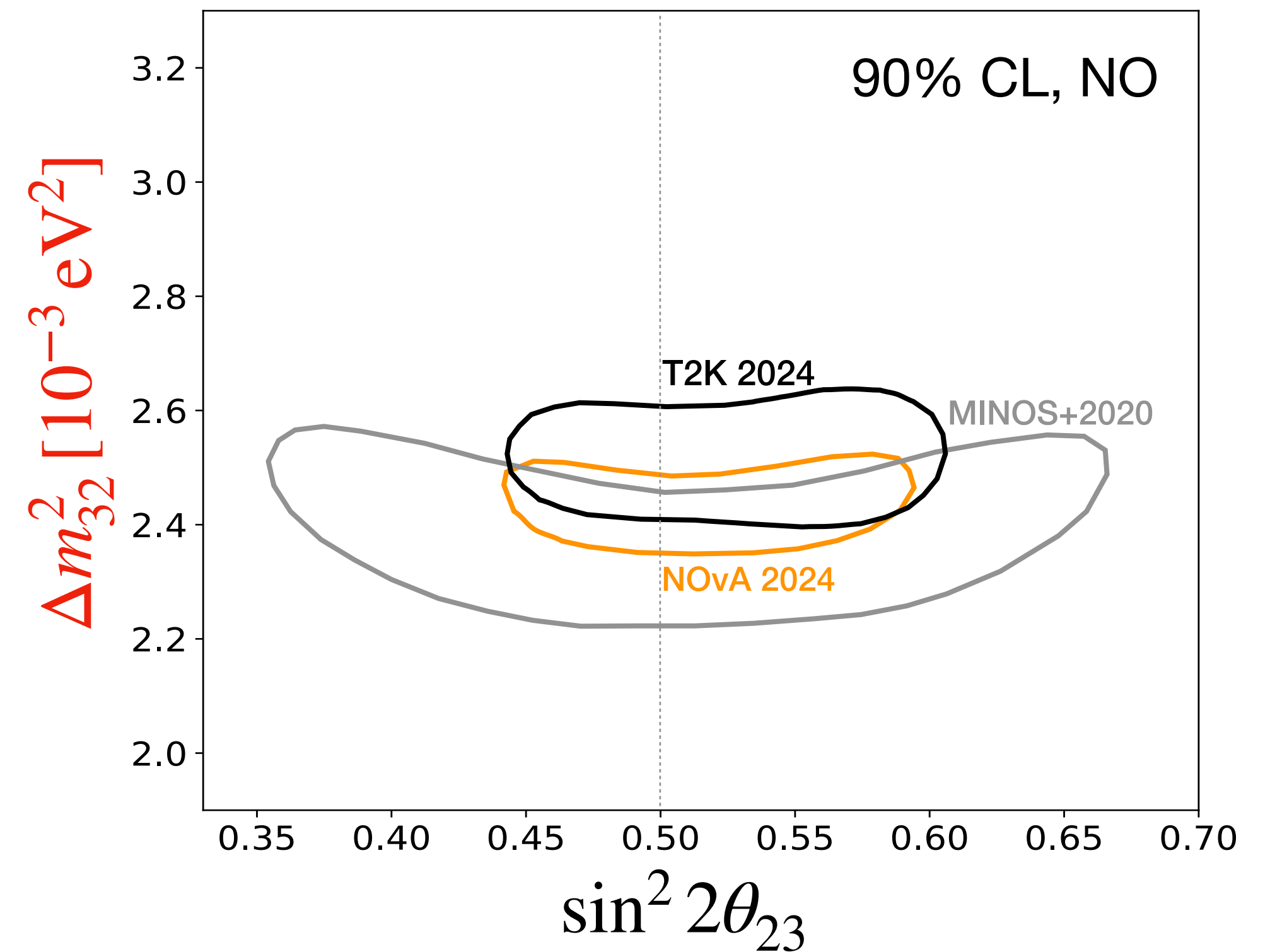
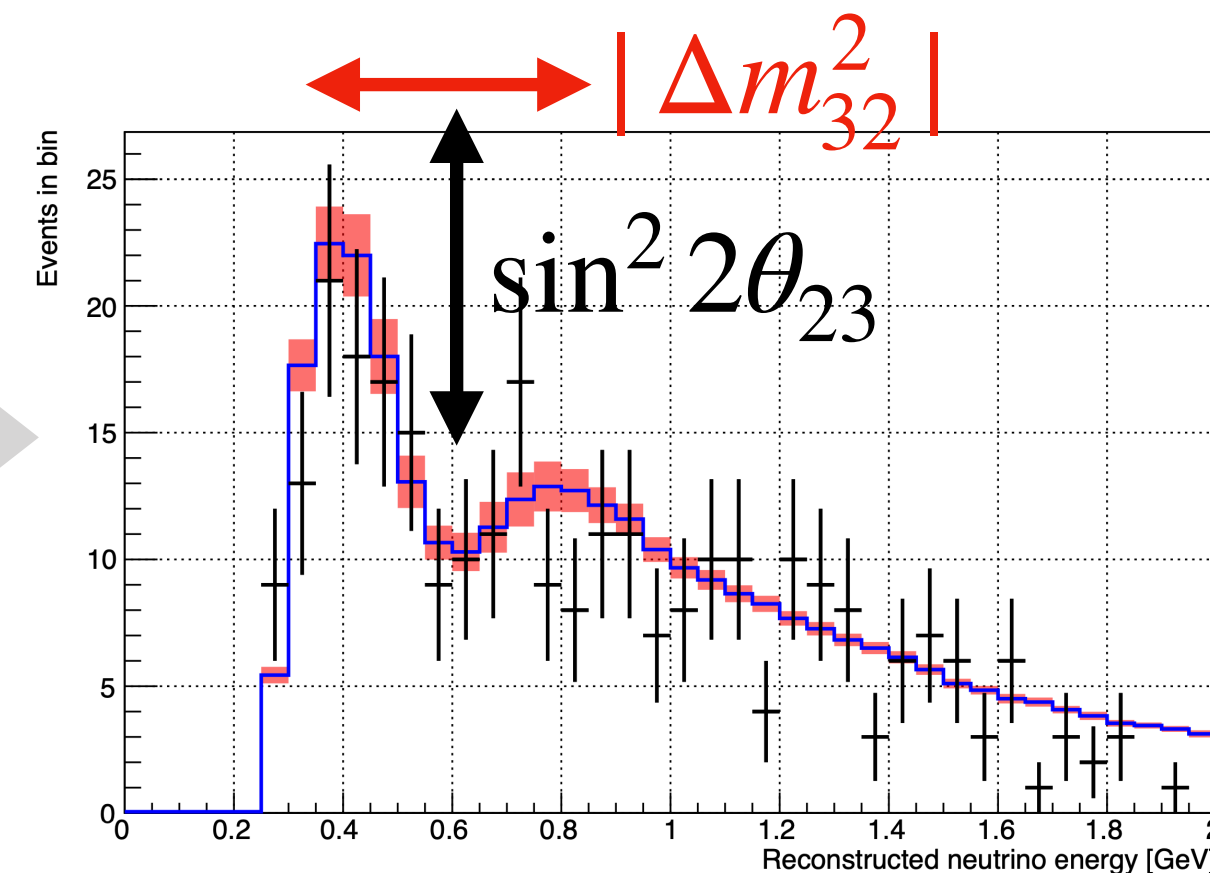
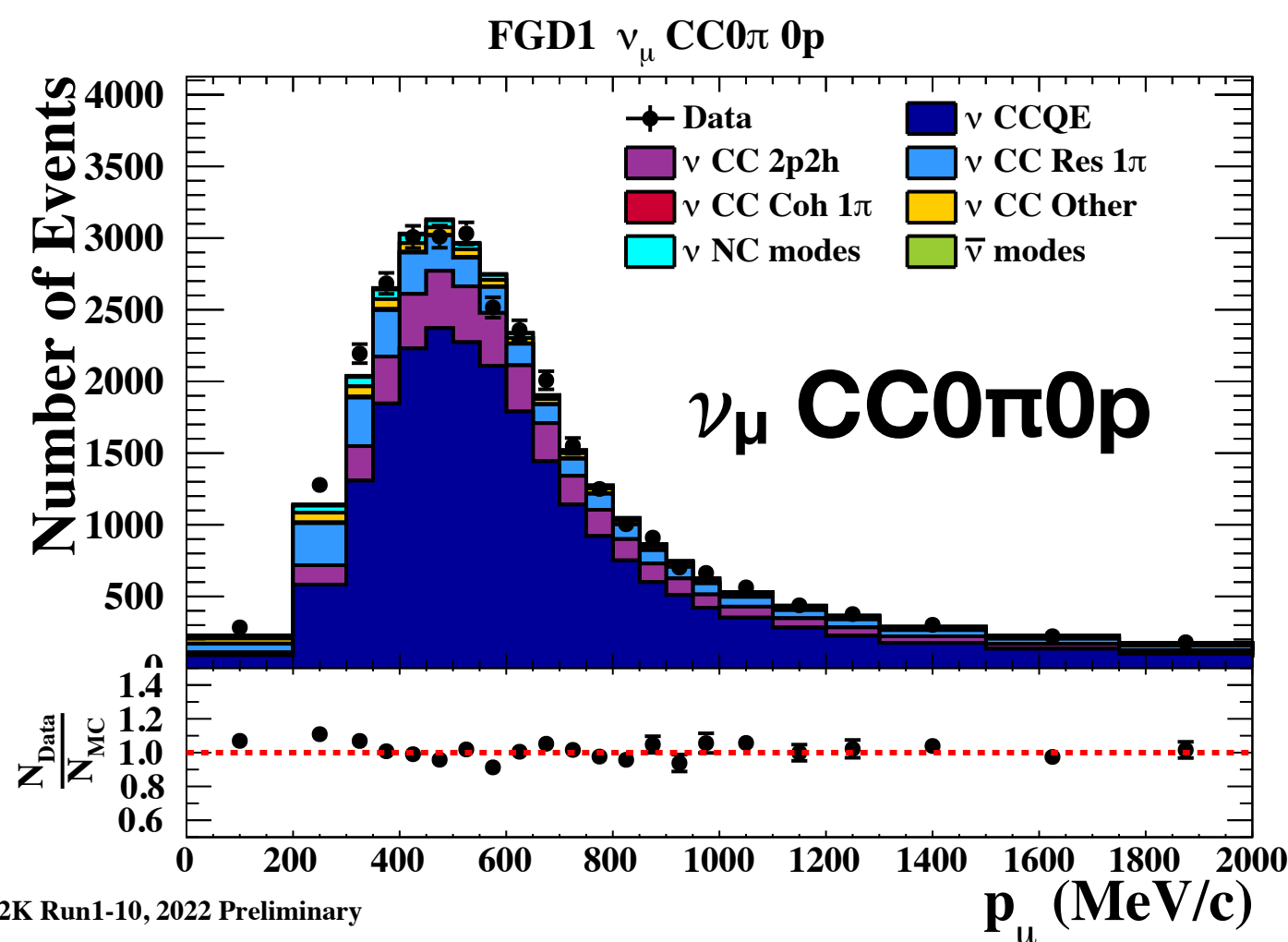
FD

NOvA doubled ν -mode statistics \rightarrow leading

T2K 10% more ν -mode compared to 2022 and reduced FD detector systematics



384 ν_μ data candidates
(11.3 background)



Consistent with maximal mixing

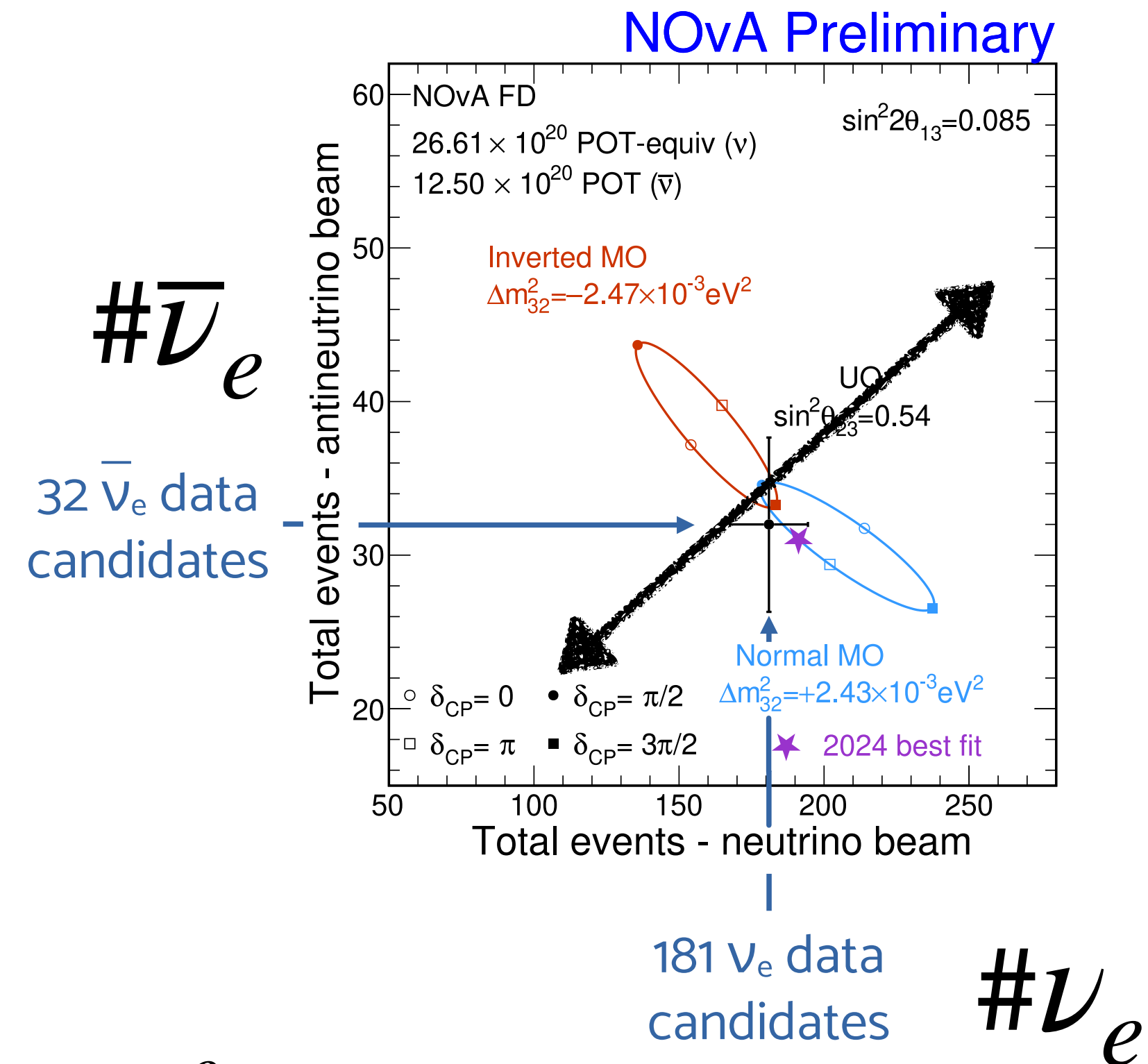
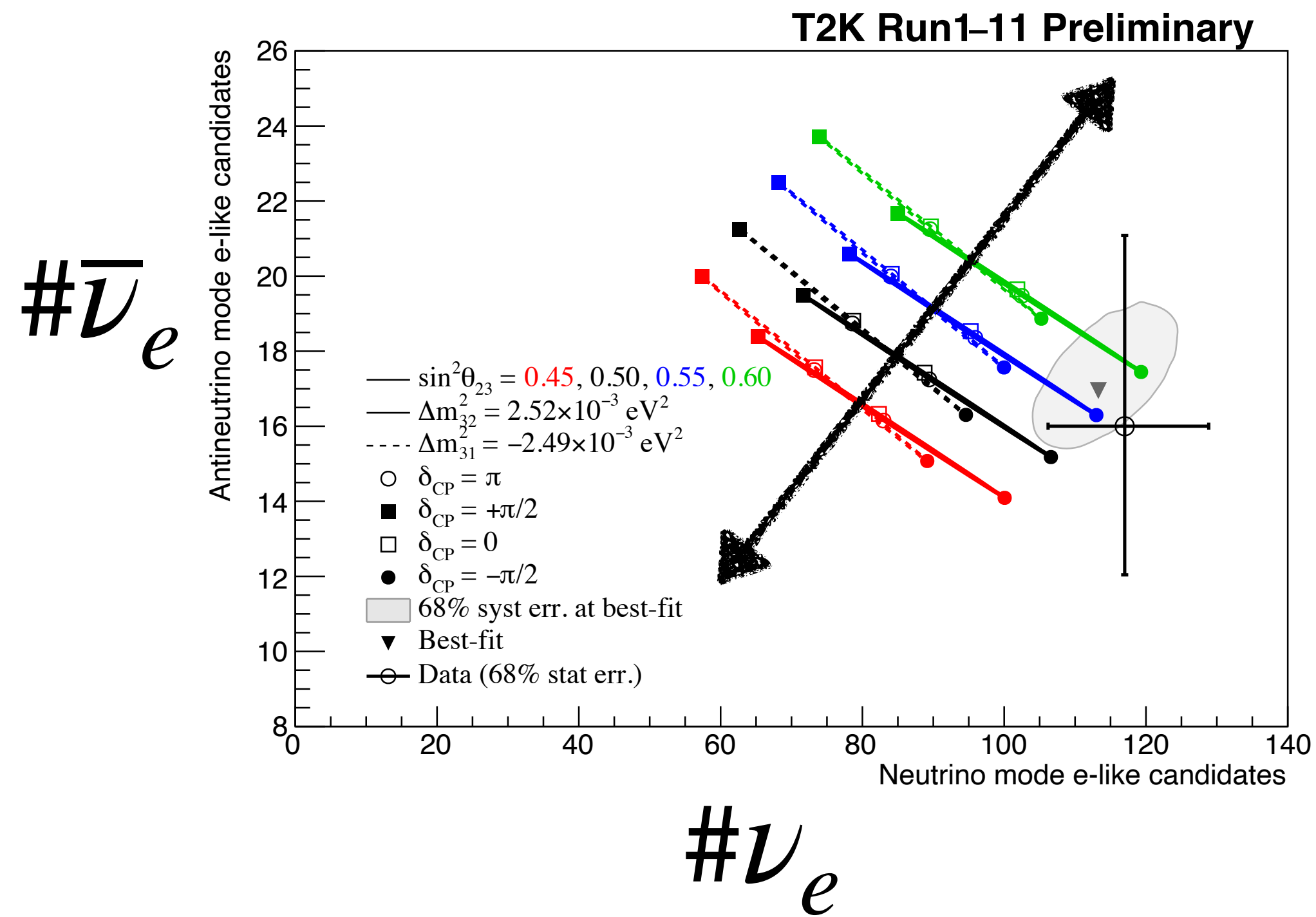
T2K has been upgrading beam-line and ND.
 \rightarrow expect improvement in future.

$\nu_e/\bar{\nu}_e$ appearance

Currently mostly a rate measurement

T2K

NOvA



Accelerator only: θ_{13}

Correlated change

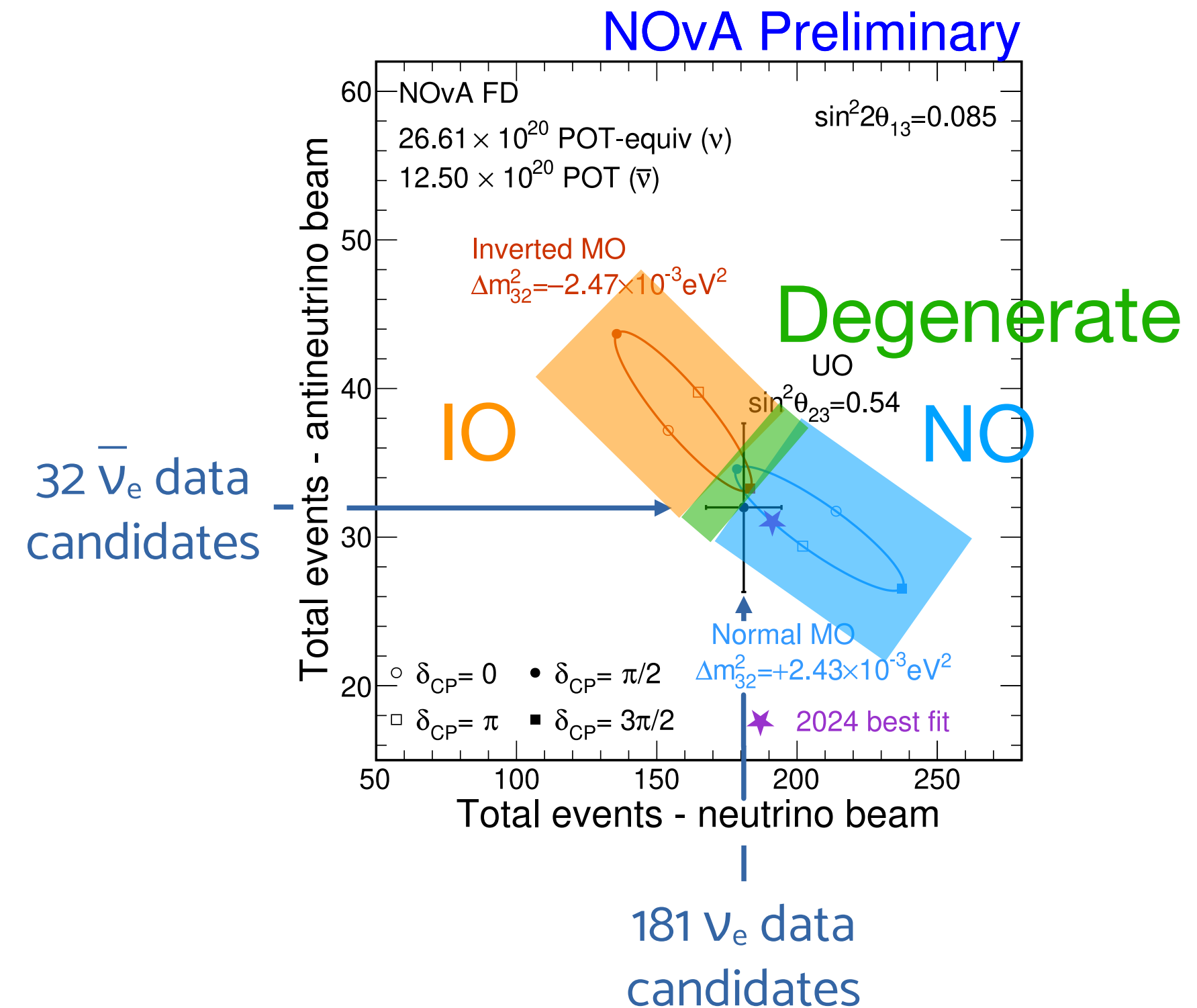
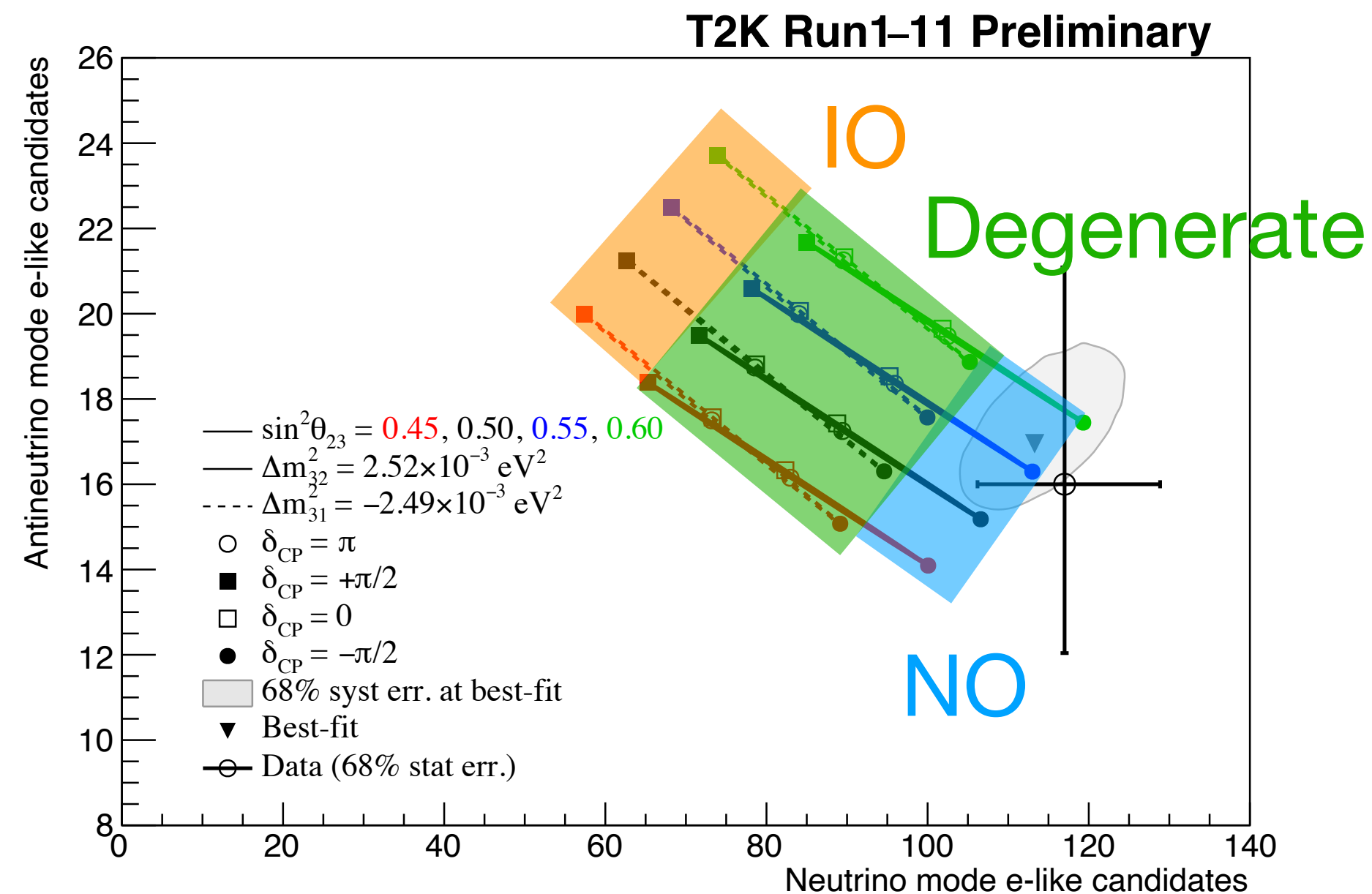
With reactor θ_{13} :
 θ_{23} octant, $\cos \delta_{\text{CP}}$

$\nu_e/\bar{\nu}_e$ appearance

Currently mostly a rate measurement

T2K

NOvA



Preference for NO

More sensitivity to MO than T2K due to higher energy (longer baseline).

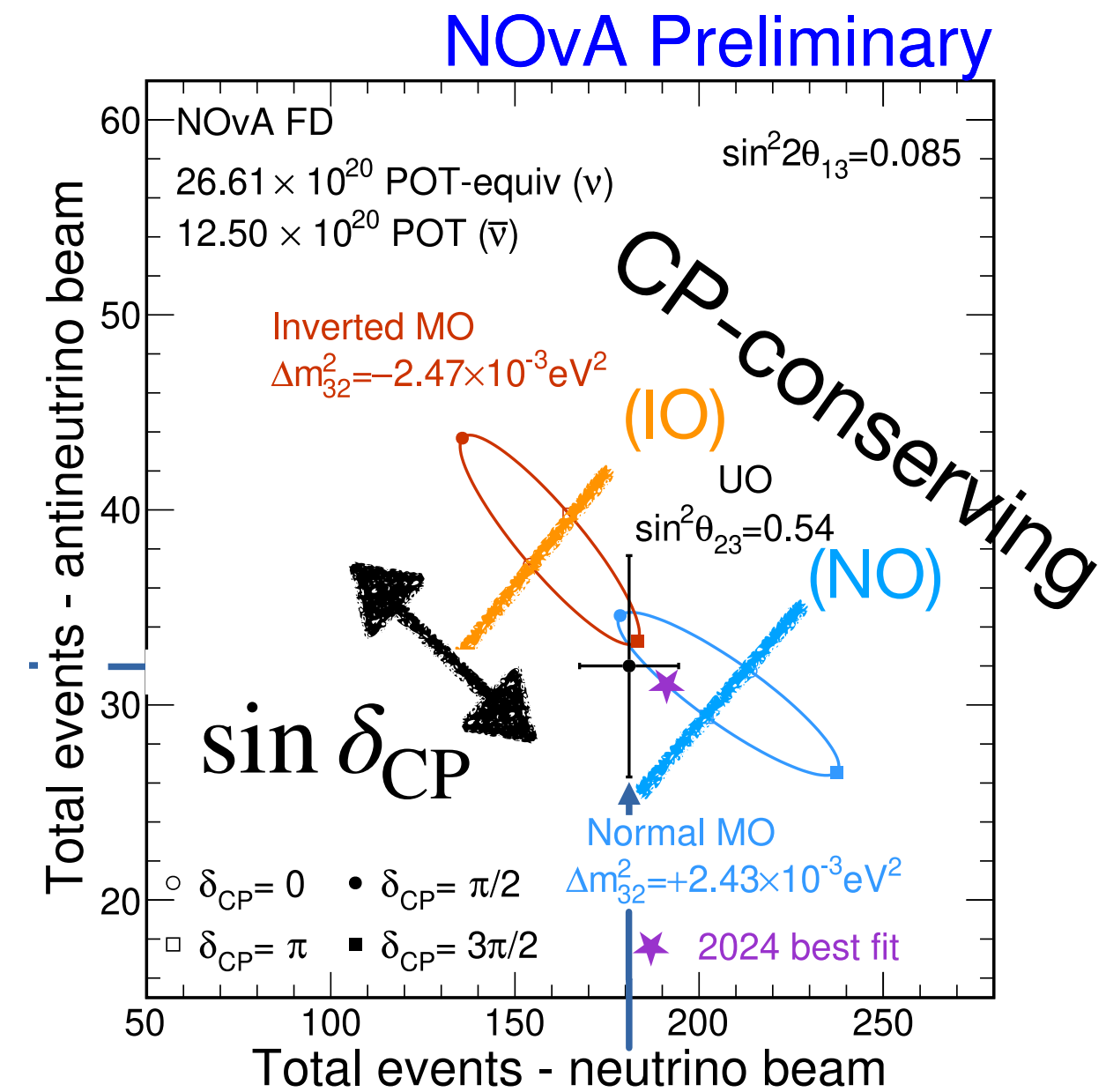
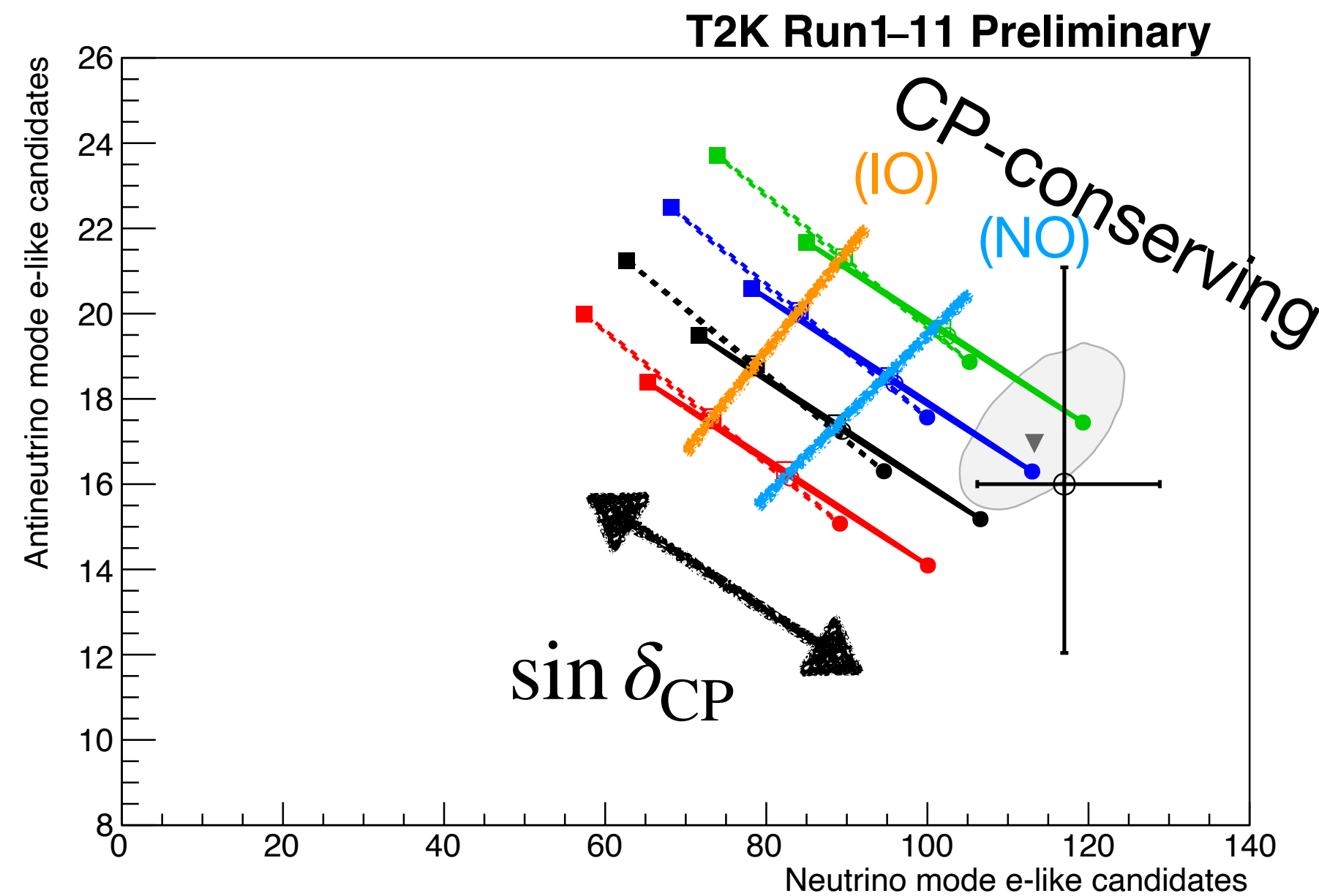
But data prefers MO-degenerate region

$\nu_e/\bar{\nu}_e$ appearance

Currently mostly a rate measurement

T2K

NOvA



Preference for $\delta_{CP} \approx -\pi/2$

Data outside of MO-degeneracy
= stronger CP constraint

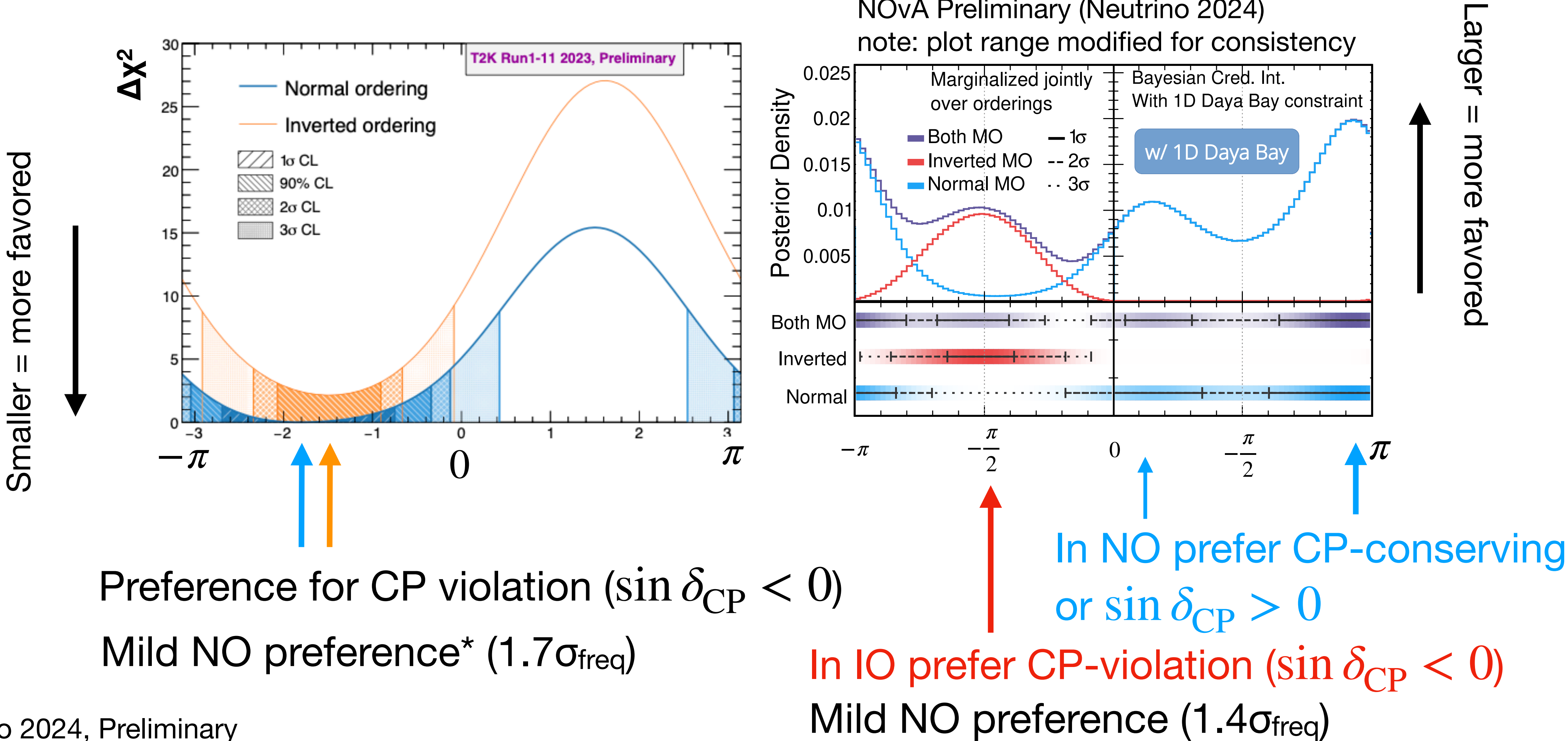
δ_{CP} preference depends on MO

Using reactor θ_{13} constraint
(different values are used)

warn: difference in freq vs. bayesian

T2K

NOvA



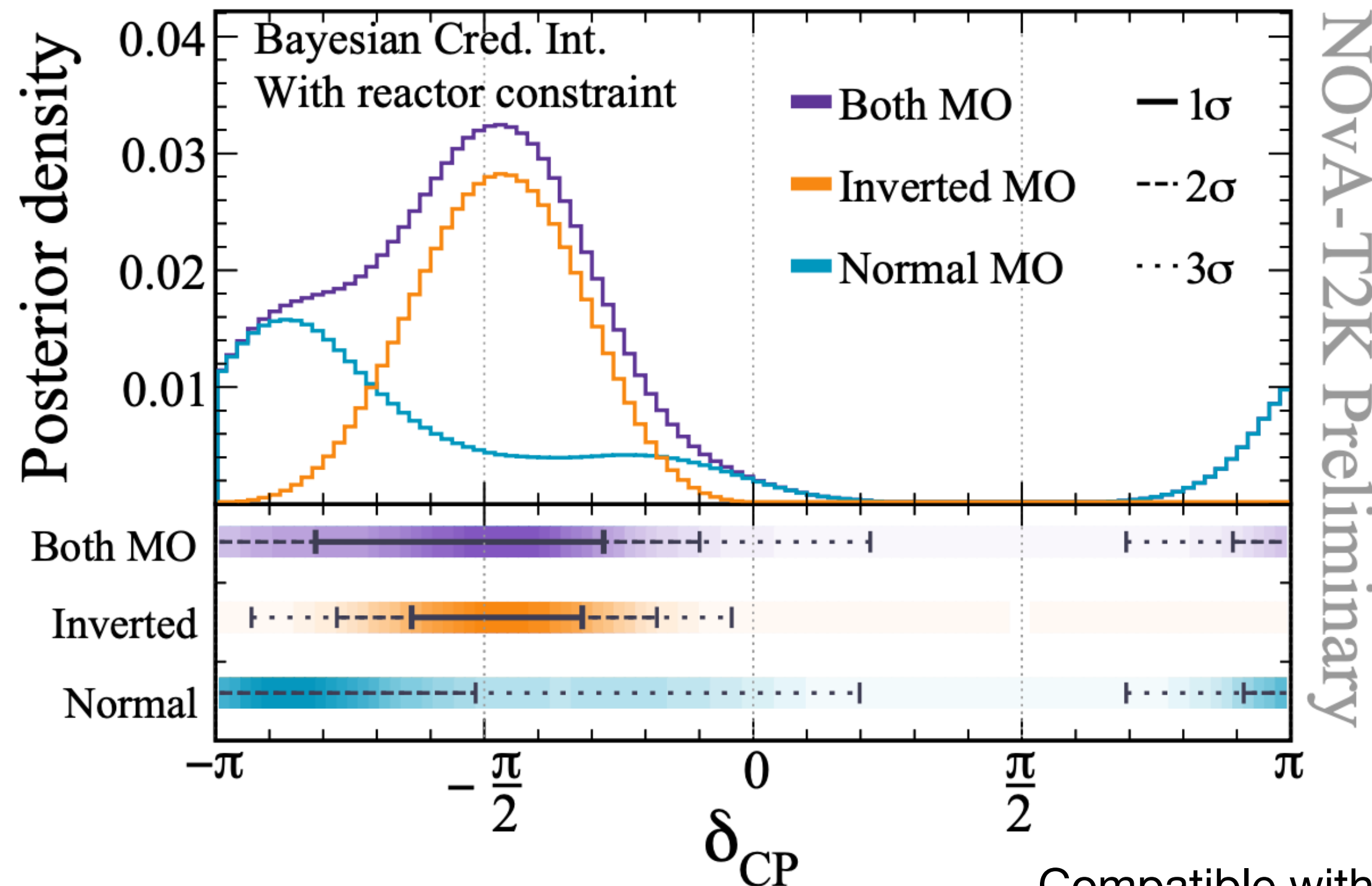
Neutrino 2024, Preliminary
* Carabadjac Denis, ICHEP 2024, T2K Preliminary

T2K + NOvA

NOvA only: [Phys. Rev. D106, 032004 \(2022\)](#)

T2K only: [Eur. Phys. J. C83, 782 \(2023\)](#)

Input analyses are not latest results, but first shown in 2020



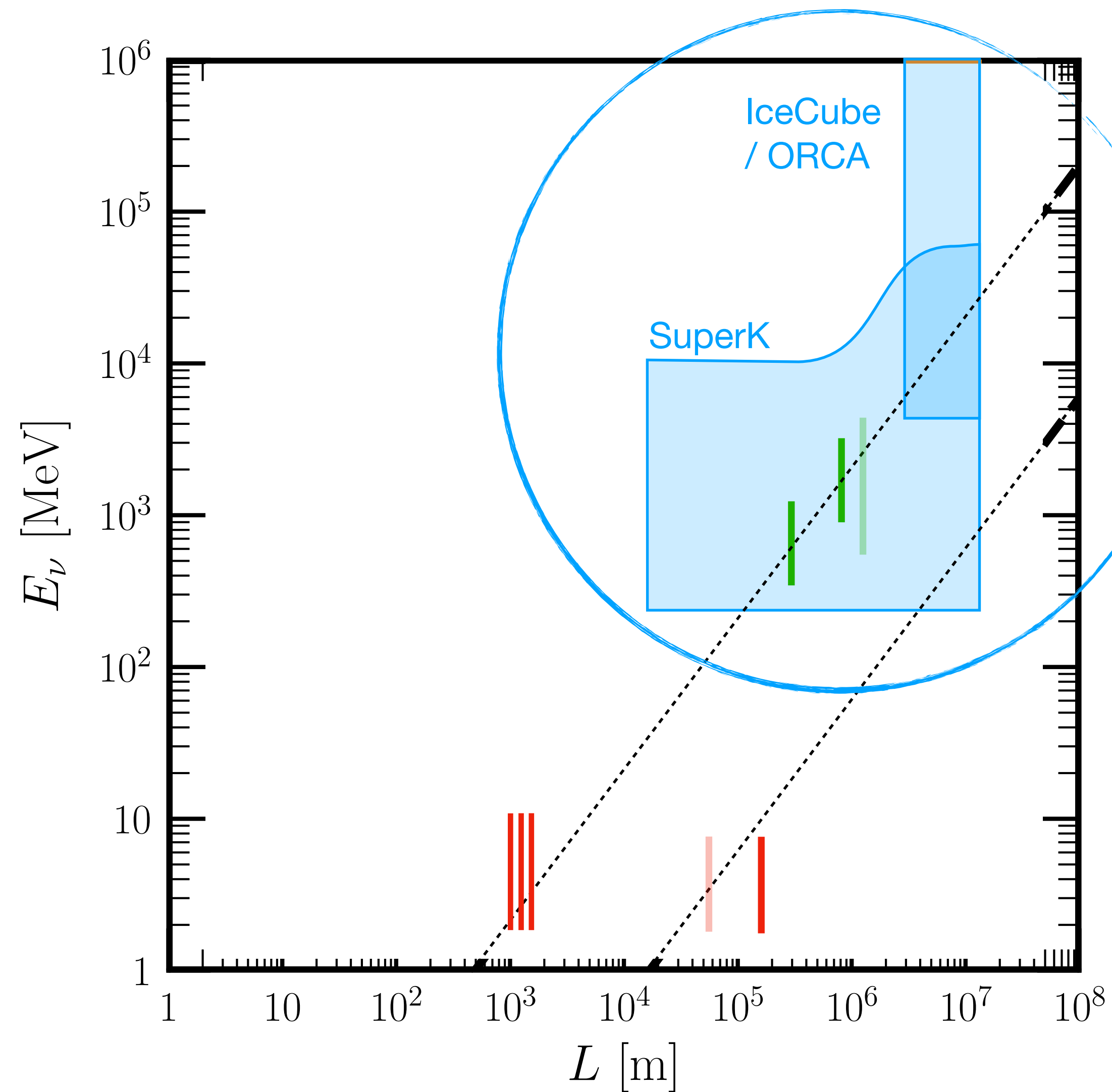
- Different degeneracy of δ_{CP} , MO, and θ_{23} octant → synergy
- A first joint fit was performed using the analyses first shown in 2020 (publication in preparation)
- Candidate for systematic correlations: ν interactions
 - No trivial mapping between parameters (except $\nu_e/\bar{\nu}_e$ systematics which were correlated)
- At current statistics omitting correlations found to not affect result
- Studied impact of interaction model differences, all tests pass pre-set criteria

- If IO, CP violated at 3σ
(Above plot is normalized over both MO, but conclusion also holds when conditioned on IO)
- If NO, consistent with CP conservation

Compatible with both MO, posterior influenced by reactor constraint

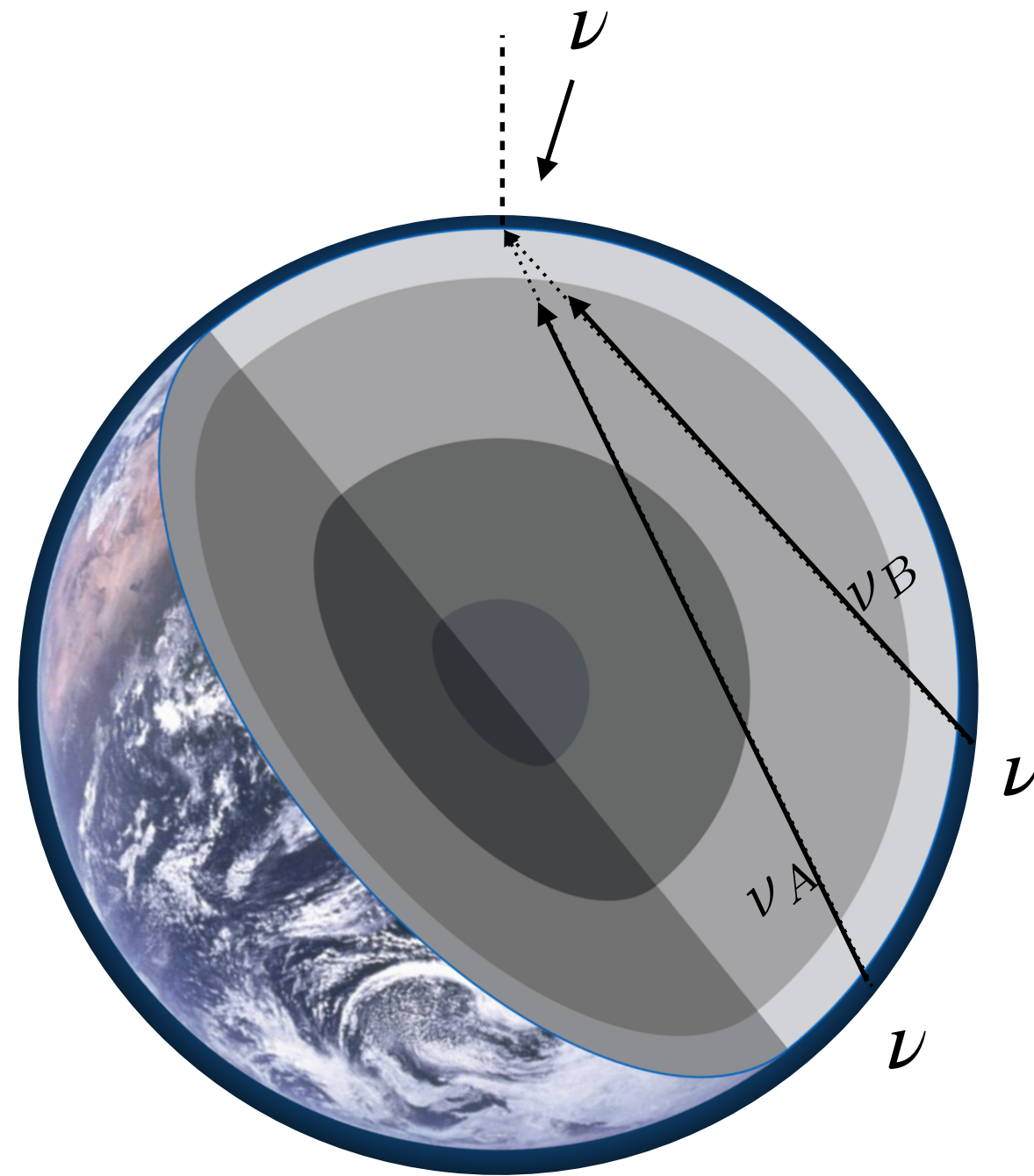
NOvA+T2K only : IO (71%)
+ 1D θ_{13} : IO (57%)
+ 2D $(\theta_{13}, \Delta m_{32}^2)$: NO (59%)

Atmospheric experiments



Assuming typical densities for terrestrial exp.
Plot inspired by PRD 101, 033008 (2020)

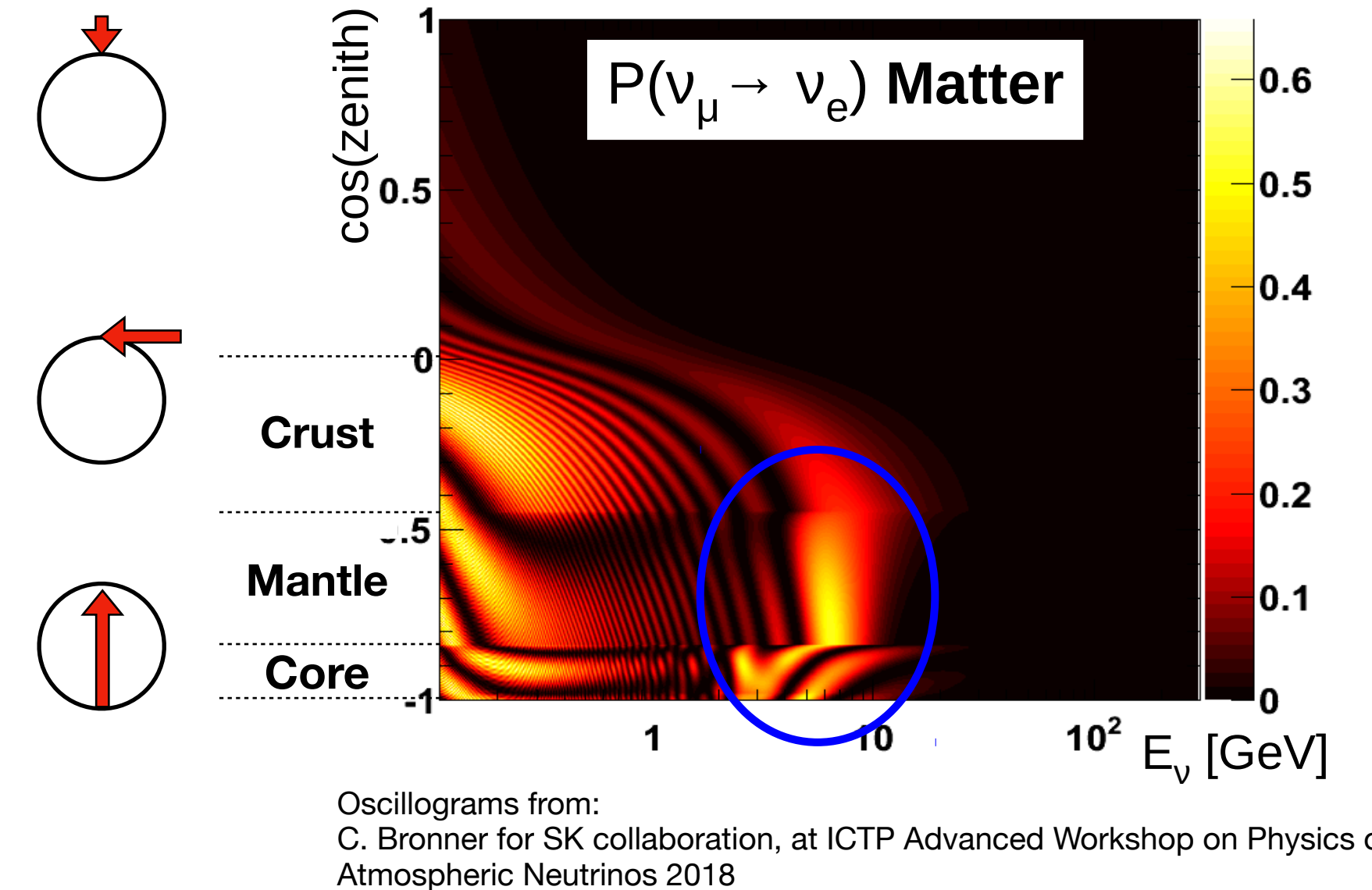
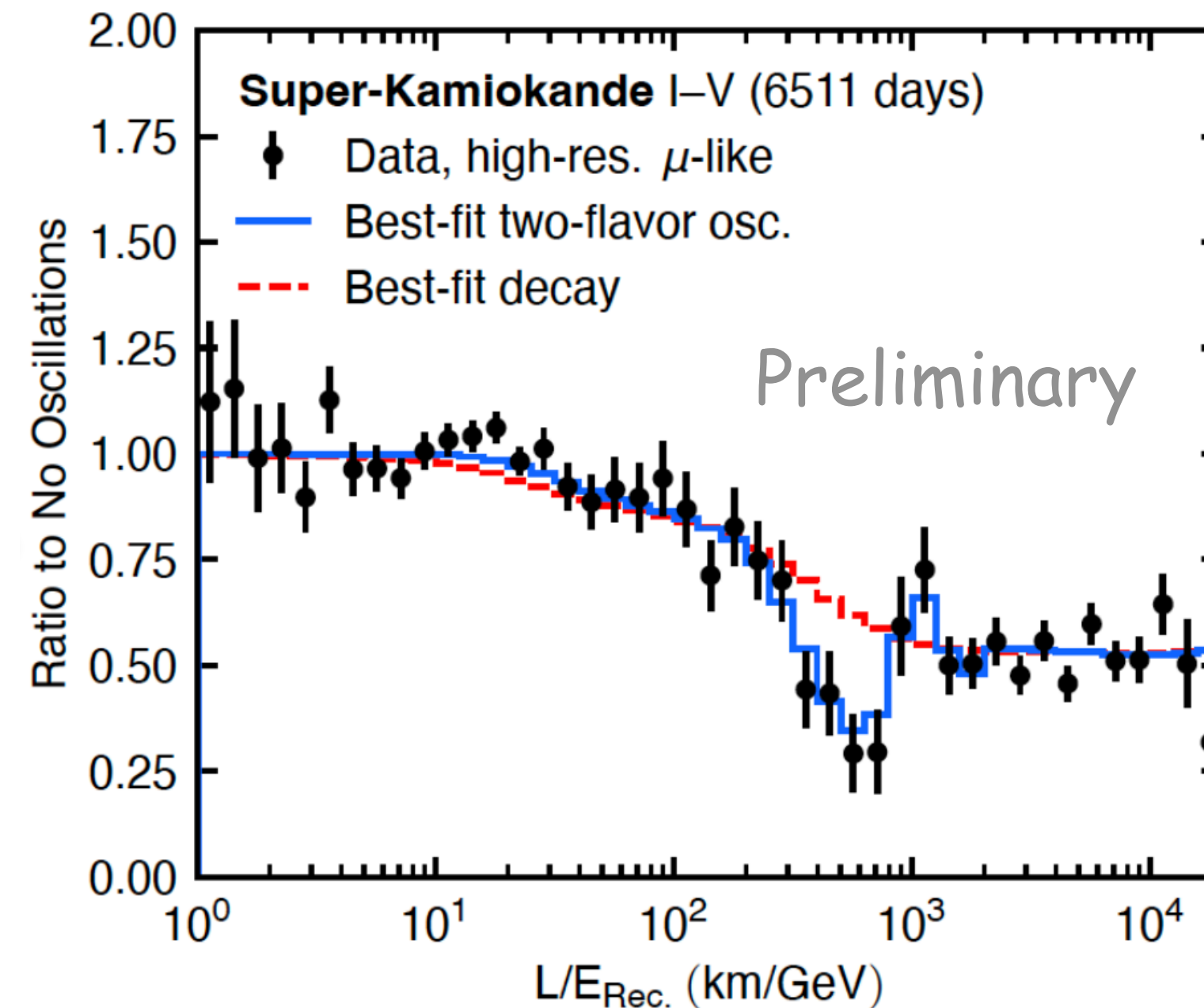
Atmospheric neutrinos



From K. Abe et al. PRD 97, 072001 (2018) + annotations

- ν_μ, ν_e produced from π^\pm, μ^\pm, K decay in atmosphere
- Neutrino zenith angle ($\rightarrow L$) inferred from measured charged lepton angle (better for $E_\nu \gg 1$ GeV)

Signals at multi-GeV



- First ν_μ oscillation dip can be resolved at high-E $\rightarrow |\Delta m_{32}^2|$
- Oscillation at larger L/E is smeared out to a ~ 0.5 reduction of ν_μ
 $\rightarrow \sin^2 2\theta_{23} \approx 1$

- Resonant enhancement of ν_e appearance in mantle/core at few GeV

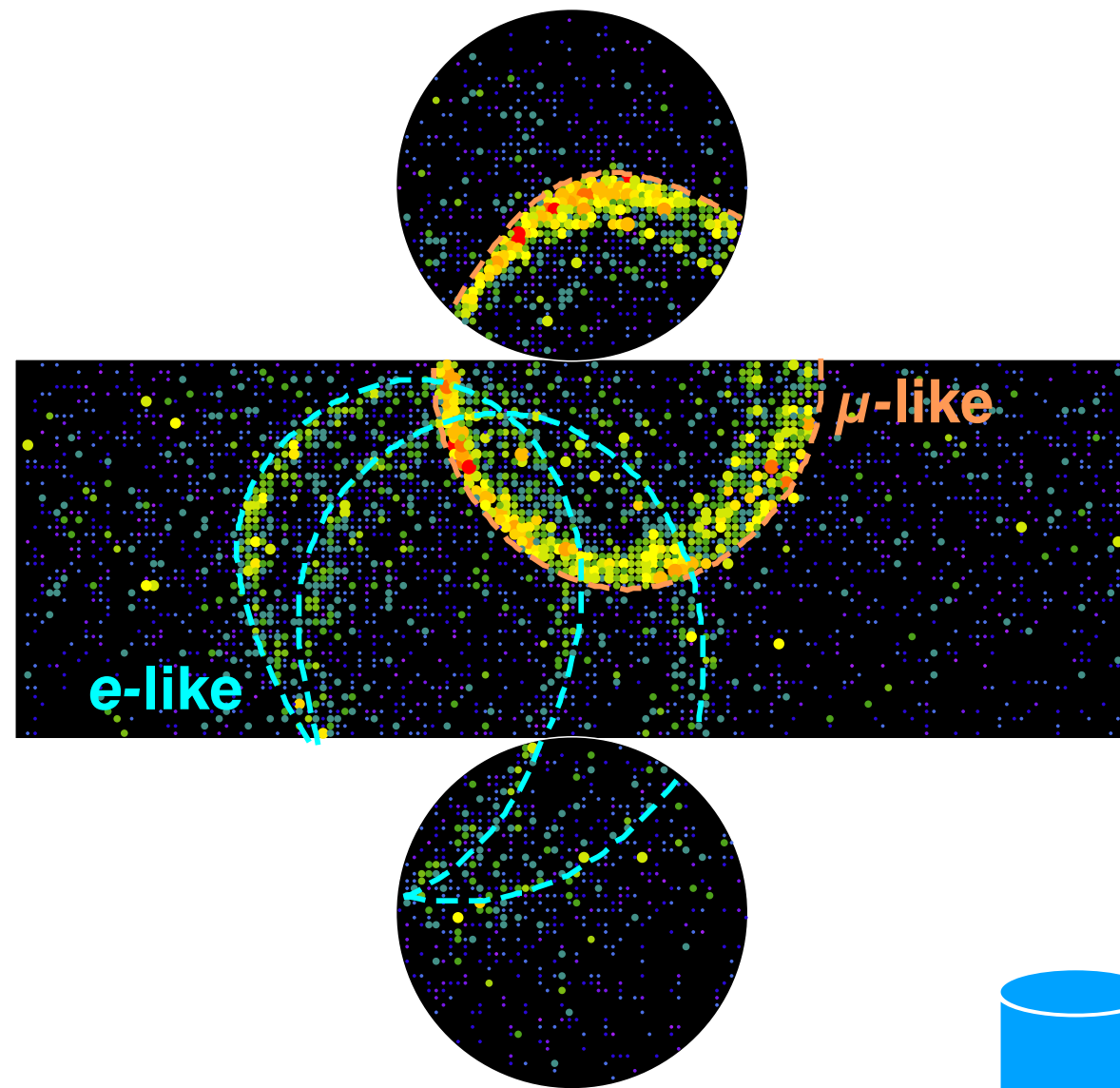
only for ν in NO \rightarrow MO
only for $\bar{\nu}$ in IO

- For flux $\nu \approx \bar{\nu}$, but ν cross-section is ~ 3 x larger \rightarrow charge ID optional

Atmospheric ν detectors

detectors drawn roughly to scale

Surface covering



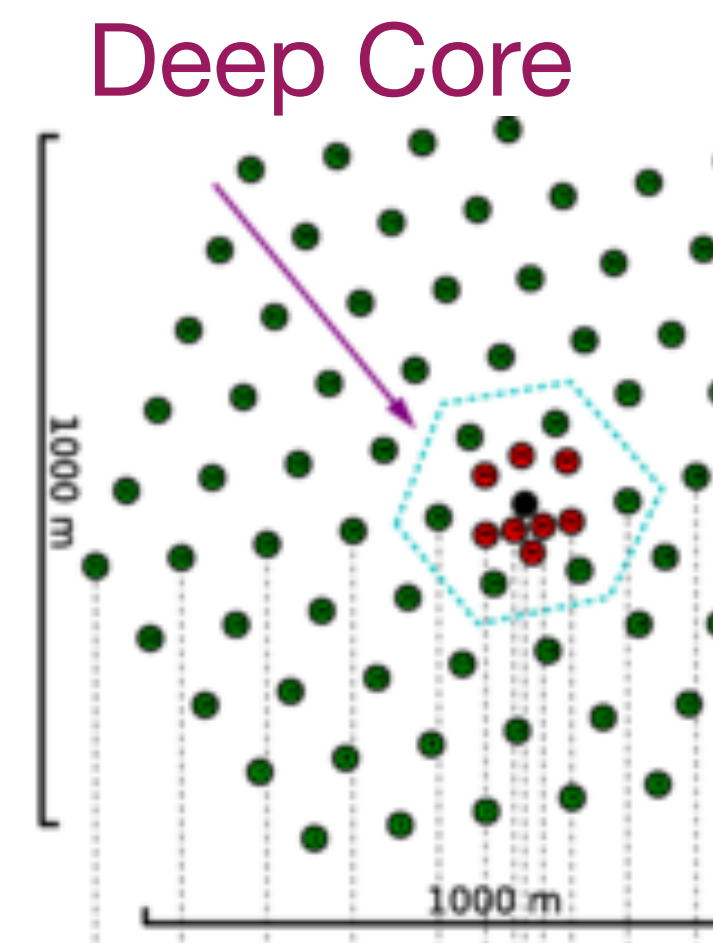
PhysRevD.109.072014

SuperK

HyperK

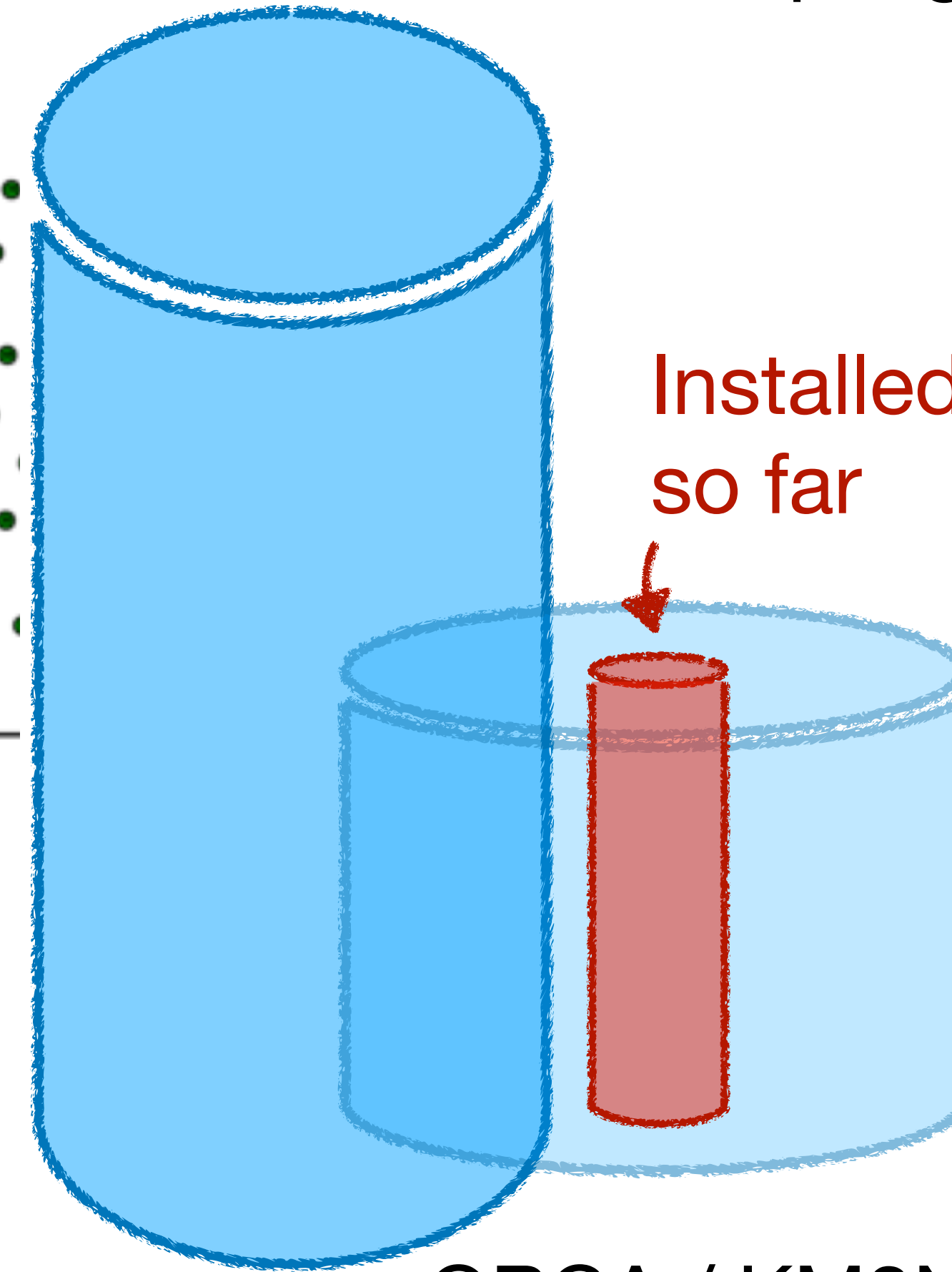
Detailed event
reconstruction over wide
 E_ν down to sub-GeV

3D sampling



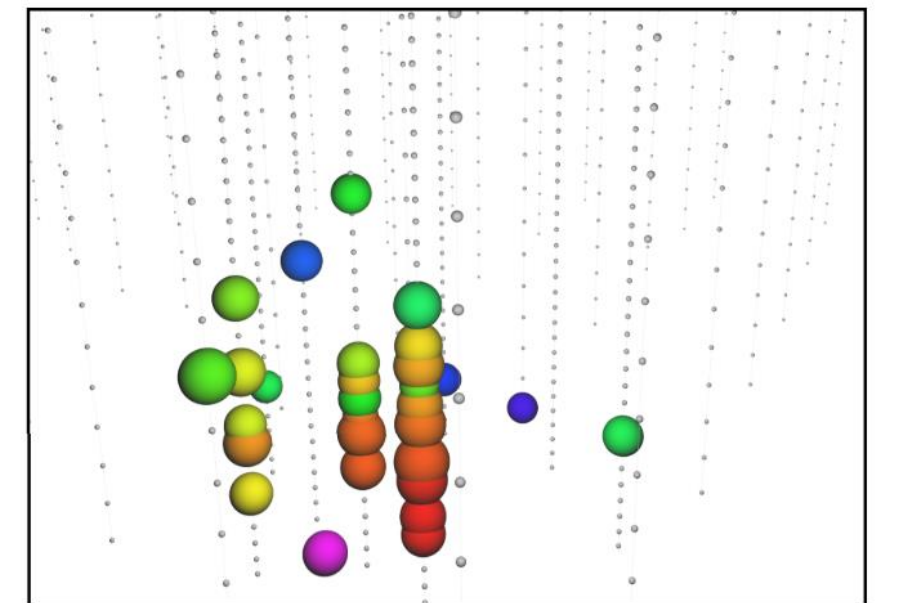
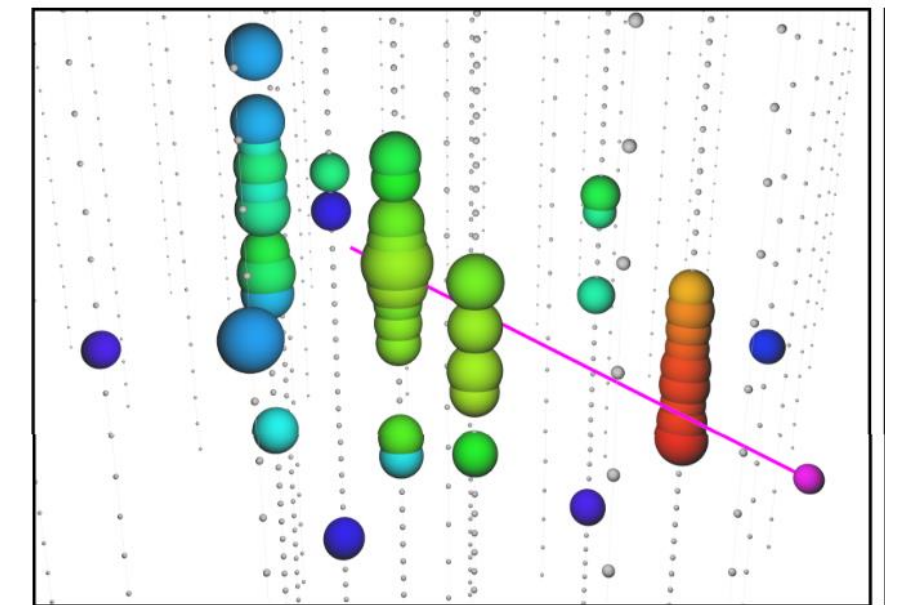
IceCube
(South Pole)

DeepCore
10 Mton



ORCA / KM3Net
(Mediterranean
Sea) 7 Mt

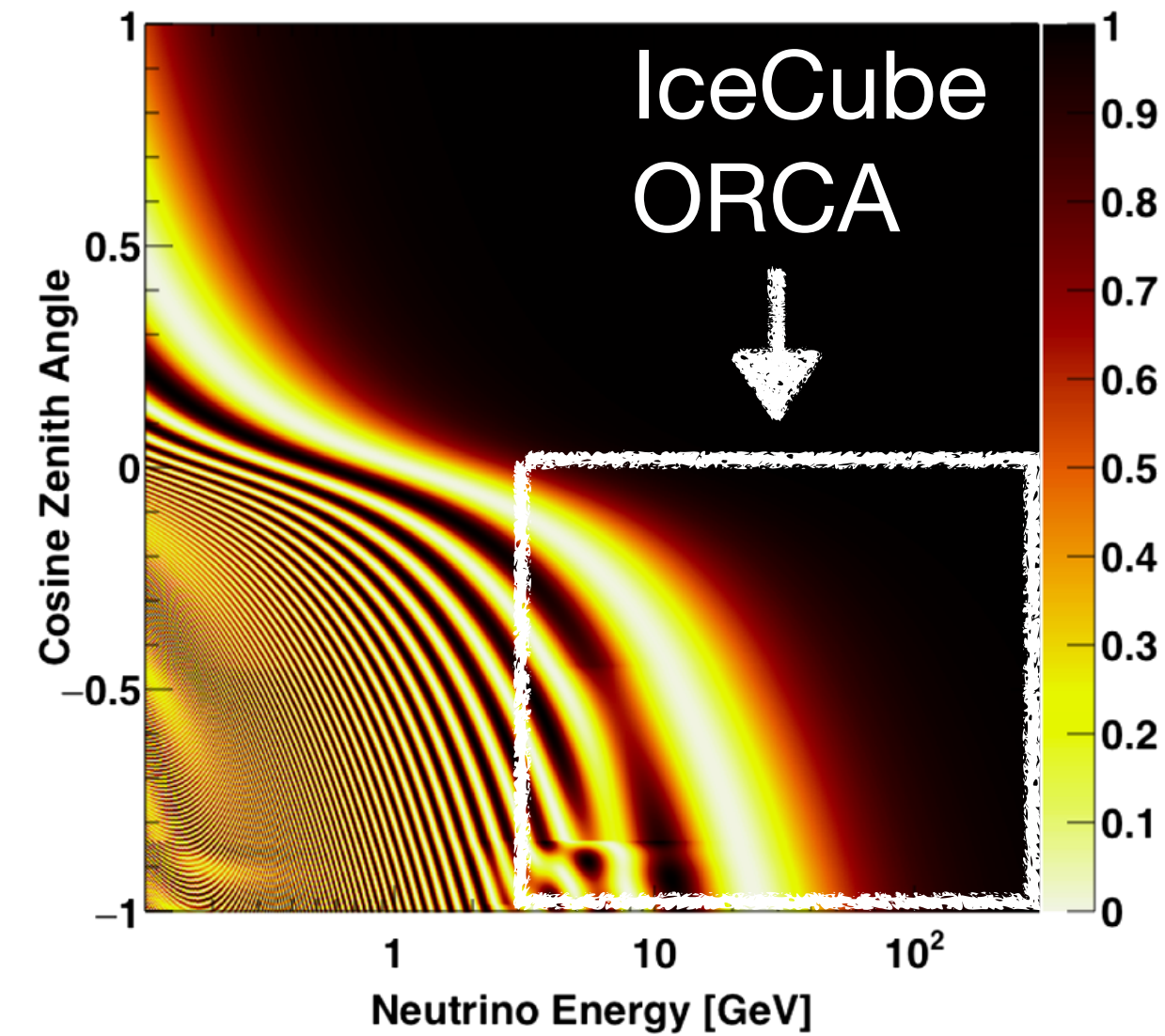
Track-like (ν_μ)



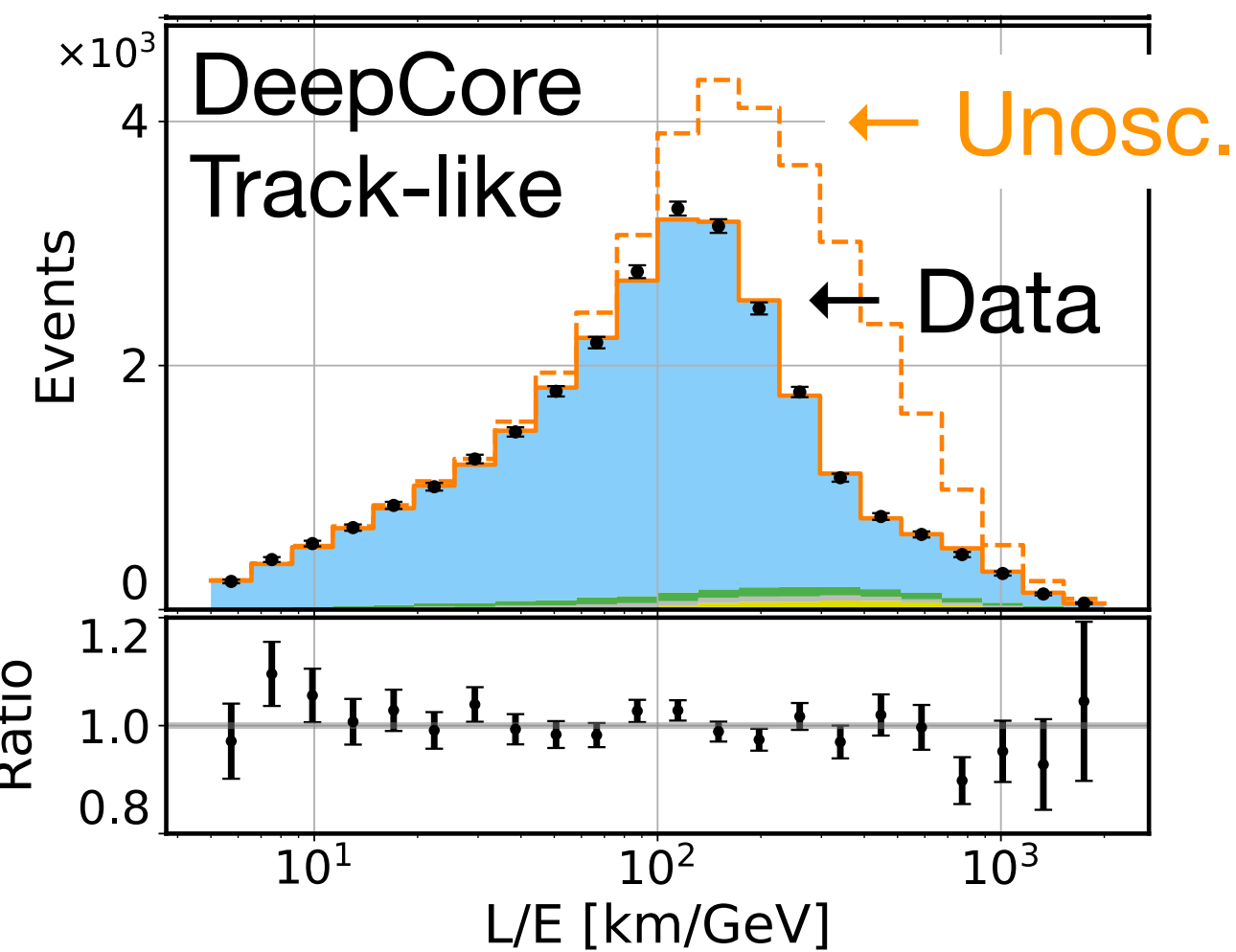
Shower-like
(ν_e , NC, ...)

Ice Cube, Neutrino 2024

ν_μ disappearance



SuperK

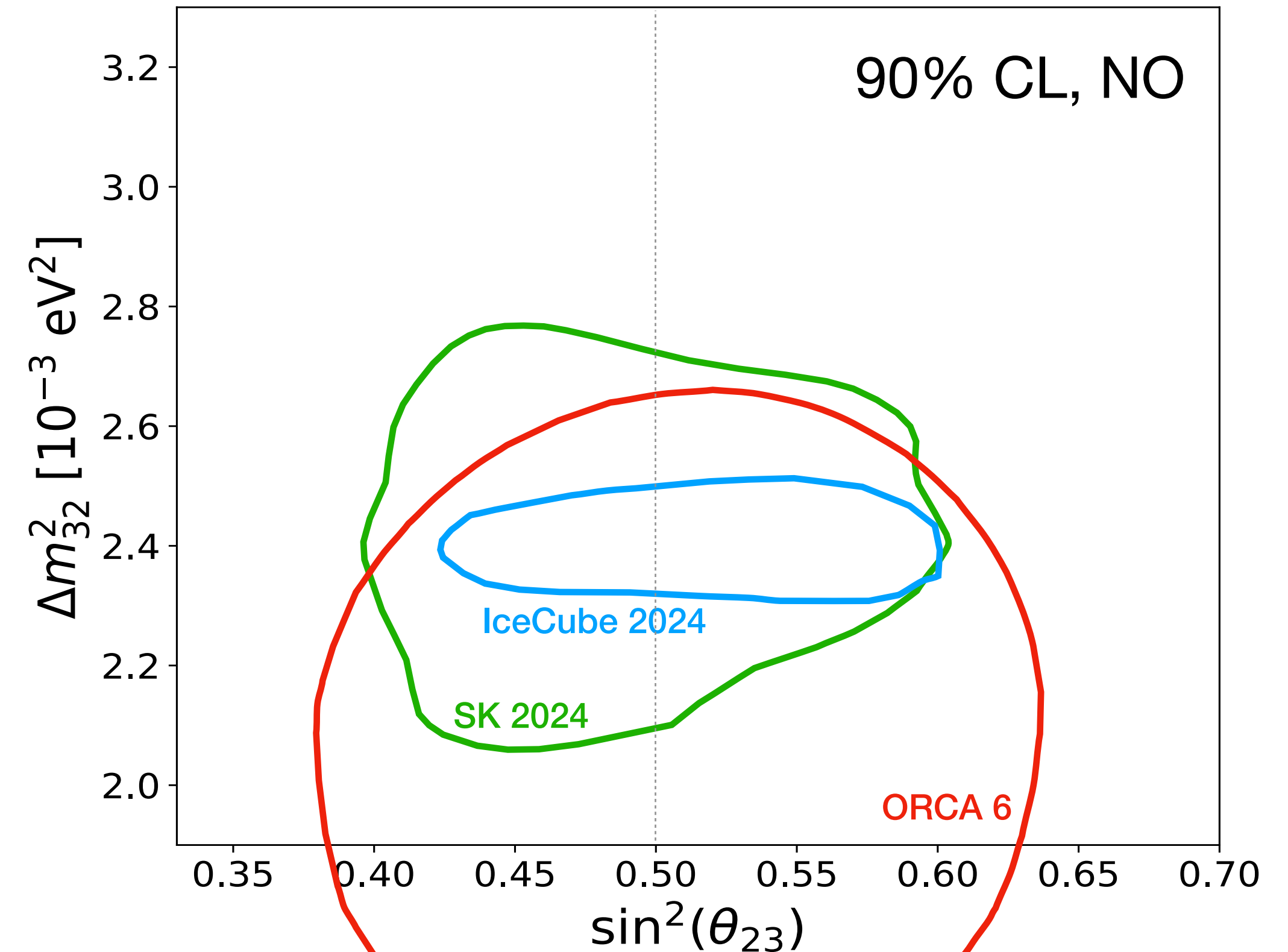
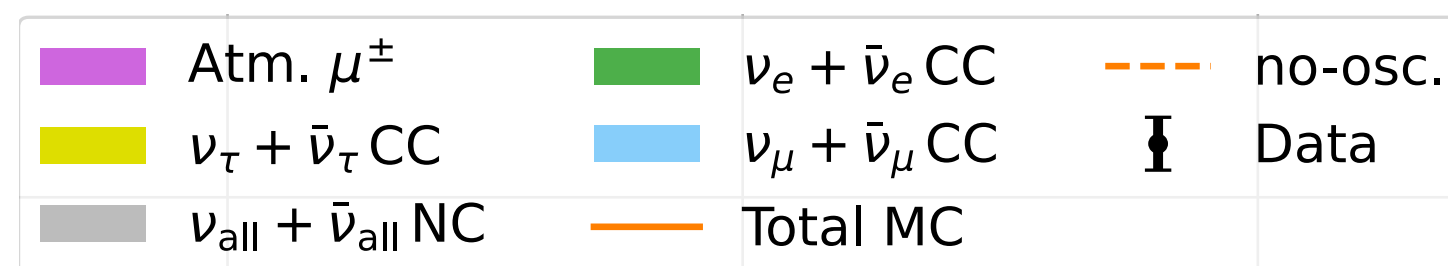


- SuperK:** 20% expansion of fiducial volume + additional years.
PRD 109, 072014 (2024)

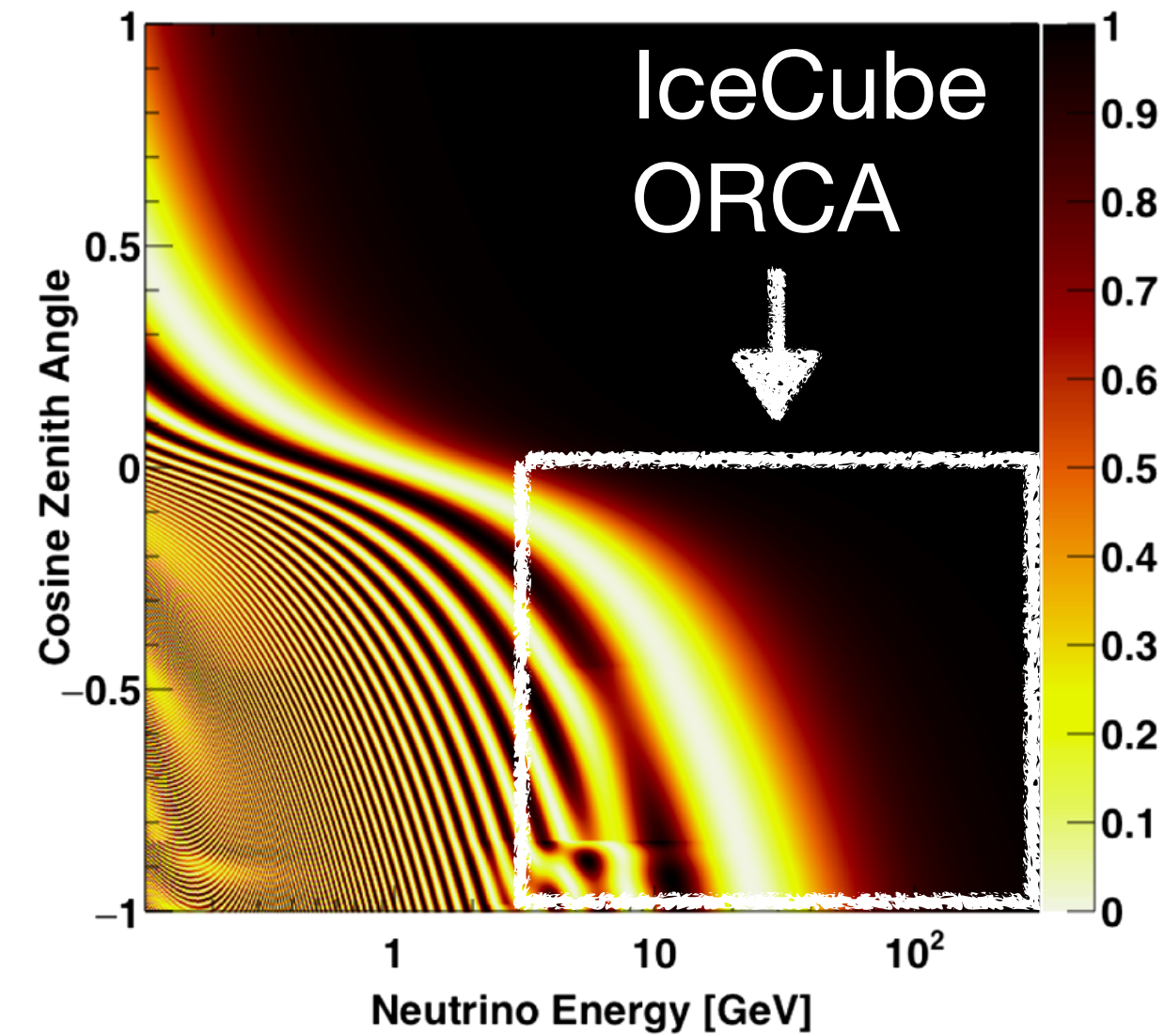
In total 48% statistics increase over previous publication

- DeepCore** moved to CNN-based reconstruction + 2 years
↓
7x increase in statistics
arXiv:2405.02163

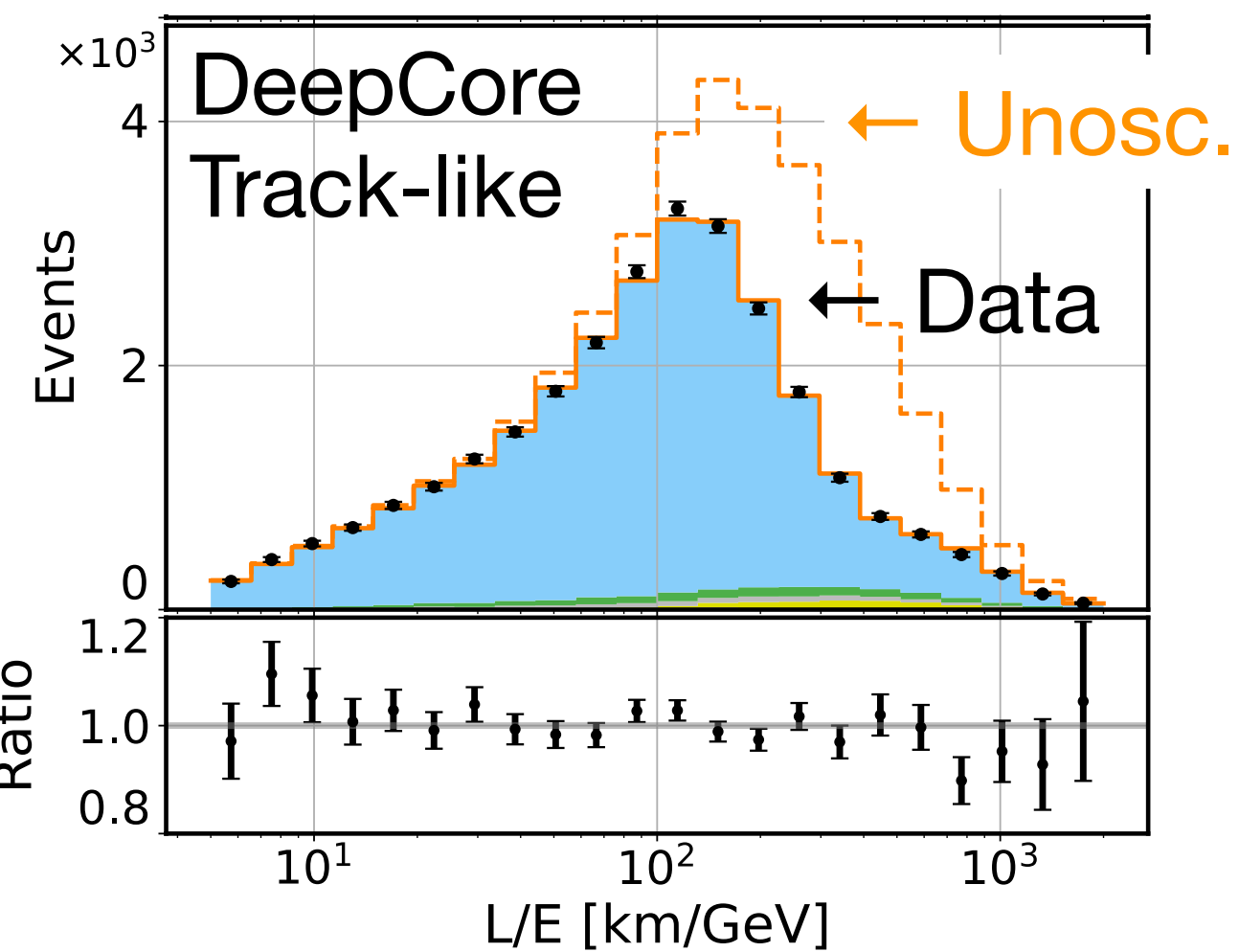
expect 2x reduction in next 4 years



ν_μ disappearance



SuperK

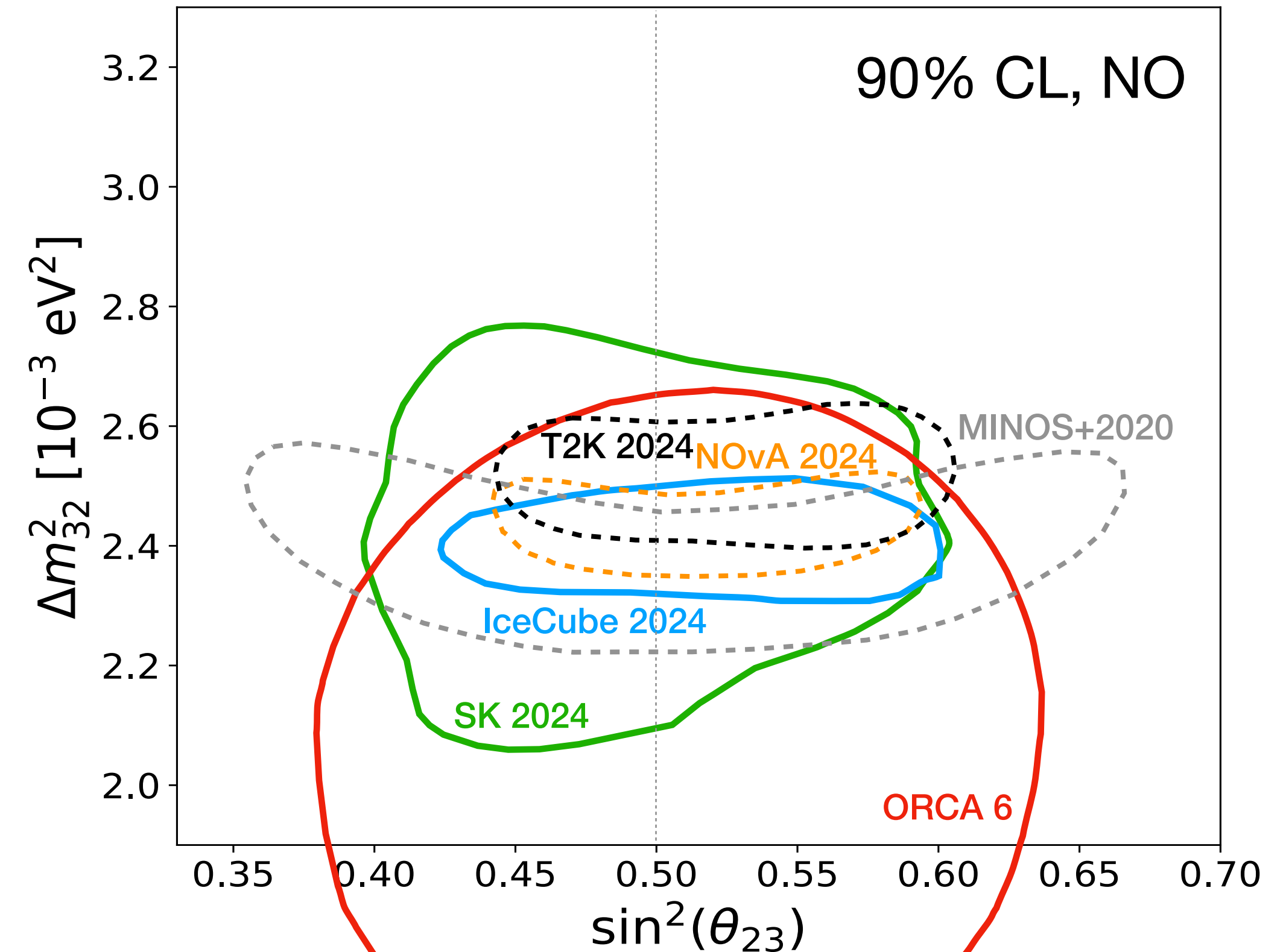


- **SuperK:** 20% expansion of fiducial volume + additional years.
PRD 109, 072014 (2024)

In total 48% statistics increase over previous publication

- **DeepCore** moved to CNN-based reconstruction + 2 years
↓
7x increase in statistics
arXiv:2405.02163

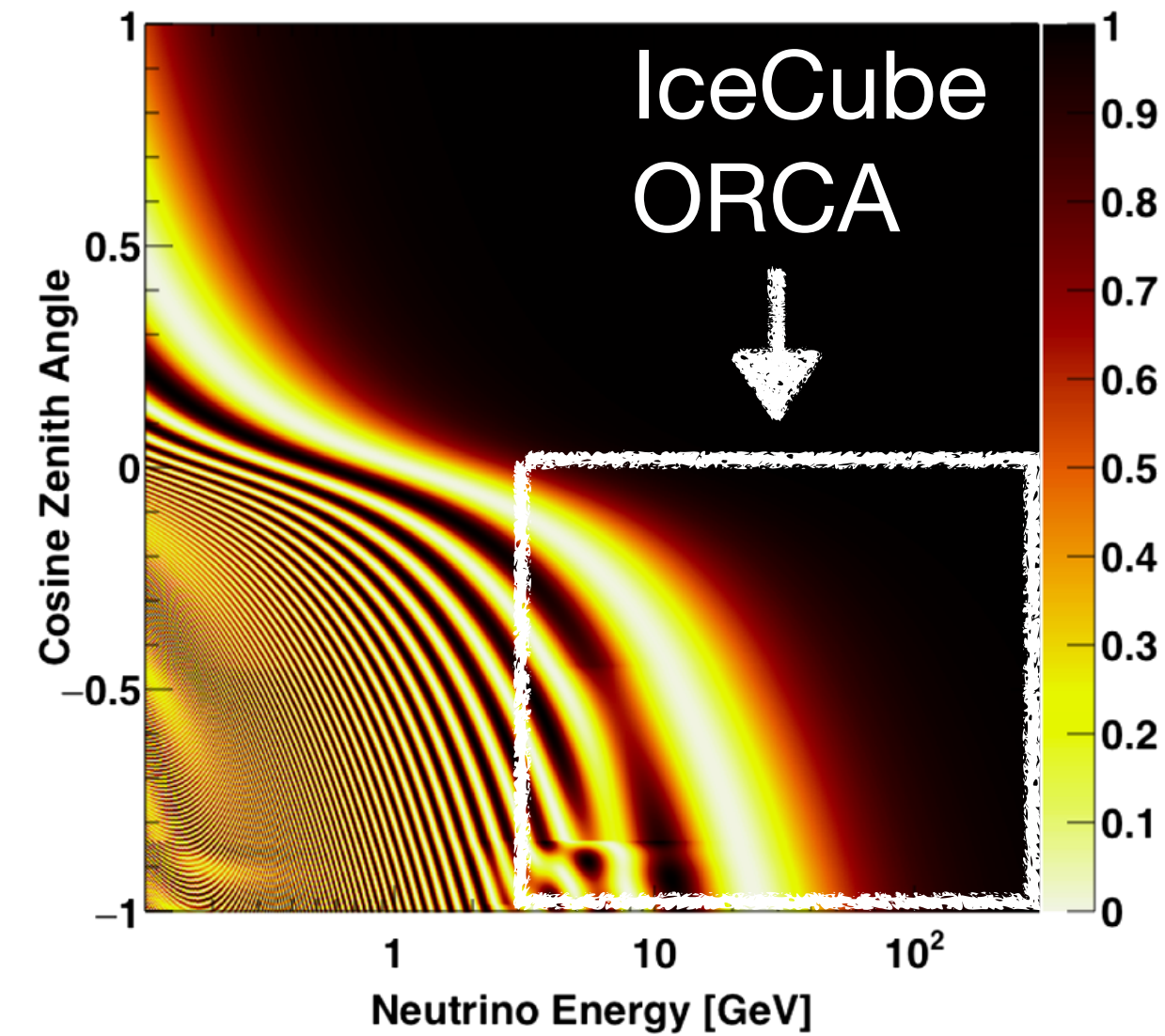
expect 2x reduction in next 4 years



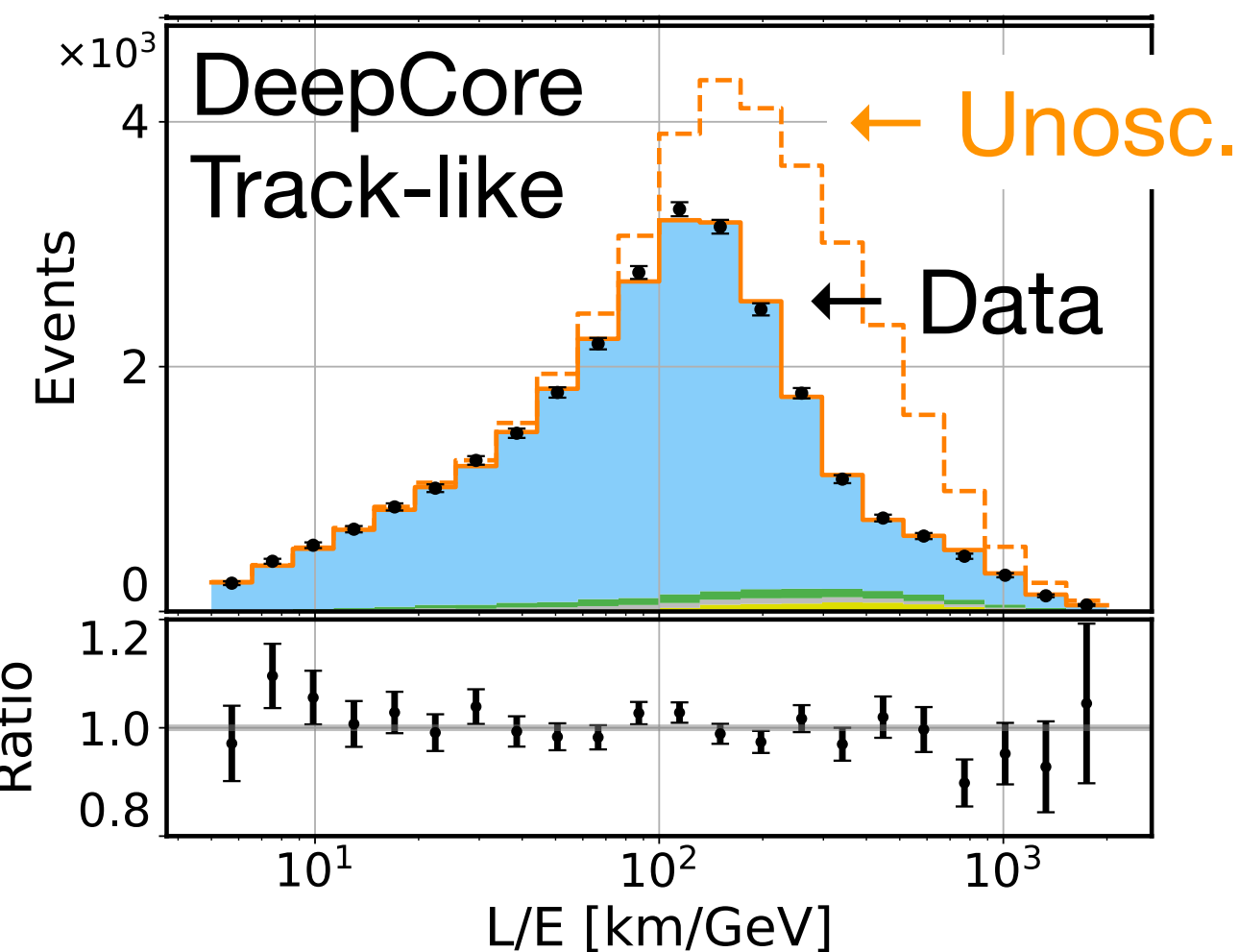
Consistent results, and now competitive (IC) with accelerator measurements

note: 20~50x longer baselines and energies

ν_μ disappearance



SuperK



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PRD 109, 072014 (2024)

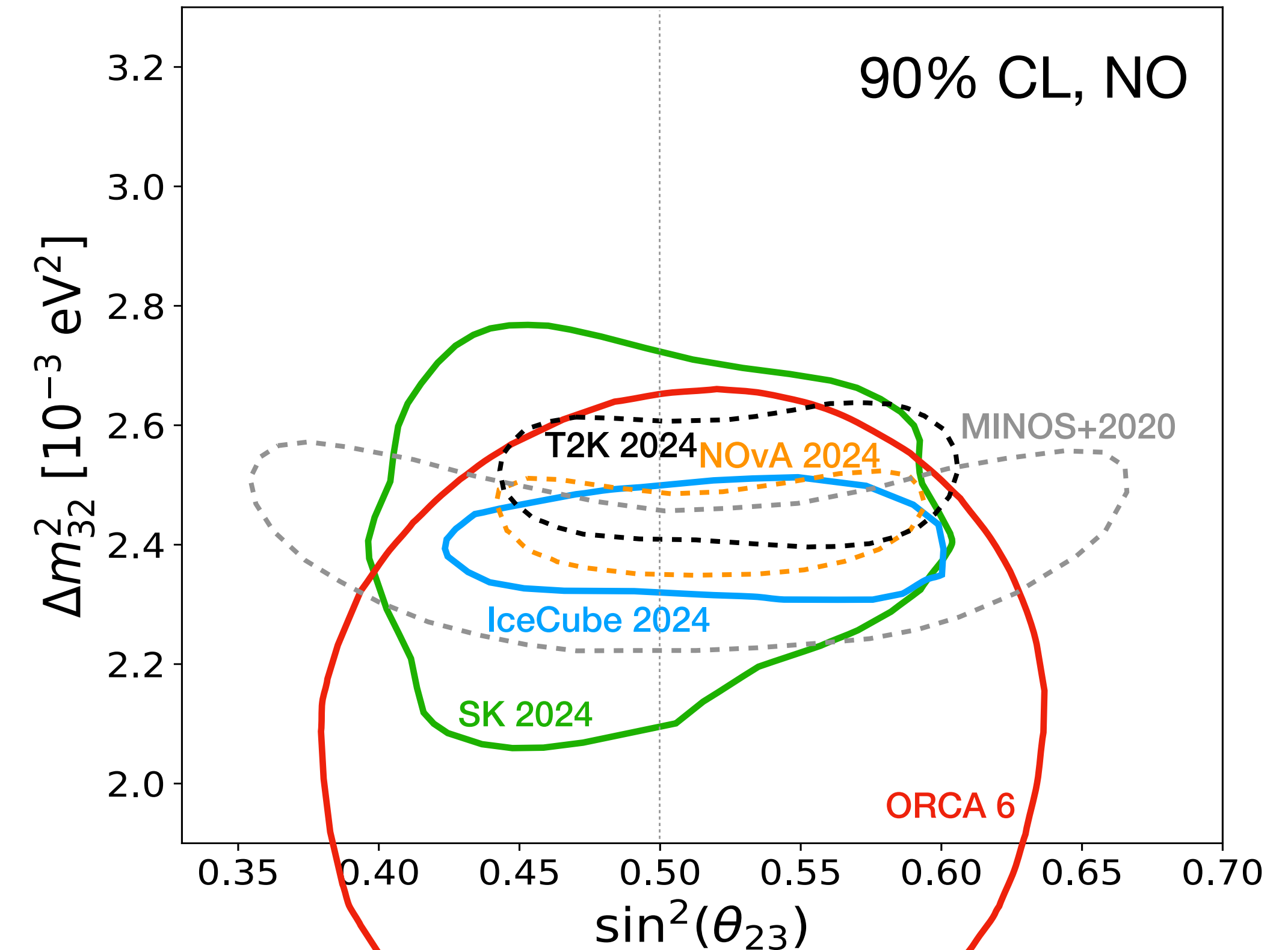
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- DeepCore** moved to CNN-based reconstruction + 2 years
↓
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expect 2x reduction in next 4 years

- ORCA** first result using ~5% of final detector size. Expect improvement in statistics and systematics as more strings get installed.

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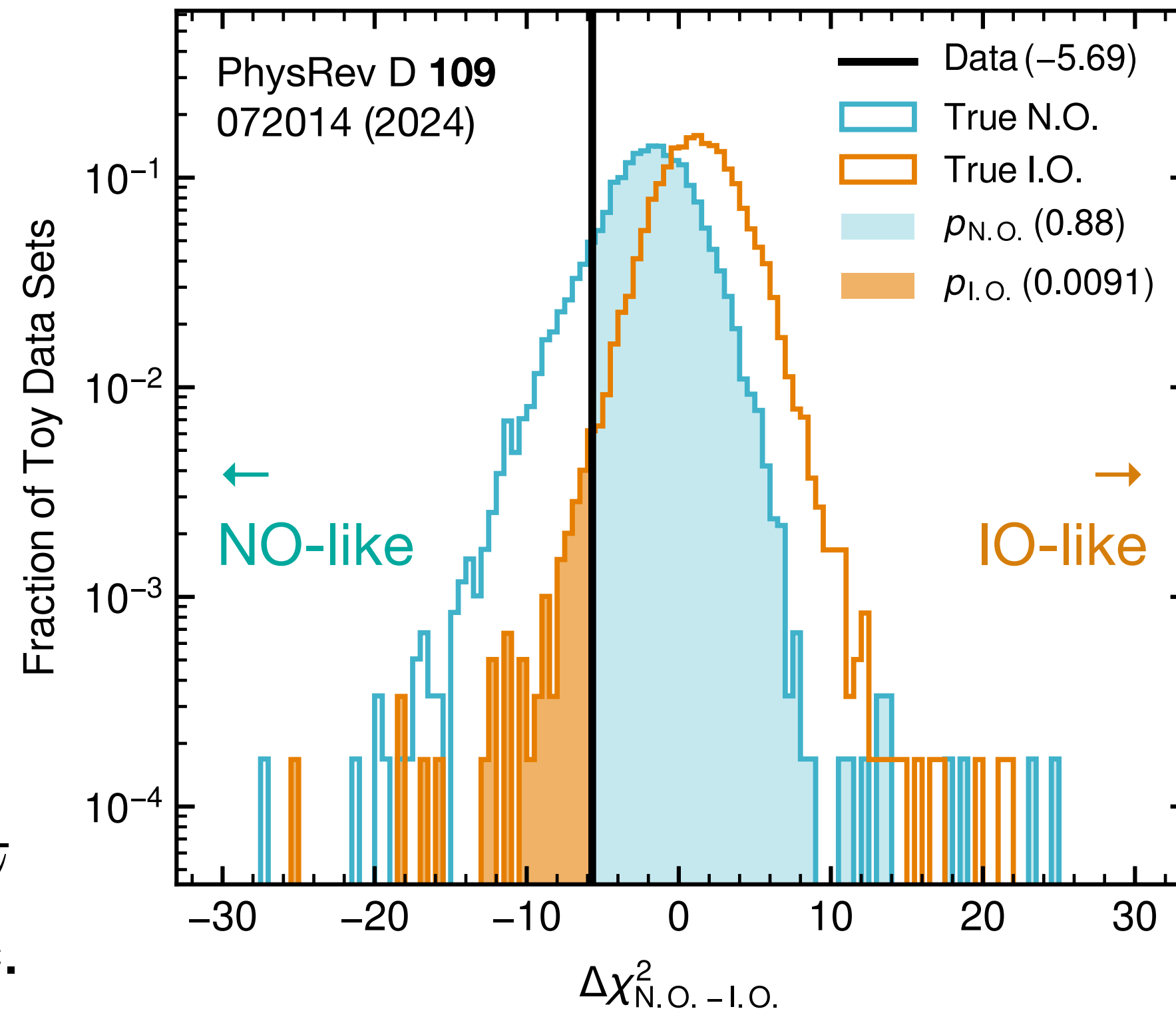
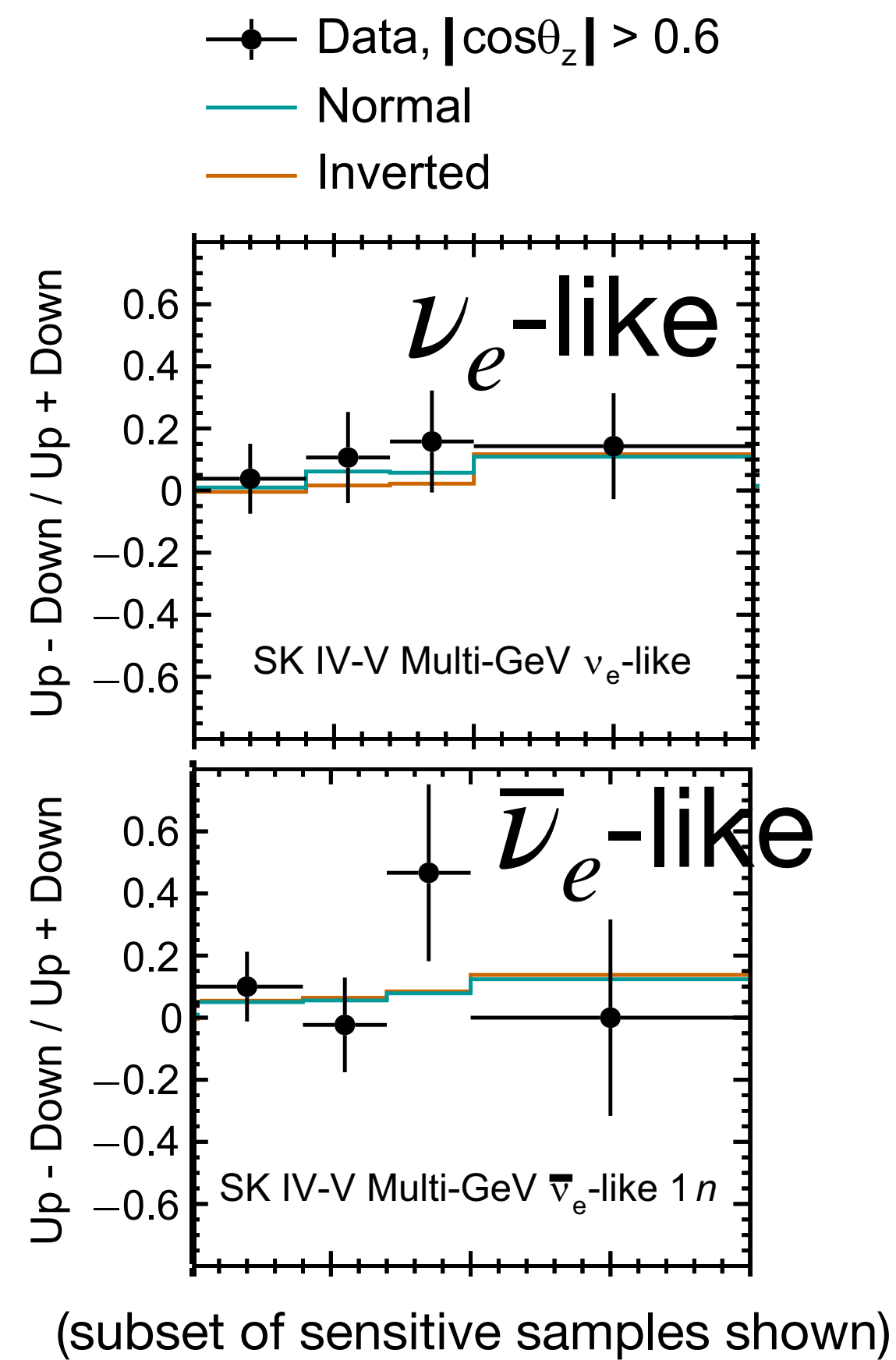
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Mass ordering

From matter resonance

SK, 484 kt years

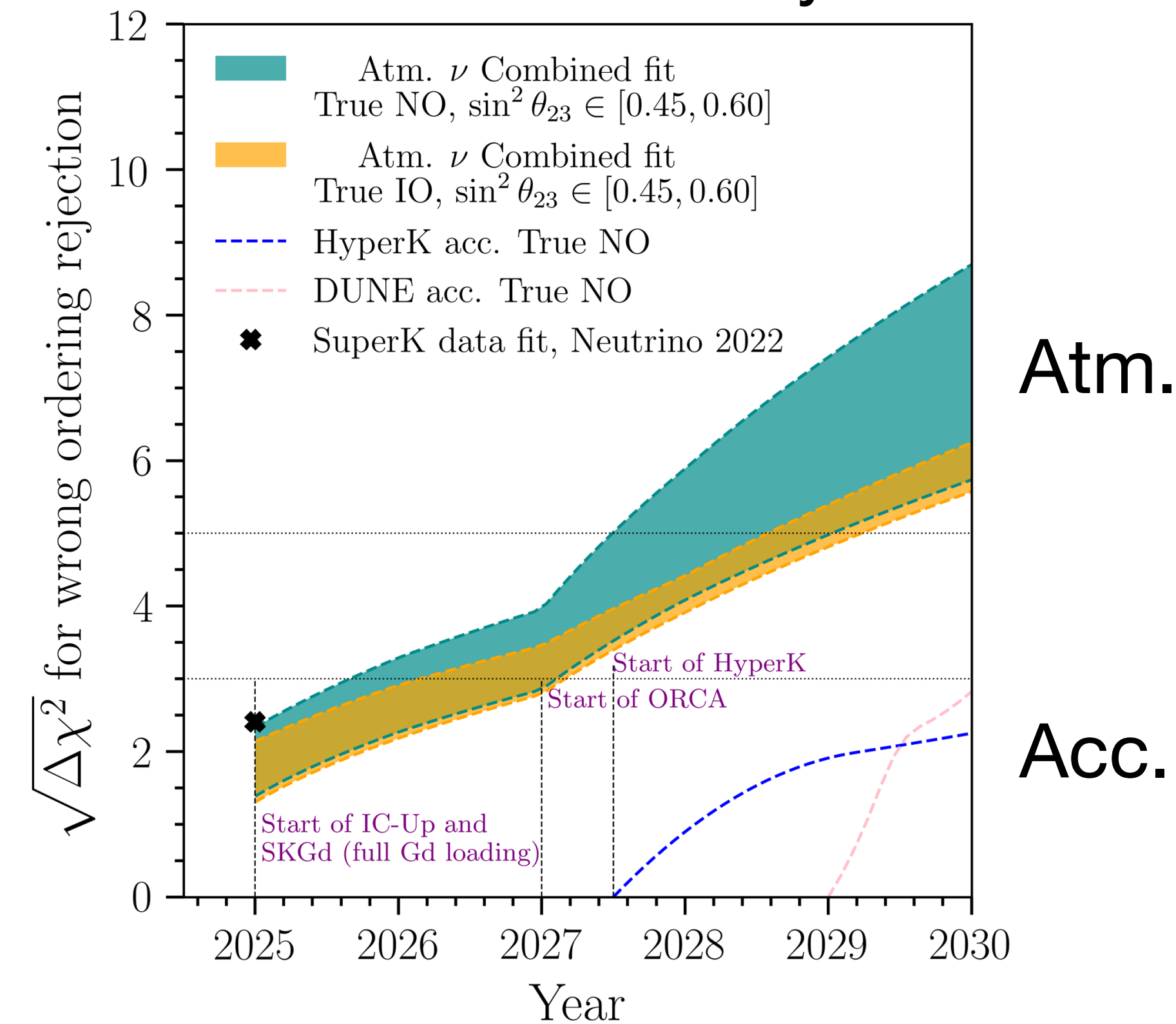


Indication of NO.

Rejection of IO at 92.3% CL_s.

Dominated by stats, then xsec

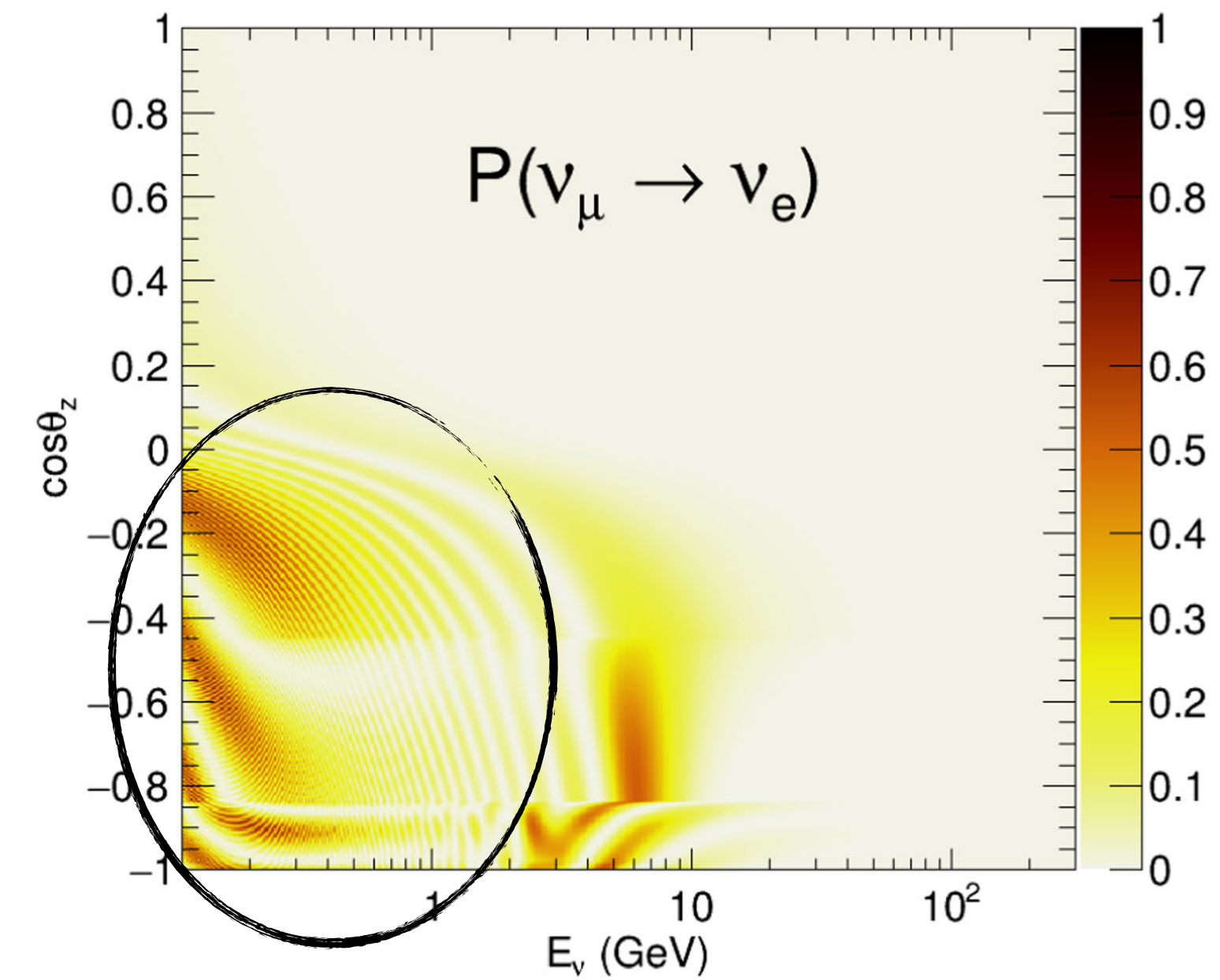
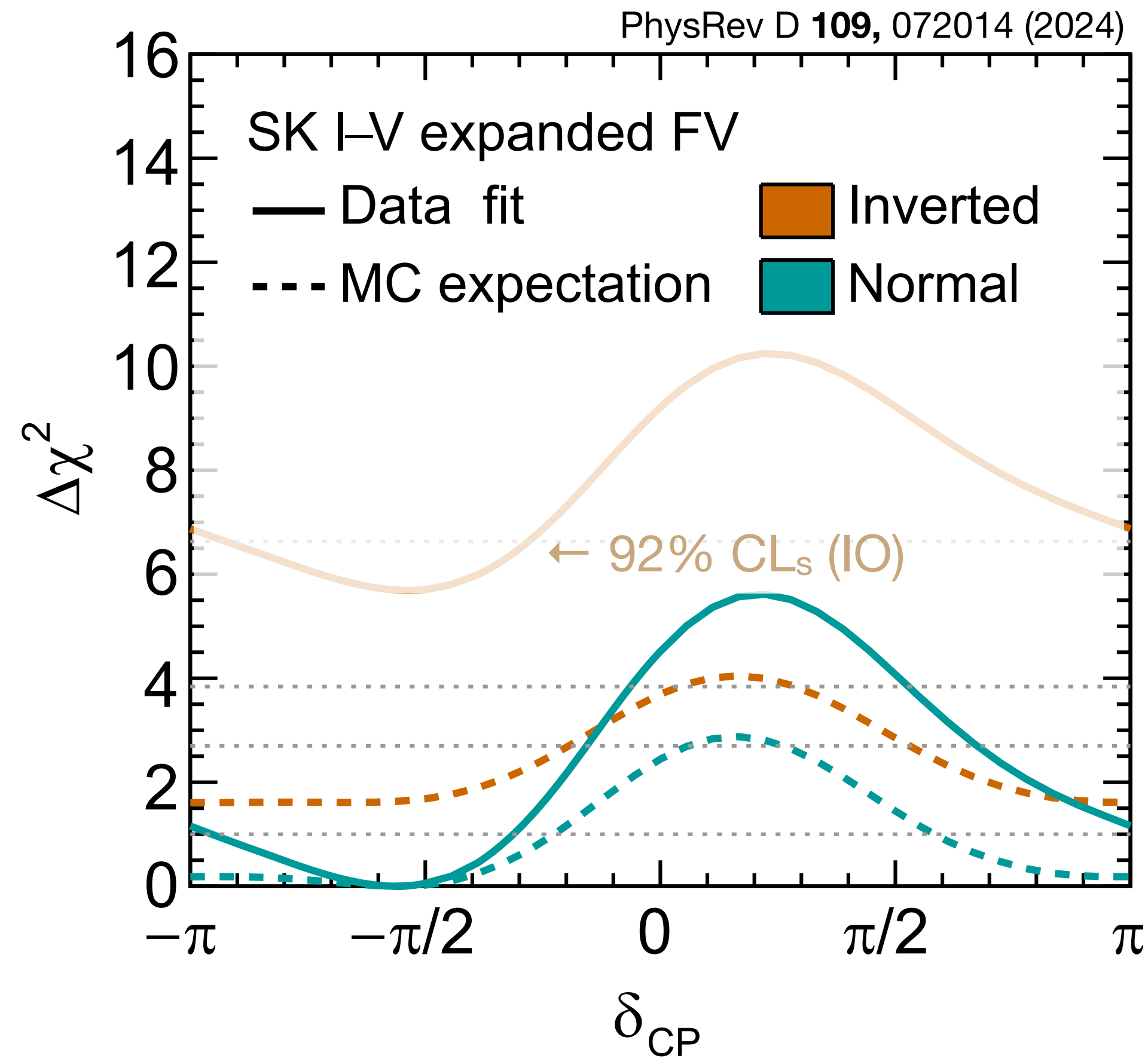
Future sensitivity



IceCube upgrade, ORCA, and HyperK expected to measure MO in the next few years. PRX 13, 041055 (2023)

- SK performs statistical $\nu/\bar{\nu}$ separation using # of π etc. to enhance purity
- Updates to selection
Multi-Ring: likelihood \rightarrow BDT
Single-Ring: + neutron tag

CP phase



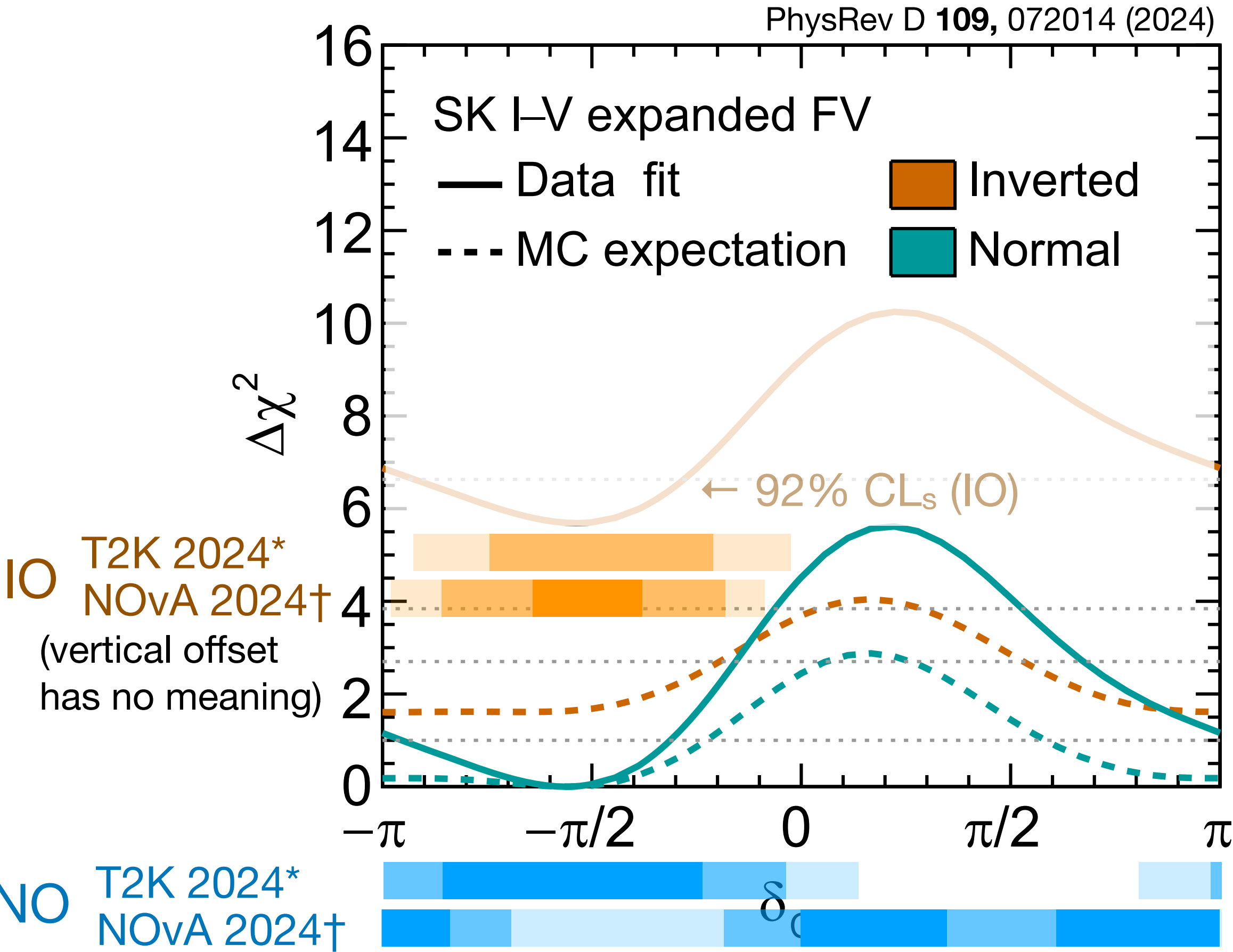
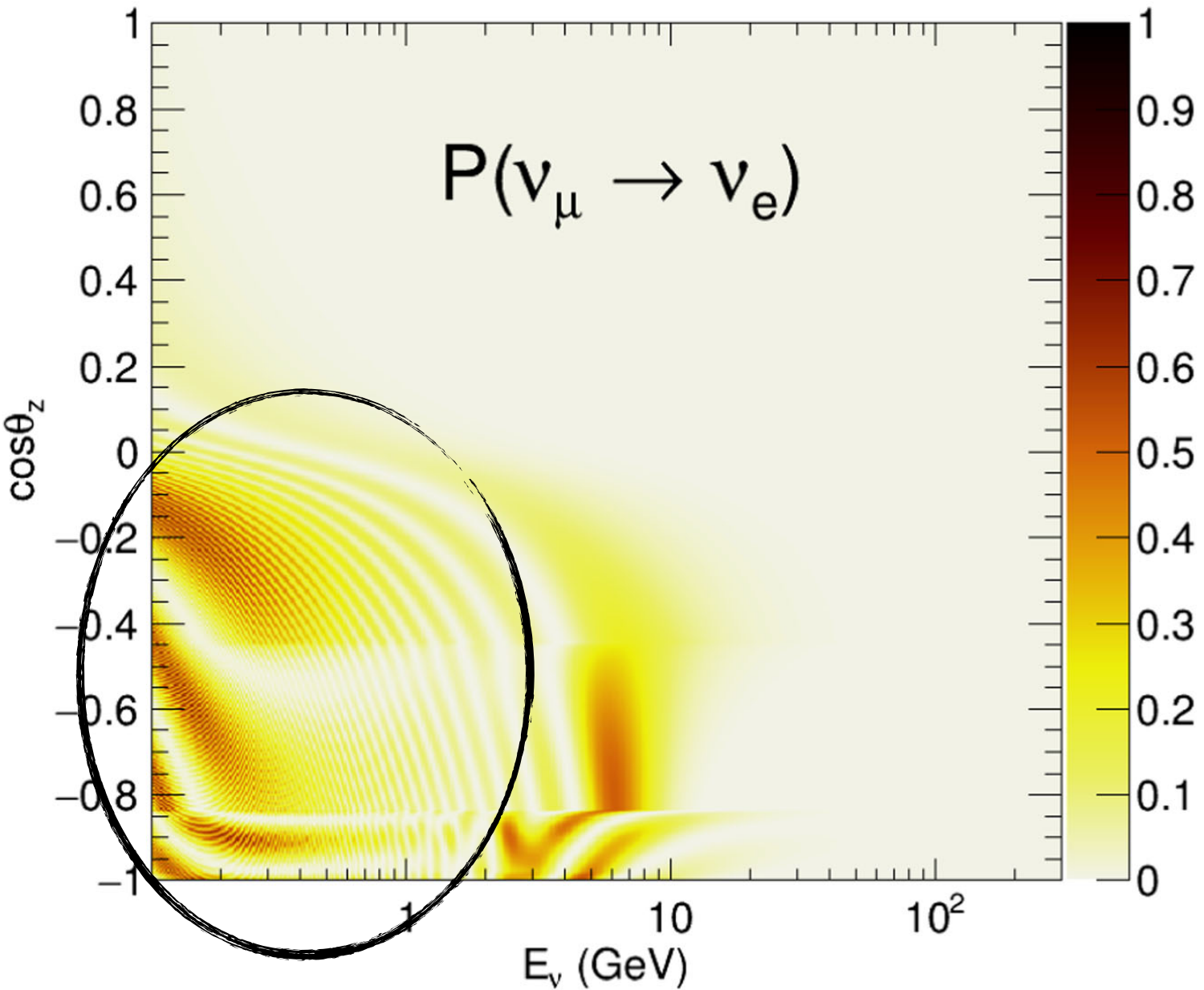
- SK also constrains δ_{CP} from normalization of sub-GeV e -like events. Unlike accelerators, decoupled from MO. Interplay of $\Delta m_{32}^2 - \Delta m_{21}^2$ interference phase-shift and flux/xsec shape, fully smeared due to resolution.
- Weak indication of maximal CP violation but CP conservation still allowed
Caveat: somewhat stronger exclusion than sensitivity. Due to parameter boundaries and degeneracies cannot take $\sqrt{\Delta\chi^2}$ as sigmas.

Using θ_{13} constraint from reactors

* Feldman-Cousins

† Bayesian

CP phase

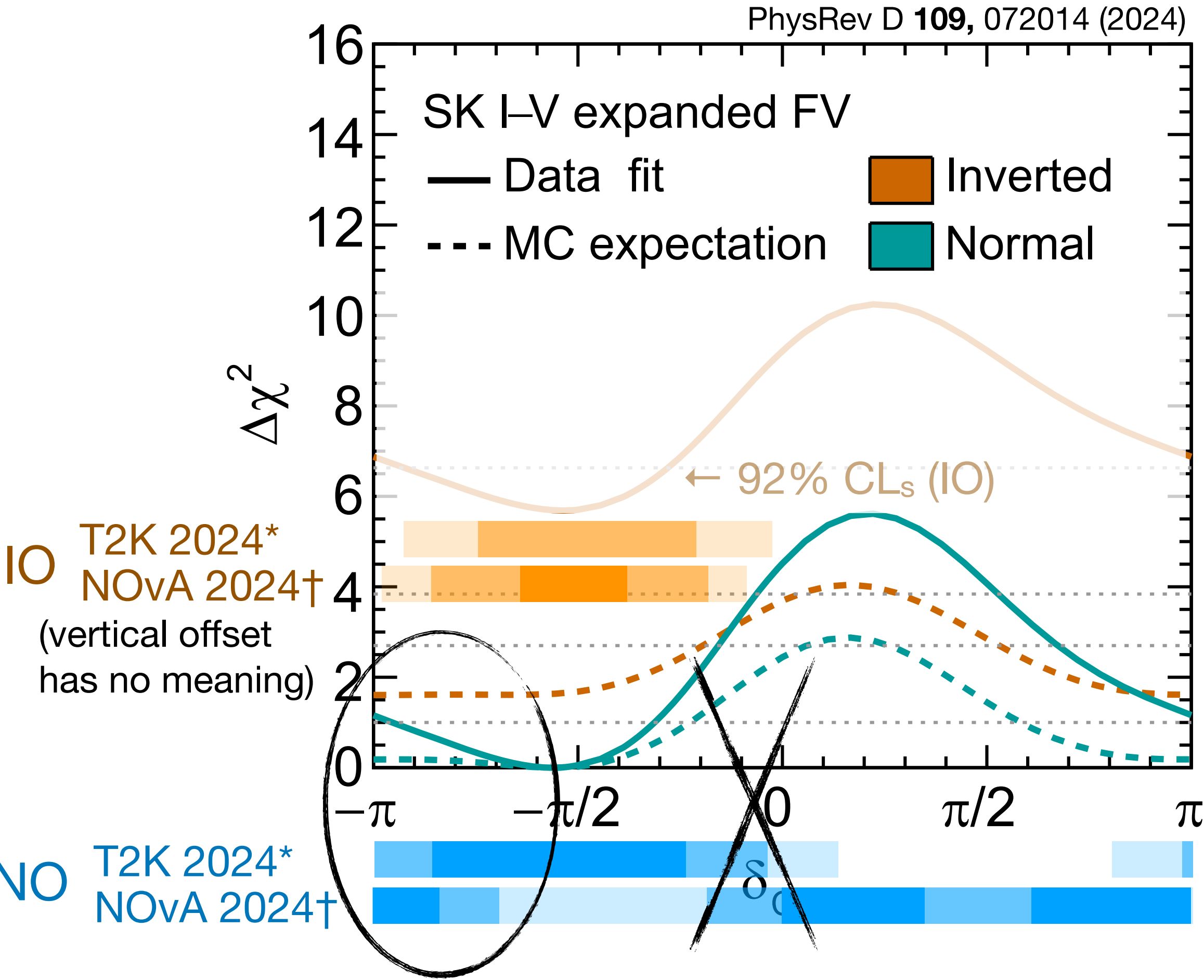
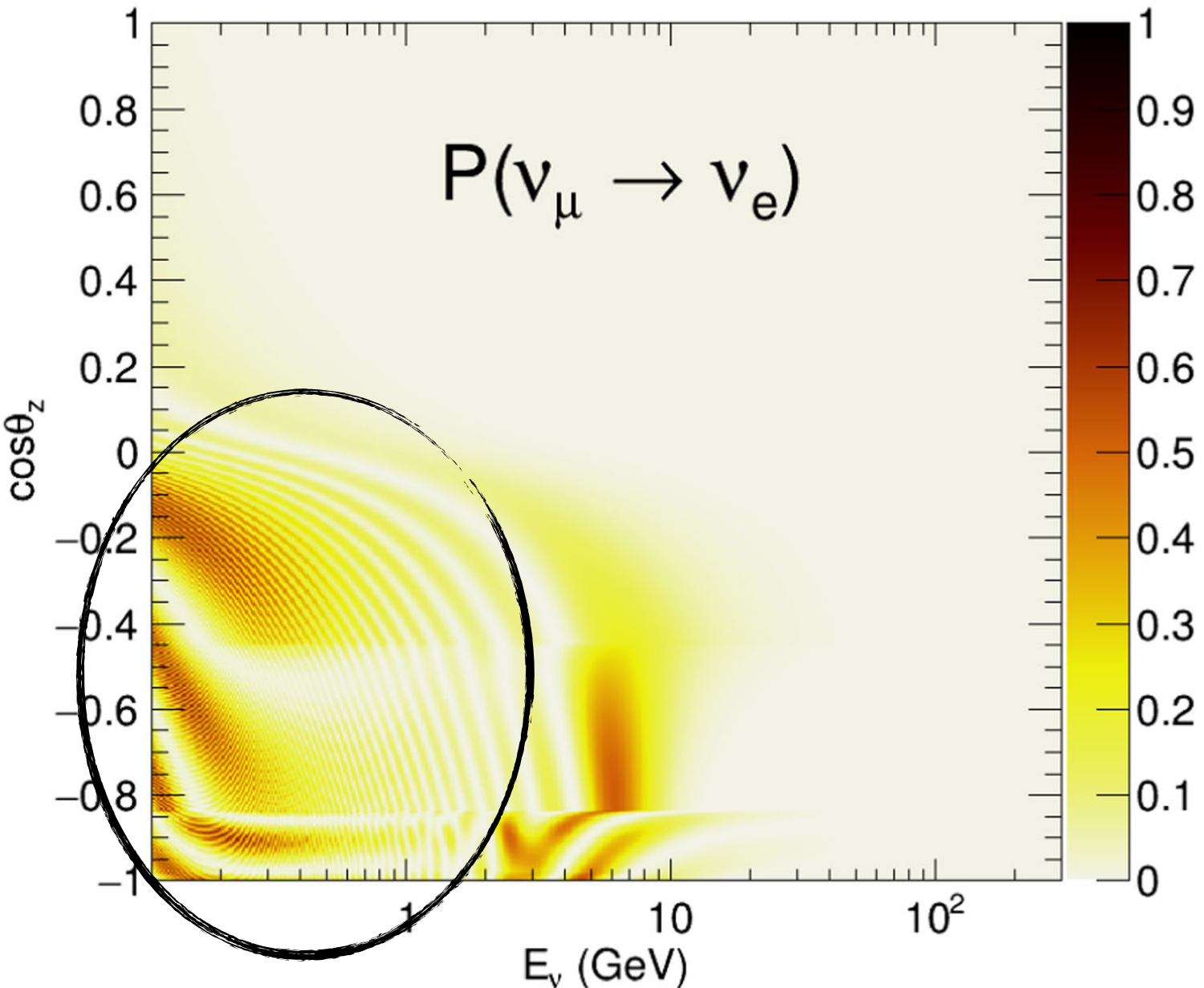


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Interesting contribution to NOvA and T2K's NO constraints.

Using θ_{13} constraint from reactors

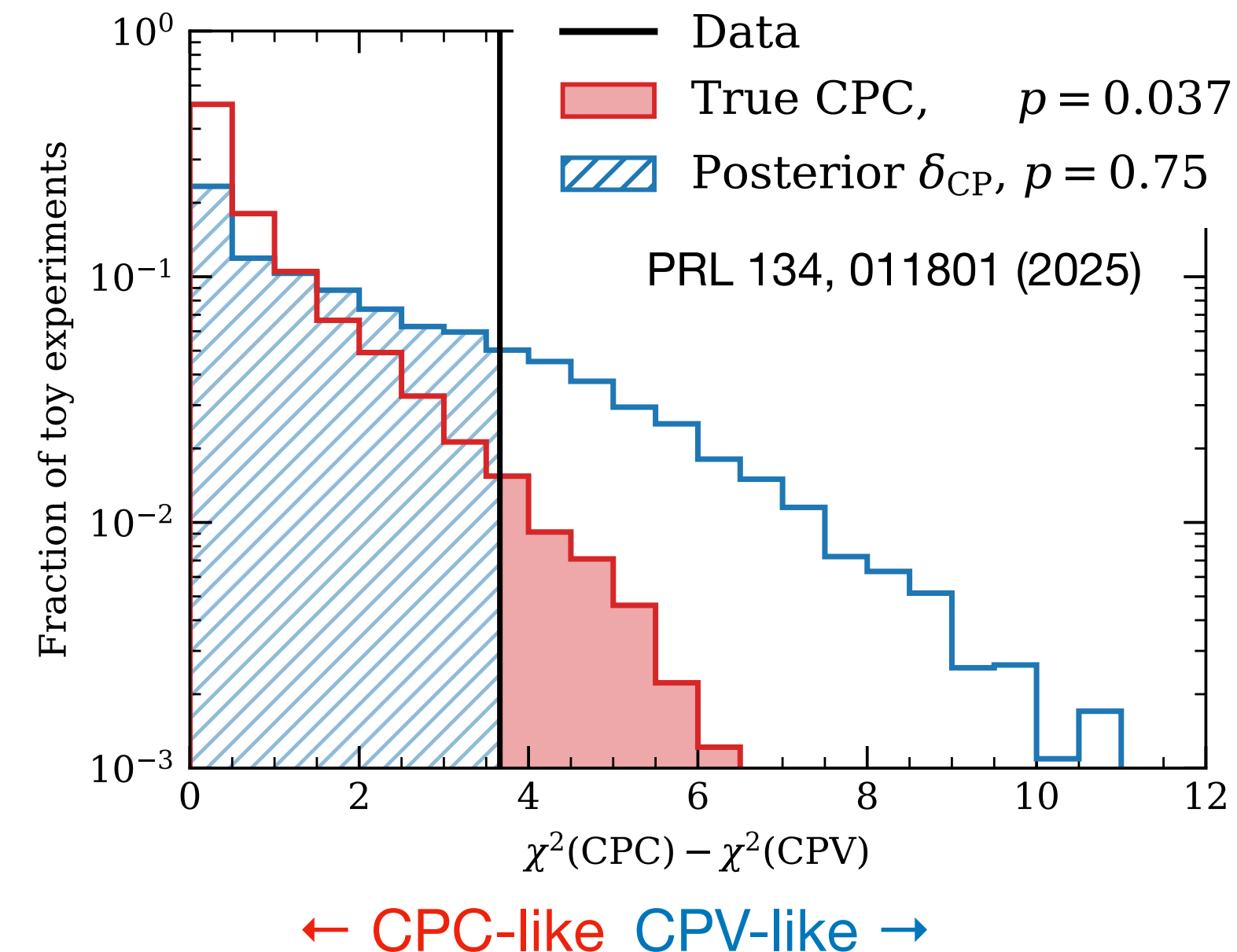
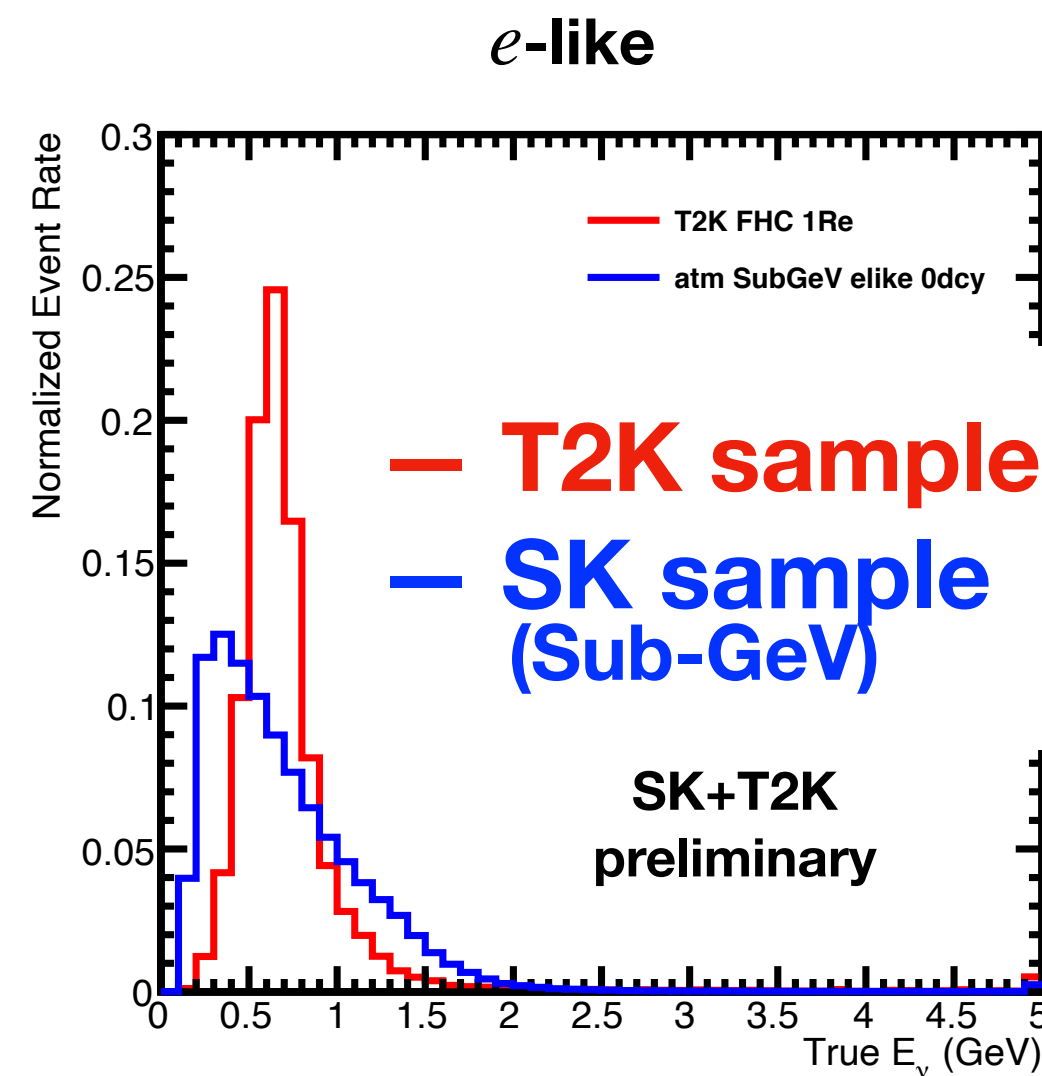
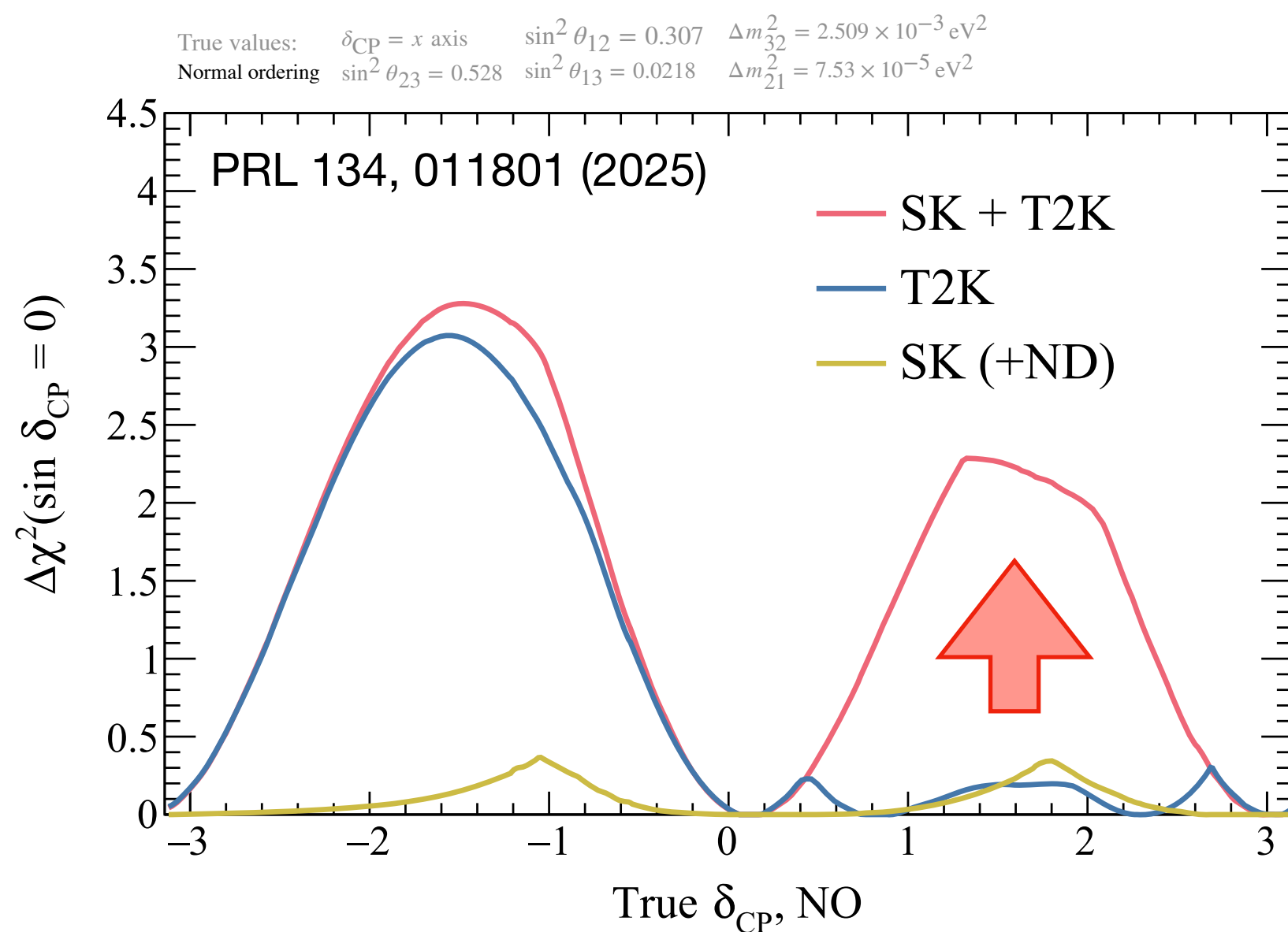
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SK IV (atm) + T2K (acc)

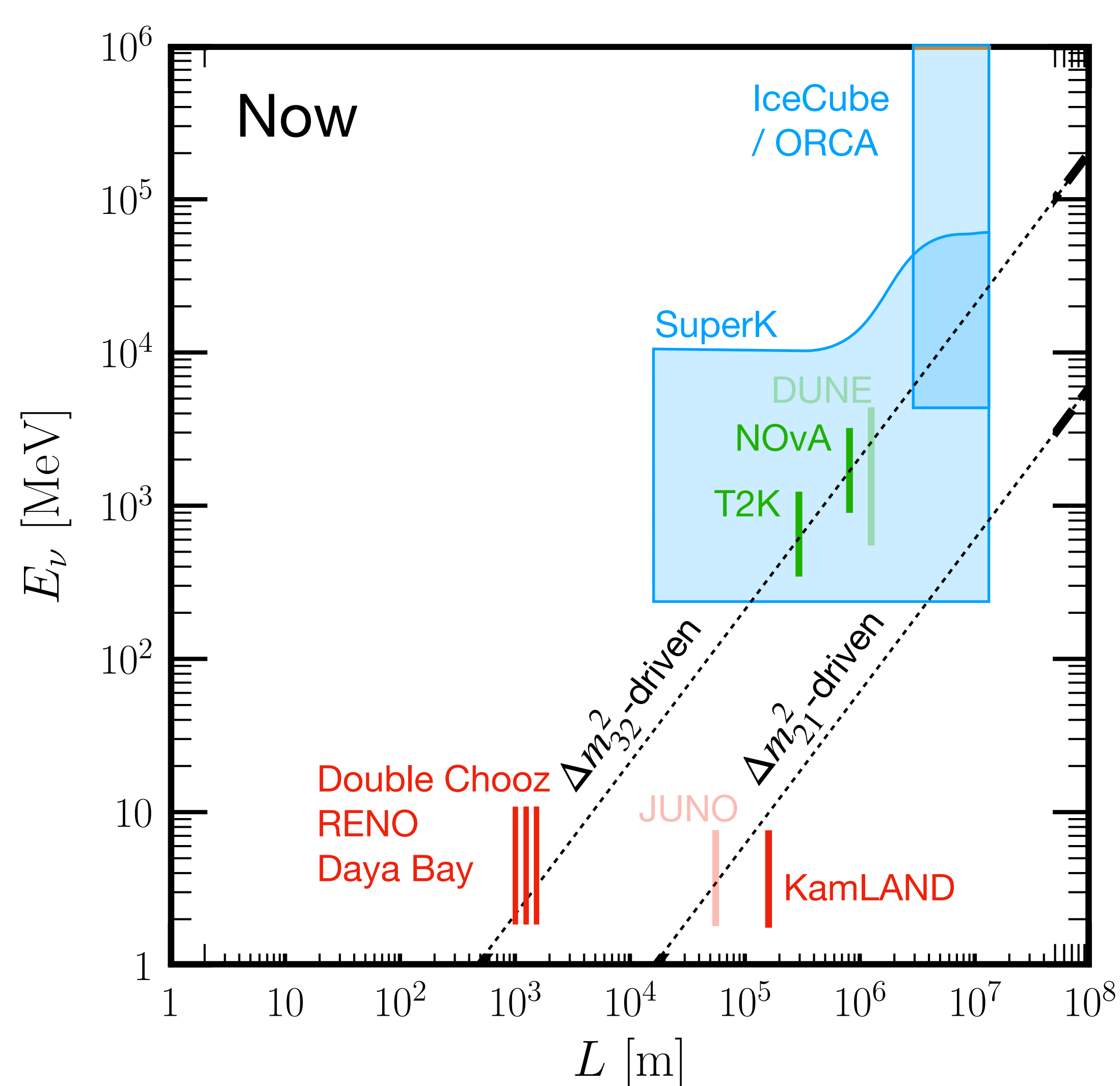


- Lifting δ_{CP} —MO degeneracy
- First joint fit based on past analyses
PRL 134, 011801 (2025)
 - SK: just SK IV period, same reconstruction and baseline interaction model as T2K
PTEP 2019, 053F01
 - T2K: results first shown in 2020
EPJ. C (2023) 83:782

- Correlate (sub-GeV) xsec and detector systematics
- Apply T2K ND xsec constraint to atmospheric fit
- Good fit using correlated systematics
 $p = 0.19$: Atm. vs. ND
 $p = 0.24$: Atm vs. acc (with ND)

- Reject CP-conservation at 1.9σ
Note: T2K data not in MO-degenerate region
- New dedicated test-statistic found to be more powerful than typical (δ_{CP}, MO) -intervals
Profiling over discrete choices ($\delta_{CP} = 0, \pi$ and MO) helps retain calibration of confidence level, especially in joint fit.

Future / Outlook



→ talks on Friday

- Reactor
 - long-baseline JUNO → θ_{12} , Δm^2_{21} , Δm^2_{32} , MO
- Accelerator
 - HyperK → θ_{13} , δ_{CP} (if given MO)
 - DUNE → θ_{13} , δ_{CP} , MO

- Atmospheric
 - IceCube upgrade, ORCA → Δm^2_{32} , θ_{23} , MO

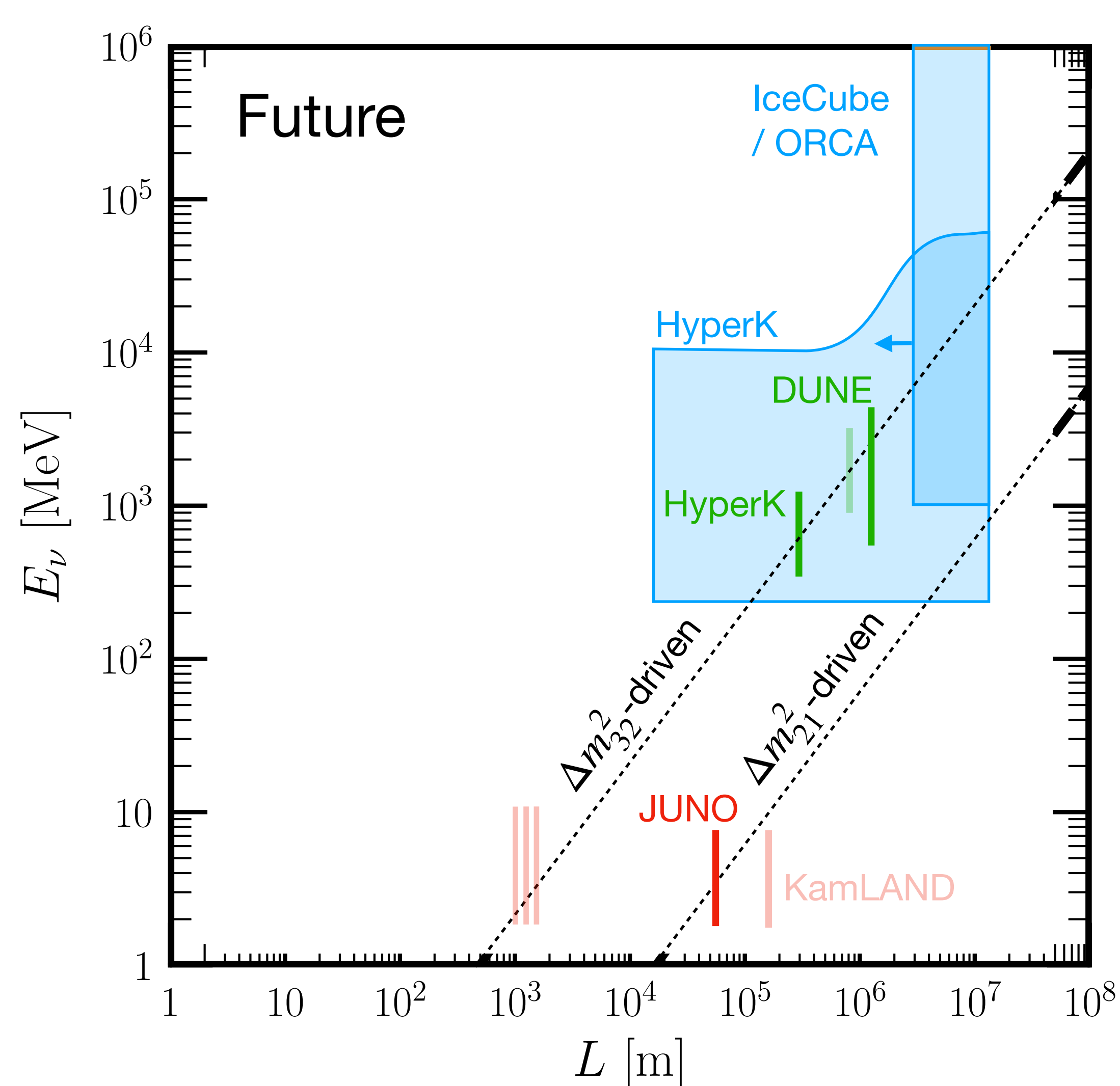
- HyperK → δ_{CP} , θ_{23} , MO → talks on Friday
- DUNE → δ_{CP} (see protons for sub-GeV)

- Cosmological constraints → MO from $\Sigma_i m_i$

- Some challenges

- Precise understanding of flux, interaction (nuclear fx, ν_e , ν_τ), detector systematics
- Effective sharing of results as analyses become more complex, but global analyses are essential if we want to stress-test 3ν -paradigm or look beyond

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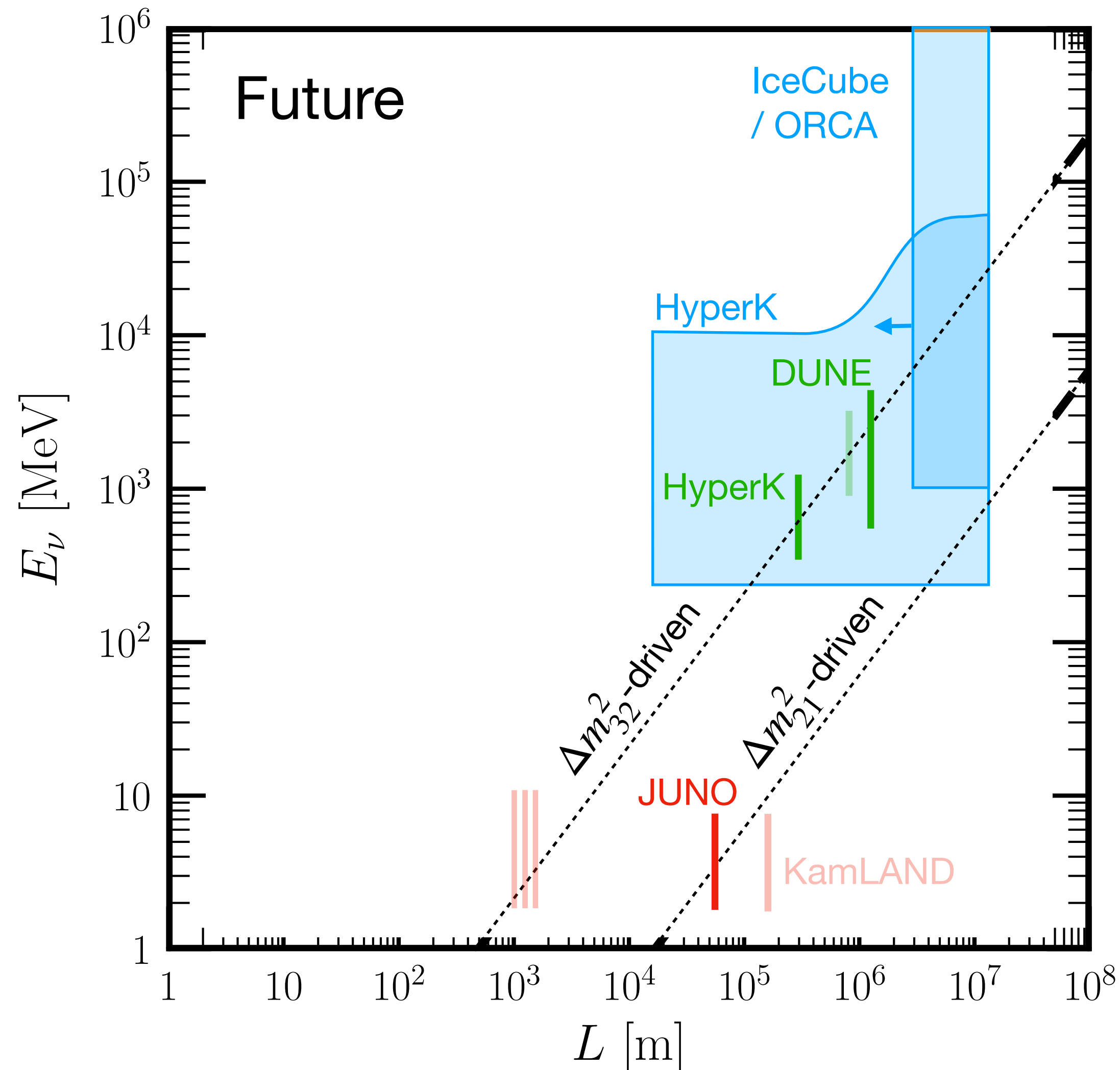
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Future / Outlook



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 - long-baseline JUNO $\rightarrow \theta_{12}, \Delta m_{21}^2, \Delta m_{32}^2, \text{MO}$
- Accelerator
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 - IceCube upgrade, ORCA $\rightarrow \Delta m_{32}^2, \theta_{23}, \text{MO}$
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$|\Delta m_{ee}^2|$
 Currently $\approx 2\%$ \rightarrow JUNO $0.X\%$
(technically value measured by medium baseline and JUNO has a slight offset, so shouldn't call Δm_{ee}^2)

$|\Delta m_{\mu\mu}^2|$
 Already $< 1\%$
(global average)

\downarrow
 Yet another way to get
 MO: from $\Delta m_{\mu\mu}^2$ vs. Δm_{ee}^2
 PRD 72, 013009 (2005)
 PRD 111, 013008 (2025)

Many ways to measure MO!

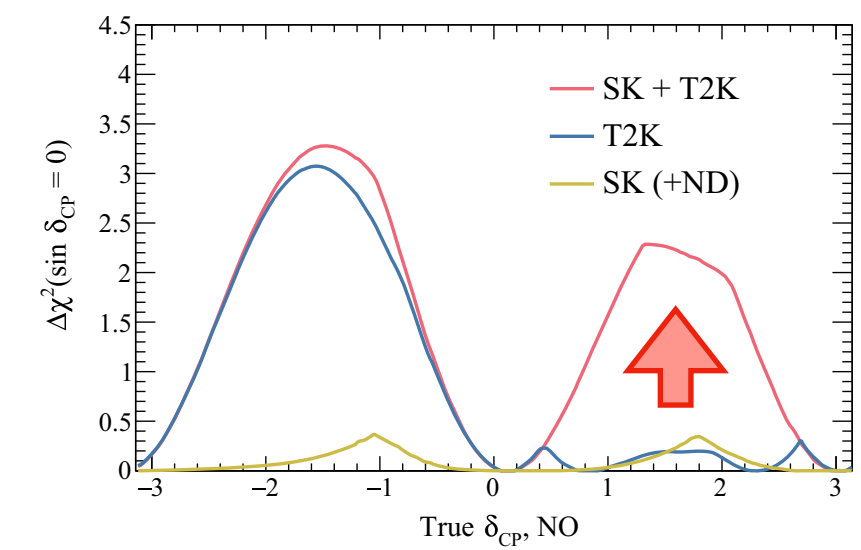
Related parallel talks

- NPA 1B: Philipp Soldin
“Double Chooz Single-Detector Physics Results”
(includes final Double Chooz oscillation results)
- NPA 1B: Zhe Wang
“The latest reactor neutrino oscillation results and reactor neutrino flux and spectrum measurement results from Daya Bay”
- NPA 1B: Antonio Marrone
“Precision Neutrino Physics: Status and Outlook in the 3ν Paradigm”
- NPA 4B: Masaki Ishitsuka
“Latest neutrino oscillation measurements from T2K”
- NPA 7A: Denise Casazza
“The SAND detector of the DUNE experiment”
- NPA 7A: Iwan Morton-Blake
“Commissioning of the JUNO detector”
- NPA 7A: Ruhui Li
“Status and Prospect of JUNO-TAO”
- NPA 7A: Akira Takenaka
“Development of the calibration sources for the JUNO experiment”

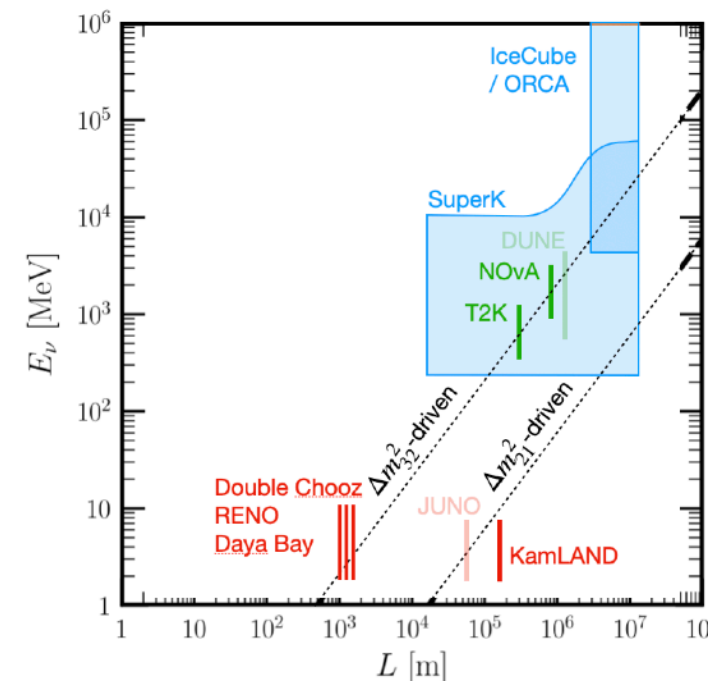
Related posters

- 48 NPA Ali Murat Guler
Results from the DsTau (NA65) experiment at the CERN-SPS
- 161 NPA Yaoguang Wang
Calibration of 20-inch Photomultiplier Tubes in JUNO
- 175 NPA Yukine Sato
Performance Evaluation of a New Sensor for the T2K Muon Monitor in Neutrino Beam Direction Measurements
- 313 NPA Laurence Cook
Reducing Systematic Uncertainties in Neutrino Detection with the Water Cherenkov Test Experiment (WCTE)
- 334 NPA Qishan LIU
Neutron Capture Information in Improving IBD Angular Resolution
- 413 NPA Rui Li
Calibration System of the JUNO experiment
- 497 NPA Jingqin Xue
The accelerated Reactor fitter and Oscillation analysis for JUNO

Summary



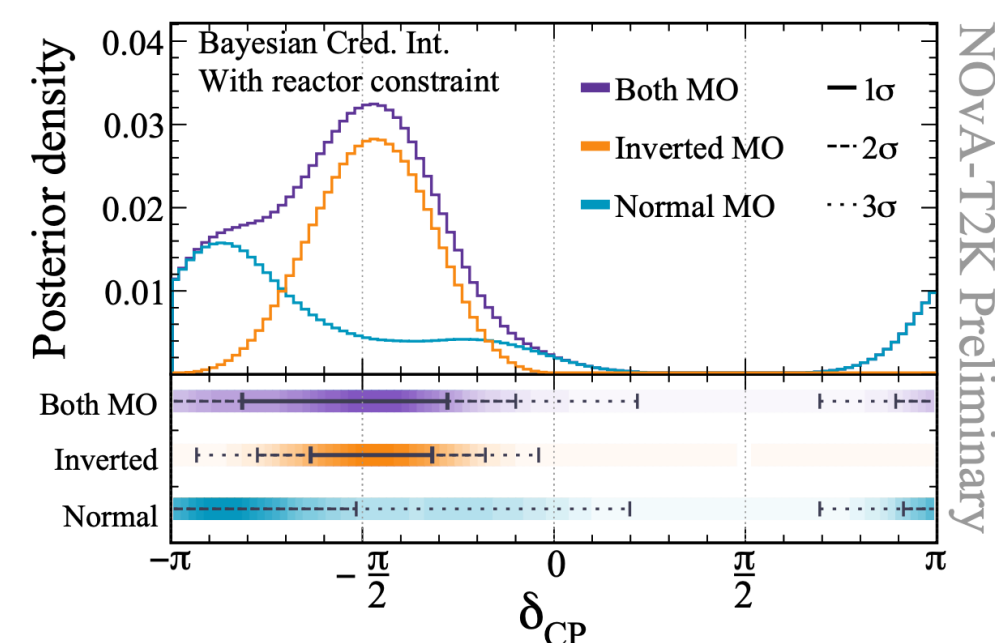
- **3ν oscillation** paradigm successfully describes many reactor/accelerator/atmospheric ν -oscillation experiments. Consistency across



- Different oscillation channels ($\bar{\nu}_e \rightarrow \bar{\nu}_e$, $\nu_\mu \rightarrow \nu_\mu$, $\nu_\mu \rightarrow \nu_e$ etc.)
- Different interaction mechanisms
- Different detector technologies
- Different contributions of matter effects
- Some parameters are entering precision era
| Δm_{32}^2 | is now measured **sub-% (global average)**

Strong contributions from recent **NOvA** and **IceCube** results

- If **IO**, we have evidence for **CP-violation**, but various weak indications of NO



- Two **joint fits** were done (T2K+NOvA, T2K+SK), potential for cancelling degeneracies, testing systematic correlations, stress-testing the model

Significance of such will increase in upcoming systematic-dominant era

- New detectors are coming online:

ORCA, IceCube upgrade, JUNO, HyperK, DUNE

High statistics experiments requiring good understanding of systematics

- Especially for MO, **many ways to measure**

Over the next 4~6 years we may not only get one, but two different measurements of MO?
→ feeds into CPV search and $0\nu\beta\beta$