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Geoneutrino Flux Prediction and Reactor Background at the JNE

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July 2026

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Earth heat budget, BSE debate, geoneutrino detection

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1. Introduction



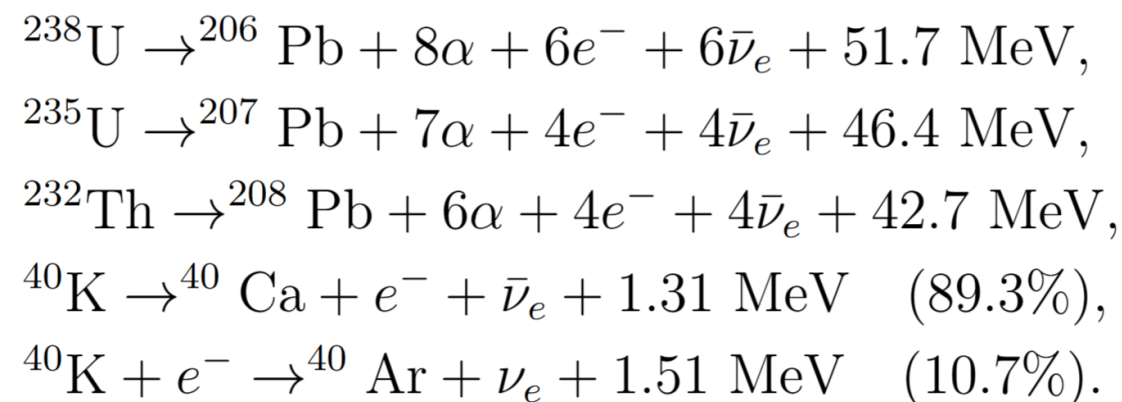
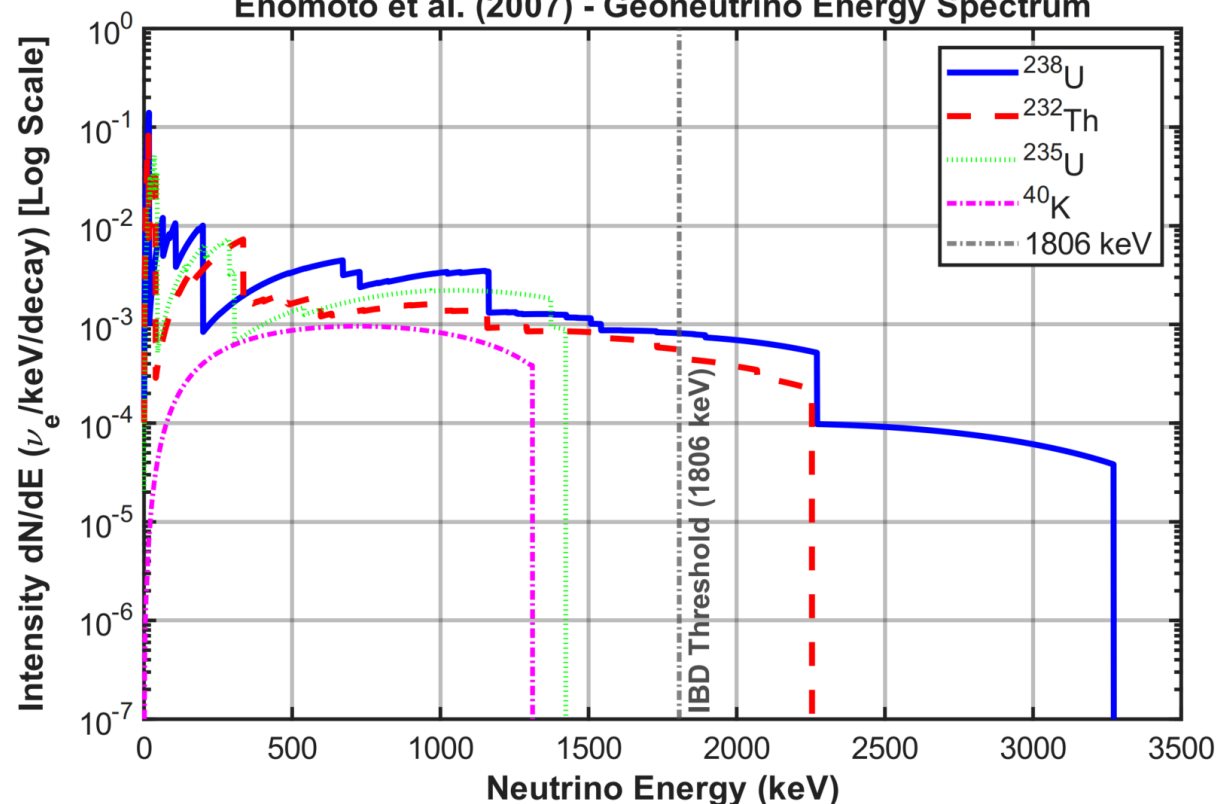
Earth surface heat flow: 46 ± 3 TW

Driven by radiogenic heat from HPEs (U, Th, K) + primordial energy.

Three competing BSE models:

Model	Basis	Radiogenic Heat
Low-Q	Cosmochemical (E-chondrites)	11 ± 2 TW
Medium-Q	Geochemical (mantle xenoliths)	20 ± 4 TW
High-Q	Geodynamical (surface heat flow)	33 ± 3 TW

Enomoto et al. (2007) - Geoneutrino Energy Spectrum



2. Flux Formula

IBD Detection Reaction



IBD Event Rate (in TNU)

$$R_i = \int \phi(E) \times \sigma_{\text{IBD}} dE \times 3.1536 \times 10^3$$

Normalization: 1 year $\approx 3.1536 \times 10^7$ seconds

ϕ : geoneutrino flux [$\text{cm}^{-2} \cdot \text{s}^{-1} \cdot \text{MeV}^{-1}$]

σ : IBD cross section [cm^2]

1 TNU = 1 event / 10^{32} protons / 1 year

Differential Geoneutrino Spectrum

$$\phi_i(E) = \frac{X_i \lambda_i N_A}{\mu_i} n_\nu(i) \int \frac{A_i(\vec{r}) \rho(\vec{r})}{4\pi L^2} P_{ee}(E, L) f_i(E) d\vec{r}$$

λ_i : decay constant N_a : Avogadro constant

μ_i : atomic mass $n_\nu(i)$: $\bar{\nu}_e$ per decay chain

$A_i(r)$: HPE abundance $\rho(r)$: density

L : distance P_{ee} : survival probability

$f_i(E)$: normalized $\bar{\nu}_e$ spectrum per decay

Neutrino Oscillation Survival Probability

$$P_{ee}(E, L) = 1 - \cos^4 \theta_{13} \cdot \sin^2 2\theta_{12} \cdot \sin^2(\Delta m_{21}^2 L / 4E) - \sin^2 2\theta_{13} \cdot \sin^2(\Delta m_{31}^2 L / 4E)$$

3 Geophysical Model

3.1 Mantle Model

Scenario	Description
① Uniform	Homogeneous composition throughout the mantle
② Layered Uniform	Depleted upper mantle + enriched lower mantle ✓
③ Single Enriched Blob	LSVPs treated as one enriched layer
④ Symmetric Blobs	Two enriched blobs beneath Pacific & Africa

→ Mantle contribution robust:
only ~10% variation across all models (Šrámek et al., 2013)

LAB (Lithosphere-Asthenosphere Boundary) Depth range: 120–250 km (crustal model constrained)
Depleted Mantle (DM) – 8 layers in total - Layer 1: 50 km thickness - Layer 2: 100 km thickness - Layer 3: 150 km thickness - Layer 4: 250 km thickness - Layer 5: 400 km thickness - Layer 6: 400 km thickness - Layer 7: 400 km thickness - Layer 8: Remaining thickness
Enriched Mantle (EM) – only in layered model Layer 9: 750 km thickness
CMB (Core-Mantle Boundary) Fixed depth: 2891 km

Crustal Model



Model	Year	Resolution	Status
CRUST5.0	~2000	5°×5°	Too coarse
CRUST2.0	~2004	2°×2°	Too coarse
CRUST1.0 ✓	2012	1°×1°	Global standard
LITHO1.0	2014	1°×1°	Includes lithosphere
ECM1 ✓	2023	1°×1°	Updated synthesis

Key Findings (Wipperfurth et al., 2020)

- The total geoneutrino signal **is insensitive** to the choice among the CRUST1.0, CRUST2.0, and LITHO1.0 models.
- Improving **near-field crustal model** is more critical for signal accuracy

Vertical column profile for cell k:

s1 (Upper sedimentary layer) – $\text{thick}(k) \times 1$ block
s2 (Middle sedimentary layer) – $\text{thick}(k) \times 1$ block
s3 (Lower sedimentary layer) – $\text{thick}(k) \times 1$ block
UC (Upper crust) – $\text{thick}(k) \times 1$ block
MC (Middle crust) – $\text{thick}(k) \times 1$ block
LC (Lower crust) – $\text{thick}(k) \times 1$ block
LM (Lithospheric mantle) – $\text{thick}(k) \times 1$ block
Mantle (LAB → CMB) – subdivided into 9 sublayers

3.2 SWChinaCVM-2.0

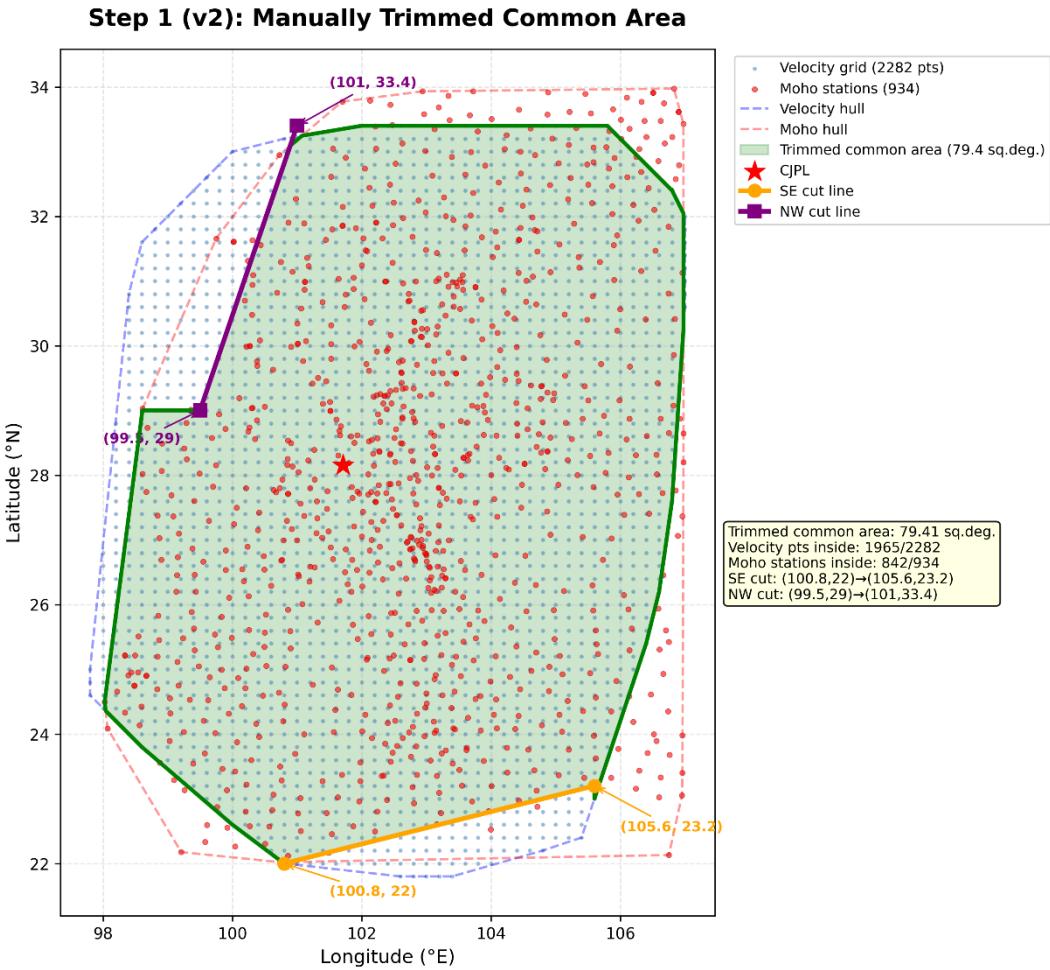
Feature	Detail
Release	2022, China Seismic Experimental Site (CSES)
Coverage	Sichuan-Yunnan region (encompasses JNE site)
Resolution	0.2° × 0.2° (upper crust) — vs 1°×1° for CRUST1.0
Data Volume	~7.06M body-wave + ~0.1M surface-wave data
Method	AI phase picking + ambient noise cross-correlation

Application: Refined Near-Field Crustal Density Model

SWChinaCVM-2.0 Vp/Vs profiles → density (Brocher 2005)
→ Merged with CRUST1.0 for a 0.2° combined near-field model

Also: SWChinaCVM-MOHO-2.0

Integrated 3 new-generation velocity models (SWChinaCVM-2.0, CSES-VM1.0, USTClitho2.0) → Public Moho interface model



3.3 CRUST1.0-SWChina-merged

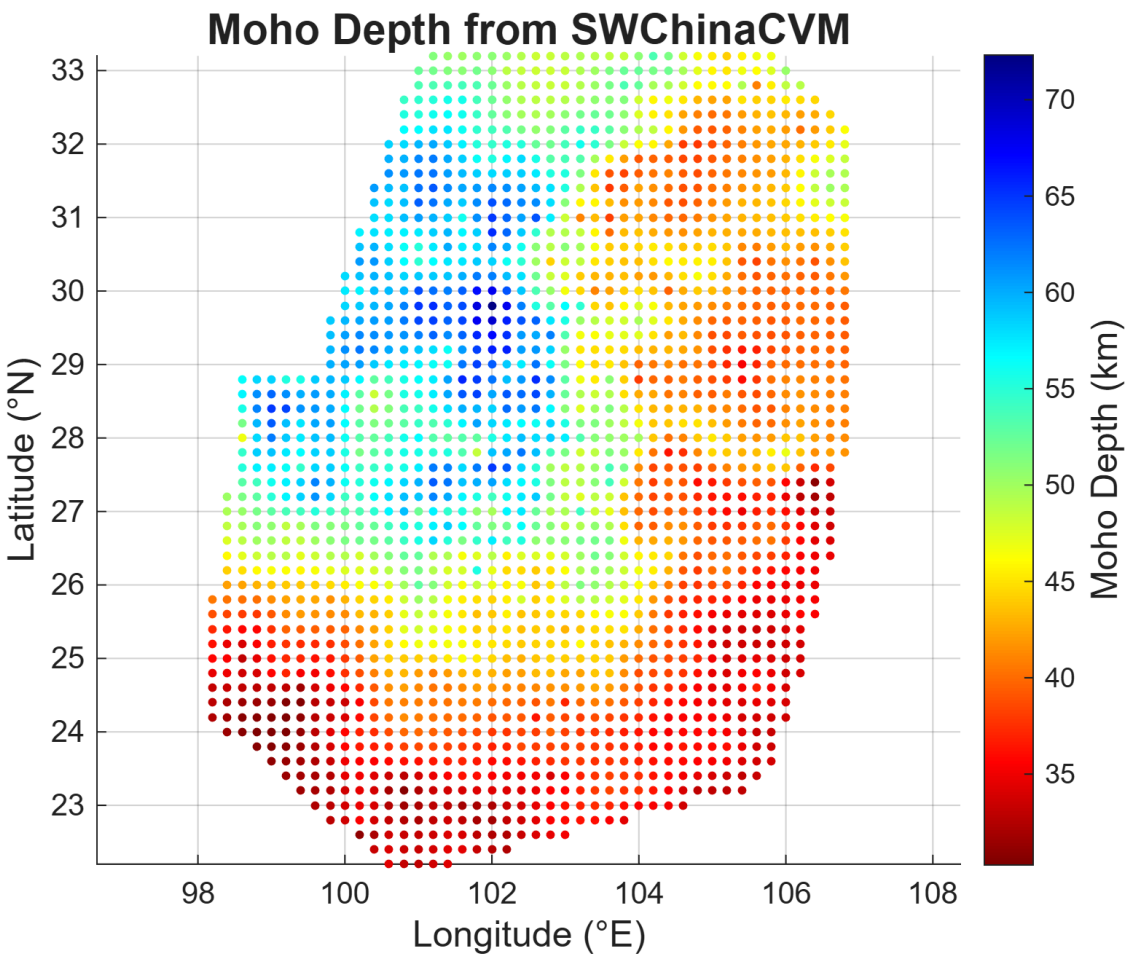
Step 1: Moho Depth for Each 0.2° Cell

Data Source	Role
SWChinaCVM-2.0	Vp/Vs profiles at each (lon,lat,depth)
SWChinaCVM-MOHO-2.0	Crustal thickness + Vp/Vs at station locations

Processing:

- 1. Identify common overlapping area between both model coverages
- 2. Kriging interpolation on MOHO-2.0 station data
- 3. Derive Moho depth at same (lon,lat) grid as CVM-2.0

→ **Output: Moho depth map at 0.2° resolution (Fig. 4 in paper)**



Step 2: Surface Elevation (ETOPO1)

CVM-2.0 issue	Reports uniform max 5 km at each grid (no real topography)
Topography source	The ETOPO1 Ice Surface data are sourced from NOAA.

Method:

For each 0.2° cell, average 9 sub-points (3×3 grid)

Replicates the averaging scheme used in CRUST1.0

→ **Compatible with CRUST1.0 surface representation**

Step 3: Sediment & Crustal Proportions

- For each 0.2° grid point (lon_i, lat_i) , identify the parent 1° CRUST1.0 grid cell that contains it, i.e., the cell centered at

$$([lon_i - 0.5] + 0.5, [lat_i - 0.5] + 0.5).$$

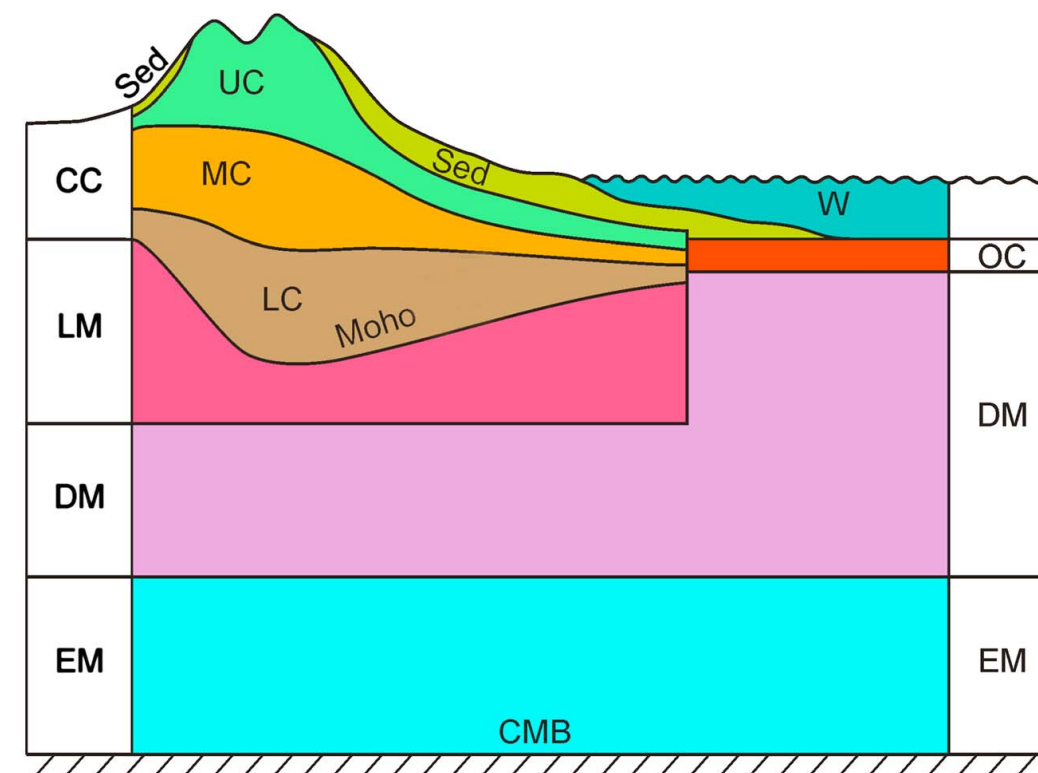
- From CRUST1.0, extract the following quantities for that parent cell:

- Sediment layer thicknesses: `s1.thick`, `s2.thick`, `s3.thick`
- Upper crustal thickness: `UC_thick`
- Middle crustal thickness: `MC_thick`
- Lower crustal thickness: `LC_thick`
- Total crystalline crustal thickness:

$$crystal_thick = UC_thick + MC_thick + LC_thick$$

- Layer fractions:

$$UC_{frac} = \frac{UC_{thick}}{crystal_{thick}}, \quad MC_{frac} = \frac{MC_{thick}}{crystal_{thick}}, \quad LC_{frac} = \frac{LC_{thick}}{crystal_{thick}}$$



Step 4: Interface Depths

For each 0.2° grid cell, using:

- Moho depth from Step 1
- Surface elevation from Step 2 (ETOPO1)
- Sediment thickness from Step 3 (CRUST1.0)
- CRUST1.0 layer fractions from Step 3

Crystalline crustal thickness:

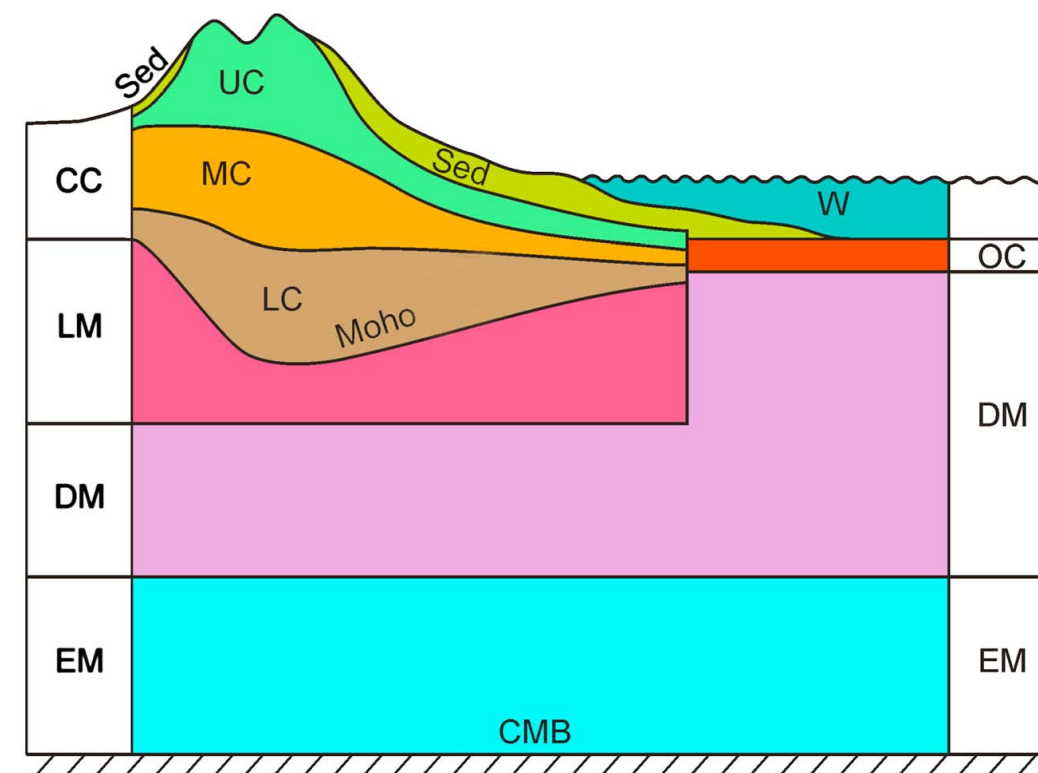
$$\text{crystal_thick} = \text{moho_depth} - \text{sed_thick}$$

Layer thicknesses:

$$\text{UC_thick} = \text{crystal_thick} \times UC_{\text{frac}}$$

$$\text{MC_thick} = \text{crystal_thick} \times MC_{\text{frac}}$$

$$\text{LC_thick} = \text{crystal_thick} \times LC_{\text{frac}}$$



Step 5: Representative Vp & Vs per Layer

For each 0.2° grid cell:

1. Retrieve CVM-2.0 depth profile (16 levels: -5 to 70 km)
2. Thickness-weighted average Vp/Vs within each layer
3. For empty layers (too thin): linear interpolation at midpoint

Velocity profile at the JNE site

Layer	Thick (km)	Vp (km/s)	Vs (km/s)	Vp/Vs
s3	0.01	5.41	3.05	1.78
UC	26.6	5.99	3.51	1.71
MC	23.6	6.24	3.54	1.76
LC	10.3	7.01	3.89	1.80

Step 6: Density Computation

Brocher (2005) empirical polynomial:

$$\rho = 1.6612V_p - 0.4721V_p^2 + 0.0671V_p^3 - 0.0043V_p^4 + 0.000106V_p^5$$

Velocity gradient: ~6.0 km/s (UC) → ~7.0 km/s (LC)
consistent with felsic→mafic crustal stratification

Step 7: Merging SWChinaCVM-2.0 with CRUST1.0

Part 1: Generate 2,325 Sub-Cells

For each of the 93 parent 1° CRUST1.0 cells:

Generate 25 sub-cells at 0.2° resolution (5×5 grid)

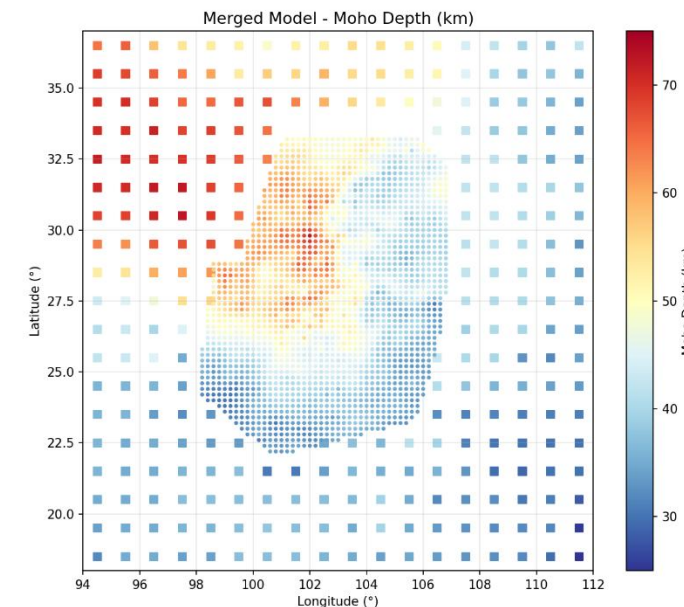
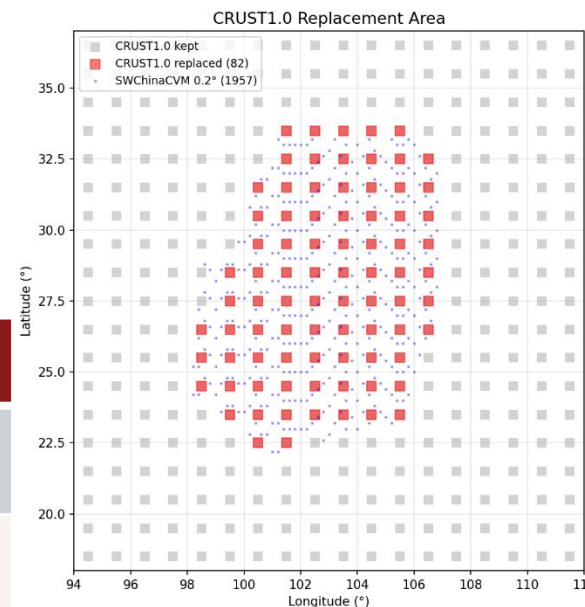
Total: $93 \times 25 = 2,325$ sub-cells

Part 2: Classify Sub-Cells

Group	Count	Assignment
A	1,957	Near-field model grid exists → use nearfieldmodel0.2deg.mat
B	368	Not covered → inherit CRUST1.0 parameters, update coordinates

Part 3: Final Model Assembly

Component	Count
Original CRUST1.0 cells	64,707
Group A (near-field, 0.2°)	1,957
Group B (CRUST1.0 sub-cells, 0.2°)	368



Final Cell Count

67,032

total cells in the merged model

4. Geochemical Model

HPE Abundance per Layer

Layer	Source / Method
Upper Crust	Rudnick & Gao (2014) meta-analysis
Middle Crust	Amphibole Vp–composition linear fit
Lower Crust	Granulite Vp–composition linear fit
Sediments	GLOSSII model (Plank, 2014)
Oceanic Crust	Global average (White & Klein, 2014)
Lithospheric Mantle	Xenolith data (Huang et al., 2013)
Mantle (DM+EM)	Depleted mantle + enriched mantle (19% by mass)

Key Coupling with Geophysical Model

Middle & lower crust HPE abundances depend on **Vp**
SWChinaCVM-2.0 Vp profiles directly control HPE prediction in these layers



Vp–Composition Linear Relationship

Mass balance between felsic and mafic endmembers:

$$V_{\text{model}} = V_f f + V_m m \tag{6}$$

$$1 = f + m \tag{7}$$

$$a_{\text{output}} = a_f f + a_m m \tag{8}$$

where V_f , V_m , a_f , and a_m are the V_P and abundance of the felsic and mafic endmembers of amphibolite (middle crust) and granulite (lower crust), V_{model} is the V_P provided by the geophysical model, and f and m are the mass proportions of felsic and mafic endmembers.

BSE Composition Calculation

Element	Method
U abundance	McDonough & Sun (1995)
Th abundance	Derived from Th/U ratio
K abundance	Derived from K/U ratio

Table 5: Geoneutrino flux from the near-field + CRUST1.0 combined model.

Layer	U (TNU)	Th (TNU)	Total Flux (TNU)
Sediment 1	0.0596	0.0189	0.0785
Sediment 2	0.2835	0.0910	0.3745
Sediment 3	0.1718	0.0546	0.2264
Upper Crust	27.9760	7.2584	35.2344
Middle Crust	6.8154	2.7482	9.5636
Lower Crust	0.6839	0.3307	1.0146
Total CC	35.9902	10.5019	46.4921
Oceanic Sediment	0.0106	0.0033	0.0139
Oceanic Crust	0.0041	0.0008	0.0050
Lithospheric Mantle	1.1349	0.3484	1.4833
Total Lithosphere	37.2796	10.8939	48.1735
Depleted Mantle	3.8246	0.8908	4.7153
Enriched Mantle	2.7104	0.5891	3.2995
Total BSE	43.8146	12.3738	56.1883

Table 6: Geoneutrino flux from the ECM1 model.

Layer	U (TNU)	Th (TNU)	Total Flux (TNU)
Sediment 1	0.3148	0.0998	0.4147
Sediment 2	0.2946	0.0934	0.3881
Sediment 3	0.0664	0.0209	0.0874
Upper Crust	26.1793	6.7602	32.9395
Middle Crust	6.3014	2.3303	8.6317
Lower Crust	1.4877	0.8085	2.2963
Total CC	34.6444	10.1132	44.7576
Oceanic Sediment	0.0458	0.0145	0.0603
Oceanic Crust	0.0103	0.0021	0.0124
Lithospheric Mantle	0.7776	0.2376	1.0152
Total Lithosphere	35.5246	10.3791	45.9037
Depleted Mantle	3.3726	0.7852	4.1578
Enriched Mantle	2.8275	0.6229	3.4504
Total BSE	41.7247	11.7872	53.5119



Table 7: Geoneutrino flux from the CRUST1.0 model.

Layer	U (TNU)	Th (TNU)	Total Flux (TNU)
Sediment 1	0.0883	0.0280	0.1164
Sediment 2	0.3499	0.1109	0.4608
Sediment 3	0.1706	0.0541	0.2247
Upper Crust	29.3805	7.6106	36.9911
Middle Crust	7.4499	3.0062	10.4561
Lower Crust	0.7605	0.3698	1.1304
Total CC	38.1998	11.1796	49.3794
Oceanic Sediment	0.0106	0.0033	0.0139
Oceanic Crust	0.0041	0.0008	0.0050
Lithospheric Mantle	1.4612	0.4472	1.9084
Total Lithosphere	39.8155	11.6704	51.4859
Depleted Mantle	4.0275	0.9378	4.9653
Enriched Mantle	2.7091	0.5888	3.2979
Total BSE	46.5521	13.1970	59.7491

bse.aU =

randist(mantle_U_total,mantle_U_total*0.2,iter,cor.bse)*10⁻⁹

5. Reactor Background

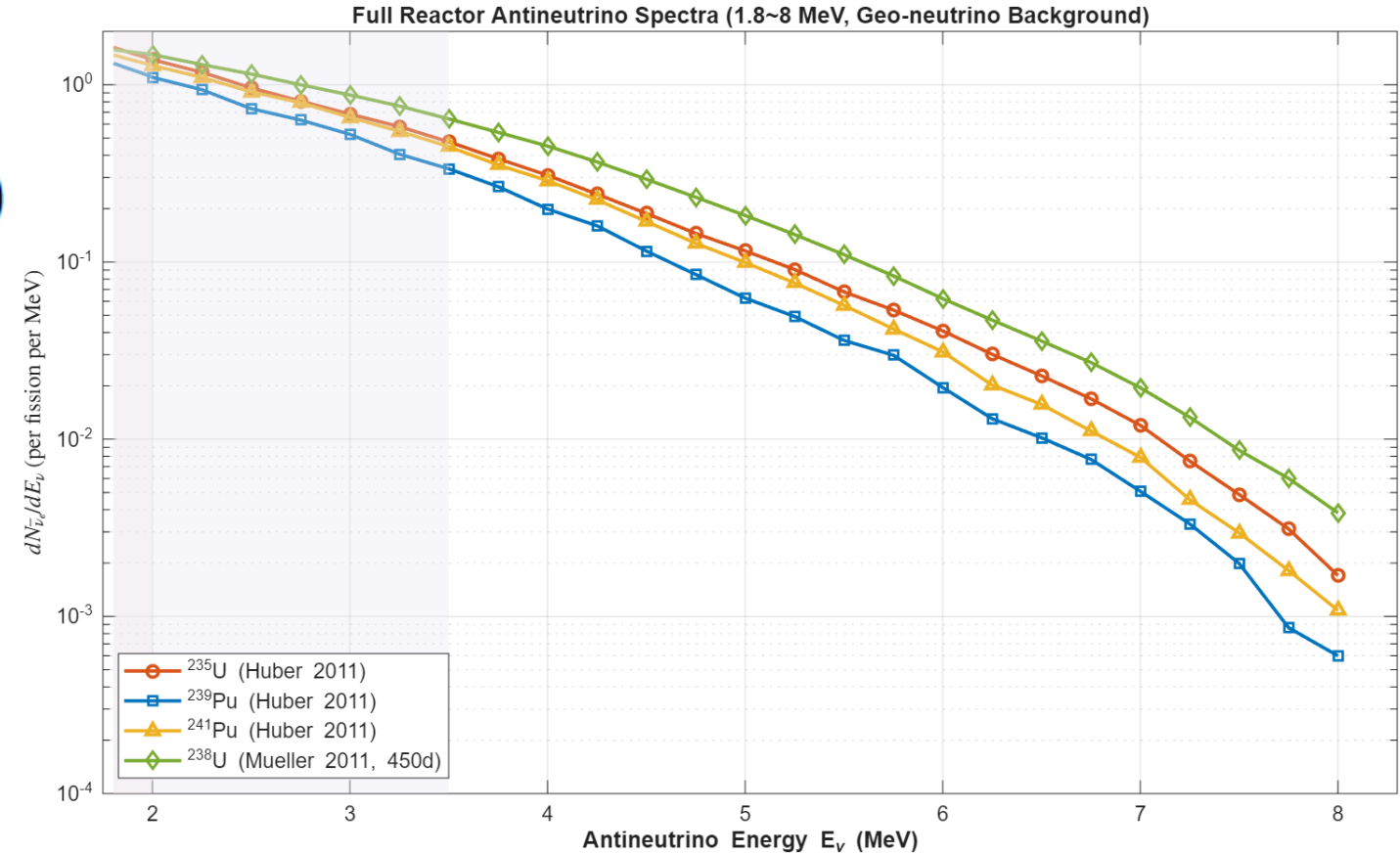


Differential flux from a reactor:

$$\phi(E_{\bar{\nu}_e}) = \frac{W_{\text{th}} \cdot LF}{\sum_i f_i e_i} \sum_i f_i S_i(E_{\bar{\nu}_e})$$

Symbol	Meaning
W_{th}	Thermal power of reactor
LF	Load factor (= 1.0 in this study)
f_i	Fission fraction of isotope i ($\sum f_i = 1$)
e_i	Energy released per fission [MeV/fission]

Isotope	f_i	e_i [MeV/fission]
^{235}U	0.58	202.36 ± 0.26
^{238}U	0.07	205.99 ± 0.52
^{239}Pu	0.30	211.12 ± 0.34
^{241}Pu	0.05	214.26 ± 0.33

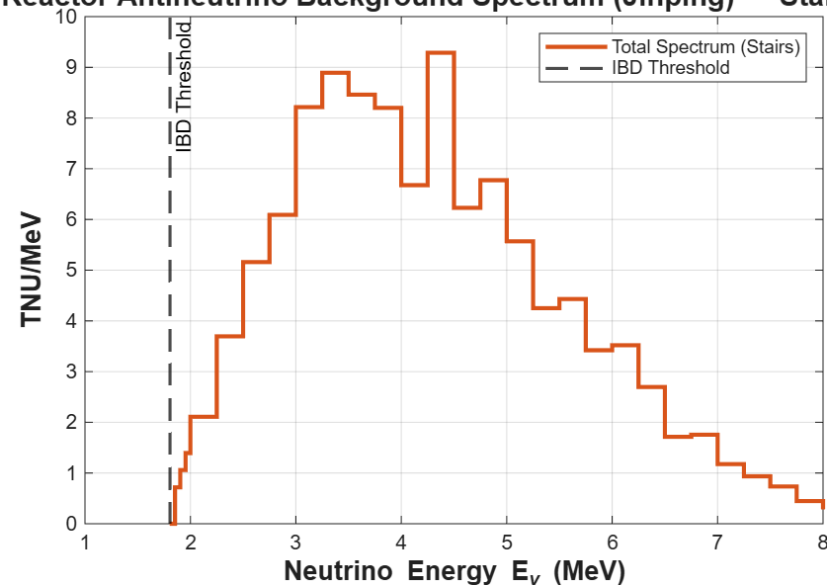


Energy Spectrum Treatment

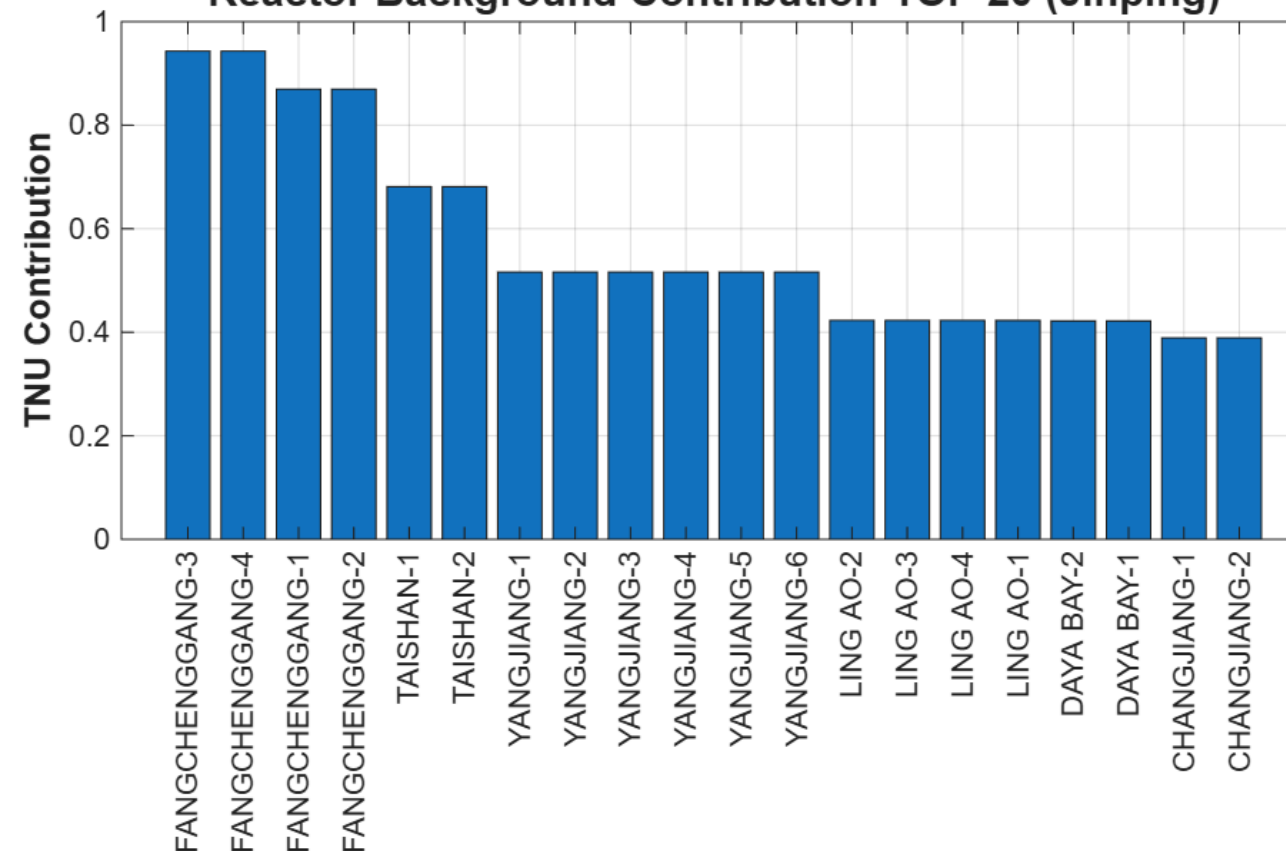
- Tabulated data start at 2.0 MeV (Huber 2011; Mueller 2011)
- Gap filled via official exponential polynomial parameterization
→ 0.05 MeV intervals from 1.806 to 2.0 MeV
- Above 2.0 MeV: original binned data (0.25 MeV resolution)



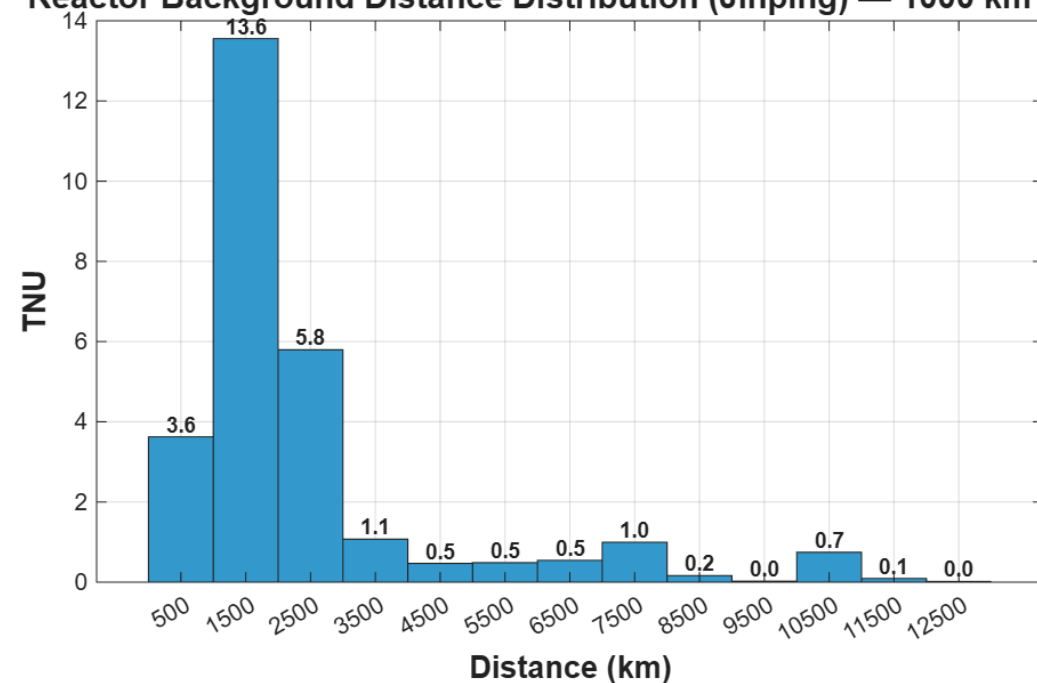
Reactor Antineutrino Background Spectrum (Jinping) — Staircase

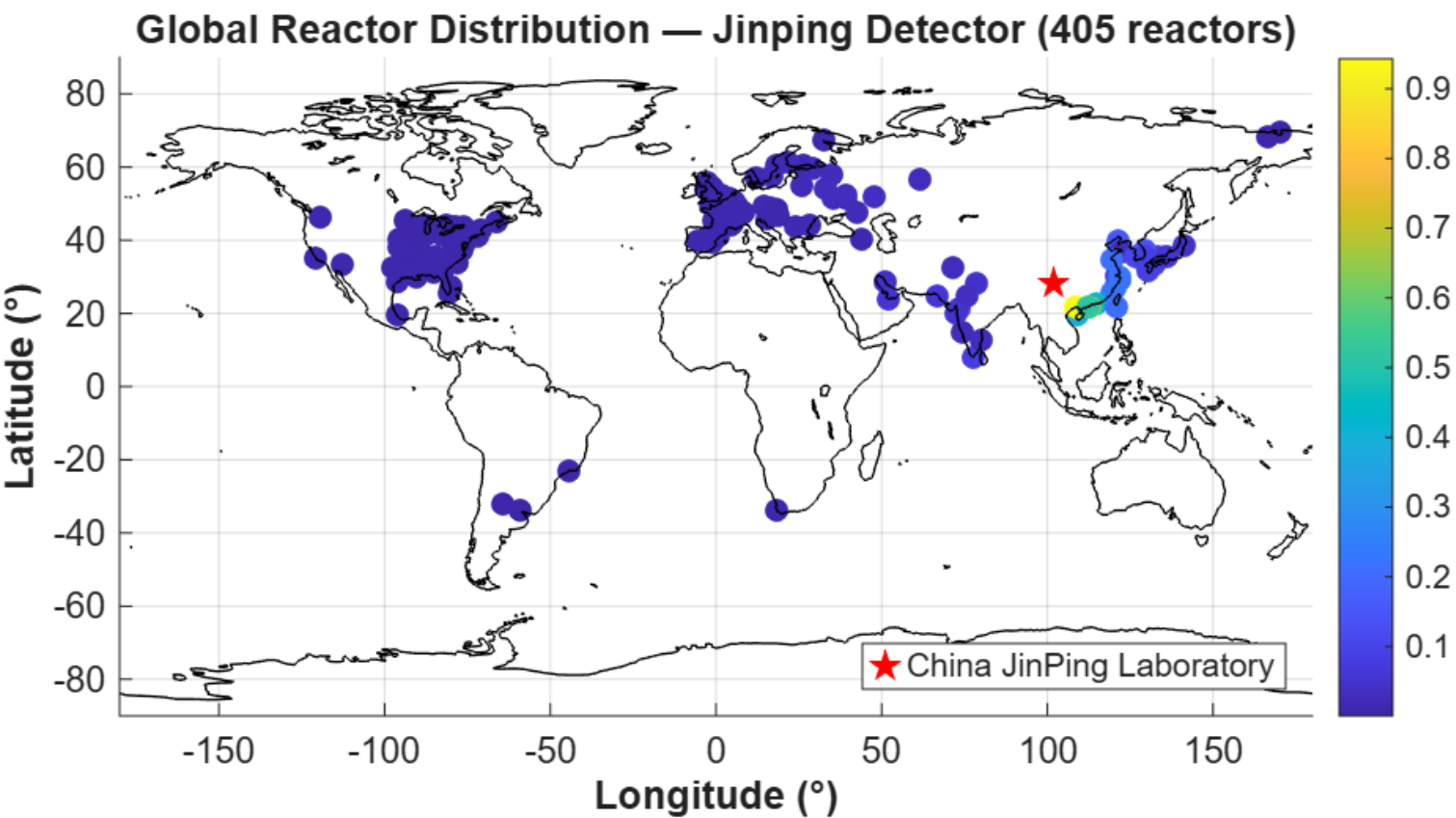


Reactor Background Contribution TOP 20 (Jinping)

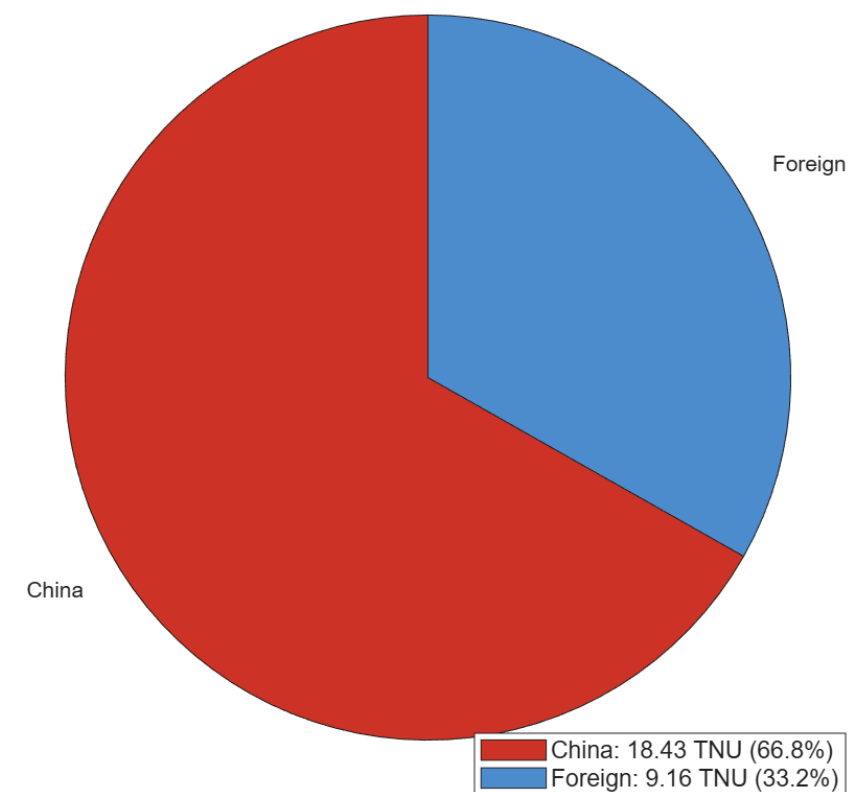


Reactor Background Distance Distribution (Jinping) — 1000 km Bins

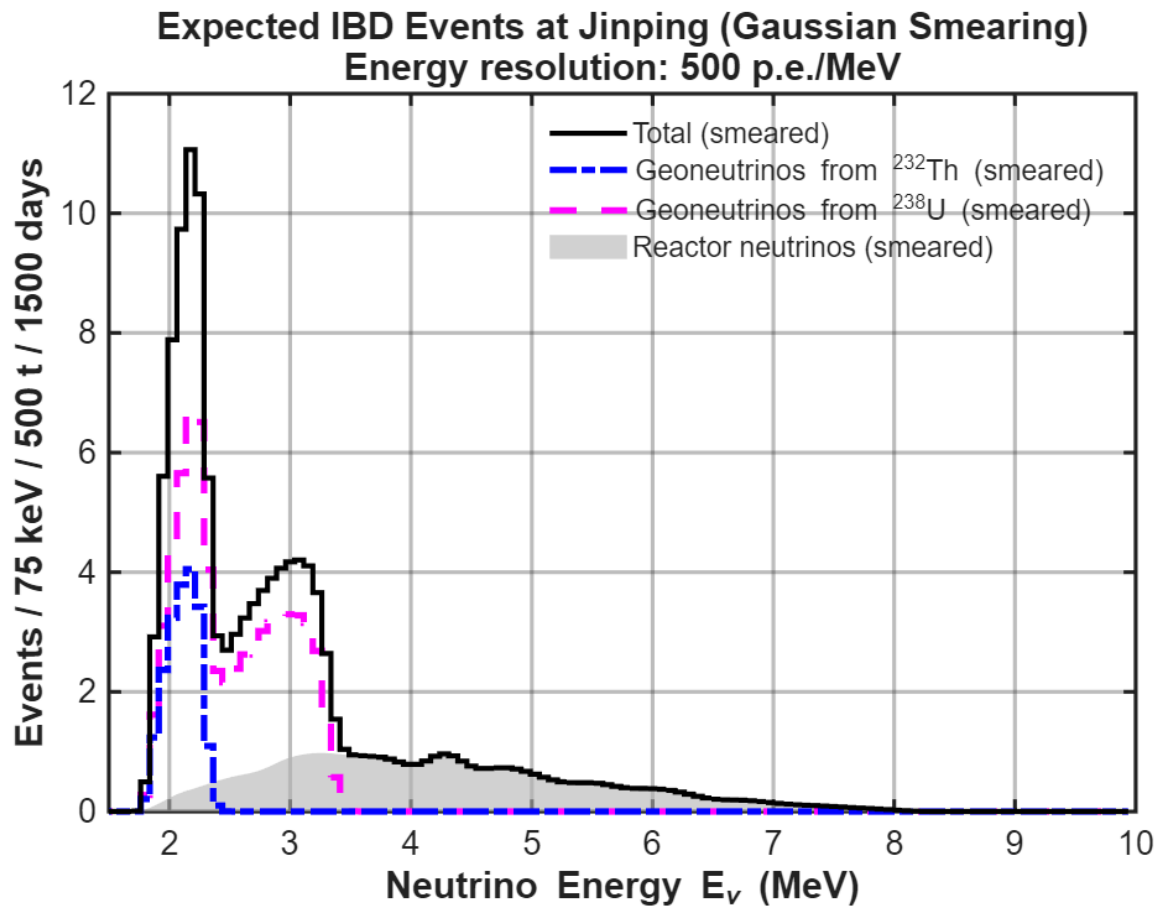




China vs Foreign Reactor TNU Contribution (Total: 27.59 TNU)



6.1 Total IBD Spectrum at Jinping



Event Rates (500 t × 1500 days)

	Geoneutrino		Reactor	
	^{238}U	^{232}Th	Total	-
Rate (TNU)	46.55	13.20	59.75	27.59
Events	68.8	19.5	88.3	40.9

6.2 Monte Carlo results

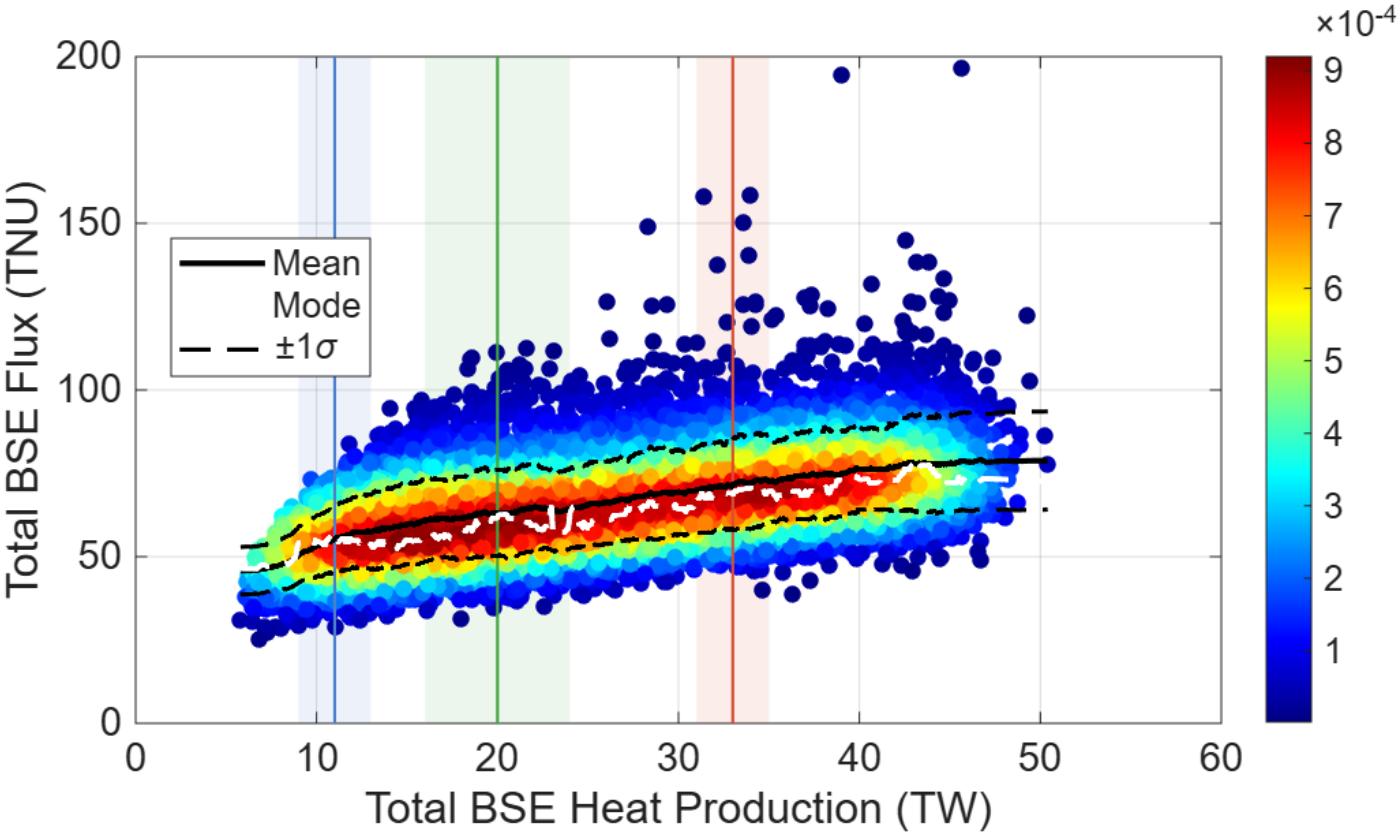
Purpose: Error Propagation

- Geophysical + geochemical inputs all carry inherent uncertainties
- [U, Th, K] abundances follow log-normal distributions
- Parameters are correlated → **analytical propagation impossible**
- Monte Carlo: sample every input PDF → recalculate full model

```
bse.aU = 6*10^-9 + 40*10^-9*rand(iter,1);
```

Table 11: Uncertainty budget for the predicted geoneutrino flux.

Source	Contributing parameters	Total contribution
Geophysical inputs	V_p ($\pm 5\%$), Thickness ($\pm 12\%$), Density ($\pm 5\%$)	$\sim 30\%$
Geochemical inputs	Layer abundances, End-member compositions, Ratios	$\sim 70\%$



7 Summary & Outlook

Jinping for Geoneutrino Detection

Advantage

Relatively low reactor background

→ Highest signal-to-background ratio of any geoneutrino detector

Challenge

Exceptionally complex crustal structure (Tibetan Plateau margin)

Thick crust → larger crustal signal uncertainty

Crustal Model Impact

Model	U (TNU)	Th (TNU)
Near-field+CRUST1.0	43.81	12.37
ECM1	41.72	11.79
CRUST1.0 (global)	46.55	13.20

Directions for Future Improvement

1. Higher-resolution near-field crustal model beyond the current 0.2° grid
2. **Field survey of U/Th abundances in the upper crust around CJPL** — primary contributor to total signal
3. More accurate estimation of U/Th in the middle and lower crusts (current Vp-based linear method is coarse)



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Thank You